

DAAD

Deutscher Akademischer Austauschdienst German Academic Exchange Service

Constraining neutrino transition magnetic moments at DUNE*

@ KSETA Plenary Workshop

Durbach

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29 September 2021

^{*}Work done in collaboration with Thomas Schwetz and Jing-Yu Zhu [arXiv:2105:09699]

• Like quarks, the mass and (weak) interaction basis of neutrinos are not the same.

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

PMNS Matrix converts mass to flavour basis

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Figure: Another way to understand neutrino oscillations, is as the coherent sum of a Feynman diagram with neutrino mass eigenstates as internal lines.

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- Unlike quarks, the quasi-degeneracy of neutrino masses allows coherent propagation of mass states.
- For $E_{\nu} \sim 1 \,\text{GeV}$, the oscillation length is $L_{\text{osc.}} \sim 1000 \,\text{km}$.
- Open question: is there a new fourth light "sterile neutrino", with a shorter oscillation baseline?

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Figure: Made by Alan Stonebraker for the American Physical Society

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- The beam is mostly muon neutrinos with a small electron-neutrino background. Measuring electron-neutrino-like events, they found a 4.8σ excess.



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- External background events and excess misreconstructed highly asymmetric π⁰ decays were disfavoured.

Hypothesis	Factor	$\chi^2/9$ ndf
NC $\Delta ightarrow N\gamma$	3.18	10.0
External Event	5.98	44.9
ν_e from K_L^0 Decay	7.85	14.8
$ u_e$ from K^\pm Decay	2.95	16.3
$ u_e$ from μ^\pm Decay	1.88	16.1
Other $\nu_e \& \bar{\nu}_e$	3.21	12.5
NC π^0 Background	1.75	17.2
Best Fit Oscillations	1.24	8.4

Table: From [arXiv:2006.16883v3]: Log-likelihood shape-only fits to the radial distribution in neutrino mode, assuming only statistical errors, with arbitrary normalisation.

Background: Possible explanations

If excess were due to new ν_μ → ν_e oscillations (via a fourth light neutrino), one should see a deficit in ν_μ flux in other beam experiments: not observed.



Figure: From [arXiv:1803.10661v1]. Right of black line is excluded region. Red islands are preferred regions from MiniBooNE

Background: Possible explanations

- If excess were due to new $\nu_{\mu} \rightarrow \nu_{e}$ oscillations (via a fourth light neutrino), one should see a deficit in ν_{μ} flux in other beam experiments: not observed.
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- Since the MiniBooNE detector only detects the Cerenkov light from a CC-produced electron, an EM shower from a photon can mimic an electron.
- Gnienko [arXiv:0902.3802] first suggested that (in addition to a fourth light neutrino) sterile neutrino to single photon decay could explain the excess.



Figure: From [arXiv:0902.3802v1]

 G. Magill, R. Plestid, M. Pospelov and Y.-D. Tsai [arXiv:1803.03262] did the first systematic study into constraints on the neutrino dipole portal. (Production and detection via dipole operator).



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- They investigated limits from existing accelerator data and the neutrino burst from SN1987A. They found SHiP would be sensitive to a model explaining the MiniBooNE excess.
- Note: purple collider bounds depend on electroweak UV completion.
- My work investigates the neutrino-dipole signal at DUNE.

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Effective operator:

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- We consider effective operator with arbitrary coupling.
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- We ignore these model-building aspects.
- See however [K.S. Babu, S. Jana and M. Lindner arXiv:2007.04291].





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To compare with tracks from NOvA ND (courtesy of A. Himmel's 2018 presentation)

Neutrino Candidates from ND Data



- DUNE: 70kt of cryogenic (88 K or -185 °C) liquid Argon: good resolution and discrimination.
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From [arXiv:2002.02967v3]. Cartoon illustrating the configuration of the LBNF beamline at Fermilab, in Illinois, and the DUNE detectors in Illinois and South Dakota, separated by 1300 km.



- DUNE: 70kt of cryogenic (88 K or -185 °C) liquid Argon: good resolution and discrimination.
- Difficult to produce intense tau-neutrino flux. Therefore consider oscillated flux at far detector.
- Can also use intrinsic beam flux at near-detector.





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Signal (types)

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Signal (types)

- **Outside events:** Up-scattering occurs in the Earth; the signature is a single-photon event.
- Inside events, coherent: The coherent scattering on the nucleus leaves a nuclear recoil of low energy, which is difficult to observe in the detector. The decay leaves a single-photon signature.
- **Inside events, incoherent:** The incoherent scattering on nucleons leads to a signature similar to NC neutrino events, whereas the scattering on electrons results in a single electron. In addition, there is a coincidental displaced single-photon event from the heavy-neutrino decay.

Signal (approximations)

Outside events (no detector geometry):

cross-section beamaxis



Signal (approximations)

Outside events (no detector geometry):



Inside events (only detector geometry, collimated beam):



Signal for d_{τ} (example differential spectra)

Spectra for inside, outside events at various masses, normalised so that the peak is 1. At low energies (dashed), we replace the oscillation probability with 1/2 to account for the averaging of fast oscillations. This is purely cosmetic for inside events.



 E_4 [GeV]

Inside spectra with arbitrary normalisation

Results (decomposition by target and event-type)



6-events/year curves for inside (solid) and outside (dashed) events at the DUNE FD for coherent scattering on nuclei (red), incoherent scattering on nucleons (blue) as well as electrons (purple). Our approximations for outside events break down at the upper curve, as decays occur very close to the detector (cyan line). This effect is negligible as inside-events will dominate. In grey is indicated when the decay-length is of the order of DUNE's spatial resolution.

Global picture (d_{τ})



Global picture (d_{μ})



Global picture (d_e)



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- The intrinsic flux can constrain $d_{\mu,e}$ at the near detector.
- Meson decay ignored though.
- Background studies needed for proper sensitivity curve.



Thanks for your attention!

Signal (equations)

Outside events (no detector geometry):

$$\begin{split} \frac{\mathrm{d}N}{\mathrm{d}E_4} &= N_{\mathrm{mod}} \frac{\rho_N}{2\pi} \int_0^{\theta_b^{\mathrm{max}}} \sin \theta_b \mathrm{d}\theta_b \int_{r_{\mathrm{min}}}^{r_{\mathrm{max}}} \mathcal{L}_{\mathrm{ND}}^2 \mathrm{d}r_p \\ &\sum_{M_{\tau}} \left[\frac{\mathrm{d}^2 \Phi}{\mathrm{d}\Omega_b \mathrm{d}E_{\nu}} \frac{\mathrm{d}E_{\nu}}{\mathrm{d}E_4} P_{\mathrm{osc}} \cdot \varepsilon_{\varphi_b} \cdot P_{\mathrm{decay}}(\ell) \frac{\mathrm{d}\sigma}{\mathrm{d}\cos\theta_s} \Delta\Omega_s \cdot \varepsilon(p_4) \right]_{\tau}. \end{split}$$

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Inside events (only detector geometry, collimated beam):

$$\begin{split} \frac{\mathrm{d}N}{\mathrm{d}E_{\nu}} = & N_{\mathrm{mod}} \frac{L_{\mathrm{ND}}^{2}}{L_{\mathrm{FD}}^{2}} \rho_{N} A_{\mathrm{det}} \left. \frac{\mathrm{d}\Phi}{\mathrm{d}\Omega \mathrm{d}E_{\nu}} \right|_{\theta_{b}=0} P_{\mathrm{osc}} \left(\frac{L_{\mathrm{FD}}}{E_{\nu}} \right) \\ & \sum_{M_{T}} \int_{0}^{L_{d}} \mathrm{d}z \int_{-1}^{1} \mathrm{d}\cos\theta_{s} \frac{\mathrm{d}\sigma_{T}}{\mathrm{d}\cos\theta_{s}} \Pi\left(\ell_{d}^{0}\right) P_{\mathrm{dec}}(\ell_{d}^{0}) \varepsilon(p_{4}) \,. \end{split}$$