

Towards a R&D station

for large area detectors in harsh environments

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HAP Workshop Advanced Technologies
24–25 January 2013 @ KIT

Outline

- > The Concept
- > Current Status
- > Ideas for Future Use



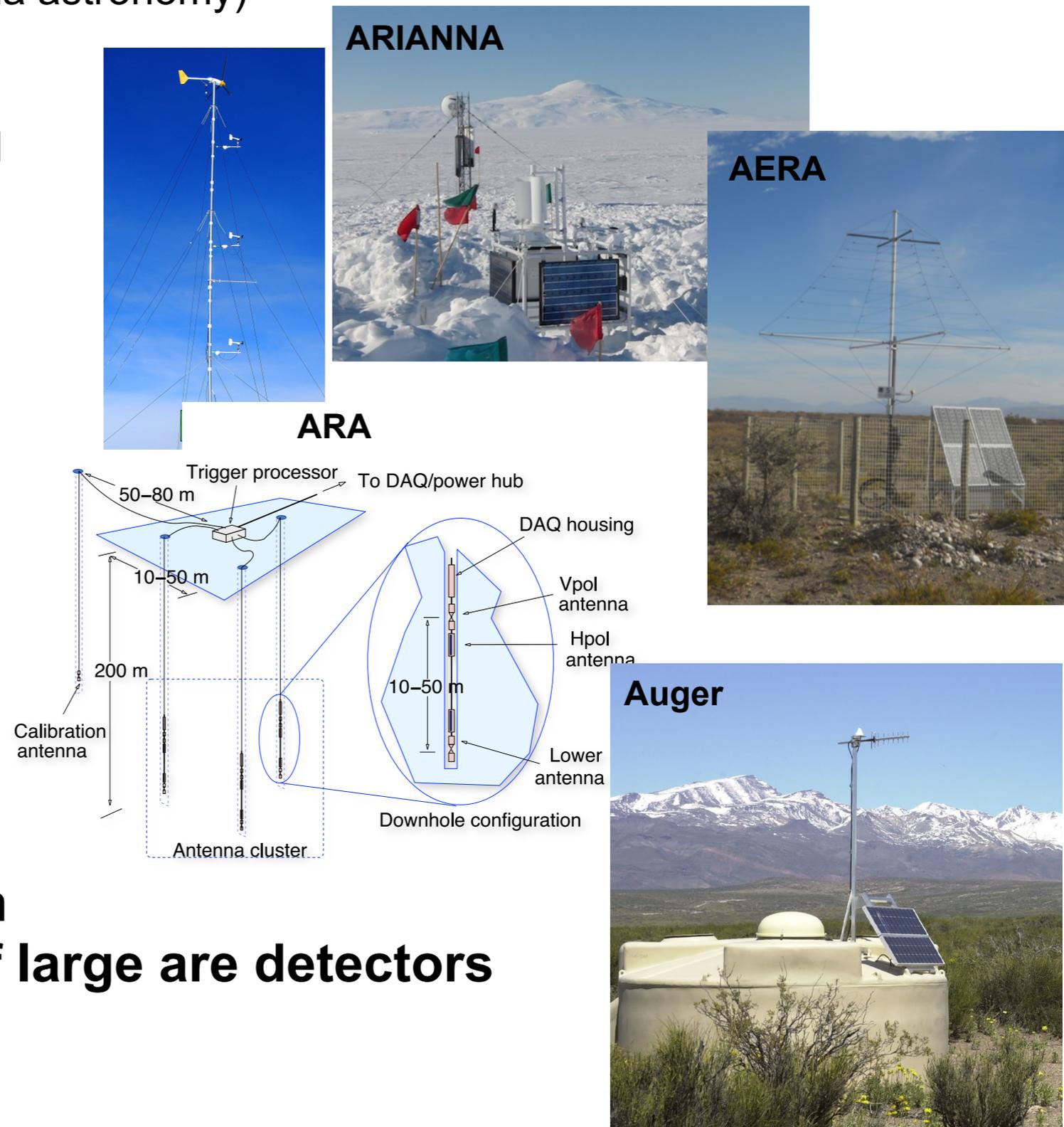
The Concept



The Idea

> Common “feature” of many astroparticle projects at the highest energies:
(UHECR, neutrinos, (non-imaging) gamma astronomy)

- Small signal fluxes:
 - Large detection areas required
- Very similar infrastructure:
 - capture of an analogue signal
 - trigger for distributed stations
 - communications
 - power distribution
 - clock distribution



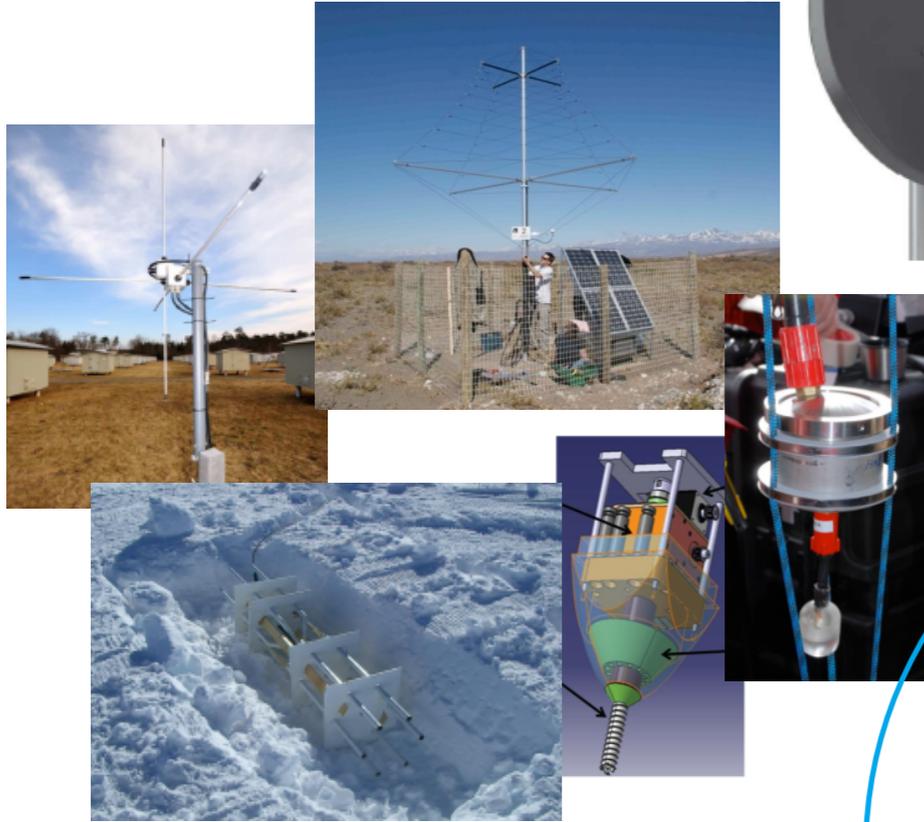
> **Development of a R&D system for testing different aspects of large are detectors**

Input signals

- > Air shower surface detector with PMTs (scintillators or water-Cherenkov)
 - (unipolar) PMT pulse (order of 10 ns)
- > Radio air shower detection
 - Waveform ≤ 100 MHz
- > Microwave air shower detection
 - (unipolar) pulse (order of 10 ns) after power detector
- > Radio neutrino detection
 - (unipolar) pulse (order of 10 ns) after power detector
- > Non-imaging Cherenkov telescopes
 - (unipolar) PMT pulse (order of 10 ns)
- > Acoustic neutrino detection
 - Waveform ≤ 1 MHz
- > **Very similar requirements:
single R&D station for different projects possible**



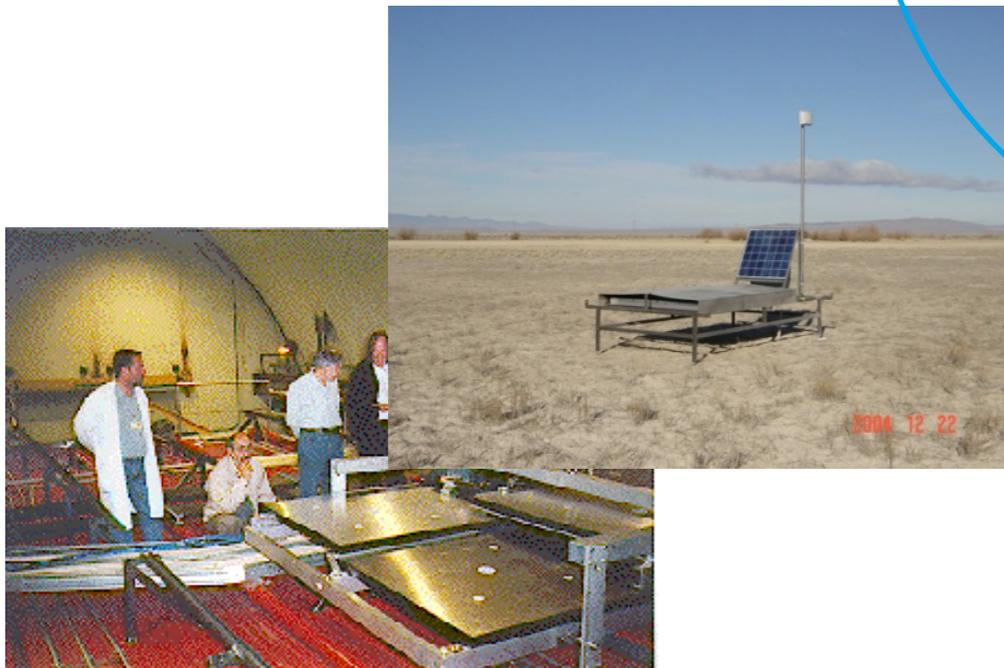
Sensors



Communication



Reference Detector



Power Source



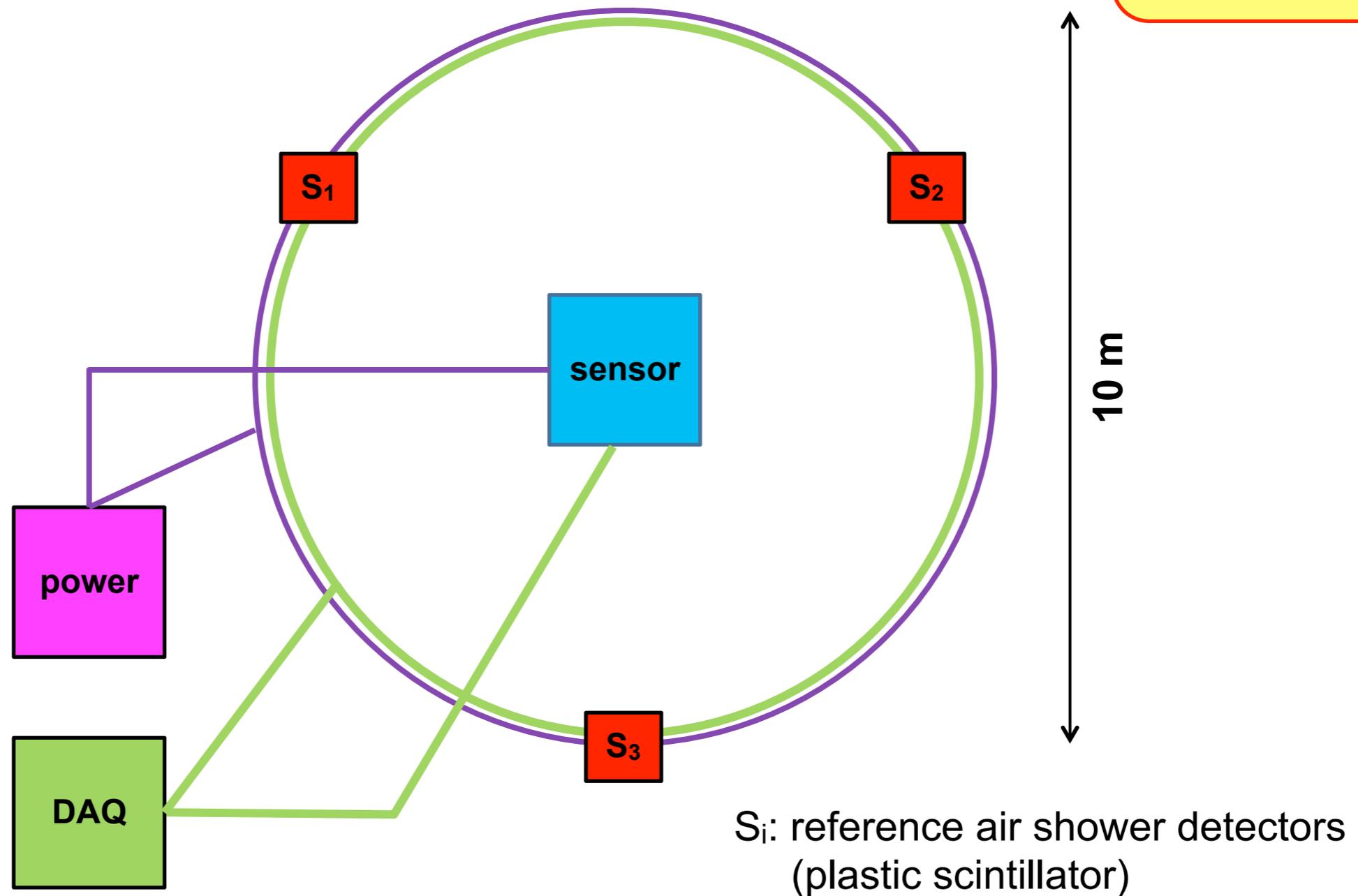
R&D
station



First Step: Single Station

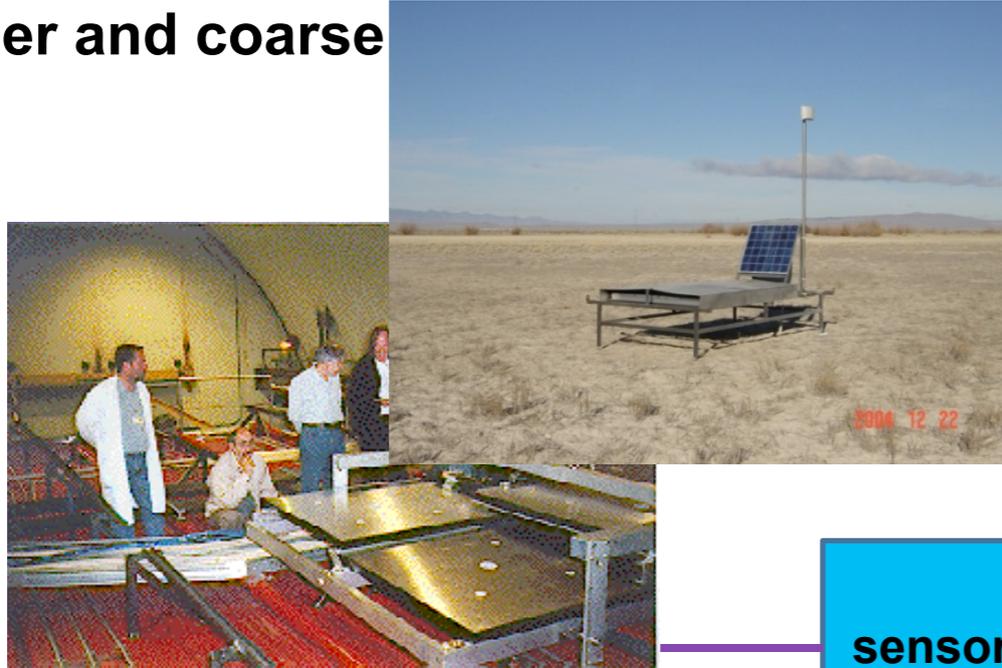
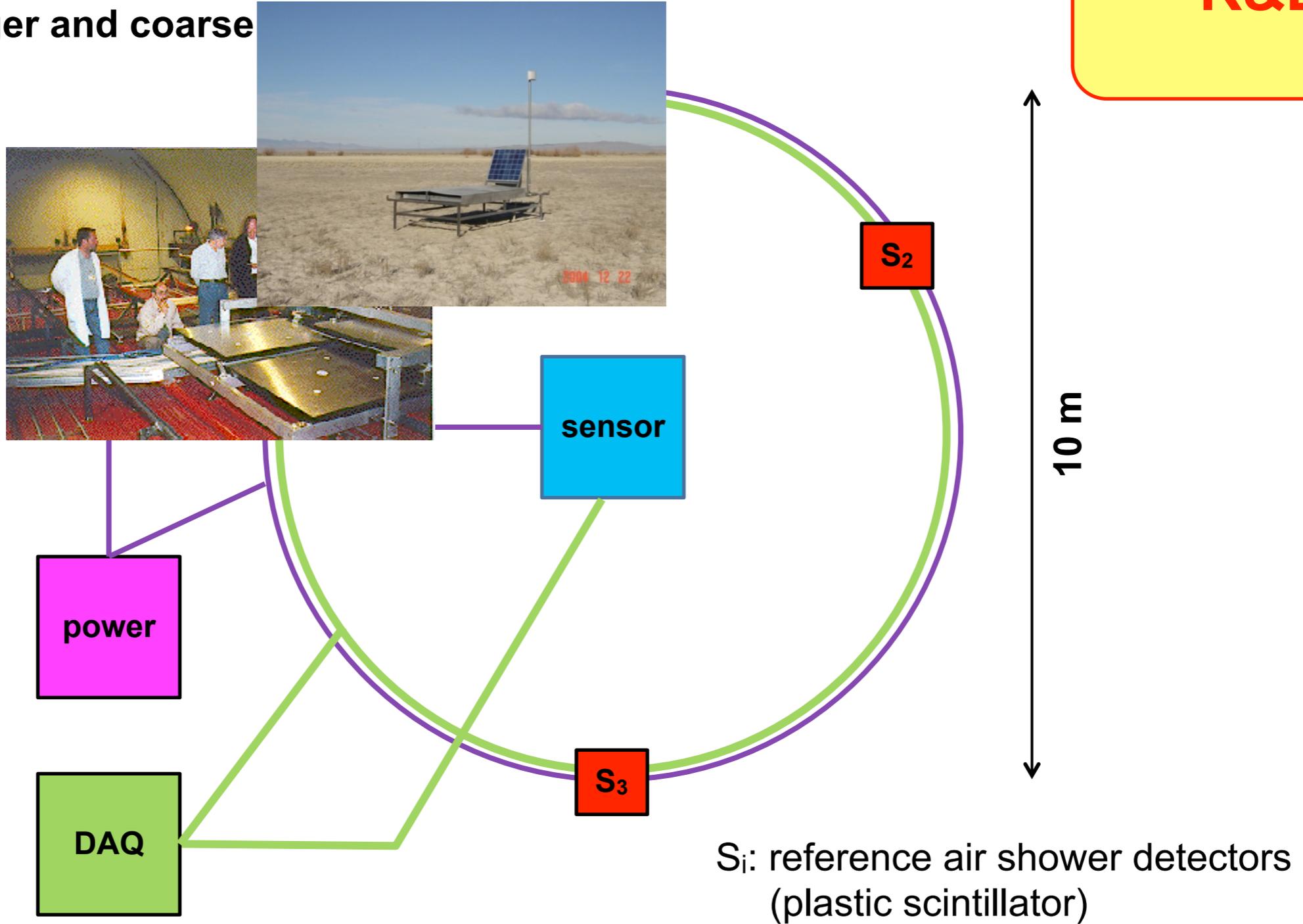
Idea: Use a simple reference air shower detector for trigger and coarse reconstruction

**Sensor
R&D**



First Step: Single Station

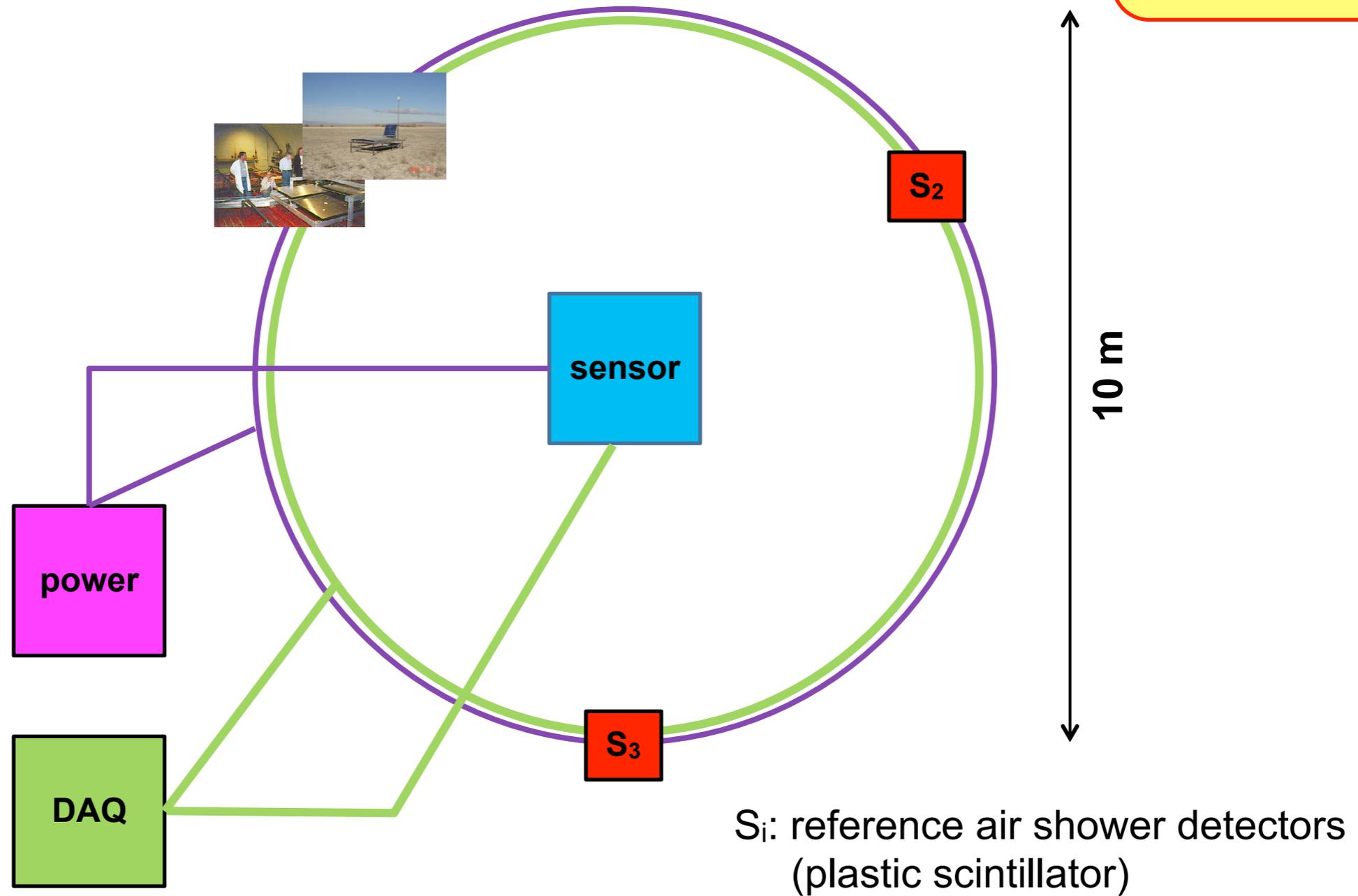
Idea: Use a simple reference air shower detector for trigger and coarse



First Step: Single Station

Idea: Use a simple reference air shower detector for trigger and coarse reconstruction

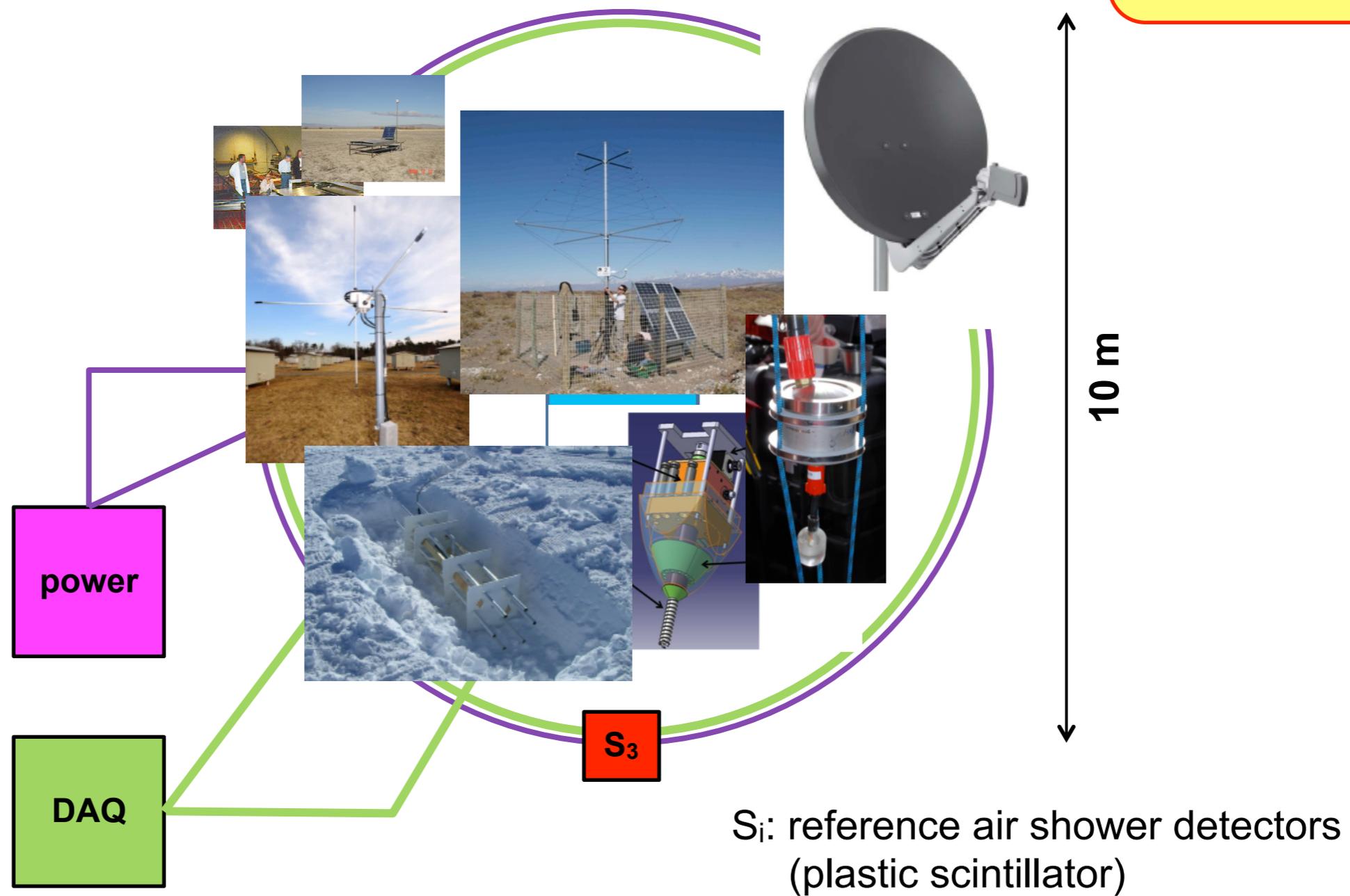
**Sensor
R&D**



First Step: Single Station

**Sensor
R&D**

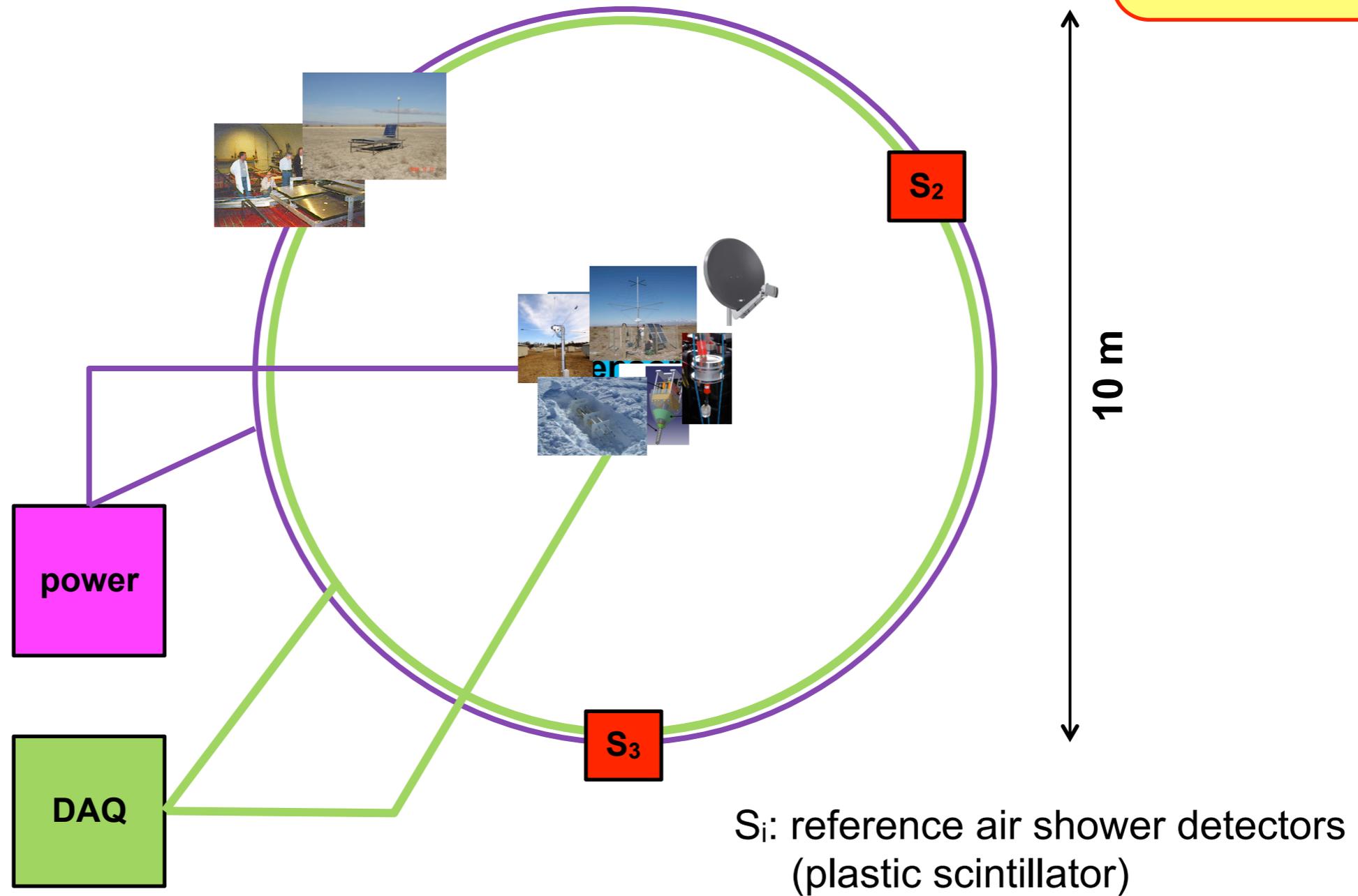
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First Step: Single Station

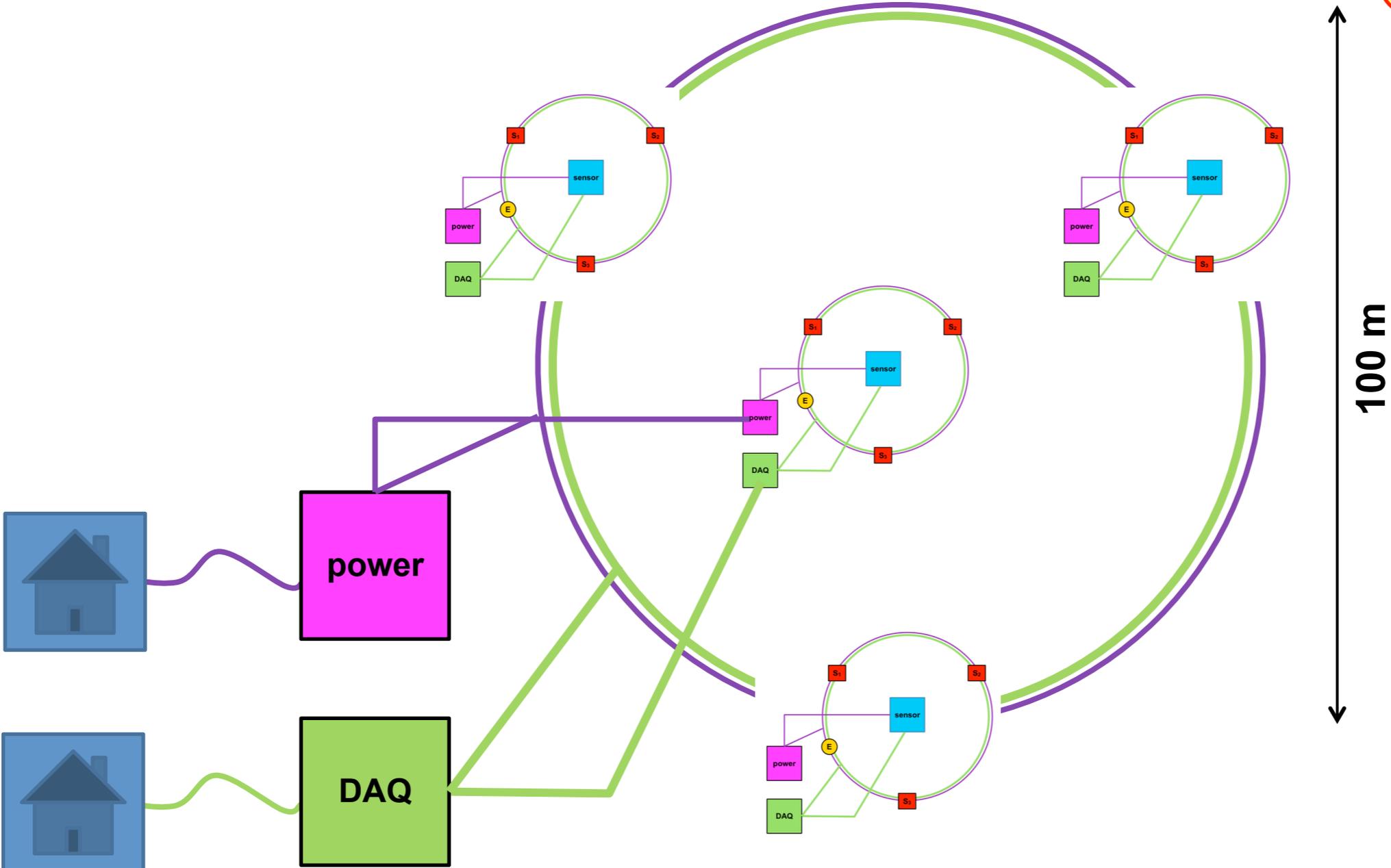
Idea: Use a simple reference air shower detector for trigger and coarse reconstruction

Sensor R&D



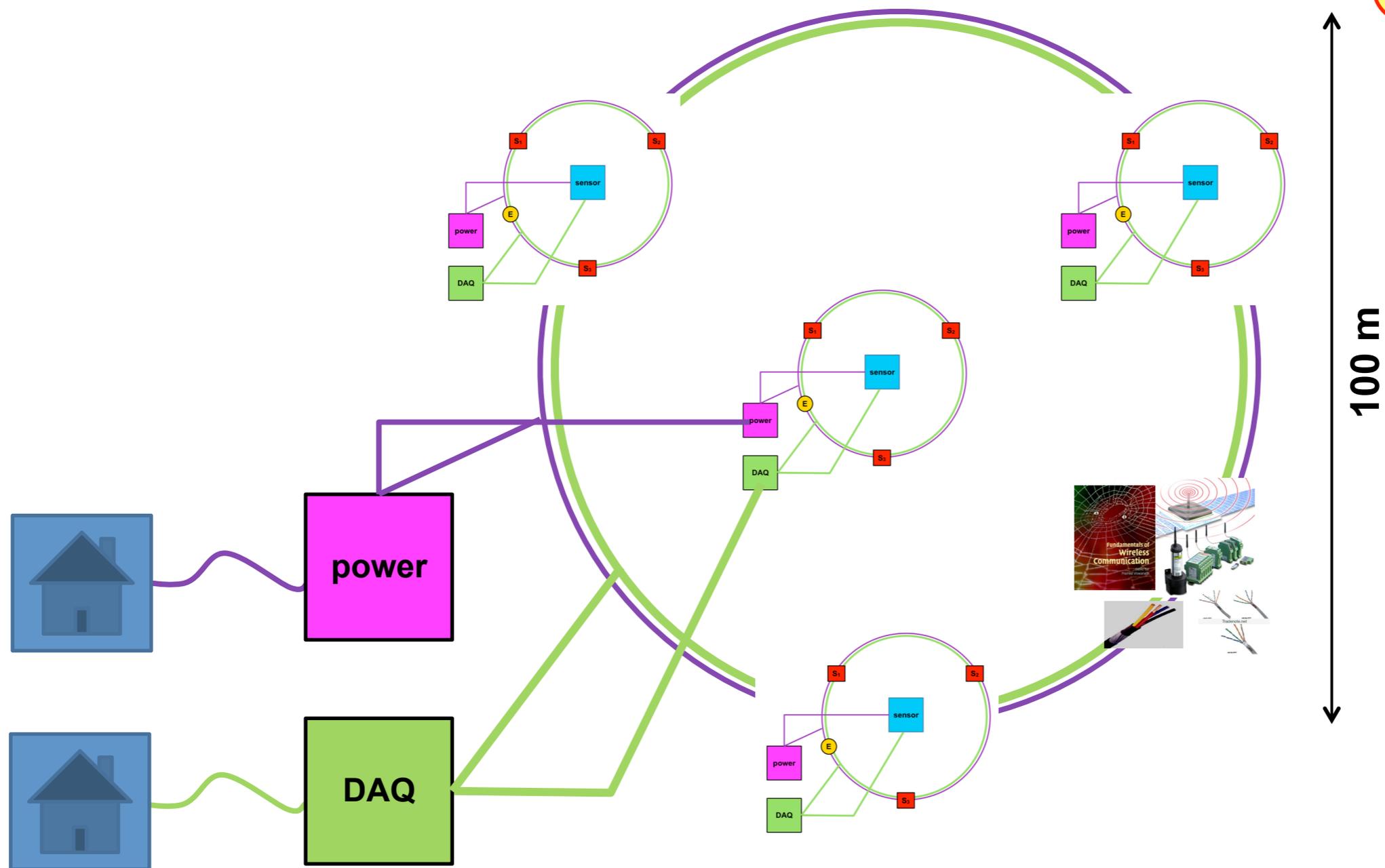
Second Step: Cluster (4 Stations)

Array
R&D



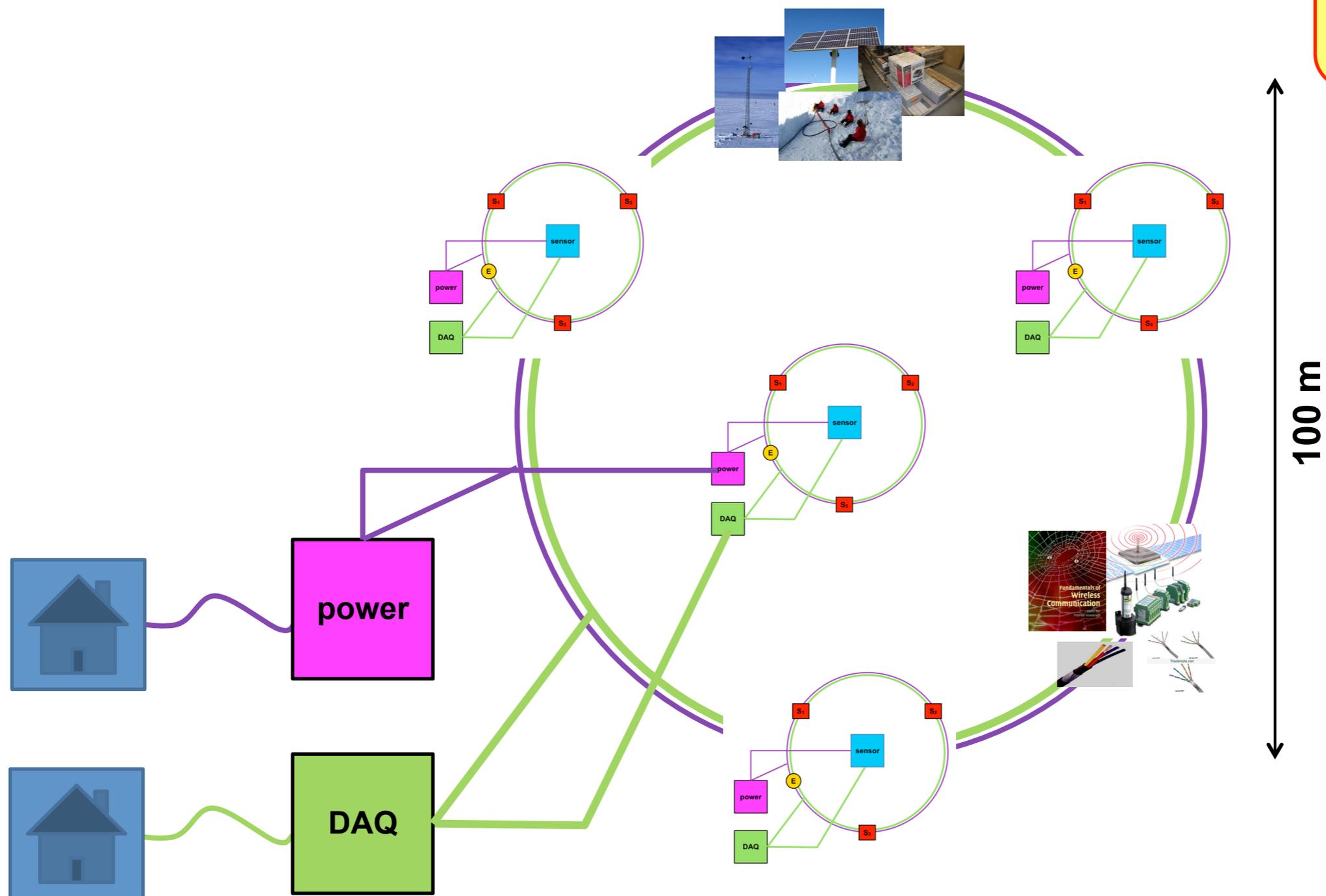
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Array
R&D



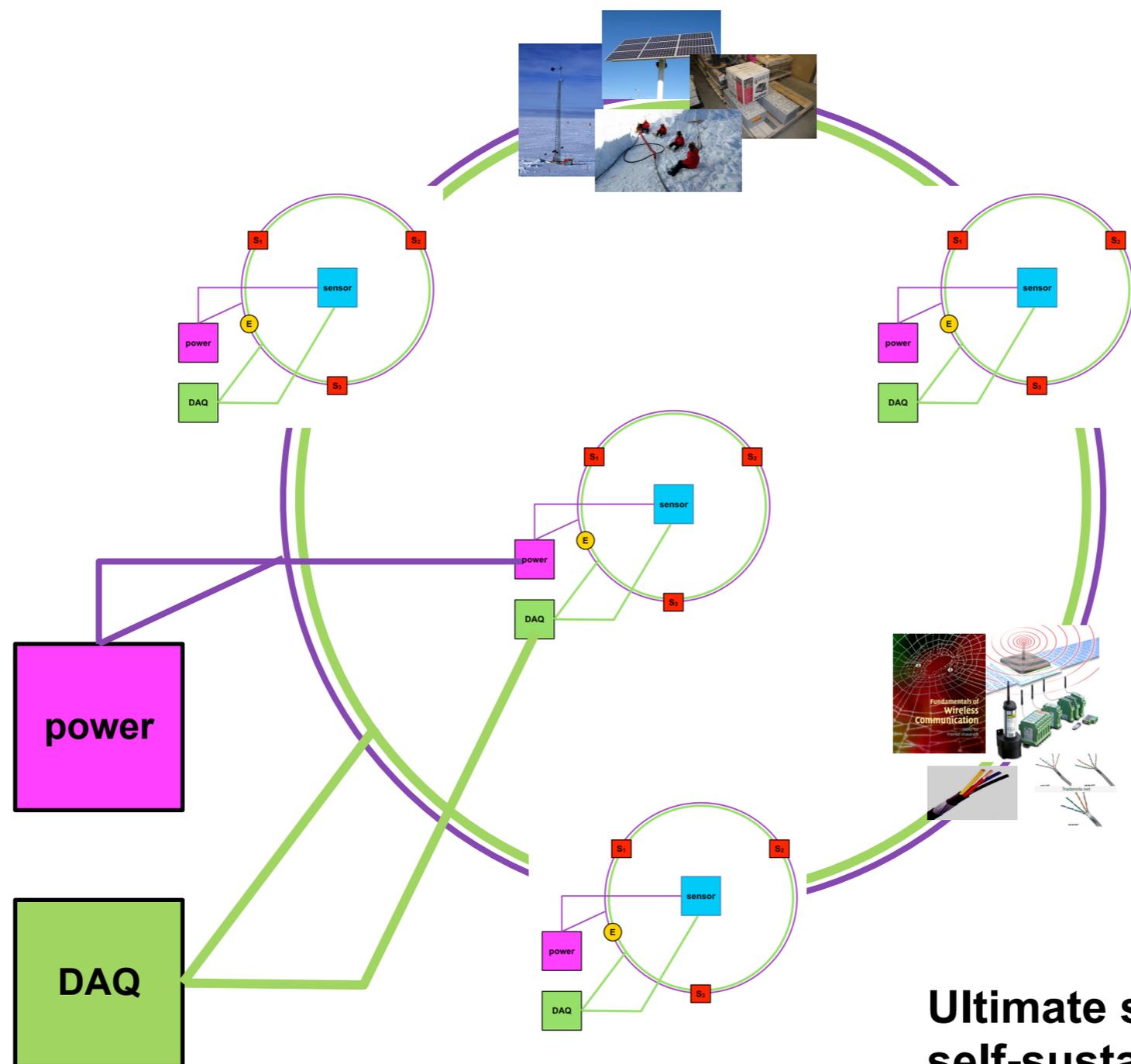
Second Step: Cluster (4 Stations)

Array
R&D



Second Step: Cluster (4 Stations)

Array
R&D



100 m

Ultimate step:
self-sustaining power and DAQ



Requirements

- Highly modular system that allows easy interchange of components
 - R&D for different system components in well defined environment

- Easy transport and setup: site studies for future projects
 - long term background measurement and monitoring
 - signal propagation studies (signal speed, attenuation, refraction, ...)

- Operation at isolated sites
 - low power, self-sustained power supply
 - environmental range from Antarctica to hot climate

- Scalability



Synergy Within the Helmholtz Alliance

> during design and construction

- experience and components for hard- and software (Aachen, Bonn, DESY, ..., KIT, ..., Wuppertal,)

> during application

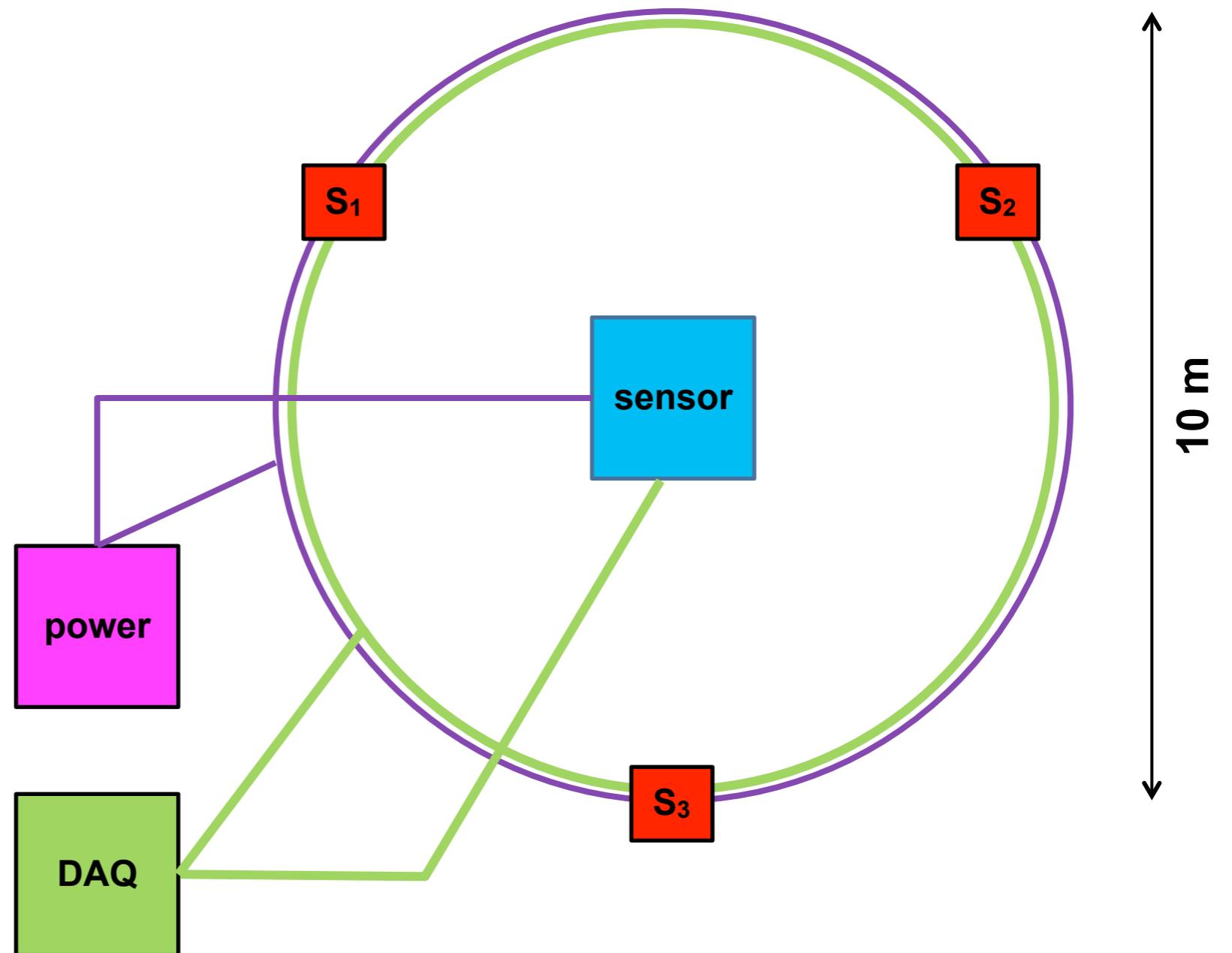
- possible examples: radio/hybrid detector in Antarctica, Auger ~~North~~ next, Tunka-Hiscore, Tunka-Rex



Current Status

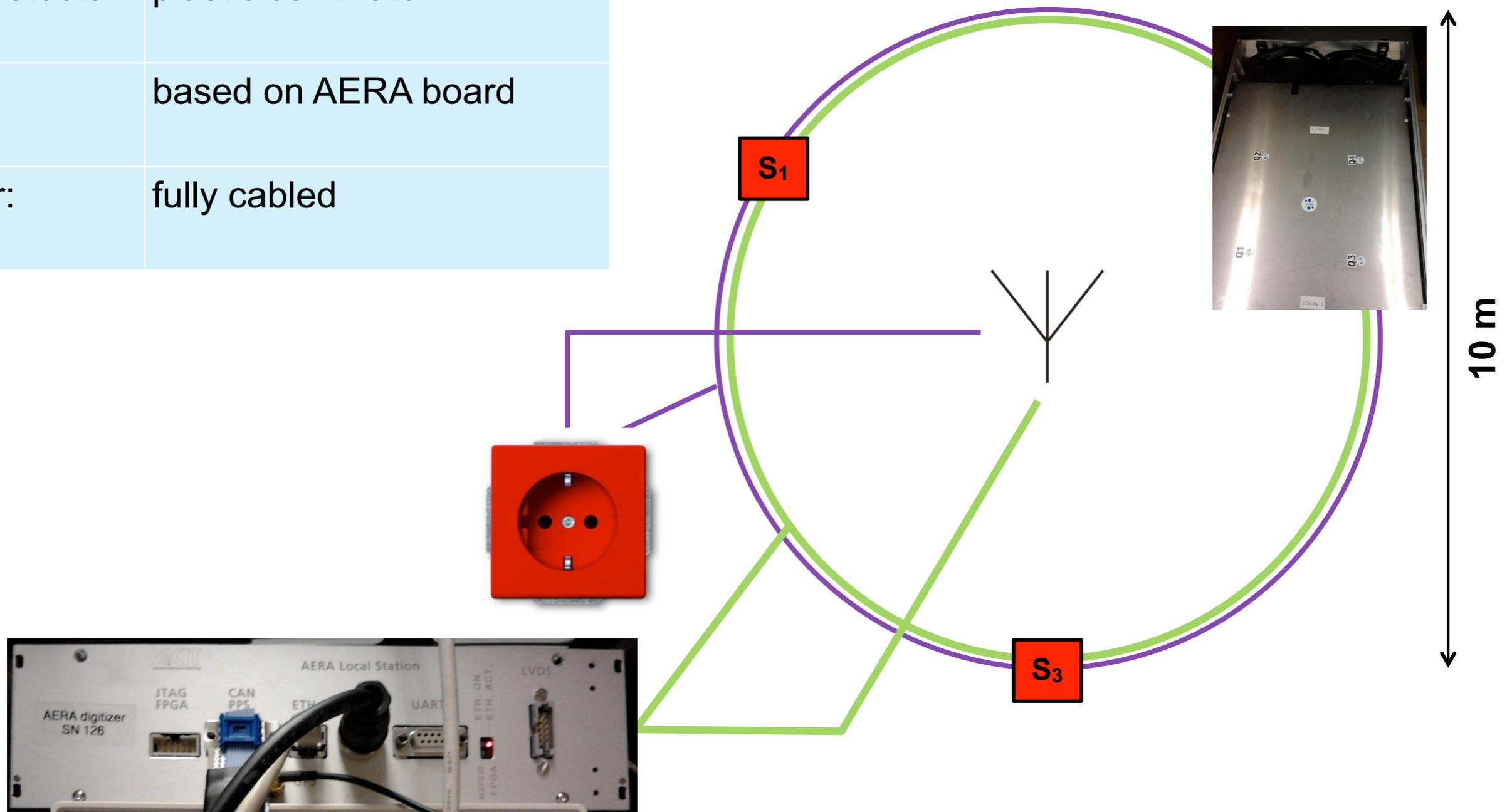


Towards Single Station #1

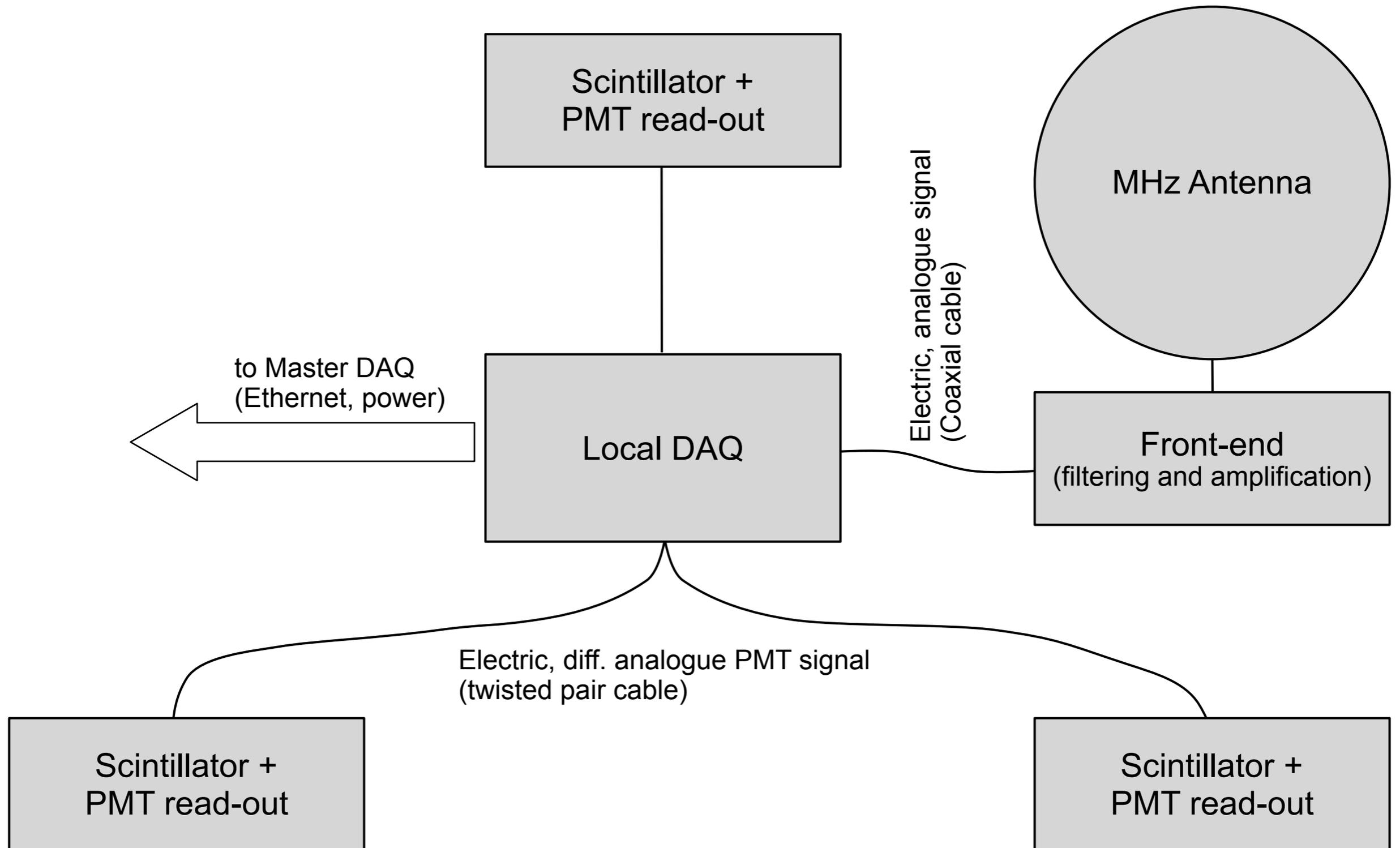


Towards Single Station #1

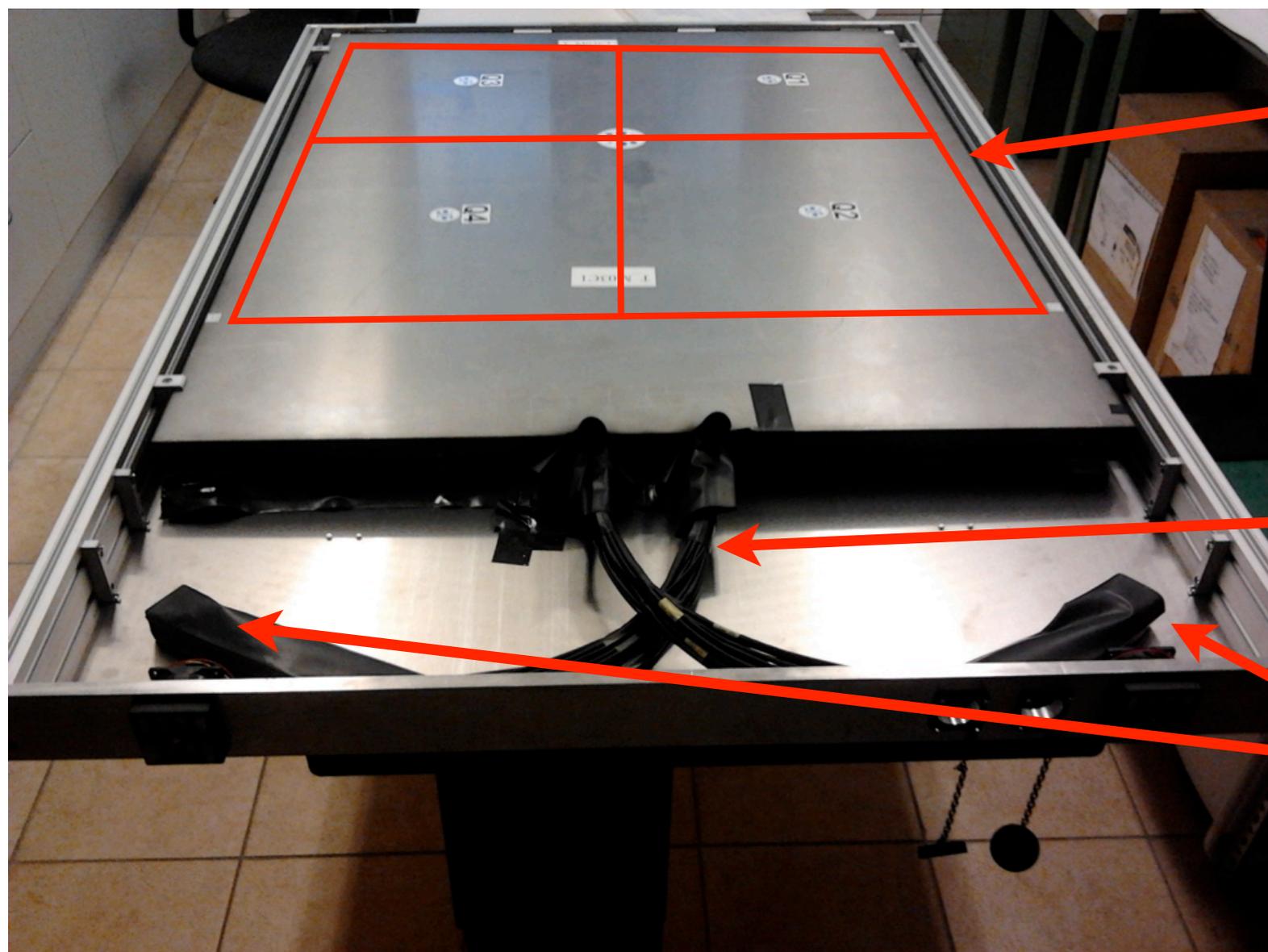
Sensor:	radio antenna (MHz air shower detection)
Ref. Detector:	plastic scintillator
DAQ:	based on AERA board
Power:	fully cabled



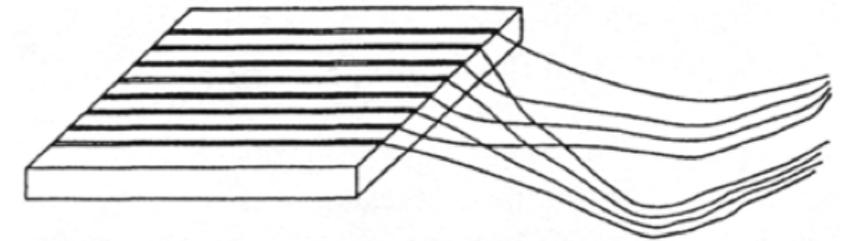
Local Station Overview



Scintillator Reference Detector



1 m² tiled plastic scintillator
4 tiles, 50 × 50 cm each



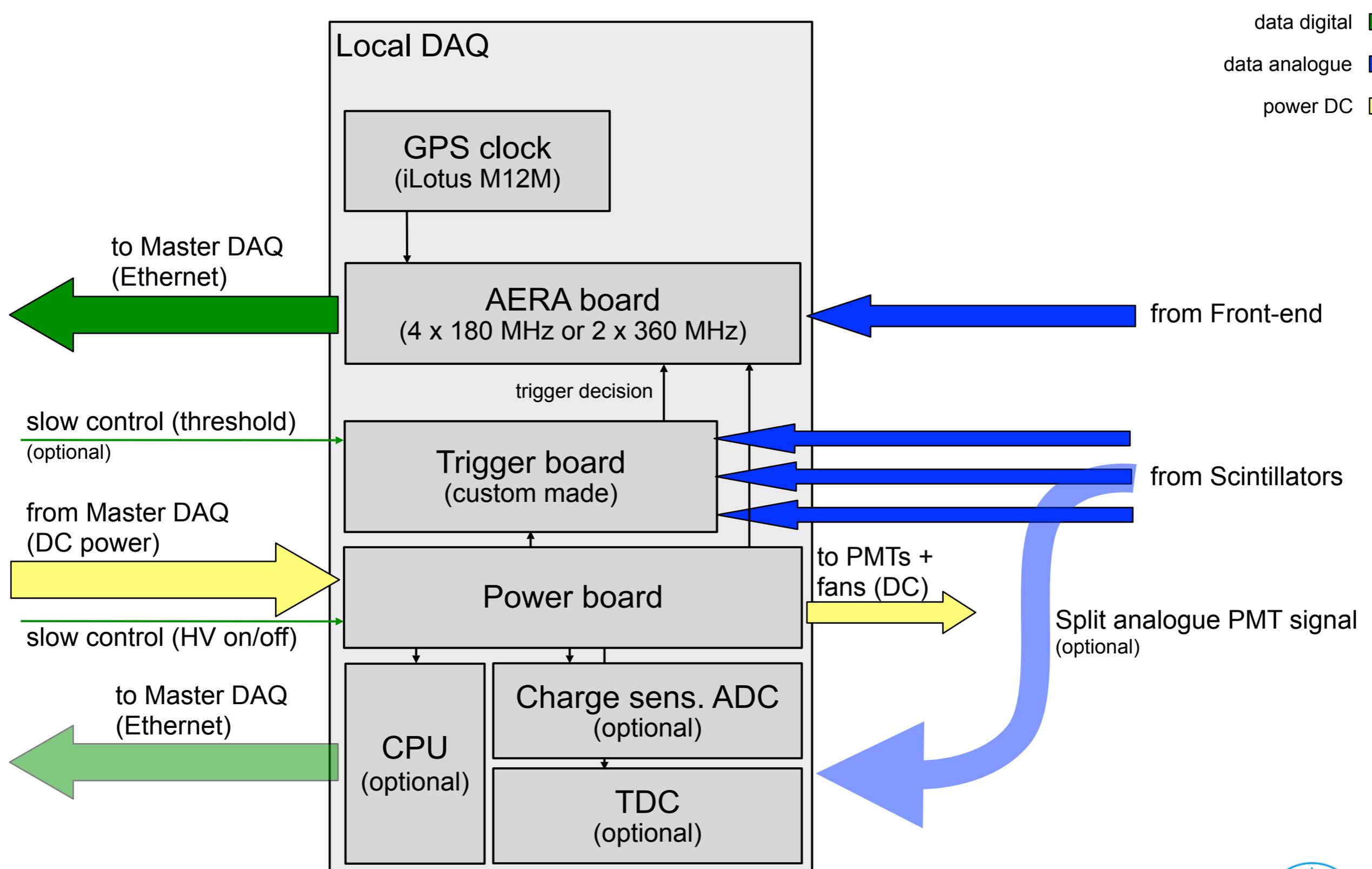
optical fibers
each tile read out by 2 sets of fibers

Hamamatsu R 5900-3-M4
2 × 2 multi-anode PMT

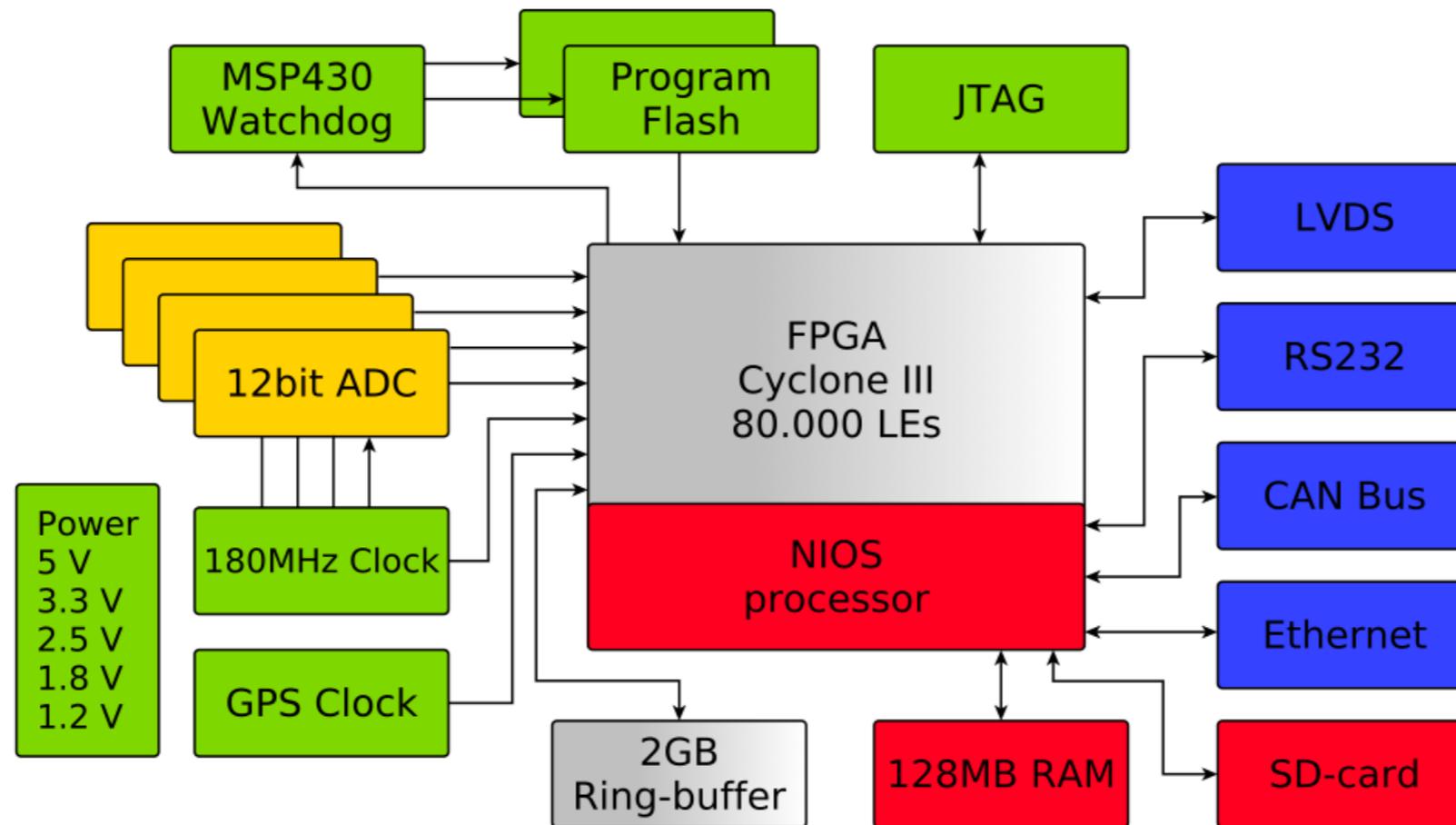
➤ Input: ± 12 V

➤ Output: analogue differential PMT signal (8 channels)

Local DAQ Overview



AERA Digitizer Board



Block diagram:
A. Schmidt, PhD Thesis, KIT (2012)

- Developed at KIT (IPE, IKP) for the Auger Engineering Radio Array (AERA)
- Four digitizers (180 MHz, 12 bit; can be interlaced to 2×360 MHz)
- deep ring buffer
- powerful FPGA for real-time signal processing
- possibility for external trigger (will be used for R&D station)

Local Air Shower DAQ

- > Analogue sum of PMT signals from same scintillator tile (noise suppression)
- > discrimination of sum signal (12 channels)
- > trigger decision in FPGA
 - sent as external trigger to AERA board
- > TDC and QDC for the 12 channels currently as VME modules
 - allows coarse estimate of air shower geometry
 - **will be implemented in trigger FPGA in next version**



- > C++ with heavy use of boost libraries
 - works on NIOS processor; performance to be studied
- > Boost.Asio: asynchronous I/O via TCP/IP between different components
 - Master DAQ, Local air shower DAQ, AERA board
- > Boost.Thread: POSIX threads
- > Boost.Serialization: data storage
- > Boost.Python: python bindings for analysis



Location and Timeline

> Step 0:

- station setup and operation
- location: DESY in Zeuthen
(station may later be used for education and outreach)
- time: first half of 2013

> Step 1:

- cluster setup and operation
- location: convenient site near to DESY in Zeuthen
- time: 2013 + first half of 2014

> Step 3:

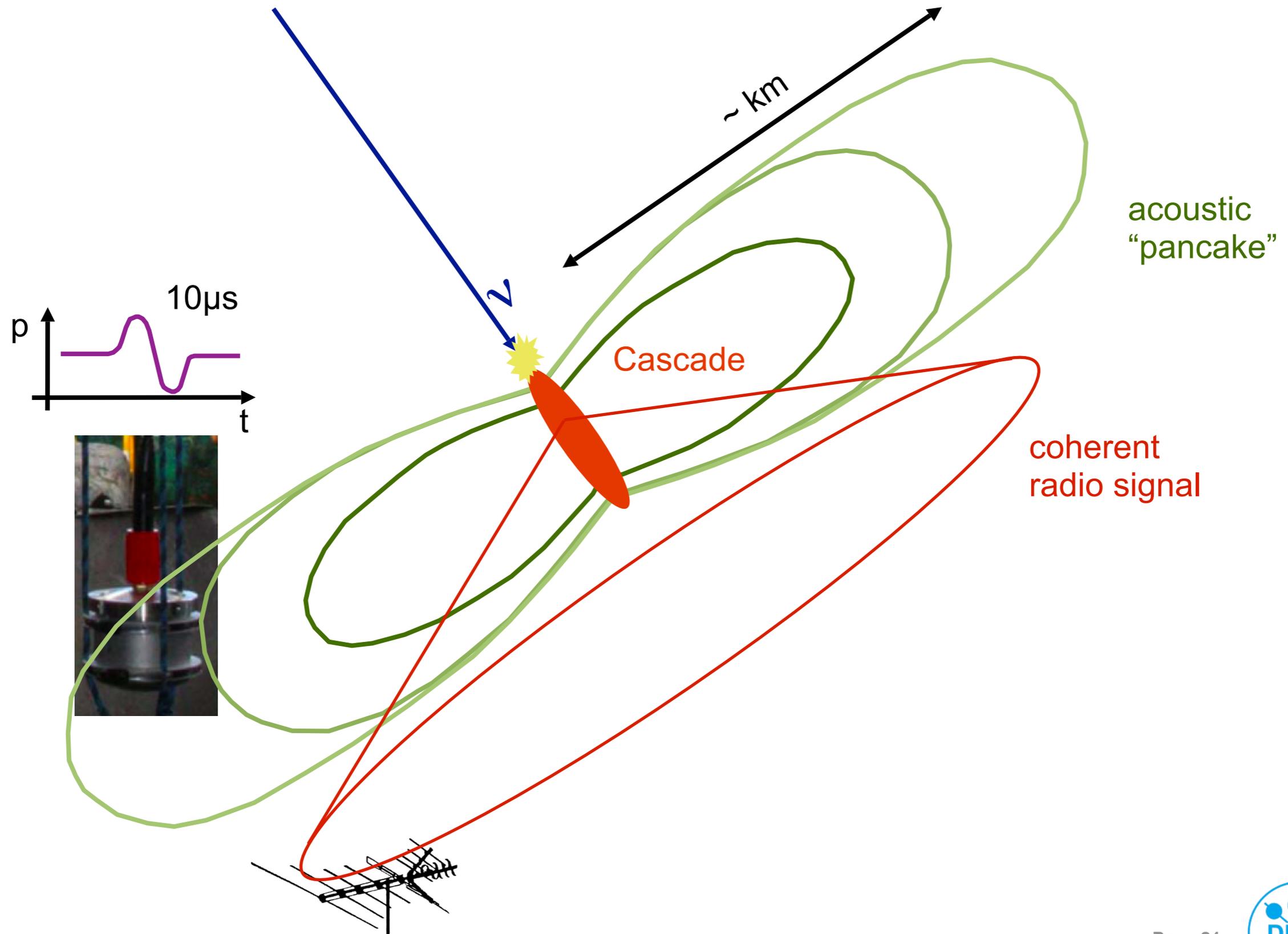
- ready for use in first field testing: > July 2014



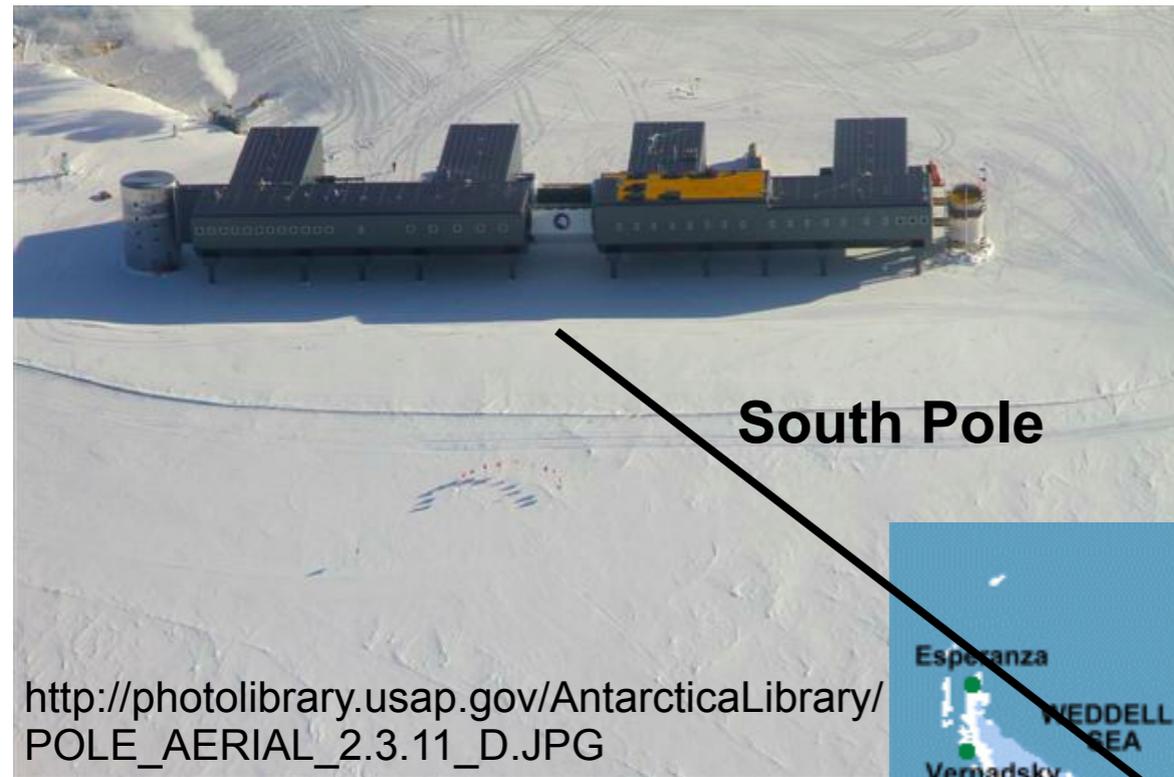
Ideas for Future Use



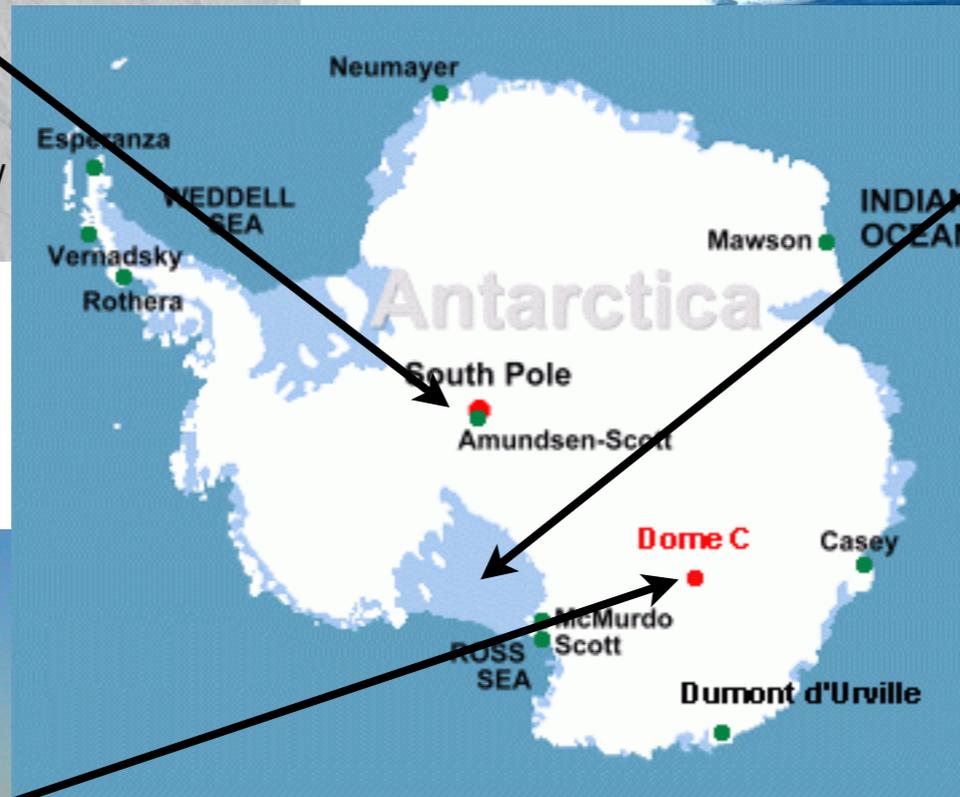
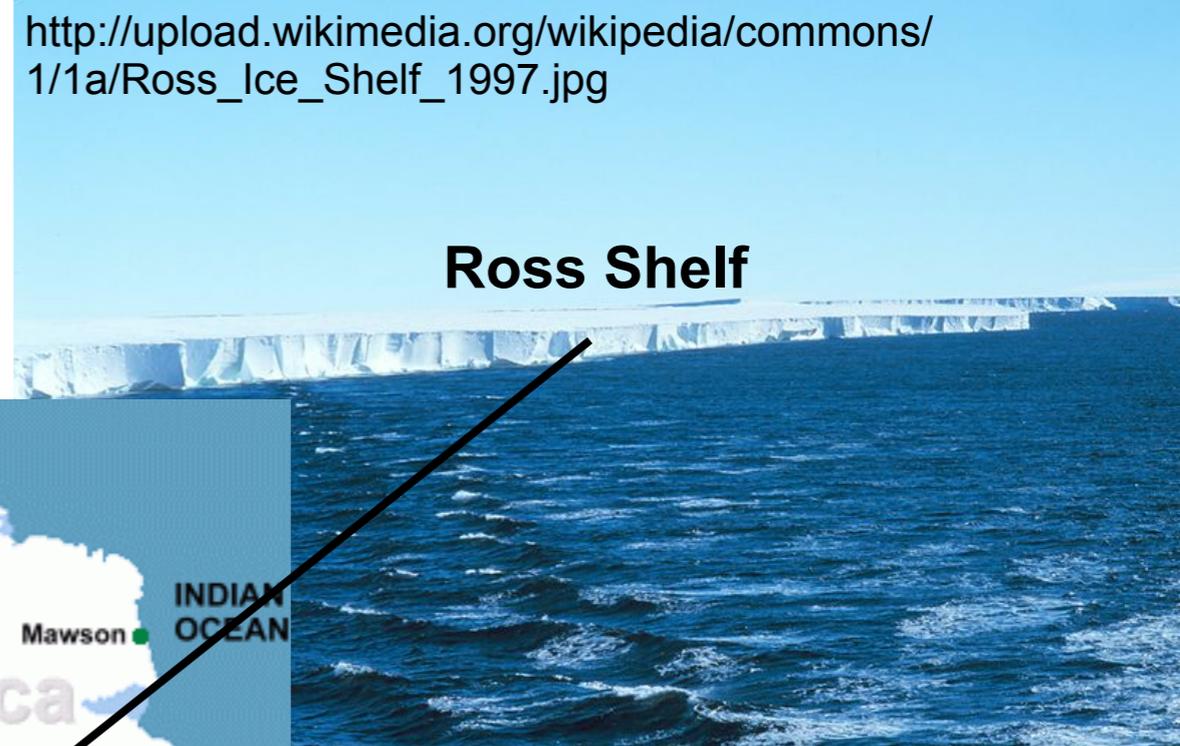
Hybrid UHE Neutrino Detector



Possible Sites to Be Studied



http://upload.wikimedia.org/wikipedia/commons/1/1a/Ross_Ice_Shelf_1997.jpg



http://www.phys.unsw.edu.au/nature/antarctica_map2.gif





**Next generation
UHECR detector array?**

Summary and Conclusions

- All large area detectors in astroparticle physics face similar challenges
 - Signal capture
 - Distributed trigger
 - Clock synchronization
 - Power distribution and communication

- Development of a modular cluster for research & development on different aspects of arrays
 - Single station with external air shower trigger
 - test and calibration of sensors and read-out
 - Four station cluster
 - development and test of clock synchronization, trigger, communication, and power distribution
 - easily transportable: exploration of prospective sites of future detectors

