

Third order corrections to the semi-leptonic $b \rightarrow c$ and the muon decays

CRC Anual Meeting 2021

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TTP KARLSRUHE

[based on: Fael, Schönwald, Steinhauser (arxiv:2011:13654)]





Outline











Calculation

Results
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ightarrow c and the muon decays



- $b \rightarrow c\ell\nu$ is an important ingredient in the inclusive determination of $|V_{cb}|$:
 - Currently there is a tension between inclusive and exclusive determination of |V_{cb}|.
 - Errors are mostly theory dominated.
 - Precise measurements of the CKM matrix elements |V_{ib}| are among main goals of Belle II and LHCb.
 - The semi-leptonic decay rate is an important ingredient in the global fit for the inclusive determination.
- $\mu \rightarrow e \nu \nu$ is the most precise way to determine G_F .



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Method of Calculation

- We calculate the inclusive decay rate to third order via the optical theorem, i.e. we consider the imaginary part of 5-loop forward scattering diagrams.
- We consider massless leptons, i.e. we have two dimensionful scales, the bottom mass m_b and the charm mass m_c.
- Analytical dependence on charm and bottom mass seems out of reach:
 - \Rightarrow consider expansion in mass difference



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The Heavy-Daughter Expansion



- Perform the expansion in the limit $m_c \sim m_b$: $\delta = 1 \rho = 1 \frac{m_c}{m_b} \ll 1$
- Limit has been shown to converge well down to $m_c/m_b \rightarrow 0$ at 2-loop order. [Czarnecki, Dowling, Piclum (Phys. Rev. D 78 (2008))]

$$\Gamma(b
ightarrow c \ell
u) = \Gamma_0 \left[X_0 + C_F \sum_{n \geq 1} \left(rac{lpha_s}{\pi}
ight)^n X_n
ight]$$

with
$$\Gamma_0 = G_F^2 m_b^2 |V_{cb}|^2 / 192 \pi^3$$



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Calculation



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Asymptotic Expansion



$$\delta = 1 - rac{m_c}{m_b}$$

- We use the method of regions to perform the expansion. [Beneke, Smirnov (Nucl. Phys. B (1998))]
- Loop momenta can either scale hard k_i ~ m_b or ultra-soft k_i ~ δm_b (regions have been cross-checked with Asy). [Pak, Smirnov (Eur. Phys. J. C (2011))]
- The momentum of the electron-neutrino loop can be integrated trivially.
- The momentum of the lepton *q* has to scale ultra-soft, otherwise no imaginary part is generated. This reduces the number of regions to be considered.
- After the asymptotic expansion the integrals over the QCD loops (k_1, k_2, k_3) have a definitive scaling in $2p \cdot q + 2m_b^2 \delta$:
 - \Rightarrow factorize out this dependence and perform the 1-loop tensor integral over *q* first.
- We are left with 3-loop integrals with integer powers in the end.

Asymptotic Expansion – Example



Look at the 1-loop integral (we already integrated out the electron-neutrino loop):

$$\sim \int \frac{\mathrm{d}q \mathrm{d}k}{[q^2]^{\alpha}[(p+q)^2 - m_c^2]^2[k^2][(p+q+k)^2 - m_c^2]}$$

case 1: q has to be ultra-soft, k is hard;

$$\rightarrow \int \frac{\mathrm{d}q}{[q^2]^{\alpha} [2p \cdot q + 2m_b^2 \delta]^2} \times \int \frac{\mathrm{d}k}{[k^2][(k+p)^2 - m_b^2]}$$

• We see an explicit factorization.

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Asymptotic Expansion – Example



Look at the 1-loop integral (we already integrated out the electron-neutrino loop):



case 2: q and k are ultra-soft;



Integrals can be factorized through definite power counting in the asymptotic expansion.

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Details on the Calculation



We can always perform the integrations over the electron-neutrino loop and lepton momenta analytically via 1-loop tensor reduction. The remaining loop integration have the following scalings:

	scaling	n. regions
$\mathcal{O}(\alpha_s)$	h, u	2
$\mathcal{O}(\alpha_s^2)$	hh, hu, uu	4
$\mathcal{O}(\alpha_s^3)$	hhh, huu, hhu, uuu	8

- In case a single region with either hard or ultra-soft scaling remains we can also integrate it out analytically.
- The remaining two- or three-loop integrals have integer powers of the propagators and can be reduced to master integrals via IBP reduction.
- Since we expand up to O(δ¹²) we have to reduce about 25M three-loop integrals with positive and negative indices up to 12. We used FIRE together with LiteRed for this task. [Smirnov,Chuharev (2020), Lee (2013)]

Details on the Calculation



Different kinds of master integrals appear in hard or ultra-soft regions:

- hard regions: up to three-loop on-shell master integrals.
 [Melnikov, van Ritbergen (Nucl. Phys. B (2000) ; Lee, Smirnov (JHEP (2011))]
- soft regions: three-loop ultra-soft master integrals with eikonal propagators
 - up to 2-loop integrals expressible in terms of Γ-functions
 - 3-loop integrals computed for the $m^{OS} m^{kin}$ relation at $O(\alpha_s^3)$ [see talk by Matteo Fael]. [Fael, KS, Steinhauser (2020); hep-ph/2011.11655]

Renormalization:

- For the renormalization of the decay width the wave function and mass renormalization constants with two massive quarks need to be known in the expansion $m_c \sim m_b$ up to $O(\alpha_s^3)$.
- Previously they were only known in the expansion $m_c \ll m_b$ and numerically for larger values of m_c . [Bekavaz, Grozin, Seidel, Steinhauser (JHEP (2007))]
- We computed them analytically and expanded around the equal mass limit to obtain the needed quantities. [see talk by Matthias Steinhauser]

Results

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$$\Gamma(b \to c \ell \nu) = \Gamma_0 \left[X_0 + C_F \sum_{n \ge 1} \left(\frac{\alpha_s}{\pi} \right)^n X_n \right], \qquad \qquad X_n = \sum_{j=5}^{\infty} \delta^j X_{n,j}$$

$$\begin{split} \mathcal{C}_{F}X_{3} &= \delta^{5} \bigg[\frac{533858}{1215} - \frac{20992a_{4}}{81} + \frac{8744\pi^{2}\zeta_{3}}{135} - \frac{6176\zeta_{5}}{27} - \frac{16376\zeta_{3}}{135} - \frac{2624l_{2}^{4}}{243} + \frac{5344\pi^{2}l_{2}^{2}}{1215} \\ &+ \frac{179552\pi^{2}l_{2}}{405} - \frac{39776\pi^{4}}{6075} - \frac{1216402\pi^{2}}{3645} \bigg] + \mathcal{O}(\delta^{6}), \end{split}$$

with $I_2 = \ln(2)$, $a_4 = \text{Li}_4(1/2)$ and $\zeta_i = \sum_{j=1}^{\infty} 1/j^j$.

• We have calculated the expansion up to δ^{12} (for general color factors).

• A subset of color factors has been independently been computed up to δ^9 . [Czakon, Czarnecki, Dowling (2021)]

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Convergence – Quark Decays



- We see a good convergence at the physical point of ρ = m_c/m_b ≈ 0.28.
- We find:

 $X_3(
ho=0.28)=-68.4\pm0.3$

- We use the difference of the last two expansion terms to estimate the uncertainty.
- For $\rho \rightarrow 0$ we can extract values for $b \rightarrow u\ell\nu$:

$$X_3^u = -202 \pm 20$$



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Convergence – Muon Decays



• Specifying the color factor to QED and setting $\rho = m_e/m_\mu \approx 0$ we get the 3-loop contributions to the muon decay.

We find:

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 $X_3^\mu = -15.3 \pm 2.3$

This leads to the shift:

 $\Delta au_{\mu} = (-9 \pm 1) \cdot 10^{-8} \, \mu s$

The current experimental value reads:

$$au_{\mu} =$$
 (2.1969811 \pm 0.0000022) μs



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Different Renormalization Schemes



• The total decay rate of quarks expressed in terms of on-shell masses converges poorly:

$$\Gamma_{\rm sl} \sim 1 - 1.72 \frac{\alpha_s(m_b)}{\pi} - 13.1 \left(\frac{\alpha_s(m_b)}{\pi}\right)^2 - 163 \left(\frac{\alpha_s(m_b)}{\pi}\right)^3$$

Also the $\overline{\mathrm{MS}}$ scheme usually behaves poorly, since the scale has to be chosen rather low.

- Different threshold masses like the PS [Beneke (1998)], 1S [Hoang, Ligeti, Manohar (1998)] or kinetic mass [Bigi, Shifman, Uraltsev, Vainshtein (1996)] have been proposed to improve the convergence.
- We see a much better behavior in the convergence for the schemes used for the global fits of inclusive quantities.
- E.g. for the kinetic mass:

 $m_b^{
m kin}, m_c^{
m kin}: \qquad \Gamma(b
ightarrow c \ell
u) / \Gamma_0 = 0.633 \left(1 - 0.066 - 0.018 - 0.007
ight) \quad pprox 0.575$

$$m_b^{
m kin}, \overline{m}_c(3~{
m GeV}): \qquad \Gamma(b
ightarrow c \ell
u) / \Gamma_0 = 0.700 \left(1 - 0.116 - 0.035 - 0.010
ight) \quad pprox 0.587$$

 $m_b^{\mathrm{kin}}, \overline{m}_c(2~\mathrm{GeV}): \qquad \Gamma(b
ightarrow c\ell
u) / \Gamma_0 = 0.648 \left(1 - 0.087 - 0.018 - 0.0003\right) \ pprox 0.580$

Different Renormalization Schemes



BLM and non-BLM part

$$\Gamma(b \to c\ell\nu) = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} X_0 \begin{bmatrix} 1 + C_F \sum_{n \ge 1} \left(\frac{\alpha_s}{\pi}\right)^n Y_n \end{bmatrix} \qquad \qquad Y_2 = \beta_0 Y_2^{\beta_0} + Y_2^{\text{rem}} \\ Y_3 = \beta_0^2 Y_3^{\beta_0} + Y_3^{\text{rem}} \end{bmatrix}$$

	<i>Y</i> ₁	$Y_2^{\rm rem}$	$\beta_0 Y_2^{\beta_0}$	$Y_3^{ m rem}$	$\beta_0^2 Y_3^{\beta_0^2}$
$m_b^{ m OS}, m_c^{ m OS}$	-1.72	3.08	-16.17	48.8	-212.1
$m_b^{ m kin}, m_c^{ m kin}$	-0.94	0.33	-4.08	-5.4	-15.4
$m_b^{\rm kin}, \overline{m}_c(3 { m GeV})$	-1.67	-3.39	-3.85	-97.7	69.1
$m_b^{ m kin}, \overline{m}_c$ (2 GeV)	-1.25	-1.21	-2.43	-68.8	67.9
$\overline{m}_b(\overline{m}_b), \overline{m}_c(3 \text{ GeV})$	3.07	-21.81	35.17	-56.7	119.4
$m_b^{ m PS}, \overline{m}_c(2~{ m GeV})$	-0.47	-6.10	-2.31	-93.1	-7.19
$m_b^{\mathrm{1S}}, \overline{m}_c(2 \mathrm{GeV})$	-3.59	-0.98	-19.39	-39.83	-80.22
m_b^{1S}, m_c via HQET	-1.38	0.73	-7.05	5.04	-38.09

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Different Renormalization Schemes



BLM and non-BLM part

$$\Gamma(b \to c\ell\nu) = \frac{G_F^2 m_b^5 |V_{cb}|^2}{192\pi^3} X_0 \begin{bmatrix} 1 + C_F \sum_{n \ge 1} \left(\frac{\alpha_s}{\pi}\right)^n Y_n \end{bmatrix} \qquad \qquad Y_2 = \beta_0 Y_2^{\beta_0} + Y_2^{\text{rem}} \\ Y_3 = \beta_0^2 Y_3^{\beta_0} + Y_3^{\text{rem}} \end{bmatrix}$$

	Y ₁	Y_2^{rem}	$eta_{0} Y_{2}^{eta_{0}}$	$Y_3^{ m rem}$	$\beta_0^2 Y_3^{\beta_0^2}$
$m_b^{ m OS}, m_c^{ m OS}$	-1.72	3.08	-16.17	48.8	-212.1
$m_b^{ m kin}, m_c^{ m kin}$	-0.94	0.33	-4.08	-5.4	-15.4
$\mathit{m}^{\mathrm{kin}}_{\mathit{b}}, \overline{\mathit{m}}_{\mathit{c}}(3~GeV)$	-1.67	-3.39	-3.85	-97.7	69.1
$m_b^{ m kin}, \overline{m}_c({ m 2~GeV})$	-1.25	-1.21	-2.43	-68.8	67.9
$\overline{m}_b(\overline{m}_b), \overline{m}_c(3 \text{ GeV})$	3.07	-21.81	35.17	-56.7	119.4
$m_b^{ m PS}, \overline{m}_c(2~{ m GeV})$	-0.47	-6.10	-2.31	-93.1	-7.19
$m_b^{1\mathrm{S}}, \overline{m}_c(2 \text{ GeV})$	-3.59	-0.98	-19.39	-39.83	-80.22
$m_b^{1\mathrm{S}}, m_c$ via HQET	-1.38	0.73	-7.05	5.04	-38.09

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Different Renormalization Schemes – kinetic mass





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Different Renormalization Schemes – 1S mass





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Moments of Differential Distributions



The method can be used to calculate inclusive moments of differential distributions.
 For example we can calculate-q² moments:



Conclusions and Outlook



Conclusions

- We have computed the α_s^3 corrections to the width of $b \rightarrow c \ell \nu$.
- We performed an expansion in the limit $1 m_c/m_b \ll 1$ and demonstrated its good convergence.
- The result is one of the few third order corrections involving two mass scales.
- The results are also relevant for $b
 ightarrow u \ell
 u$ and the muon decay.

Outlook

The method of calculation can be applied for the calculation of moments of the differential distributions.

Backup

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Master Integrals in the ultrasoft region



Denominator structures:

- solid lines: $1/(2k_i \cdot p + 1)$, with $p^2 = 1$
- double lines: $1/(2k_i \cdot p)$, with $p^2 = 1$
- dotted lines: 1/k_i²
- Analytically calculated via direct integration, symbolic summation and differential equations using the packages Sigma [Schneider (2007-)] and HarmonicSums [Ablinger et al (2011-)], verified with PSLQ.
- Only ζ values appear in the results.
- Already needed for the calculation of the relation between the OS and kinetic mass (although to lower order in *ϵ*).



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