

Real time cosmology

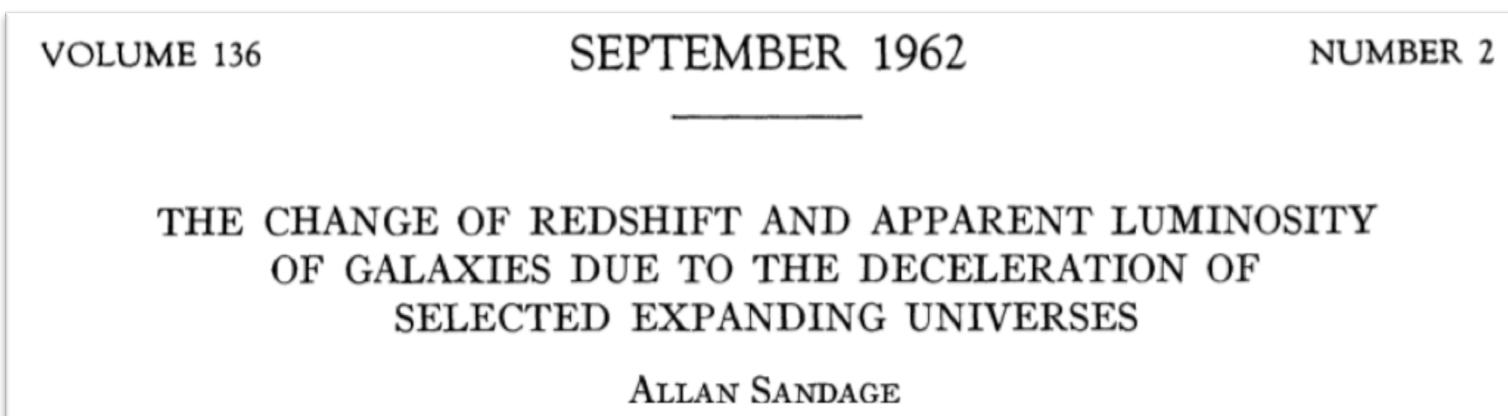
GLOWSKA 2016 KIT

Hans-Rainer Klöckner (MPIFR)

Real time cosmological experiments

Idea measure the change of observables within 10 or 30 years

- angle – cosmological parallax less than $\mu\text{as} / \text{years}$

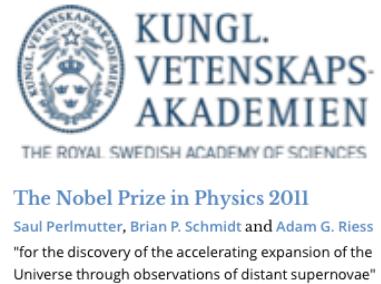


1. The foregoing considerations show that an “ideal” deceleration test exists between the exploding and the steady-state models in the sense that the *sign* of the effect is reversed. However, for the test to be useful, it would seem that a precision redshift catalogue must be stored away for the order of 10^7 years before an answer can be found because the decelerations are so small by terrestrial standards.

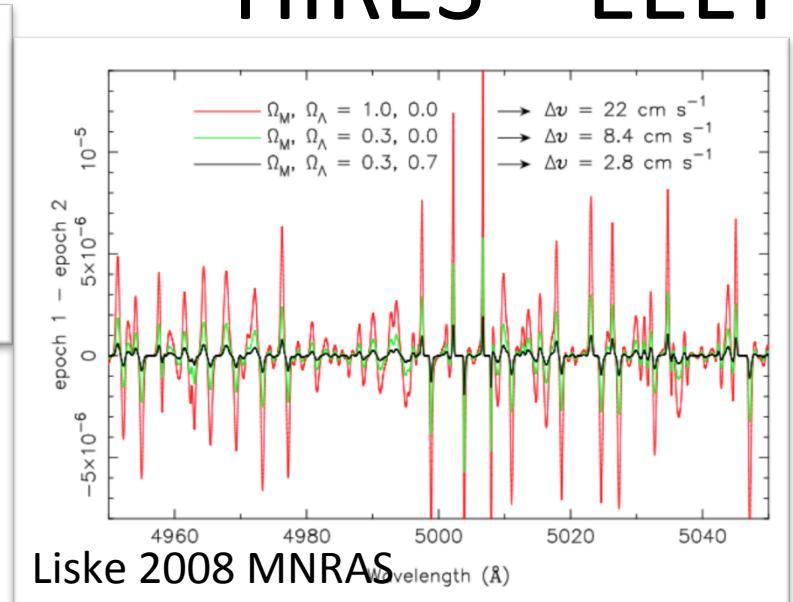
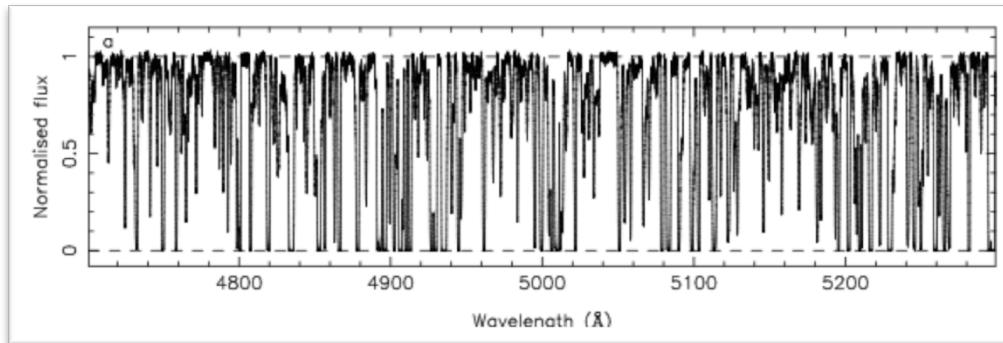
CODEX like experiment

Observational evidence from supernovae for an accelerating universe and a cosmological constant [Riess, et al., 1998, AJ 116]

Measurement of Ω and λ from 42 high-redshift supernovae
[Perlmutter et al., 1999, ApJ 517]



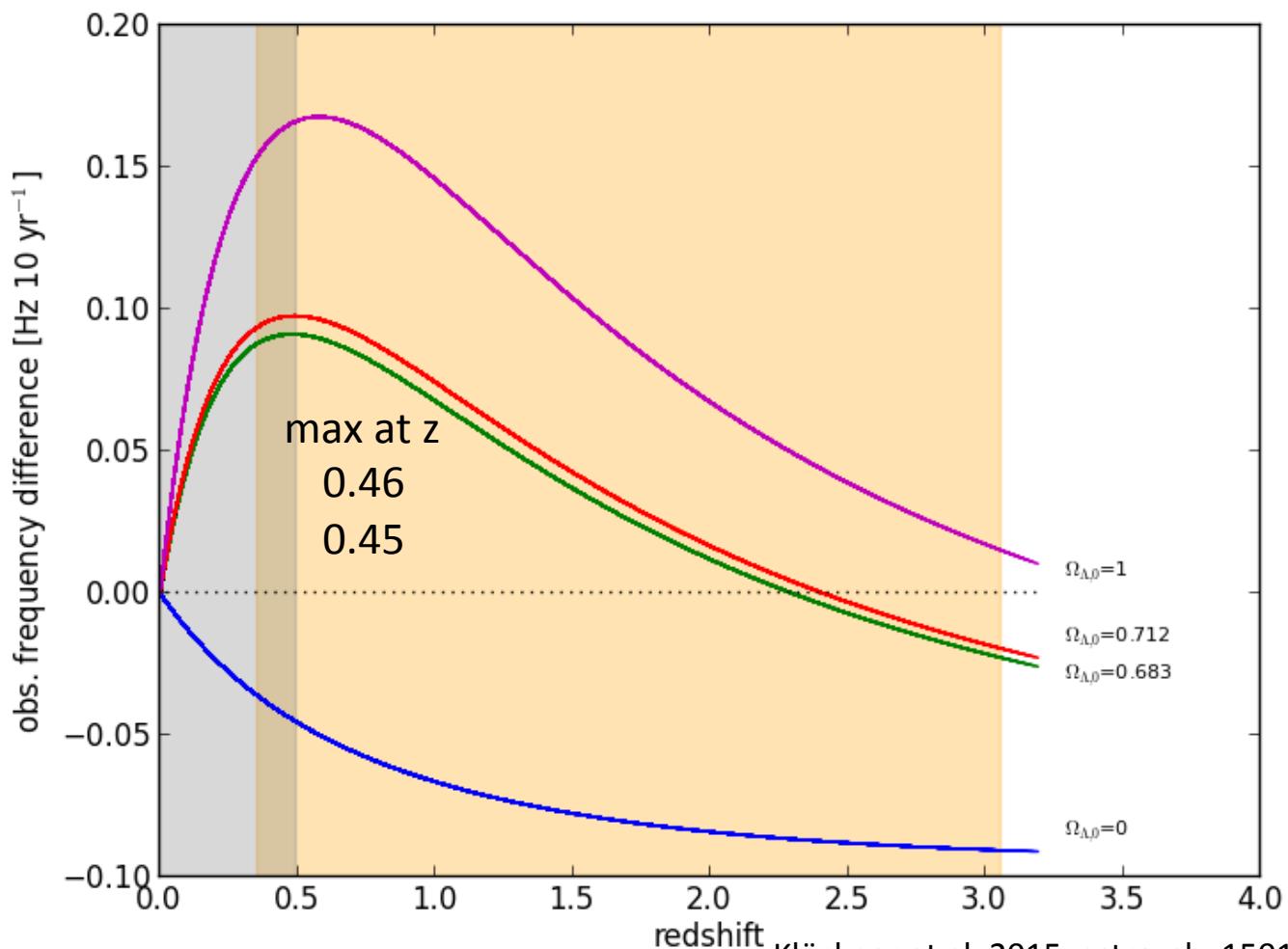
A. Loeb 1998 The cosmic signal can then be searched for through a cross-correlation analysis of the Ly α forest template in the spectrum of these quasars, taken at two different times with the HIRES instrument on the Keck 10 meter telescope.



Key figure in the frequency regime (10 yrs)

from theory $\dot{z} = H(z) - (1+z)H_0$

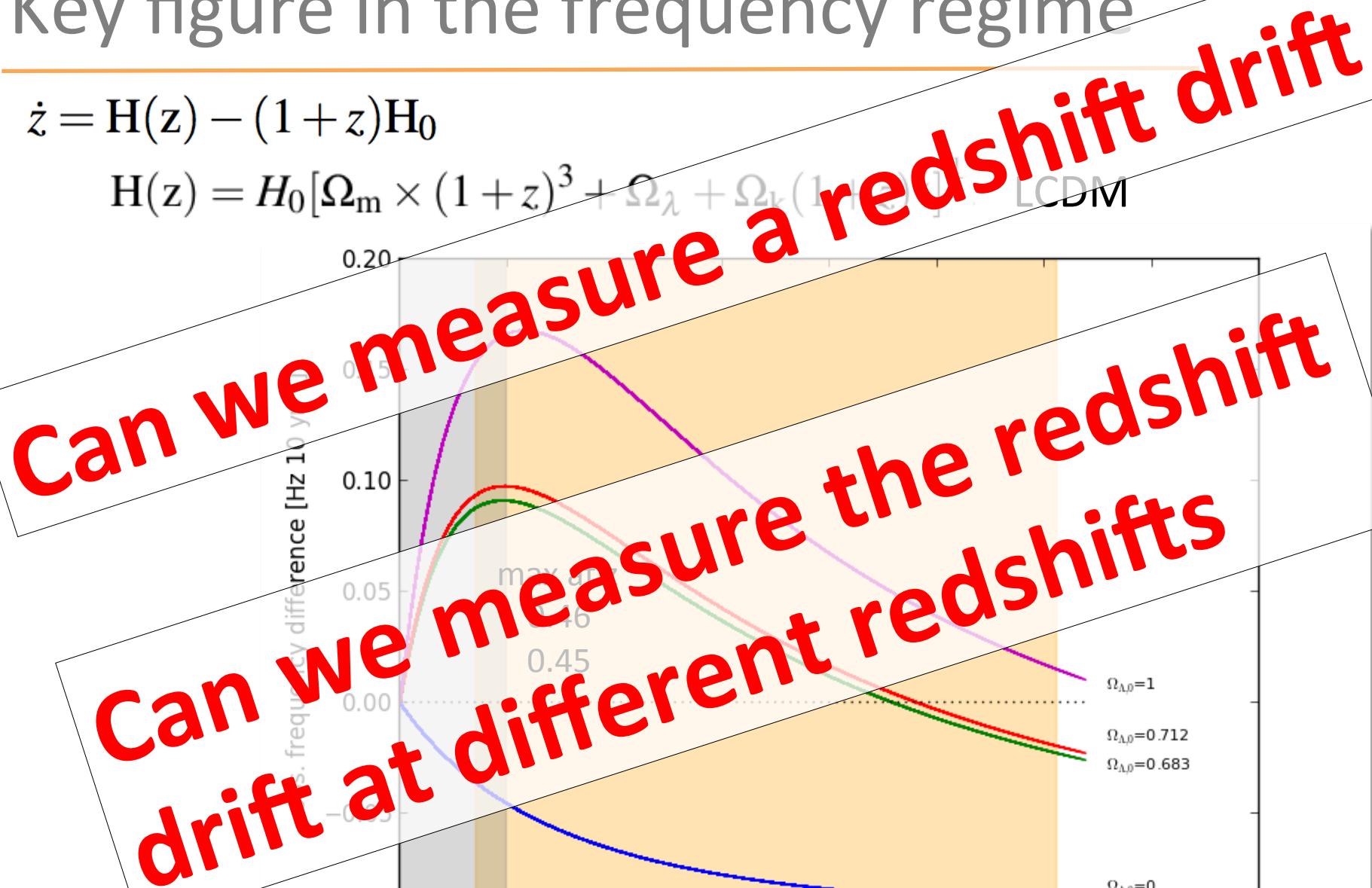
$$H(z) = H_0 [\Omega_m \times (1+z)^3 + \Omega_\lambda + \Omega_k (1+z)^2]^{1/2} \quad \text{LCDM}$$



Key figure in the frequency regime

$$\dot{z} = H(z) - (1+z)H_0$$

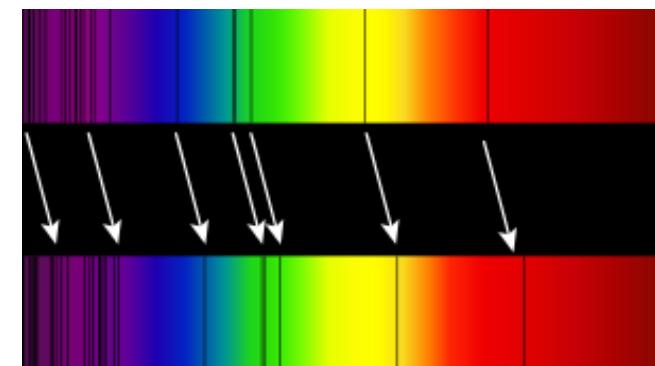
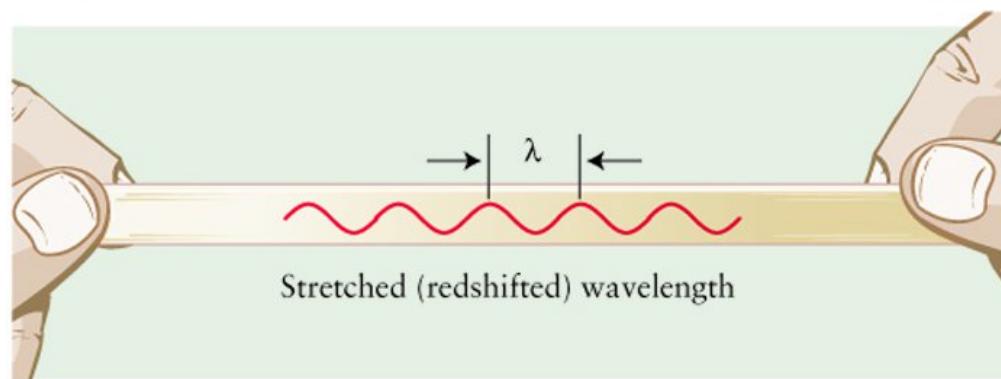
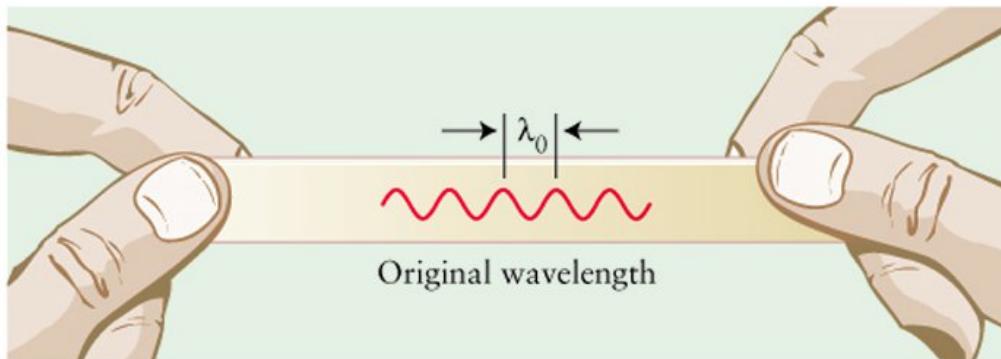
$$H(z) = H_0 [\Omega_m \times (1+z)^3 + \Omega_\lambda + \Omega_k (1+z)^2] \quad \text{LCDM}$$



cosmological redshift

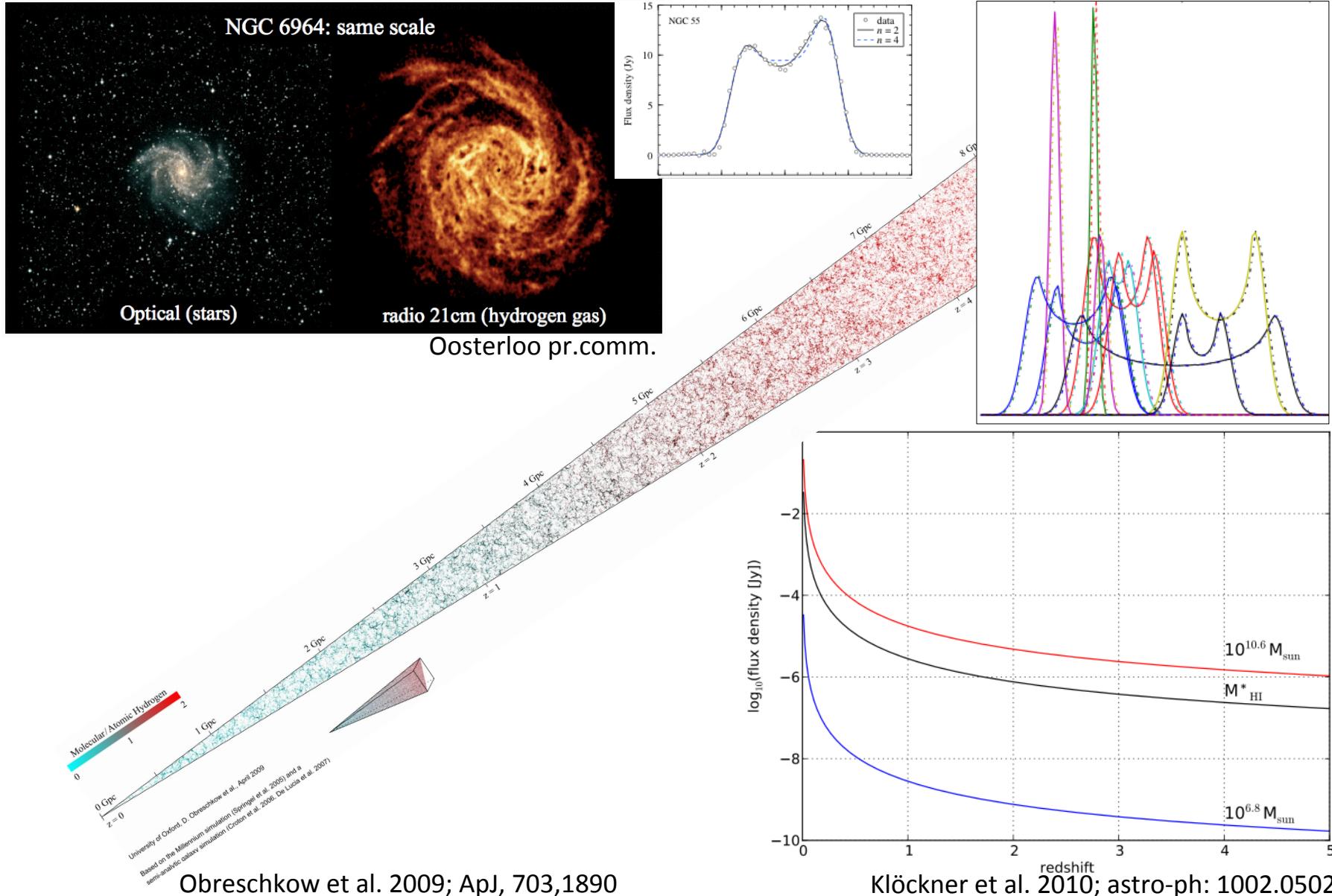
$$\frac{R(t_0)}{R(t_e)} = (1 + z)$$

$$ds^2 = (c dt)^2 - R^2(t) \left[\frac{dr^2}{1 - kr^2} + r^2(d\theta^2 + \sin^2 \theta d\phi^2) \right]$$



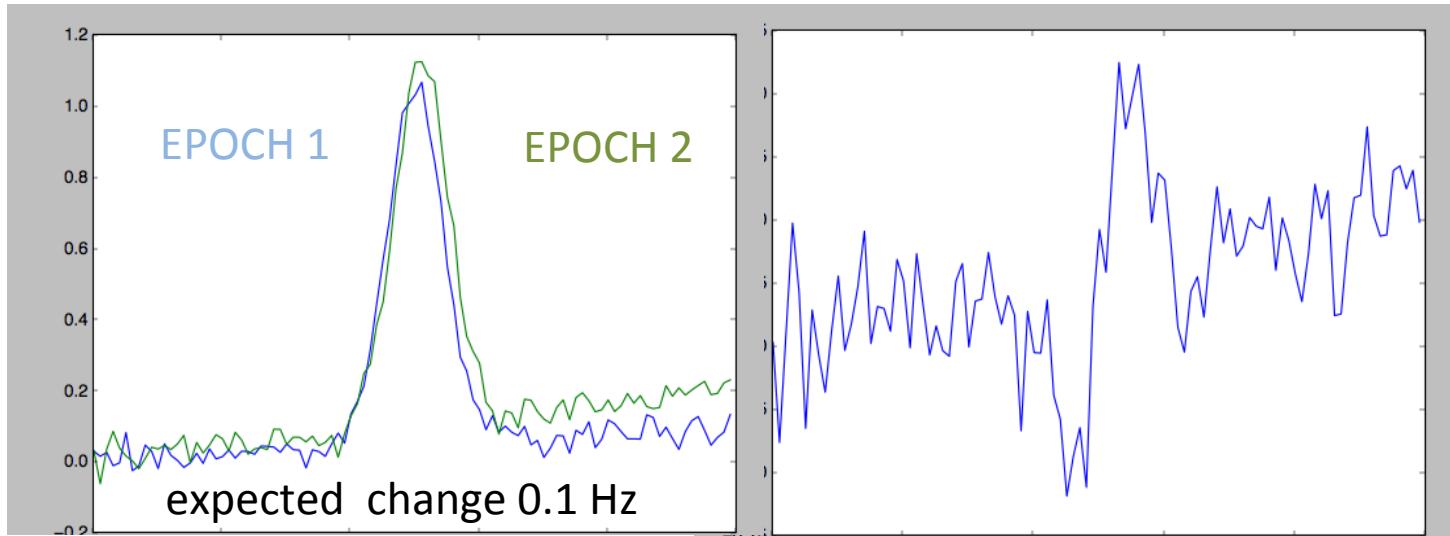
$$z = \frac{\nu_{\text{restframe}} - \nu_{\text{obs}}}{\nu_{\text{obs}}}$$

finding the right tracer

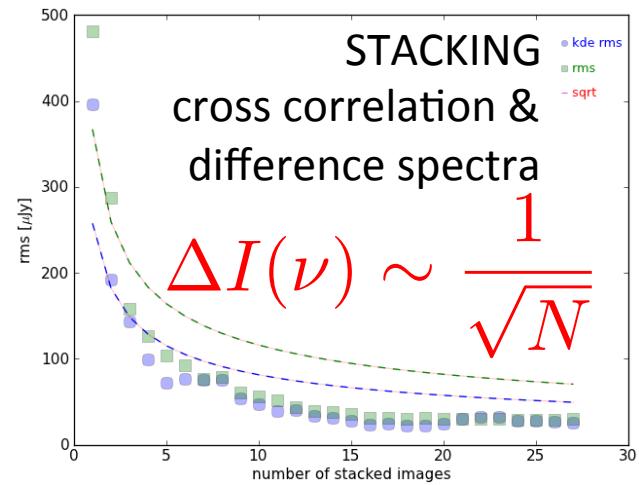
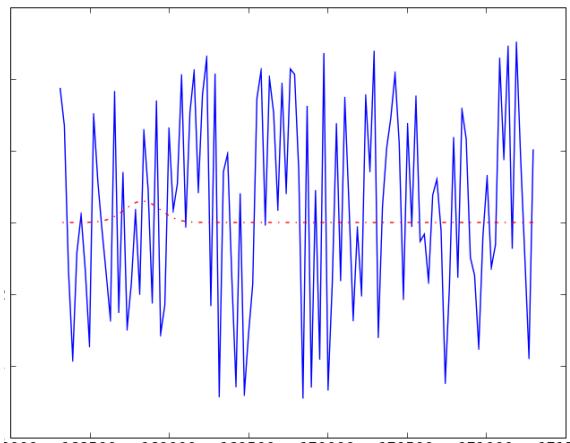


The basics - measure acceleration

back to the roots re-visiting Sandage idea in the radio regime



$$\Delta I = \frac{\text{SEFD}}{\eta_s \sqrt{t \Delta \nu}}$$
$$\text{SEFD} = \frac{2kT_{\text{sys}}}{A_{\text{eff}}}$$

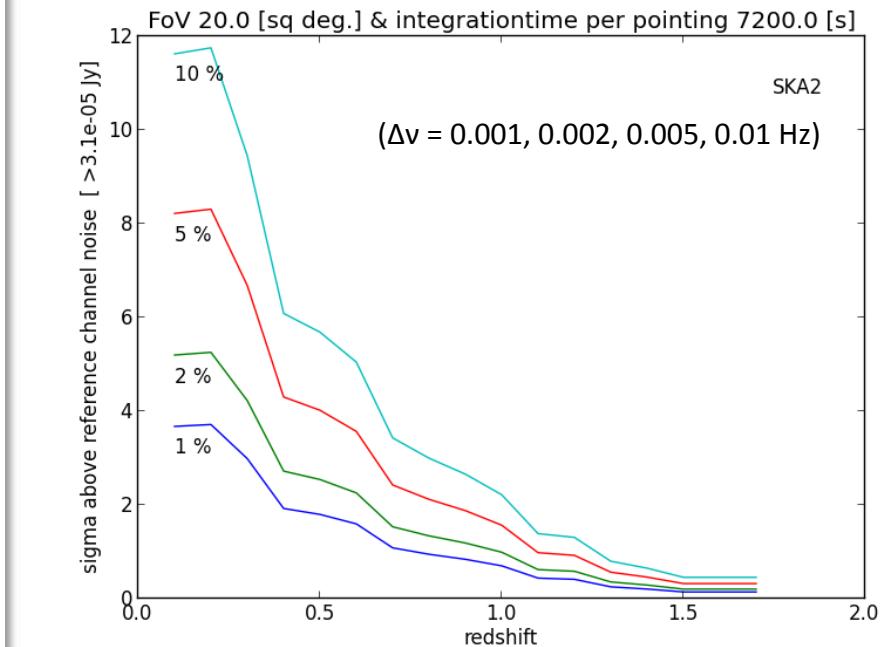
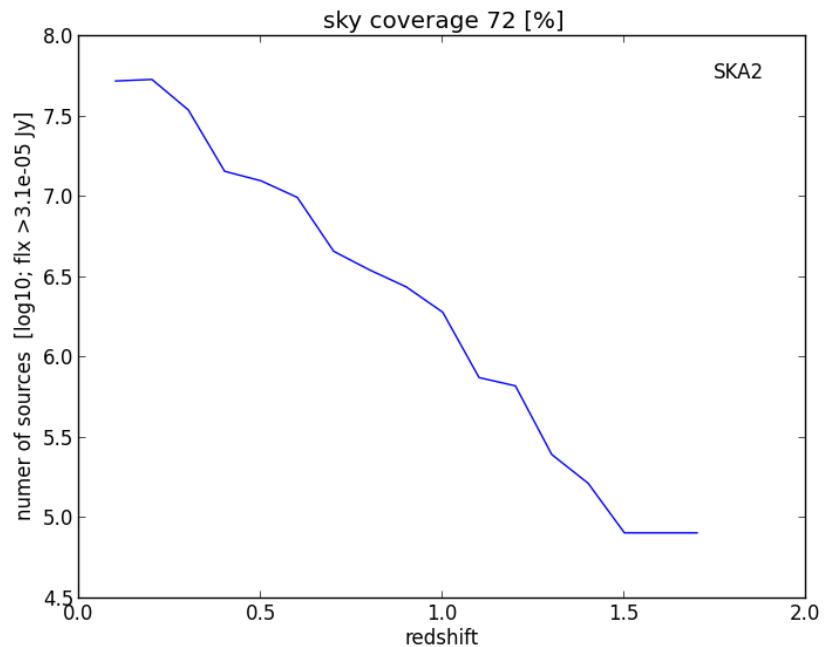


SKA baseline design 3.9 kHz \rightarrow 0.01 Hz
simulation are currently on its way

Measuring acceleration with the SKA

KEY is number density and channel resolution

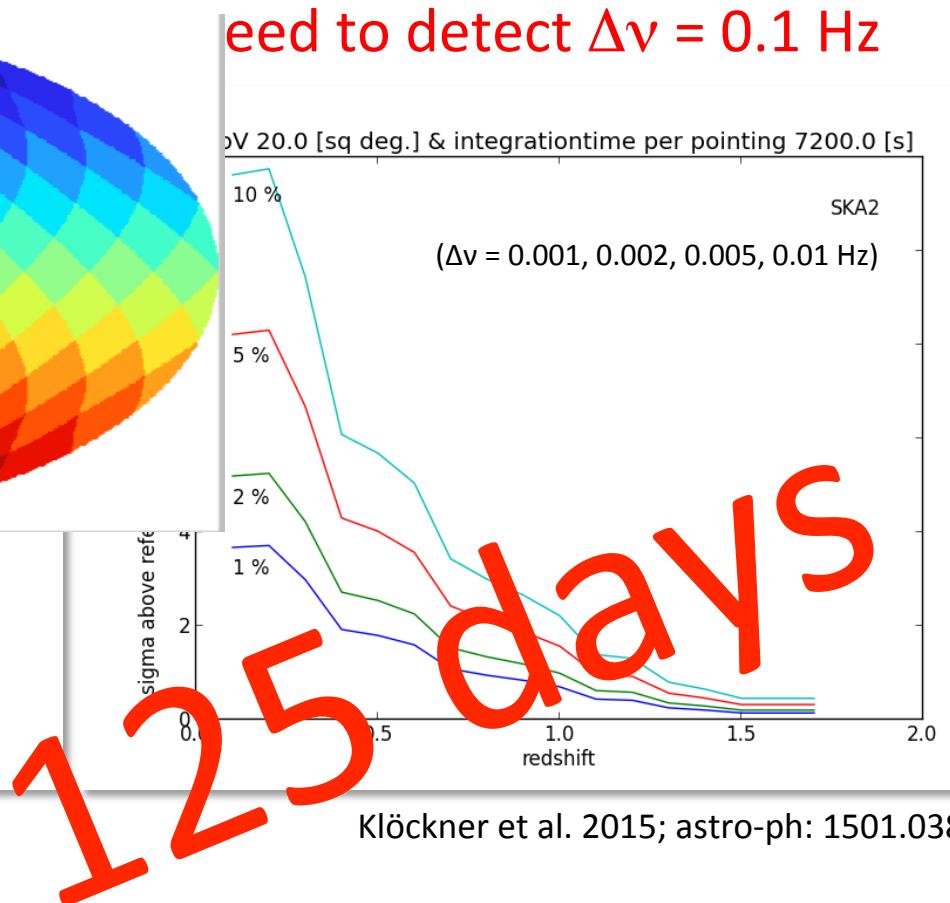
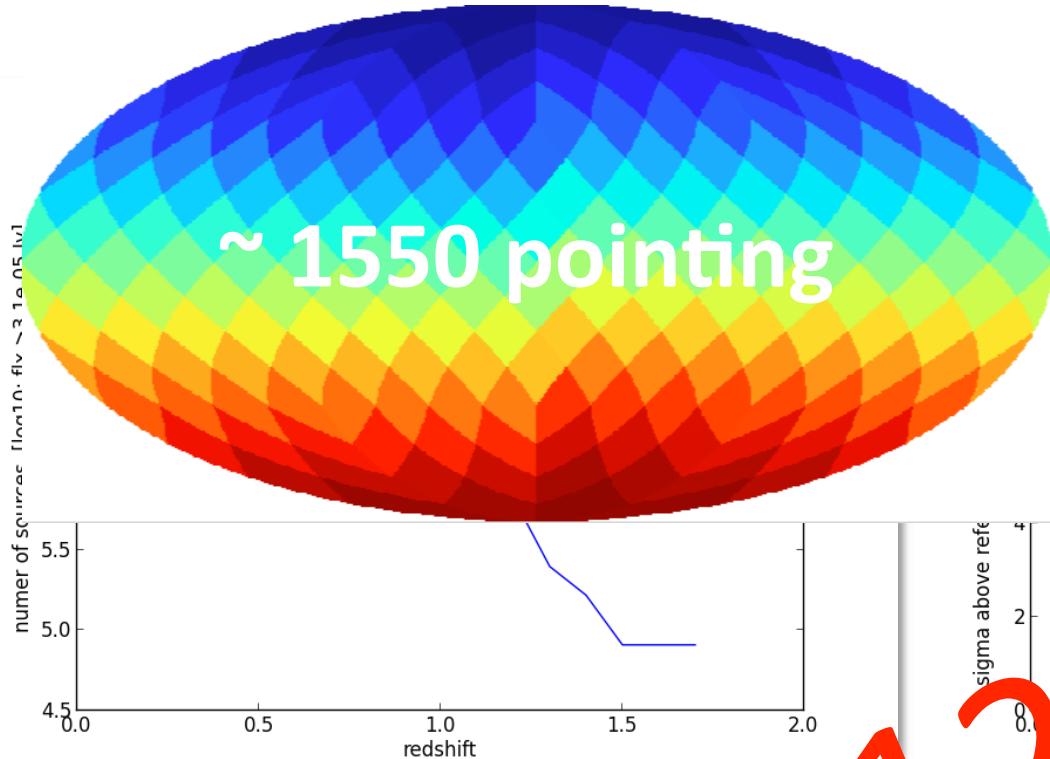
need to detect $\Delta v = 0.1$ Hz



Klöckner et al. 2015; astro-ph: 1501.03822

Measuring acceleration with the SKA

KEY is number density and channel resolution



The small print – handling the systematics



S³ The SKA Simulated Skies

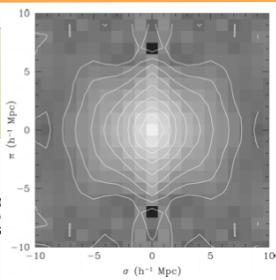
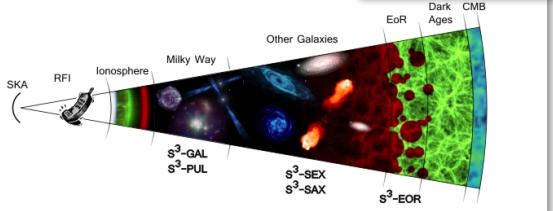
- Home
- S³-SEX
- S³-SAX
- S³-PUL
- S³-GAL
- S³-EOR
- S³-Tools
- Contacts
- Admin



Introduction

The SKA Simulated Skies (S³) are a set of computer simulations of the radio and optical skies dedicated to the preparation of the Square Kilometer Array (SKA) and its pathfinders. University of Oxford as part of the Square Kilometer Array Design Studies (SKADS Programme 6, contract #011938).

Simulation types



SAX simulation - Obreschkow et al. 2009

- Station / Source positions: different frames (ITRF, ICRF), motions
- Times: UTC; TAI, TT; UT1; TDB/TCB/TCG
- Orientation: Precession (50''/yr), Nutation (9.6'', 18yr), Polar Motion (0.6'', 1yr)
- Diurnal Spin: Oceanic friction (2ms/cy), CMB (5ms, dcds), AAM (2ms, yrs)
- Tides: Solid-earth (30cm), Pole (2cm)
- Loading: Ocean (2cm), Hydrologic (8mm), Atmospheric (2cm), PGR (mm's/yr)
- Antennas: Axis offset, Tilt, Thermal expansion
- Propagation: Troposphere (dry [7ns], wet [0.3ns]), Ionosphere
- Relativistic $\tau(t)$ calculation: Gravitational delay, Frame choice/consistency

stolen from B. Cambell

mm < 3 cm/s 10/yr



Jet Propulsion Laboratory
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+ Near-Earth Object (NEO) Program

JPL HOME

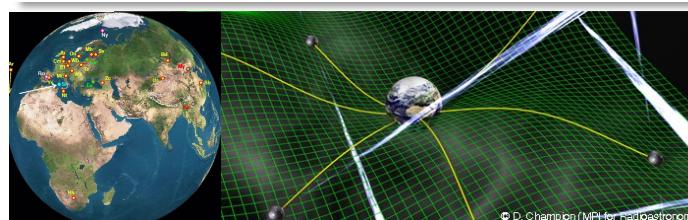
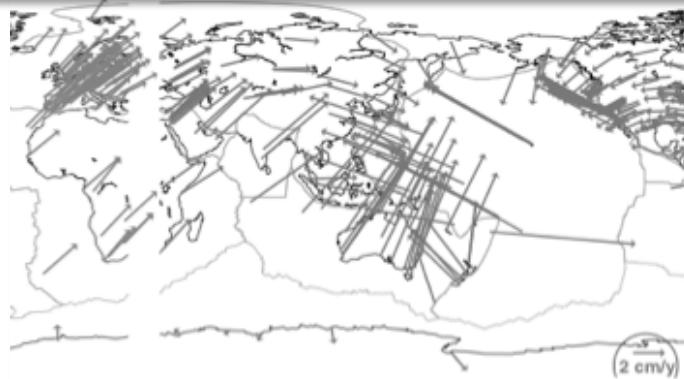
EARTH

SOLAR SYSTEM

STARS & GALAXIES

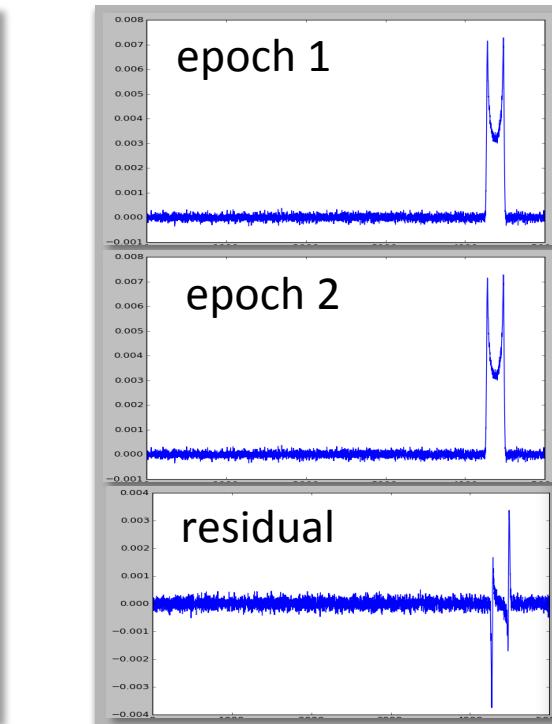
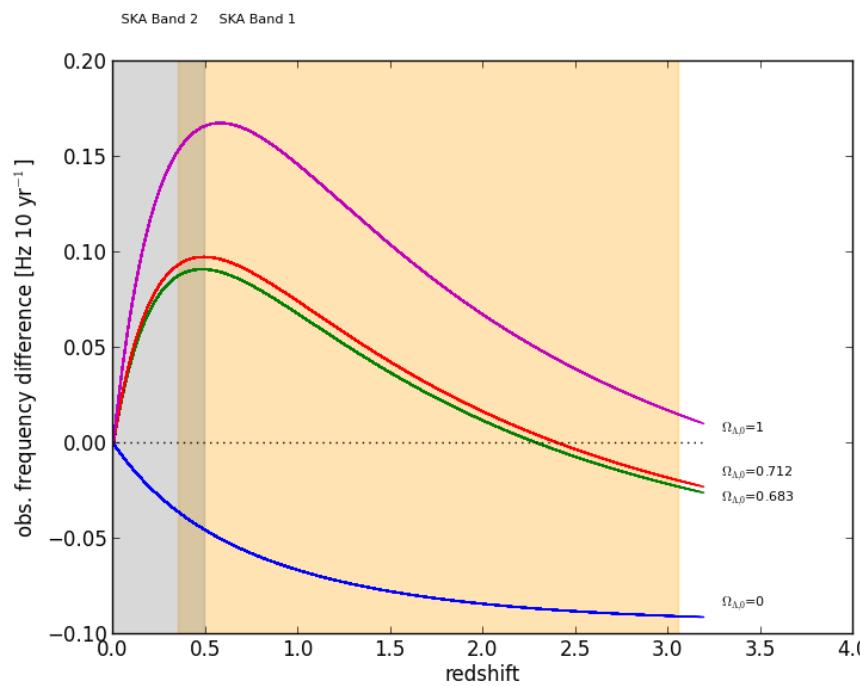
Solar System Dynamics

| BODIES | ORBITS | EPHEMERIDES | TOOLS | PHYSICAL DATA | DISCOVERY |
|--------------------------|--------|---------------------------------|-------|---|-----------|
| Rest Frame | | Corrected for | | Amplitude of Correction, km s ⁻¹ | |
| Geocentric | | Earth rotation | | 0.5 | |
| Earth-Moon Barycentric | | Effect of the Moon on the Earth | | 0.013 | |
| Heliocentric | | Earth's orbital motion | | 30 | |
| Solar System Barycentric | | Effect of planets on the Sun | | 0.012 | |
| Local Standard of Rest | | Solar motion | | 20 | |
| Galactocentric | | Milky Way rotation | | 230 | |
| Local Group Barycentric | | Milky Way motion | | ~100 | |
| Virgo-centric | | Local Group motion | | ~300 | |
| Microwave Background | | Local Supercluster motion | | ~600 | |



Real time cosmology - summary

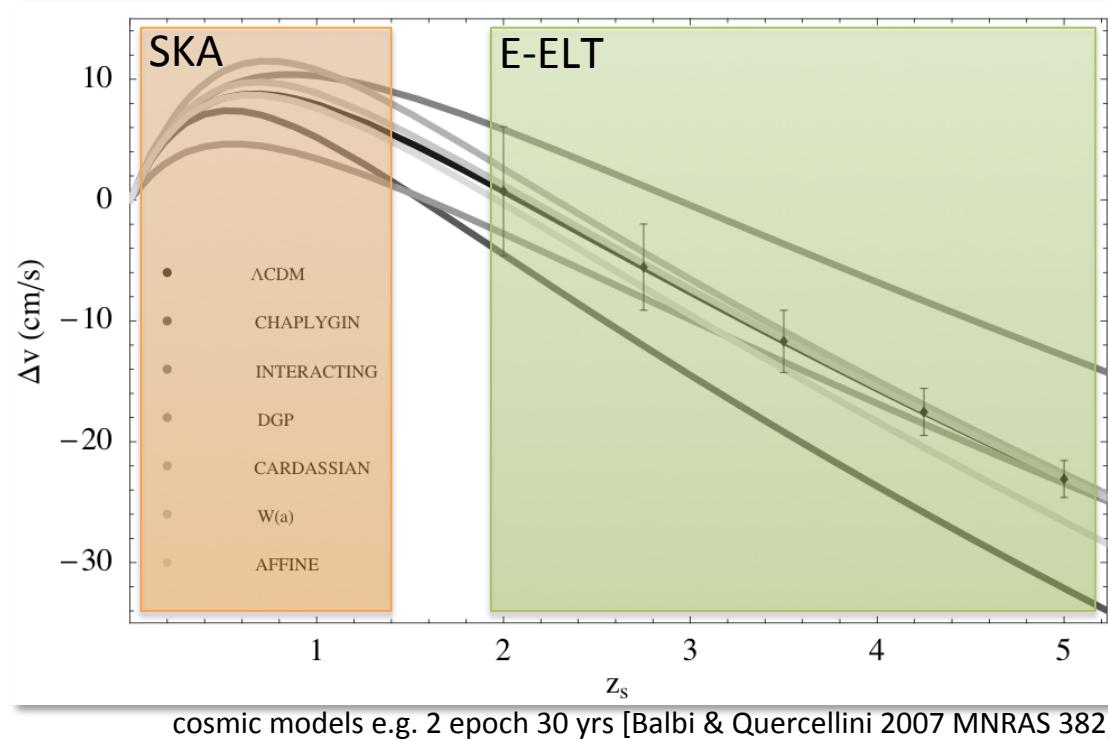
- Cosmic acceleration, a test of dark energy, will change the redshift of individual galaxies within 10 years by $dz/dt \sim 10^{-10}$. Stacking the residual or cross-correlated spectra of 10^{7-9} individual galaxies can measure this offset of ~ 0.1 Hz.



- This real time cosmology experiment does require some changes to the baseline design (e.g. channel resolution of ~ 0.001 Hz) and will challenge the overall system stability. Need at least 10^8 channels to cover HI line width (fan-out mode of sub-bands).

Real time cosmology - summary

- The SKA provides us with a model independent measure of the dynamics of the cosmic expansion within 10 years.
- The SKA will be the **only** observatory probing the cosmic acceleration and its dependency on redshift by facilitating a 3π HI galaxy survey.
- The SKA will break degeneracy with other probes allowing to check

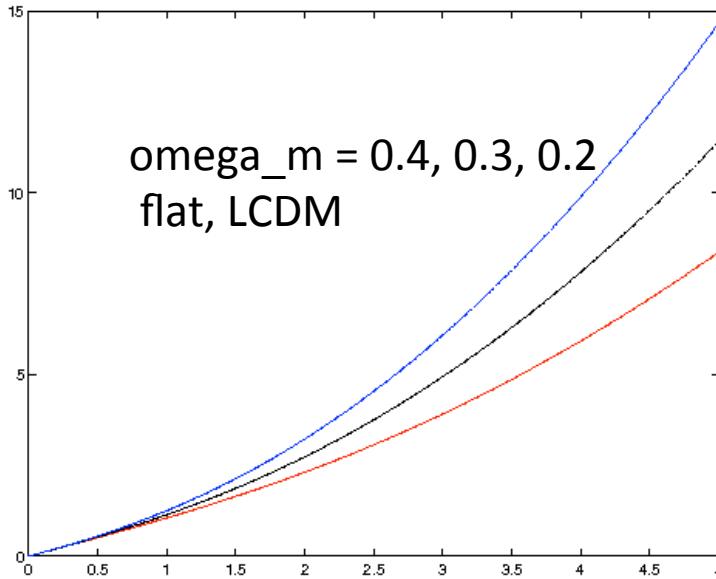


Real time cosmology - summary

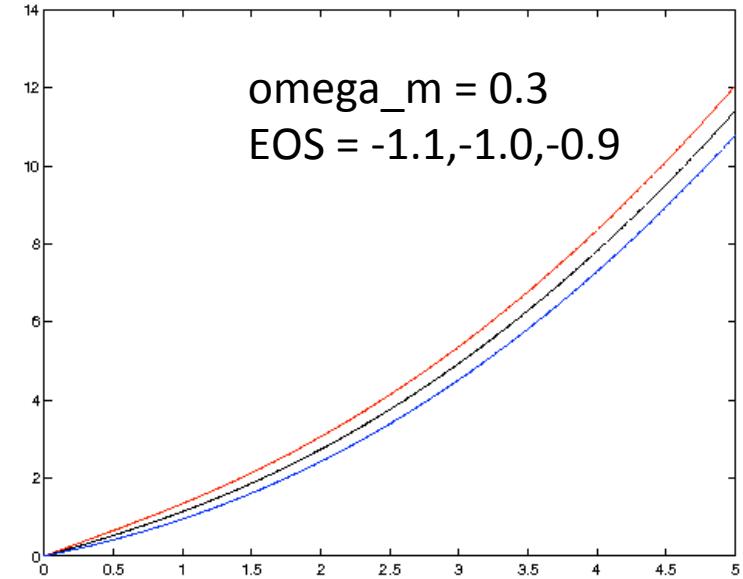
- SKA lifetime will be 50 year, e.g. redshift drifts and its evolution in time, alone and in combination with other experiments enables us to constrain cosmological models.

impact on the equation of state

cosmic jerk – term versus redshift



cosmic jerk – term versus redshift



test for modified cosmological