

Highlights of the updated hadronic interaction models DPMJET-III and SIBYLL

HAP Workshop | Non-Thermal Universe

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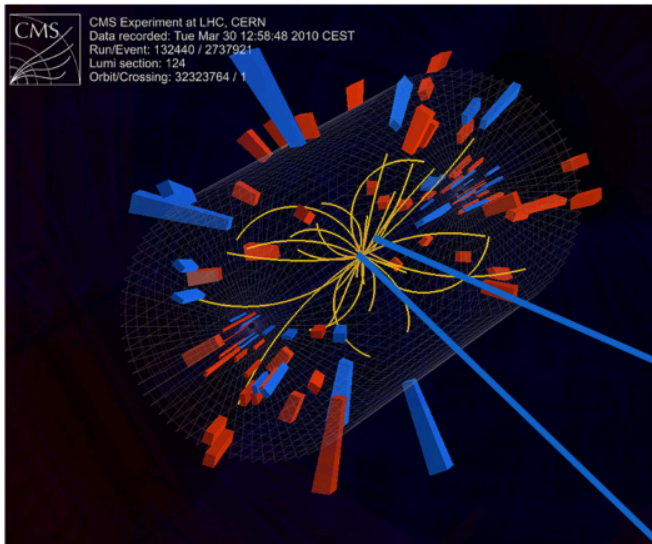
[Felix Riehn](#), Bartol Research Institute, Delaware

[Ralph Engel](#), KIT

together with Thomas K. Gaisser and Todor Stanev

Hadronic interaction models

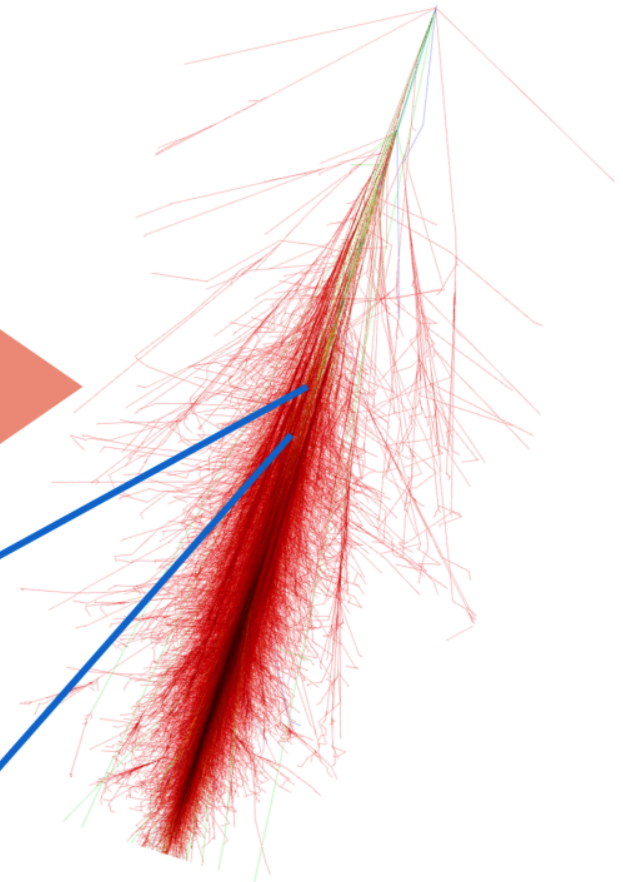
Interpretation of cosmic ray measurements **crucially** depends on hadronic interaction models



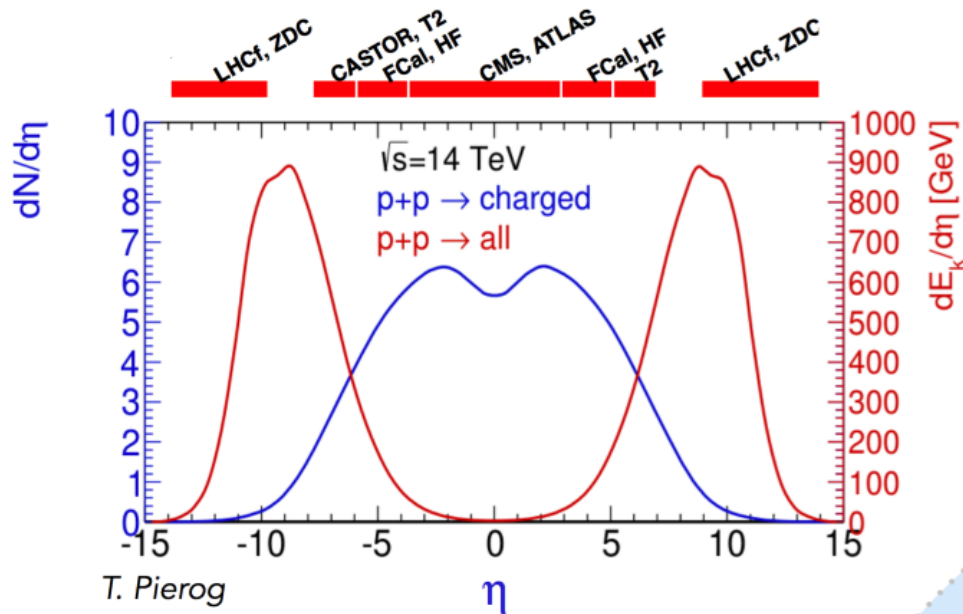
Detectors have holes!



DPMJET
EPOS
SIBYLL
QGSJet



Limited experimental coverage of cosmic ray relevant physics



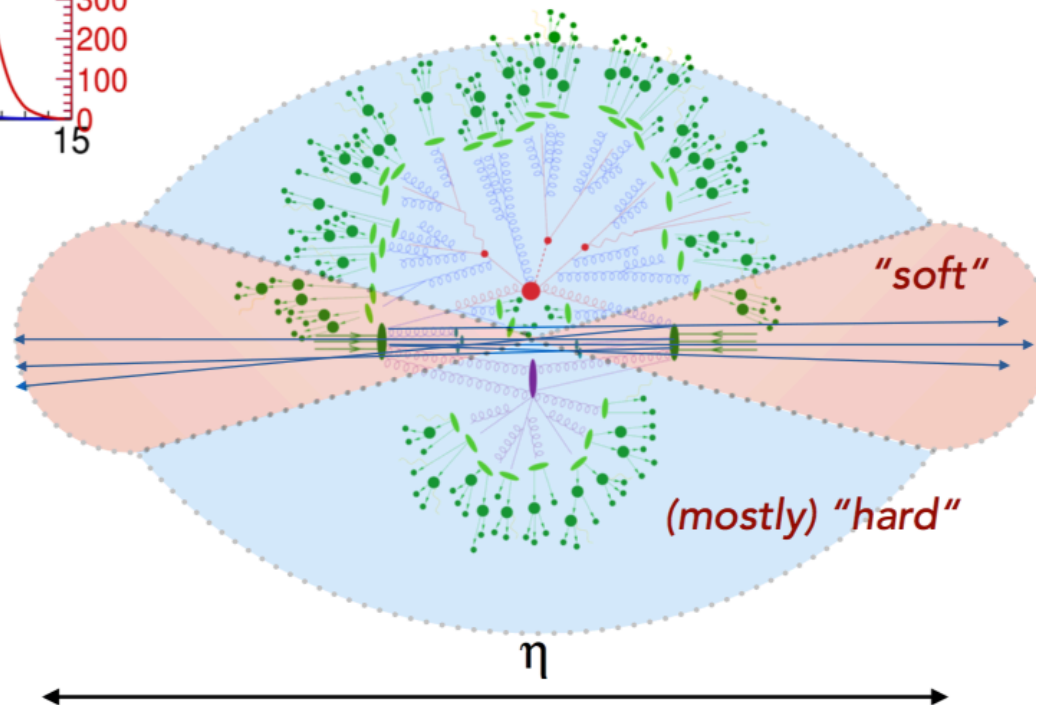
Typical acceptance at LHC

$$x_F = 2p_Z/\sqrt{s} \ll 0.1$$

Most energy is dissipated in not instrumented region!

Diagram illustrating the definition of pseudorapidity η in terms of the angle θ relative to the beam axis. The formula is $\eta = -\log \tan(\frac{\theta}{2})$.

p_T



The nuclear Monte Carlo event generator DPMJET-III

> DPMJET-III

- supports hadron-hadron, photon-hadron, photon-photon, hadron-nucleus and nucleus-nucleus
- from few GeV (lab) up to UHECR energies
- light and heavy nuclei ($>^{235}\text{U}$) via Glauber, intra-nuclear cascade,...
- is the main hadronic model for high energies and nucleus-nucleus interactions in FLUKA

> Individual hadron-/photon-nucleon interaction modeled with recently updated PHOJET minimum-bias model

> Under test in upcoming versions of FLUKA, CORSIKA, MCEq

> ... soon available as standalone at <https://dpmjetiii.hepforge.org>



DPMJET-III cascade mode

- Accepts most of the hadrons as projectiles
- Initialization in run-time, on demand
- Projectiles without known cross sections mapped to closest related particle combination (with valence exchange)
- Low minimal energy (no low-E model necessary)
- Relatively fast (wrt. EPOS-LHC or QGSJET-II-04)



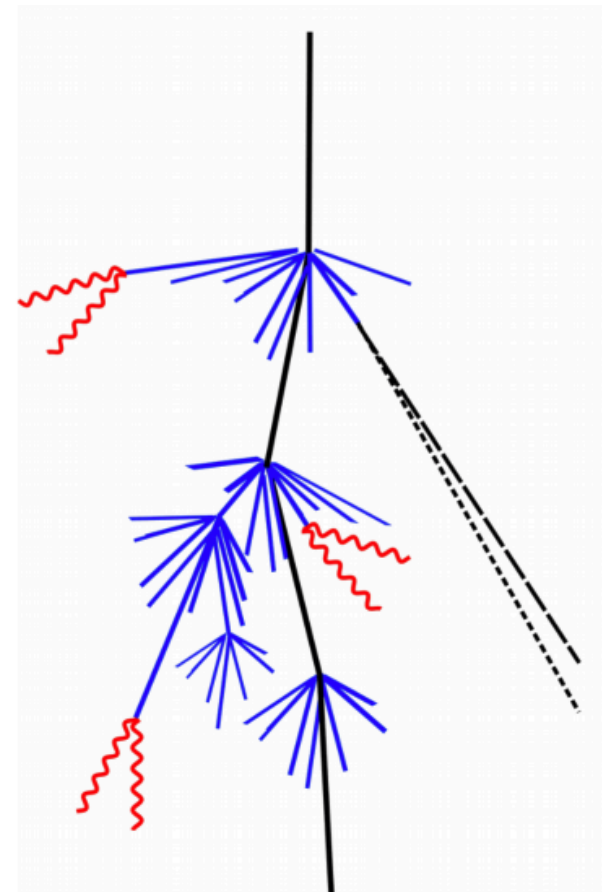
Can be used in air-shower Monte Carlo (CORSIKA)



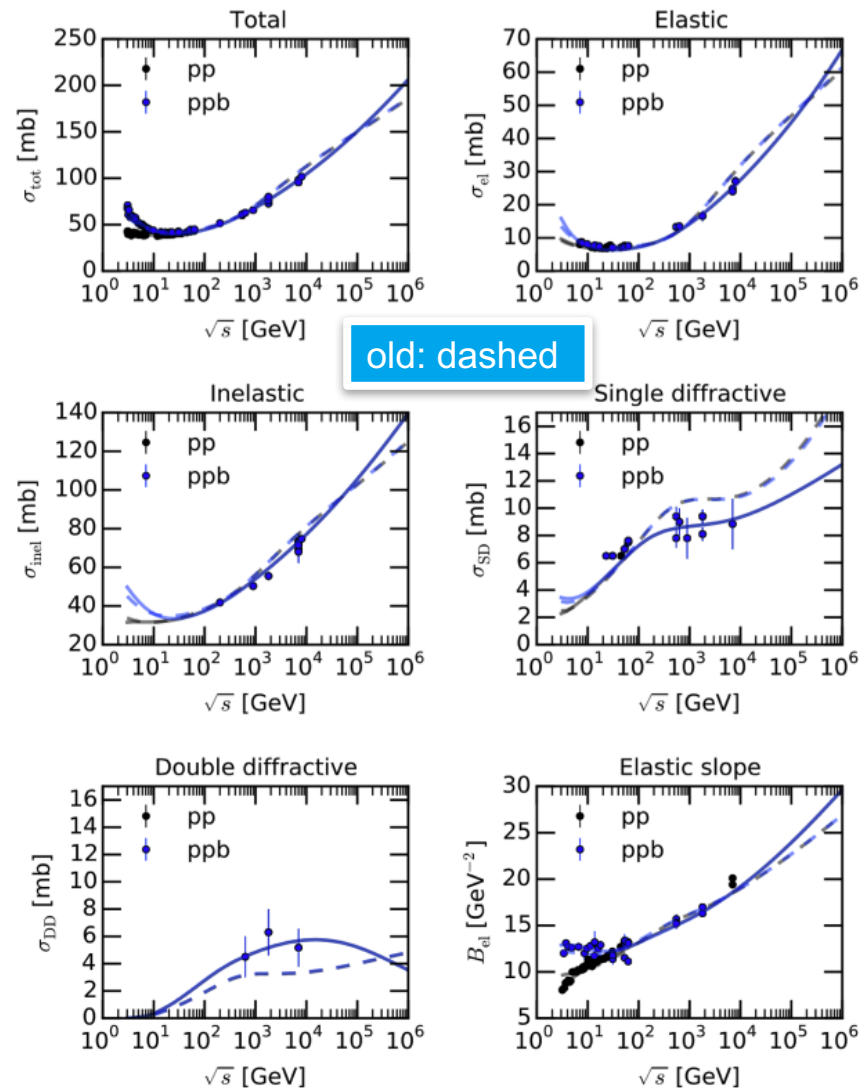
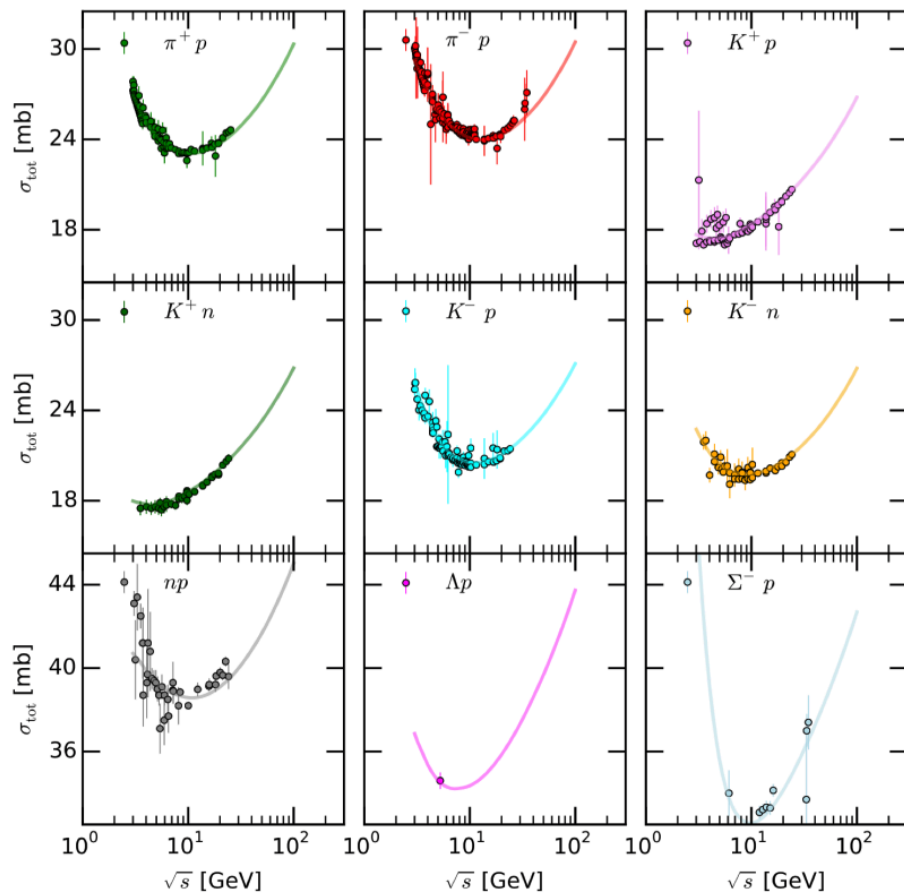
Can be used in FLUKA without workarounds



Could be implemented in, e.g., GEANT 4

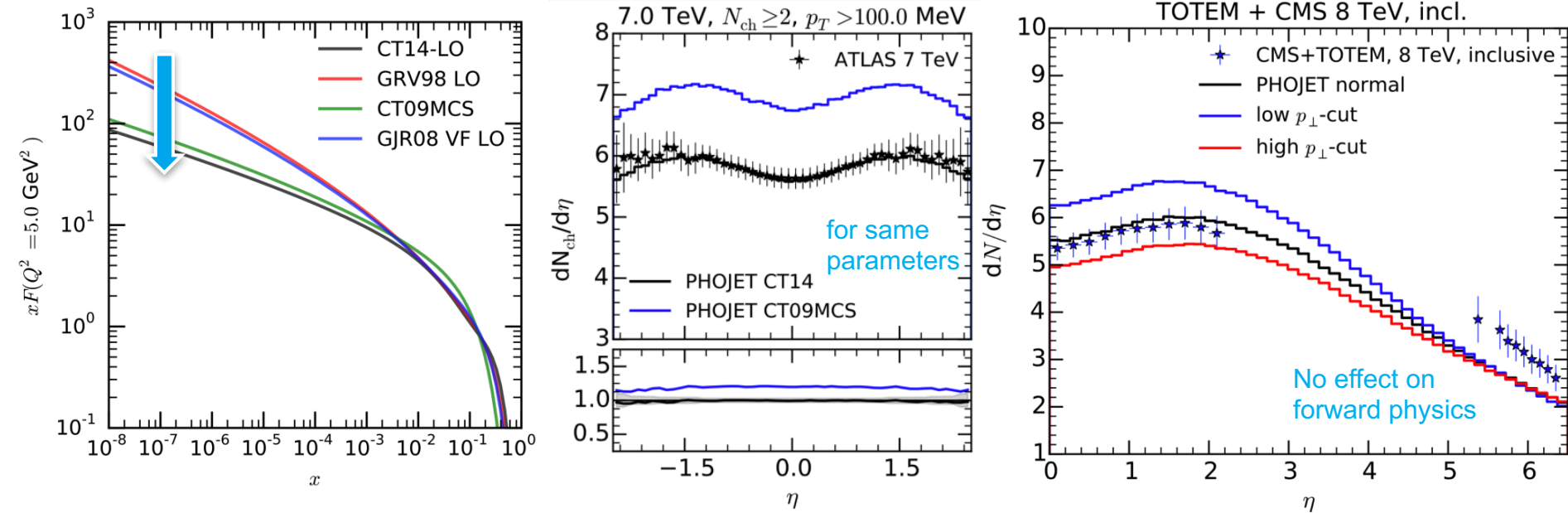


Interaction cross sections



LHC data restricts high energy behavior

Modern parton distribution functions



Hard cross section:

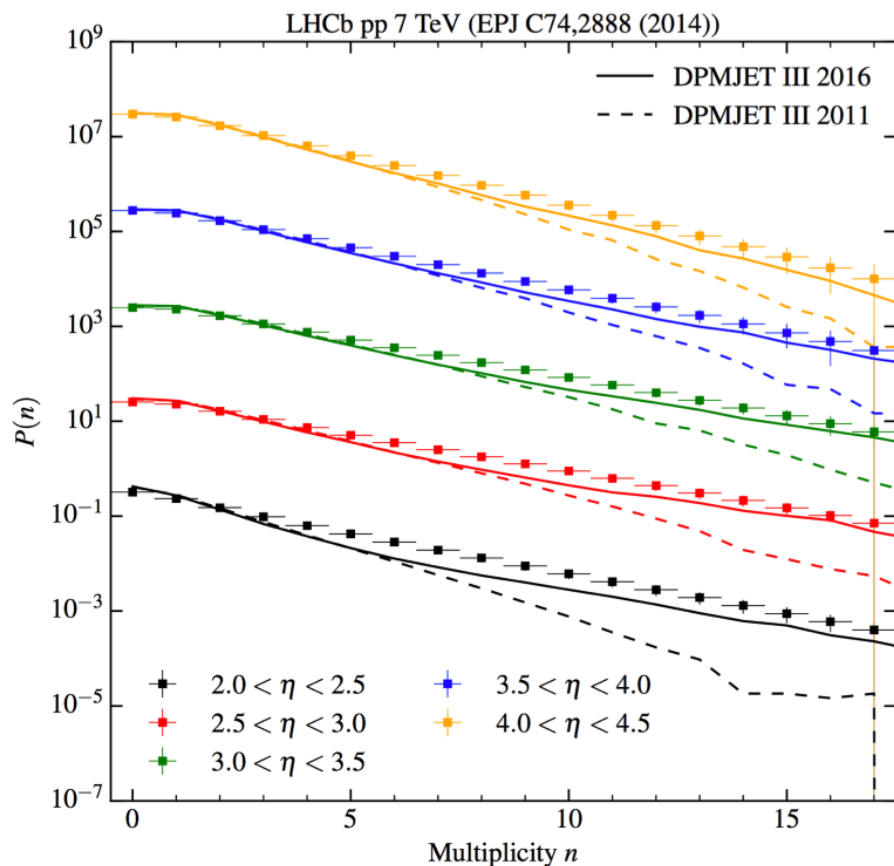
$$\sigma_{QCD} = \sum_{i,j,k,l} \frac{1}{1 + \delta_{kl}} \int dx_1 dx_2 \int_{p_{\perp}^{\text{cutoff}}} dp_{\perp}^2 f_i(x_1, Q^2) f_j(x_2, Q^2) \frac{d\sigma_{i,j \rightarrow k,l}}{dp_{\perp}}$$

- CT14LO parton distribution functions for hard scatterings (instead of GRV94/98)
- Significantly improved description of LHC distributions

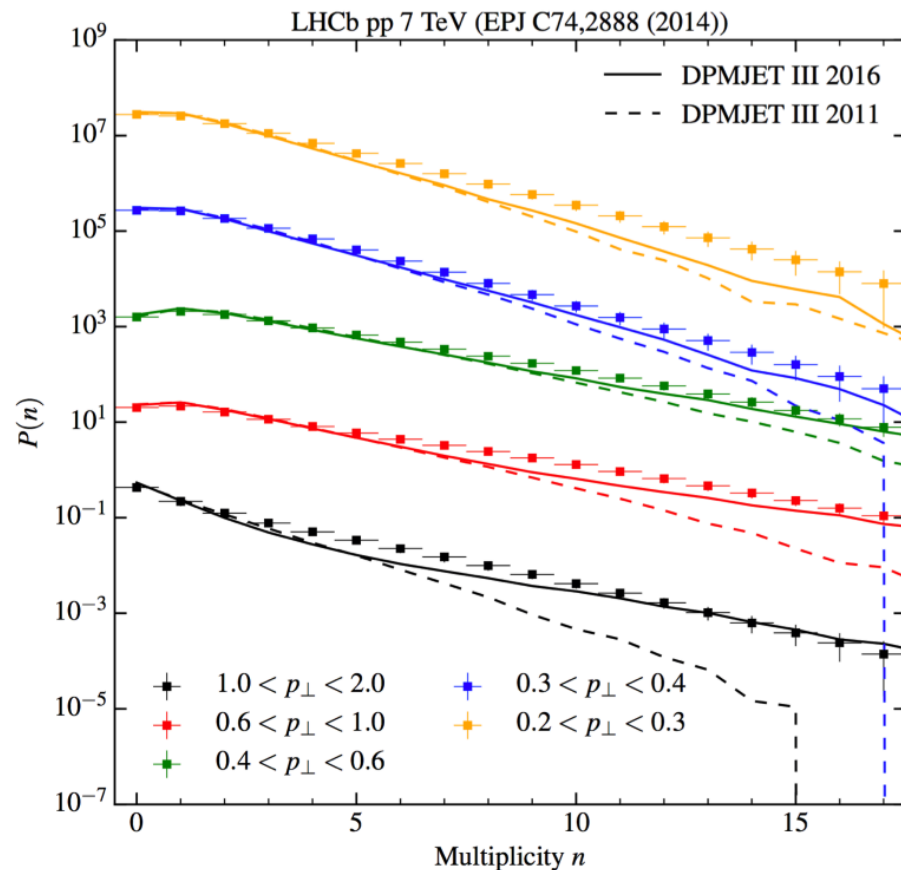


Improvement at LHC energies

angular distributions



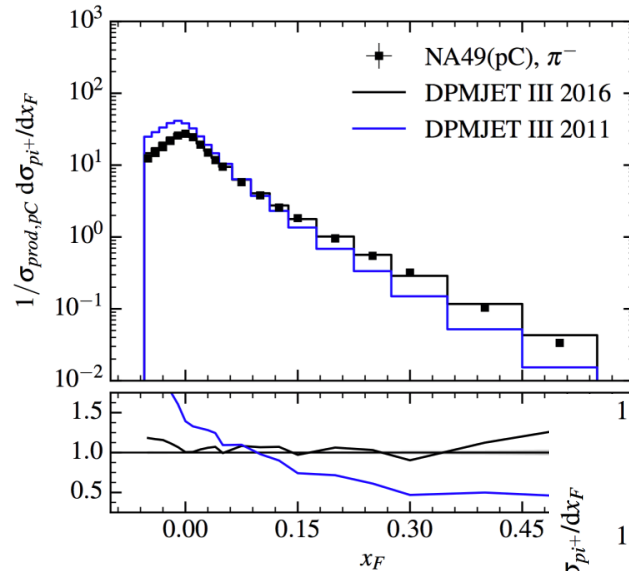
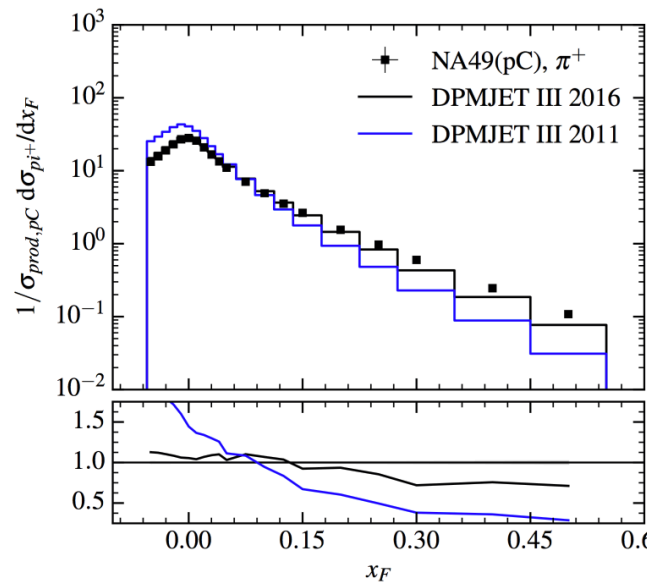
p_T distributions



...also for many other measurements

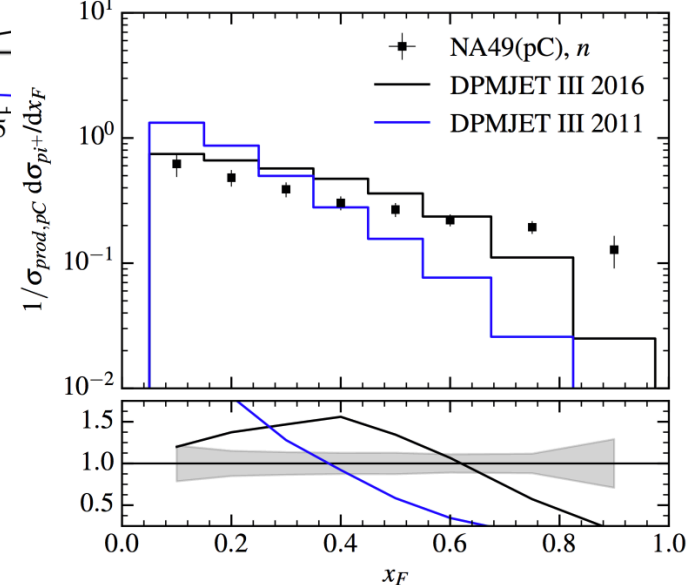


DPMJET at fixed-target energies



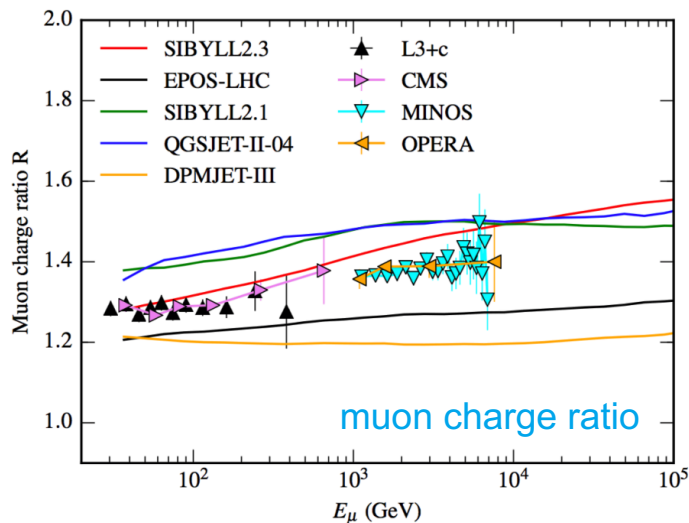
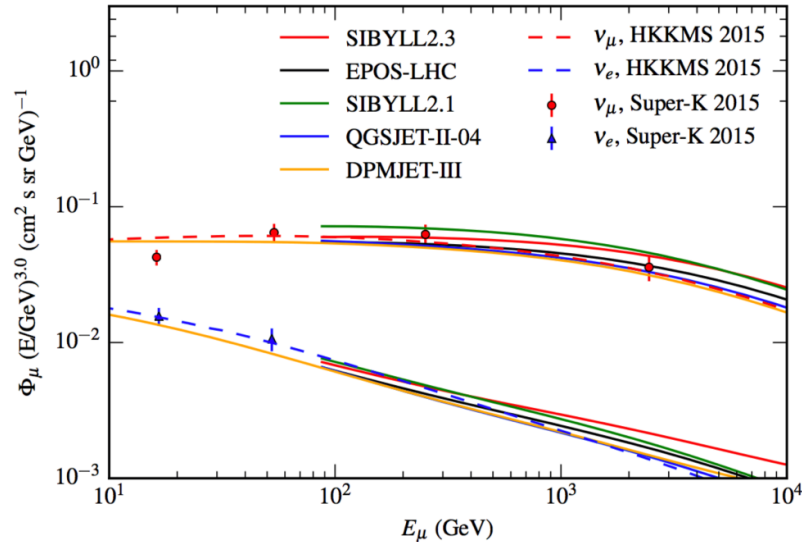
proton-Carbon
 $E_{lab} = 158 \text{ GeV}$

- Incremental improvement for $p, n, \pi, K, \nu, \Lambda$
- Baryon spectra as problematic as in SIBYLL 2.1
- “Soft” interactions need more work



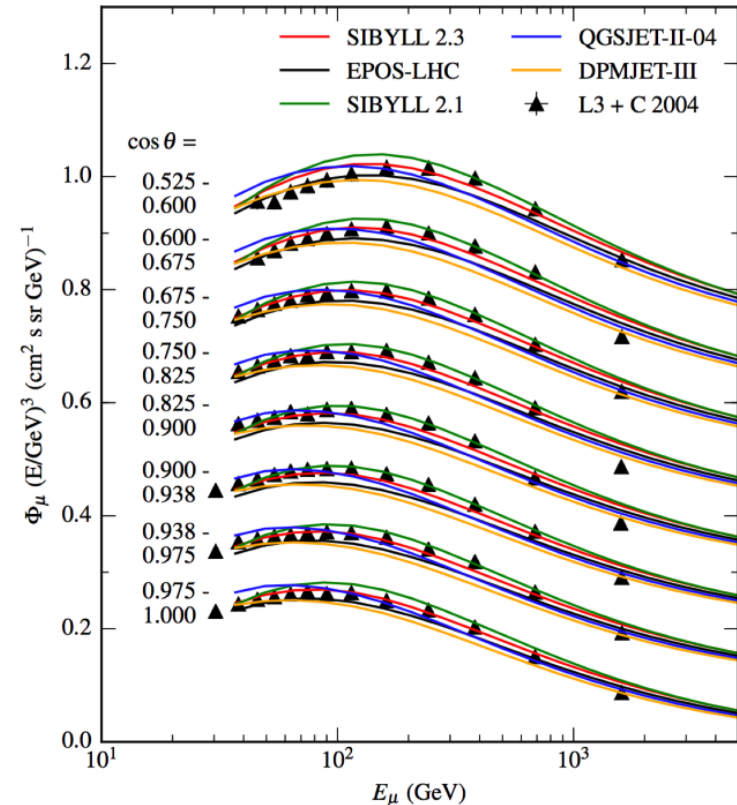
Inclusive lepton fluxes with DPMJET-III and SIBYLL

atmospheric neutrinos



muon charge ratio

atmospheric muons

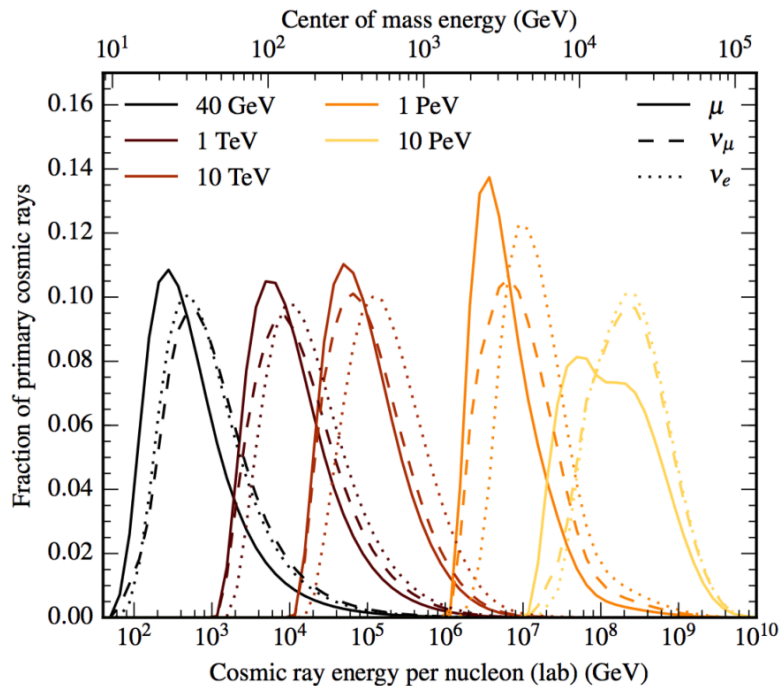


- > Just the beginning of CR studies with DPMJET
- > No air-shower predictions, yet

Explore energy extrapolation via CR “beam”

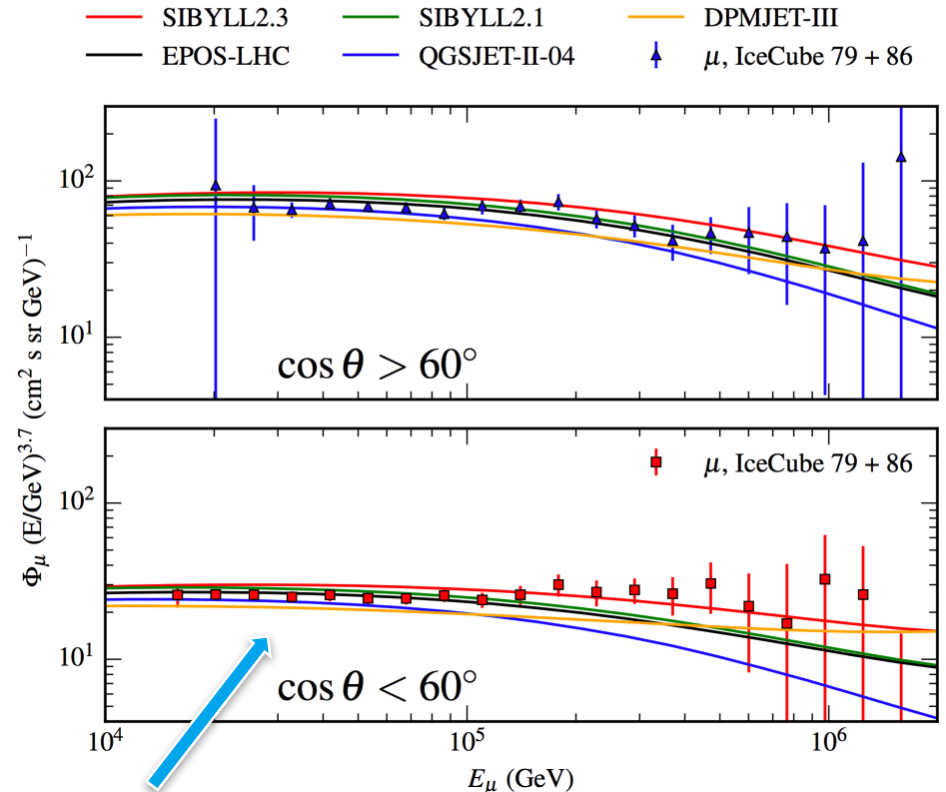
IceCube, Astropart. Phys. 78, 1 (2016)

Relation between primary nucleon energy and inclusive lepton energy



LHC energy probed by μ & ν around 100 TeV

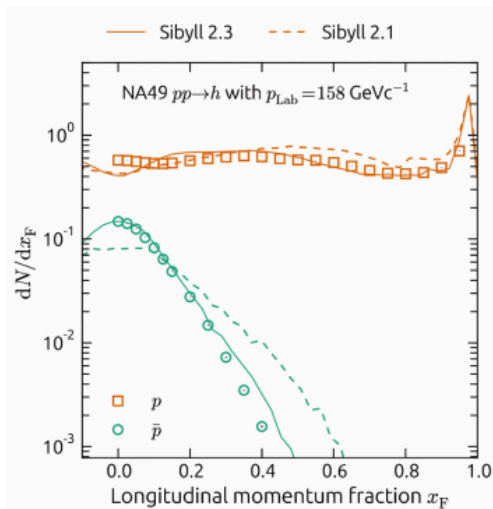
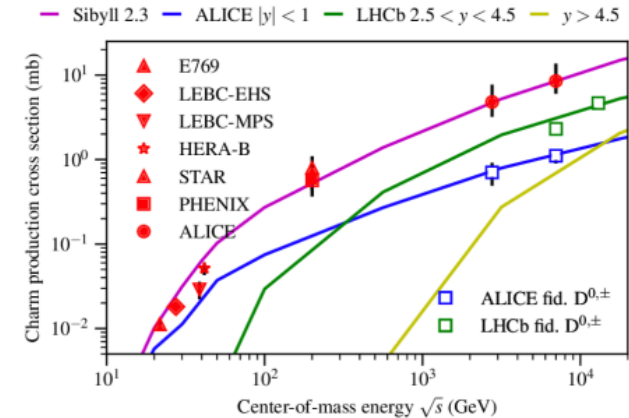
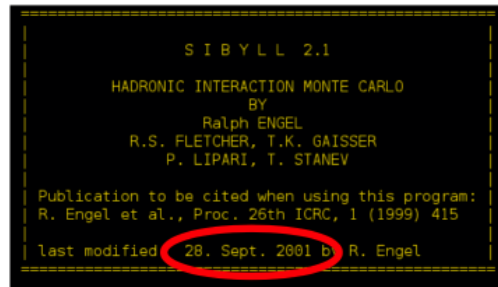
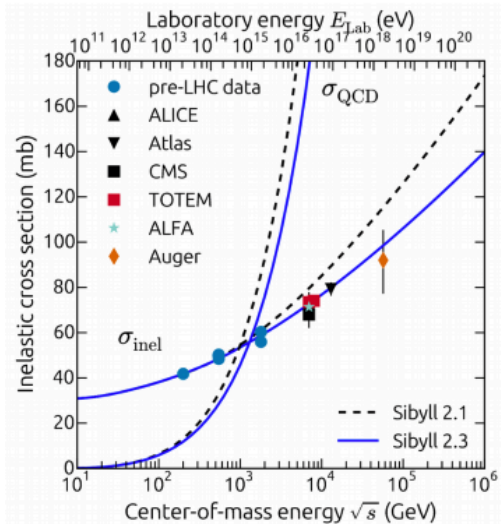
Muon flux measurement by IceCube



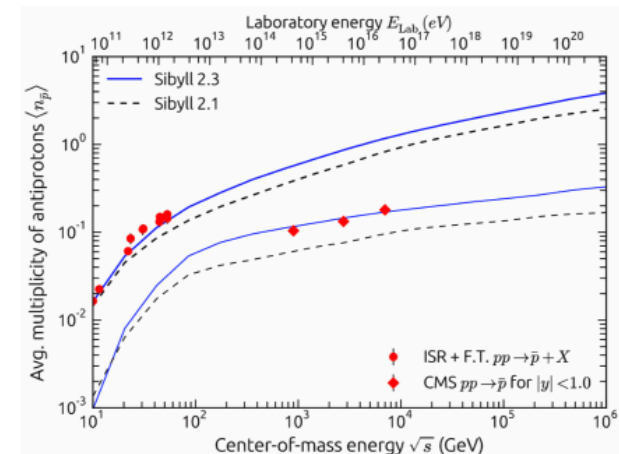
Forward particle physics beyond reach of current detectors at accelerators



Improvements in SIBYLL 2.3 at a glance



- Interaction cross sections
- Charm production
- Leading particles
- Baryon production
- ...

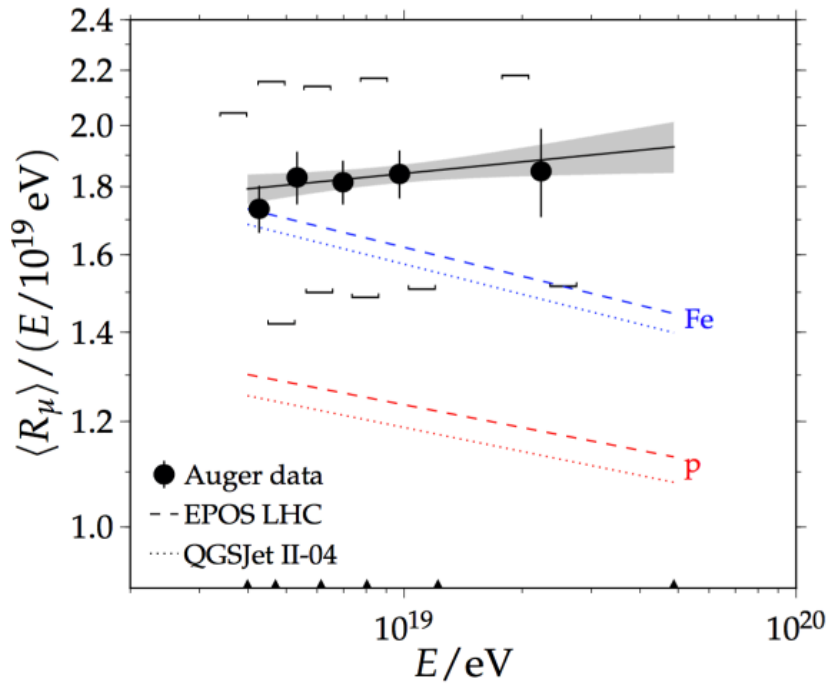


By Felix Riehn & Ralph Engel!



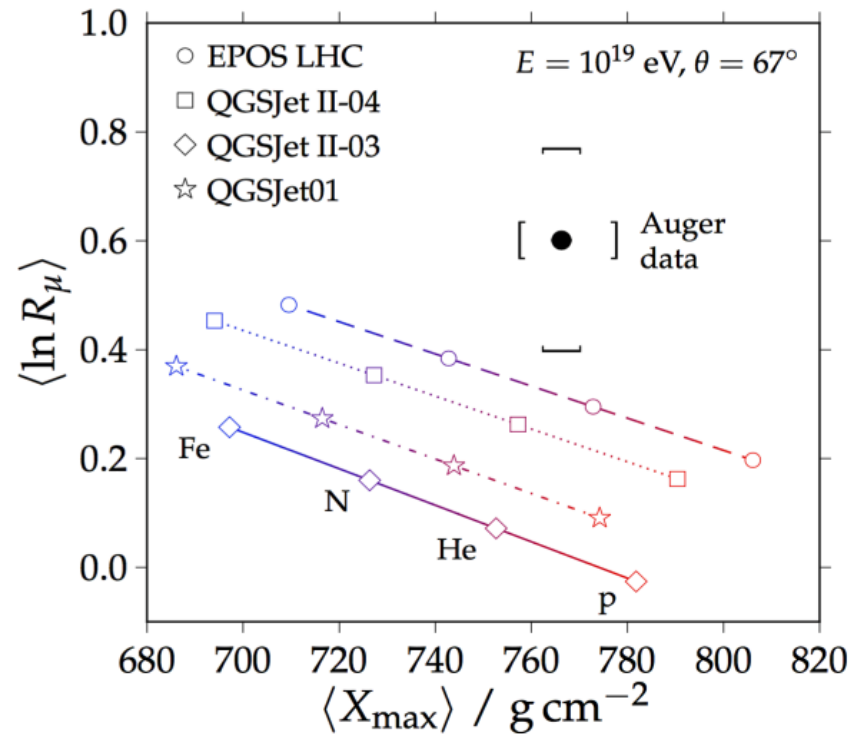
The “muon mystery”

muon number in inclined showers $\theta > 60^\circ$



Something is really inconsistent
[hadronic interaction models]

Combination of information on mean X_{max} and muon number at ground

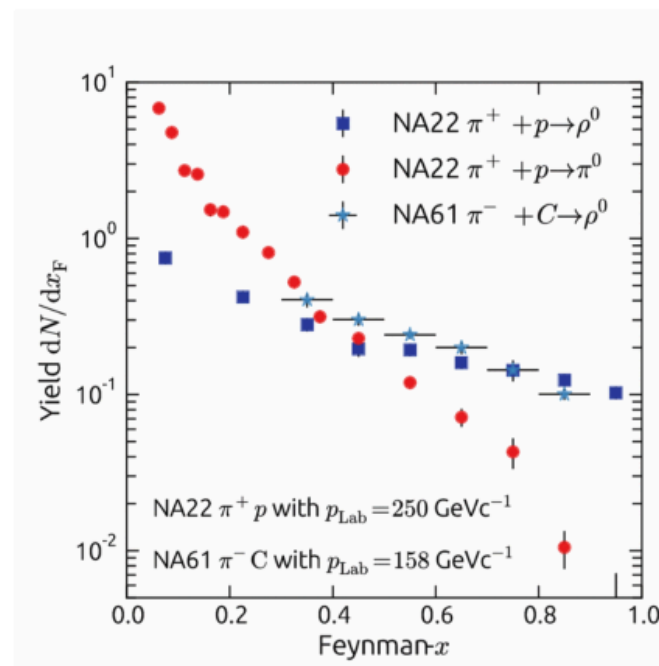
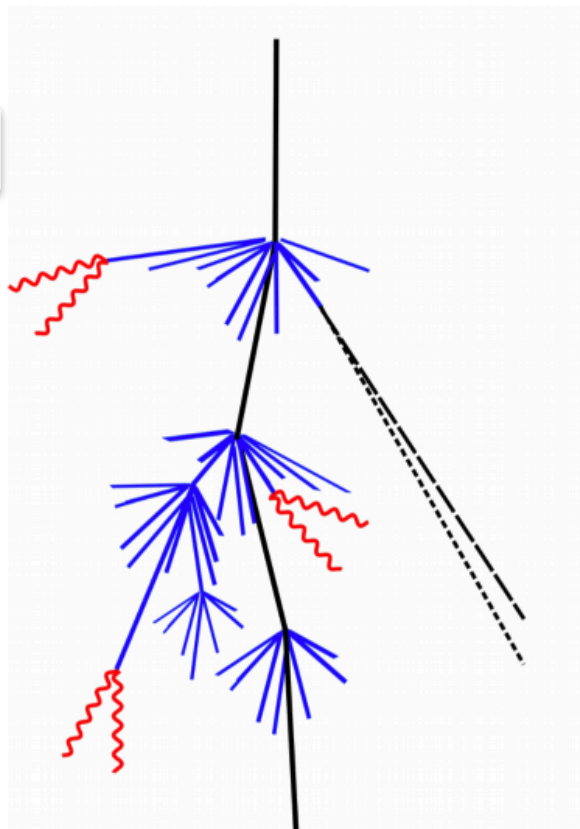


More (low-energy) muons?

Pion interactions!

$$\pi^{\pm} + \text{Air} \rightarrow \rho^0$$

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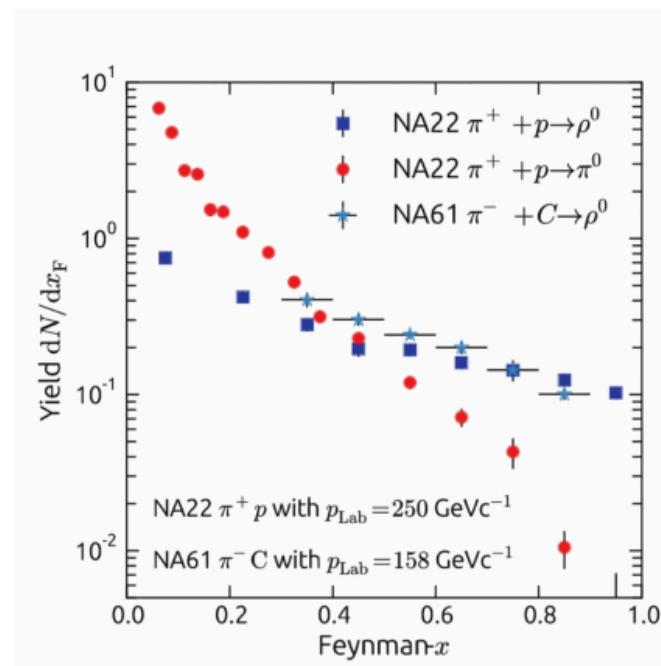
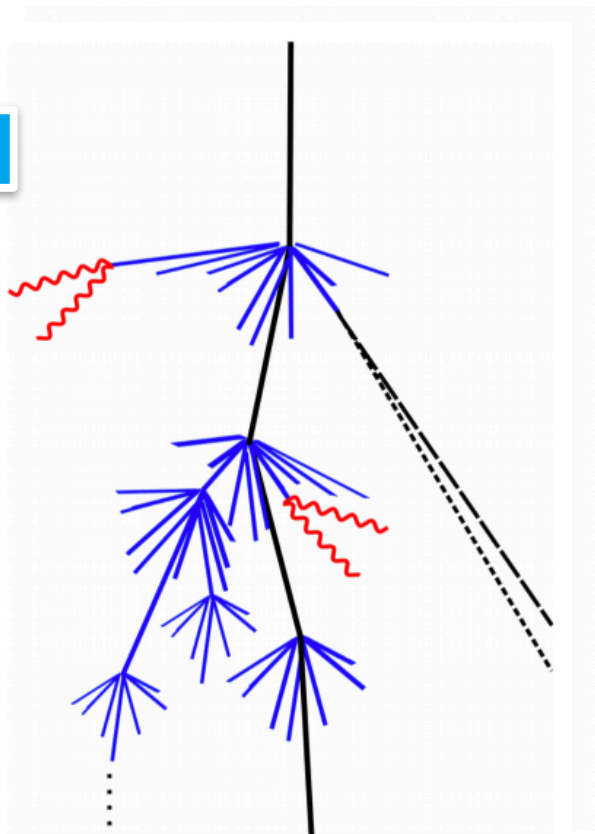


More (low-energy) muons?

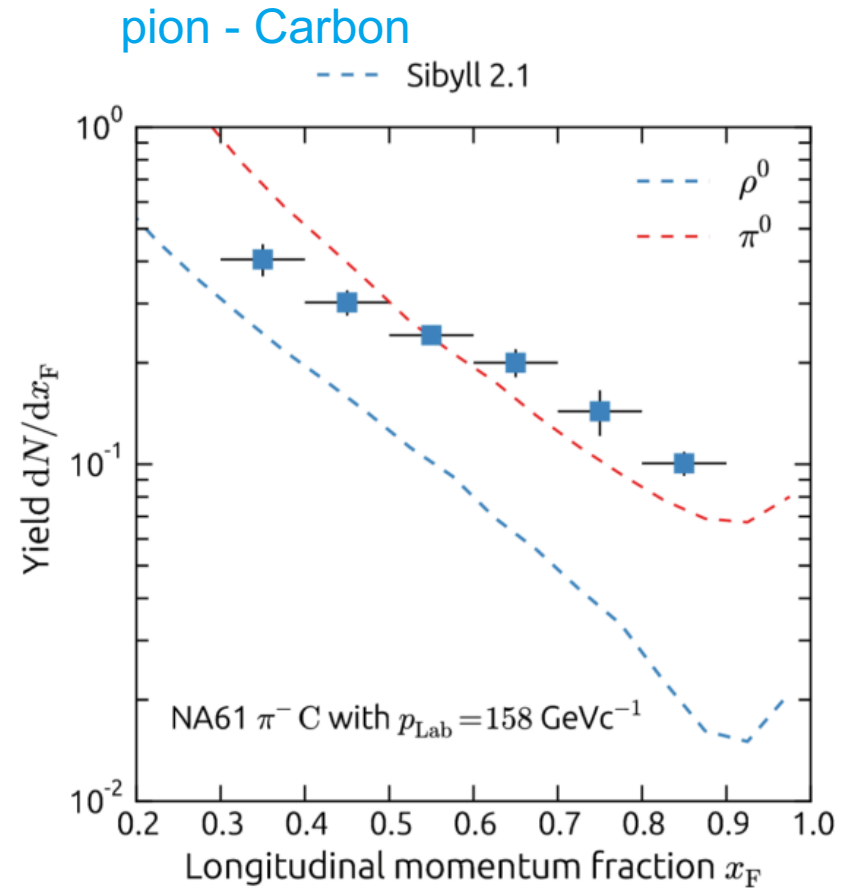
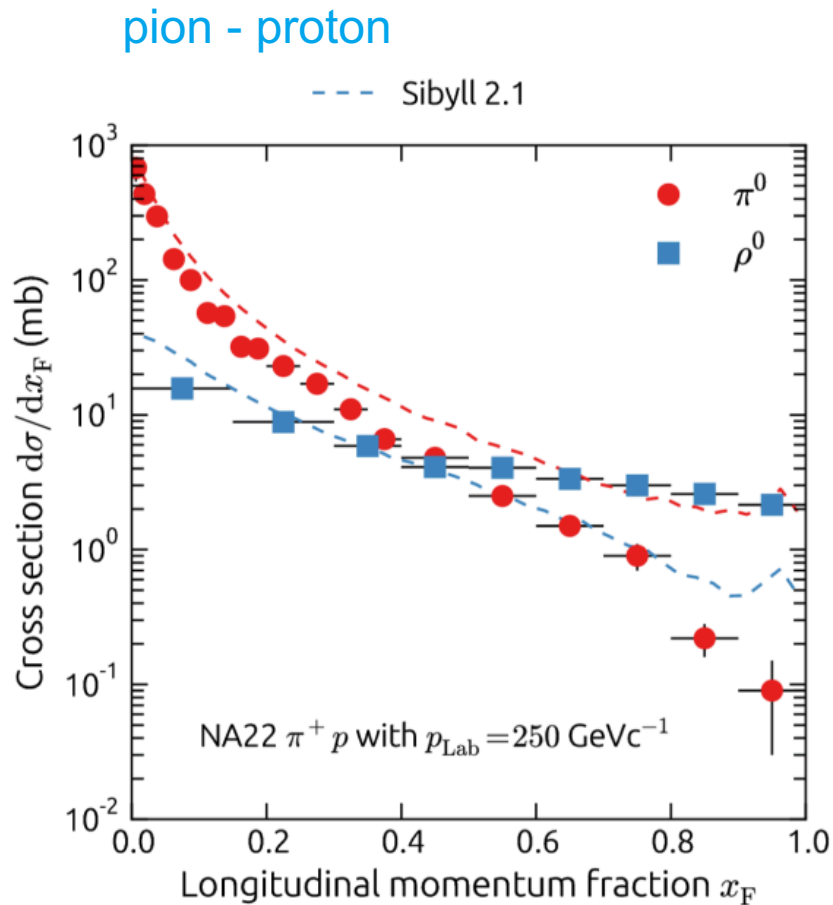
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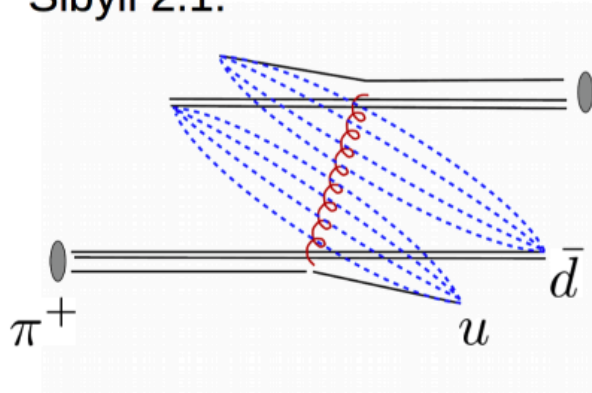
Leading vector mesons



Incorrect leading meson!

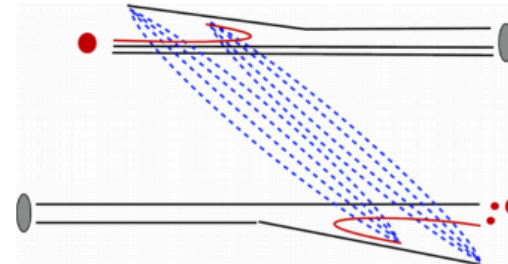
New degrees of freedom needed in microscopic model

Sibyll 2.1:

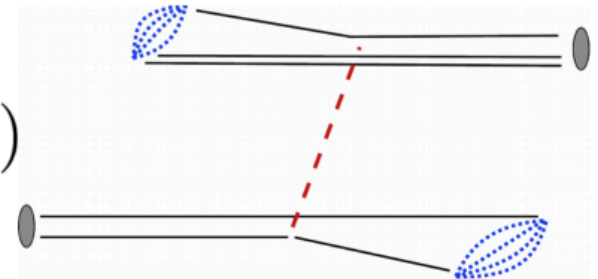


$$R_{\rho^0}/R_{\pi^0} = 0.3$$

Sibyll 2.3:

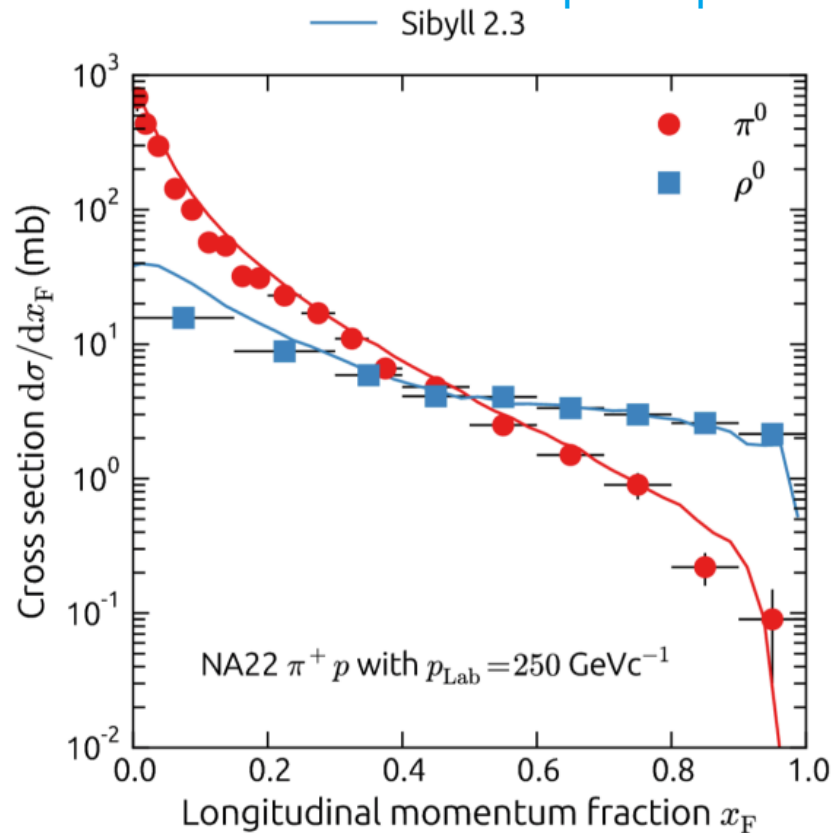


$$R_{\rho^0}/R_{\pi^0} = f(x_F)$$

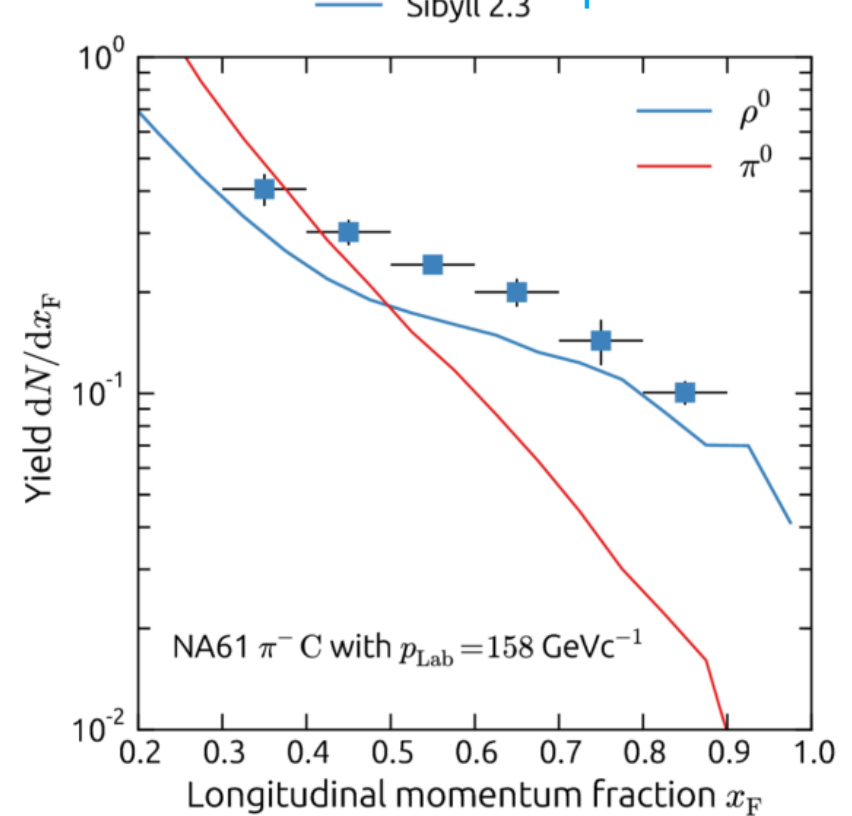


Leading vector mesons

pion - proton

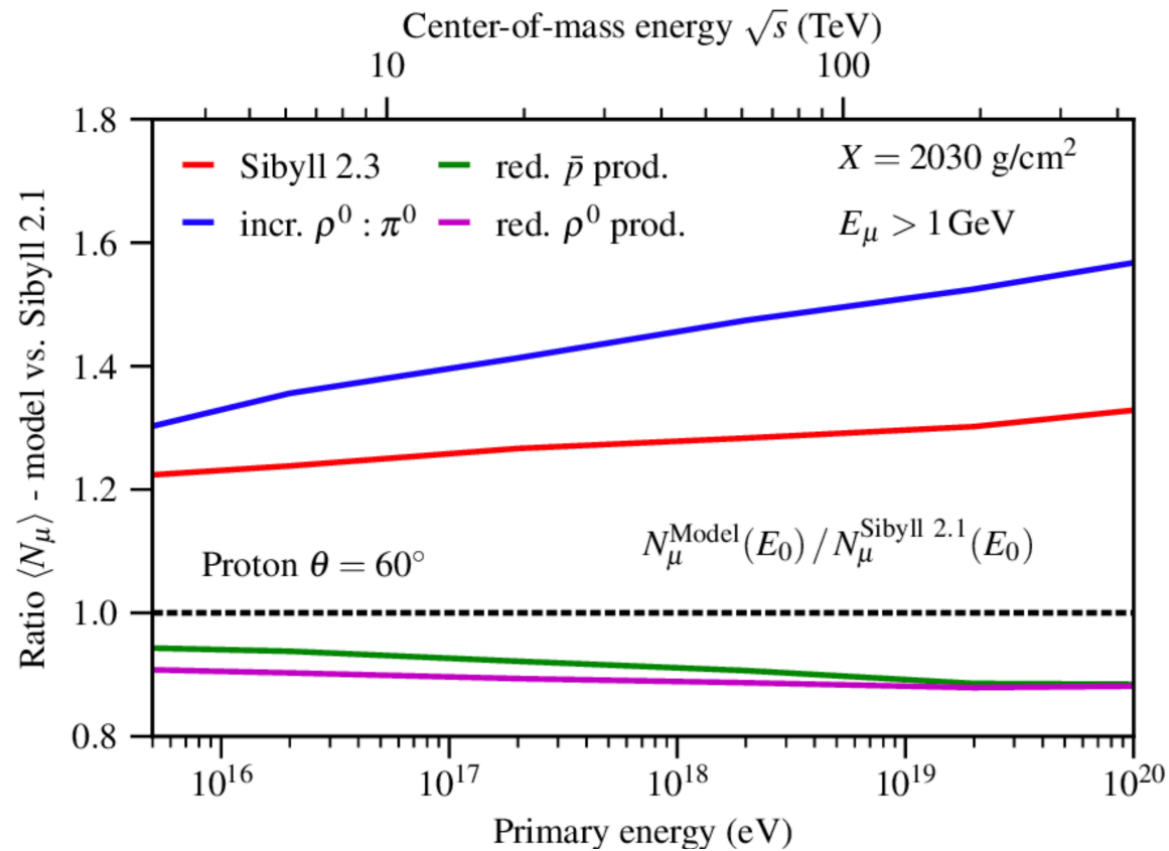


pion - Carbon



Significantly improved description!

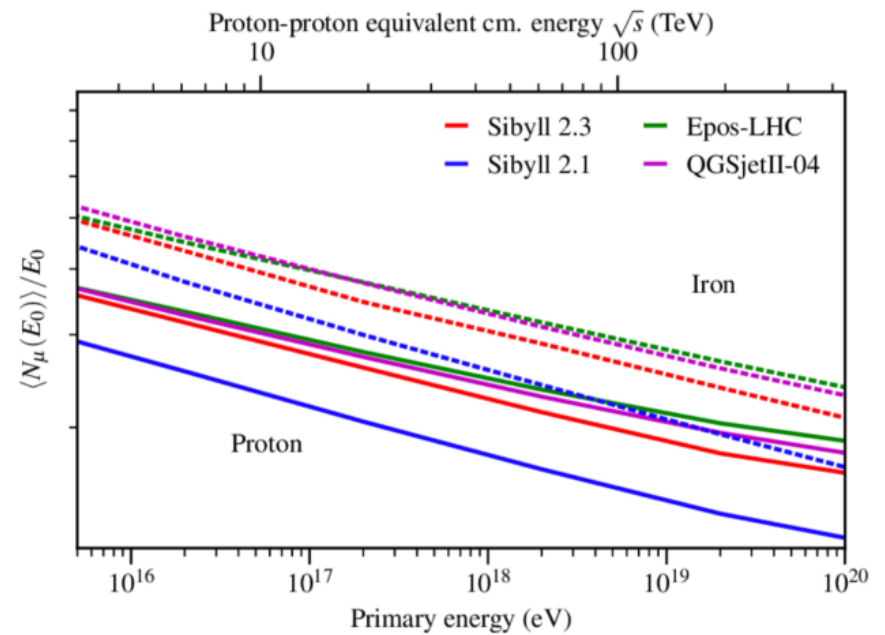
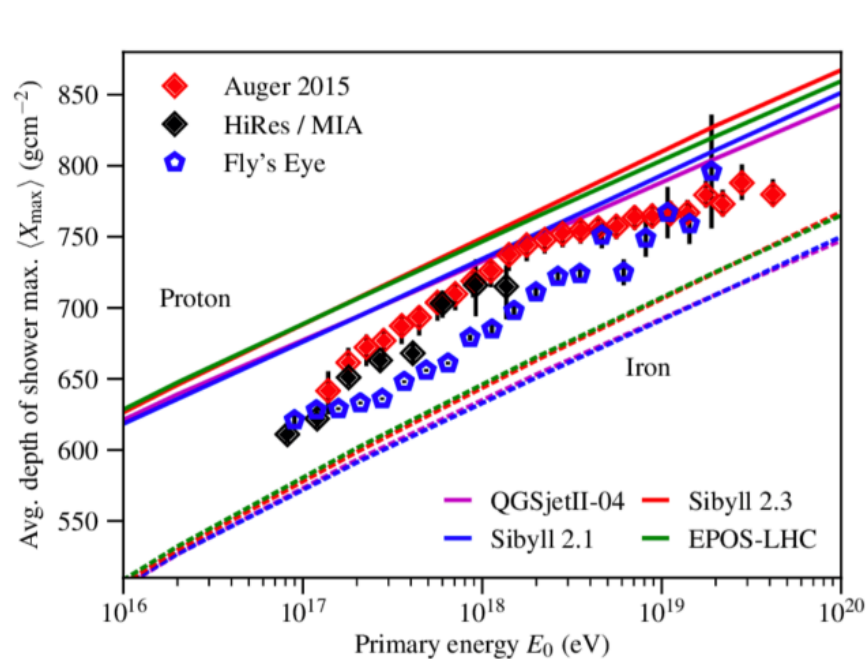
Impact on number of muons



20-30% more muons



New predictions



Deeper showers with more muons

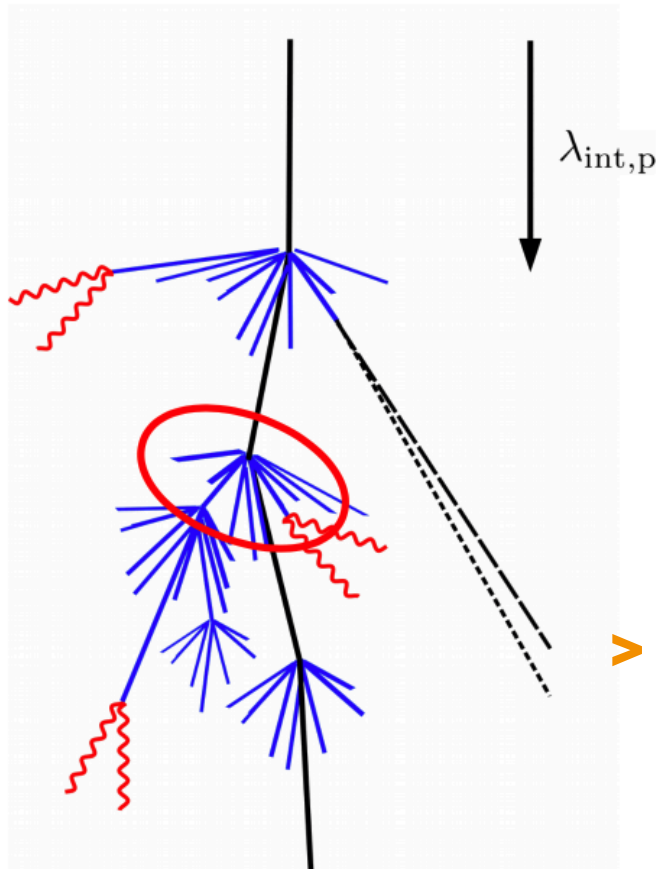
Conclusion

- Significant improvements in the description of hadronic interactions in both models DPMJET-III and SIBYLL 2.3
- DPMJET-III is technically ready for cosmic ray physics, first calculations with inclusive leptons show promising performance
- SIBYLL 2.3 changes notably air-shower predictions in muon number and X_{max}
- It also includes a new model for charm production/prompt neutrinos (not mentioned in this talk)

What's next?

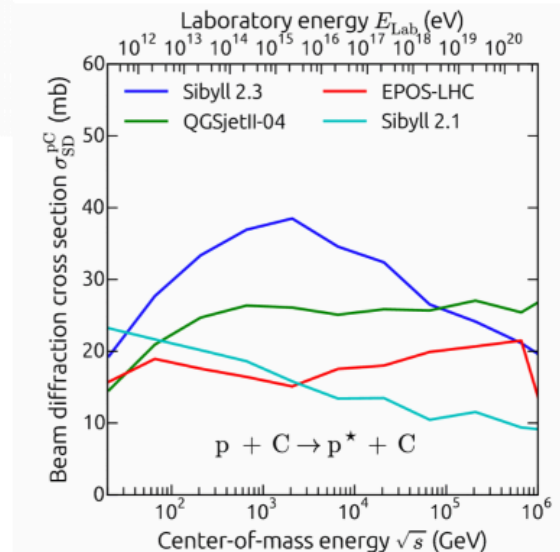
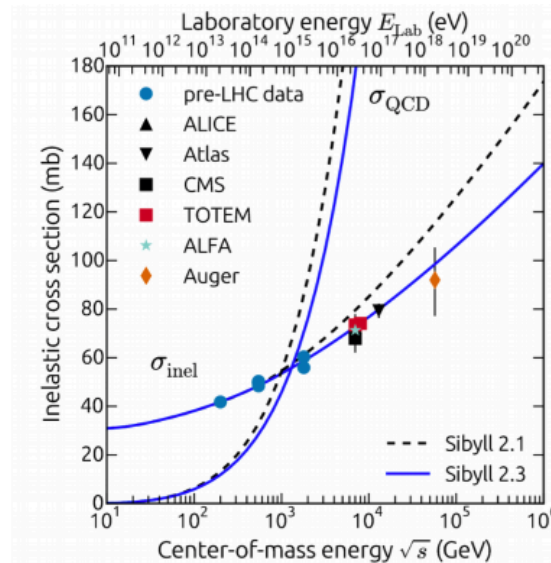
- Extrapolation of non-perturbative models to high energies?
- Re-iteration expected after deeper analysis of CR data with new models
- Nucleus-nucleus interactions, pions, kaons and baryons at fixed target energies need more attention in DPMJET-III





➤ Interactions more elastic

- nuclear Diffraction
- harder leading particles



Impact on X_{\max}

