

Strong first-order electroweak phase transition in selected extended scalar sector models

Lisa Biermann¹

¹Institute for Theoretical Physics (ITP)
Karlsruhe Institute of Technology (KIT)
`lisa.biermann@kit.edu`

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Electroweak Baryogenesis

Observation of a **non-zero baryon asymmetry** of the universe (**BAU**)

$$\eta \equiv \frac{n_b - \bar{n}_b}{n_\gamma} \simeq \frac{n_b}{n_\gamma} \simeq 6.1 \times 10^{-10} \quad [\text{Planck, 2018}]$$

What happened to the antimatter in our universe?

How can we generate a non-zero baryon asymmetry of the universe?

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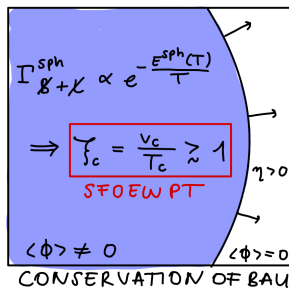
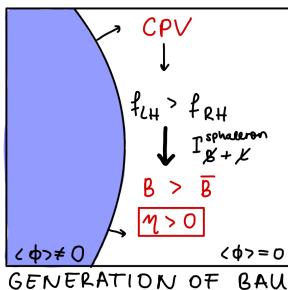
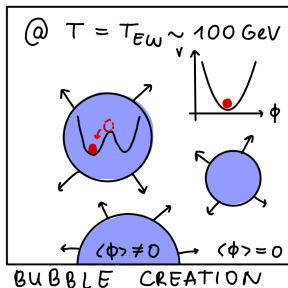
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Electroweak Baryogenesis (EWBG) [A. D. Sakharov, 1967], [D. Morrissey, M. Ramsey-Musolf, 2012]



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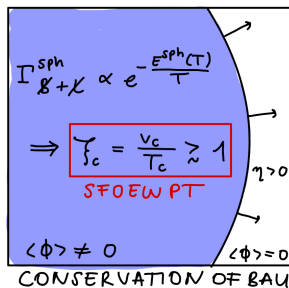
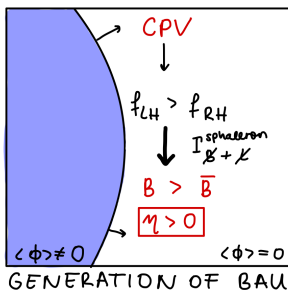
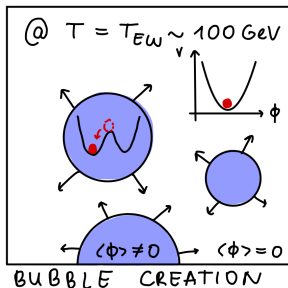
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Electroweak Baryogenesis (EWBG) [A. D. Sakharov, 1967], [D. Morrissey, M. Ramsey-Musolf, 2012]



⇒ **BSM physics** required for **additional CP violation** and **strong first-order electroweak phase transition (SFOEWPT)**

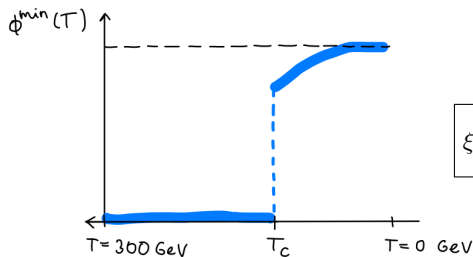
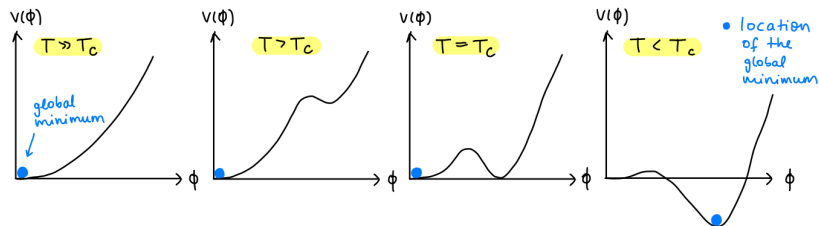
→ **Extended Higgs sector models** as promising road to BSM

Investigating SFOEWPTs with BSMPT

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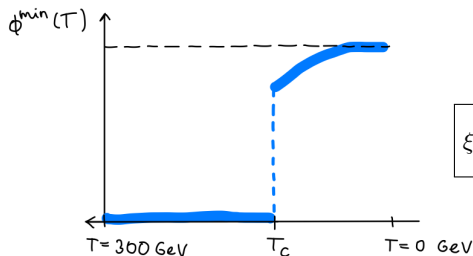
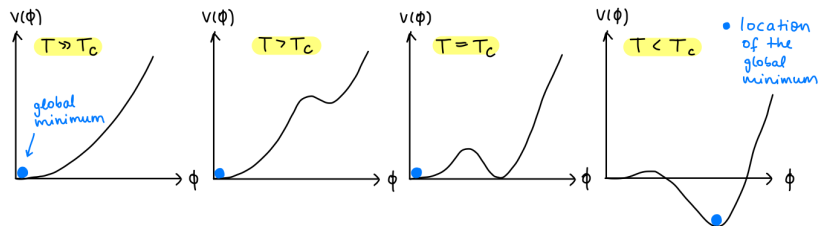


$$\xi_c \equiv \frac{\phi^{\min}(T = T_c)}{T_c} \gtrsim 1$$

SFOEWPT

Investigating SFOEWPTs with BSMPT

How do we look for *strong first-order electroweak phase transitions* (SFOEWPTs)?



$$\xi_c \equiv \frac{\phi^{\min}(T = T_c)}{T_c} \gtrsim 1$$

SFOEWPT

⇒ global minimization of the one-loop corrected effective potential @ $T \neq 0$ with **BSMPT**

[P. Basler, M. Mühlleitner, J. Müller, 2018/20] <https://github.com/phbasler/BSMPT>

Searching for an SFOEWPT in extended scalar sector models

⇒ **Extended Higgs sector models** as promising road to BSM:

Searching for an SFOEWPT in extended scalar sector models

⇒ **Extended Higgs sector models** as promising road to BSM:

- **2HDM-type II EFT** (*): $\Phi_1, \Phi_2 + \frac{O_6}{\Lambda^2}$

→ struggle to reach SFOEWPT (compared to type I-2HDMs) [P. Basler et al., 2017]

→ look@2HDM-type II parameter points with $\xi_c < 1$:

What extra dynamics are required to achieve an **SFOEWPT**?

⇒ *bottom-up* extension of the 2HDM scalar potential by purely scalar dim-6 operators in an EFT approach

What are the phenomenological implications on **Higgs-Pair production**?

(*) *Two Higgs doublets, Effective Interactions and a Strong First-Order Electroweak Phase Transition*
by Anisha, LB, Christoph Englert and Margarete Mühlleitner

[JHEP 08 (2022) 091]

Searching for an SFOEWPT in extended scalar sector models

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- **CP in the Dark** (2HDM + real singlet)^(**): Φ_1, Φ_2, Φ_S

- provides *viable* particle **DM** candidate and **explicit** CP violation *solely* in the dark sector @ $T = 0$

⇒ Can we generate an **SFOEWPT** within ‘CP in the Dark’?

⇒ Can the ‘hidden’ **CP violation** be translated to the visible sector?

^(**) *Electroweak Phase Transition in a Dark Sector with CP Violation*

by LB, Margarete Mühlleitner and Jonas Müller

[2204.13425]

- **CP-conserving 2HDM**, softly broken discrete $\mathbb{Z}_2$ symmetry: $\Phi_1 \rightarrow -\Phi_1, \Phi_2 \rightarrow \Phi_2$
[T. D. Lee, 1973], [G. C. Branco et al., 2012]

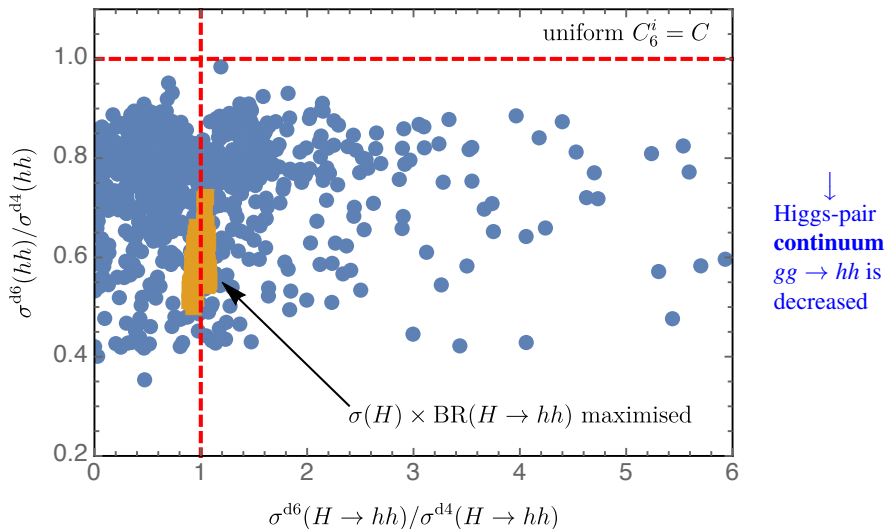
$$V_{\text{tree}}(\Phi_1, \Phi_2) = m_{11}^2(\Phi_1^\dagger \Phi_1) + m_{22}^2(\Phi_2^\dagger \Phi_2) - m_{12}^2(\Phi_1^\dagger \Phi_2 + \Phi_2^\dagger \Phi_1) + \lambda_1(\Phi_1^\dagger \Phi_1)^2 + \lambda_2(\Phi_2^\dagger \Phi_2)^2 \\ + \lambda_3(\Phi_1^\dagger \Phi_1)(\Phi_2^\dagger \Phi_2) + \lambda_4(\Phi_1^\dagger \Phi_2)(\Phi_2^\dagger \Phi_1) + \frac{1}{2}\lambda_5[(\Phi_1^\dagger \Phi_2)^2 + (\Phi_2^\dagger \Phi_1)^2]$$

- inclusion of **(purely scalar) dim-6 EFT contributions** to the Higgs potential [Anisha et al., 2019]

$$\mathcal{L}_{\text{EFT}} = \mathcal{L}_{\text{2HDM}} + \sum_i \frac{C_6^i}{\Lambda^2} O_6^i \quad \Rightarrow \quad V_{\text{dim-6}} = - \sum_i \frac{C_6^i}{\Lambda^2} O_6^i$$

O_6^{111111}	$(\Phi_1^\dagger \Phi_1)^3$	O_6^{222222}	$(\Phi_2^\dagger \Phi_2)^3$
O_6^{111122}	$(\Phi_1^\dagger \Phi_1)^2(\Phi_2^\dagger \Phi_2)$	O_6^{112222}	$(\Phi_1^\dagger \Phi_1)(\Phi_2^\dagger \Phi_2)^2$
O_6^{122211}	$(\Phi_1^\dagger \Phi_2)(\Phi_2^\dagger \Phi_1)(\Phi_1^\dagger \Phi_1)$	O_6^{122122}	$(\Phi_1^\dagger \Phi_2)(\Phi_2^\dagger \Phi_1)(\Phi_2^\dagger \Phi_2)$
O_6^{121211}	$(\Phi_1^\dagger \Phi_2)^2(\Phi_1^\dagger \Phi_1) + \text{h.c.}$	O_6^{121222}	$(\Phi_1^\dagger \Phi_2)^2(\Phi_2^\dagger \Phi_2) + \text{h.c.}$

- absorb dim-6 contributions (to scalar masses) in shifts $\lambda_i \rightarrow \lambda_i + \delta\lambda_i, m_{12}^2 \rightarrow m_{12}^2 + \delta m_{12}^2$
- ⇒ scalar mass spectrum same as for dim-4 @ LO
 ⇒ shift EFT effects into **Higgs self-couplings & multi-Higgs final states**



→ (exotic) Higgs **resonance** increased (within (future) LHC-sensitivity!)

- N2HDM-like extended scalar sector, *one* discrete $\mathbb{Z}_2$ symmetry

$$\Phi_1 \rightarrow +\Phi_1, \quad \Phi_2 \rightarrow -\Phi_2, \quad \Phi_S \rightarrow -\Phi_S$$

- $SU(2)_L \times U(1)_Y$ and $\mathbb{Z}_2$ -invariant tree-level potential:

$$V^{(0)} = m_{11}^2 |\Phi_1|^2 + m_{22}^2 |\Phi_2|^2 + \frac{m_S^2}{2} \Phi_S^2 + \left(A \Phi_1^\dagger \Phi_2 \Phi_S + h.c. \right) + \frac{\lambda_1}{2} |\Phi_1|^4 + \frac{\lambda_2}{2} |\Phi_2|^4 \\ + \lambda_3 |\Phi_1|^2 |\Phi_2|^2 + \lambda_4 |\Phi_1^\dagger \Phi_2|^2 + \frac{\lambda_5}{2} \left[\left(\Phi_1^\dagger \Phi_2 \right)^2 + h.c. \right] + \frac{\lambda_6}{4} \Phi_S^4 + \frac{\lambda_7}{2} |\Phi_1|^2 \Phi_S^2 + \frac{\lambda_8}{2} |\Phi_2|^2 \Phi_S^2$$

- general vacuum structure @ $T \neq 0$:

$$\Phi_1 = \frac{1}{\sqrt{2}} \left(\begin{matrix} \rho_1 + i\eta_1 \\ \zeta_1 + \omega_1 + i\Psi_1 \end{matrix} \right), \quad \Phi_2 = \frac{1}{\sqrt{2}} \left(\begin{matrix} \rho_2 + \omega_{\text{CB}} + i\eta_2 \\ \zeta_2 + \omega_2 + i(\Psi_2 + \omega_{\text{CP}}) \end{matrix} \right), \quad \Phi_S = \zeta_S + \omega_S$$

↖ charge-breaking VEV, $\omega_{\text{CB}} = 0$

↗ CP-violating VEV

‘CP in the Dark’ in Detail

- general vacuum structure @ $T = 0$:

$$\Phi_1 = \frac{1}{\sqrt{2}} \begin{pmatrix} \rho_1 + i\eta_1 \\ \zeta_1 + v_1 + i\Psi_1 \end{pmatrix}, \quad \Phi_2 = \frac{1}{\sqrt{2}} \begin{pmatrix} \rho_2 + i\eta_2 \\ \zeta_2 + i\Psi_2 \end{pmatrix}, \quad \Phi_S = \zeta_S$$
$$\langle \Phi_1 \rangle|_{T=0} = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v_1 \end{pmatrix}, \quad \langle \Phi_2 \rangle|_{T=0} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}, \quad \langle \Phi_S \rangle|_{T=0} = 0$$

→ $\omega_1|_{T=0 \text{ GeV}} = v_1 \equiv v = 246.22 \text{ GeV}$, SM-Yukawa sector and tree-level FCNCs prohibited

→ $\mathbb{Z}_2$ symmetry *unbroken* $\Rightarrow$ conserved quantum number: **dark charge**

- * Φ_1 (**SM-like particles** with **+1**): $G^\pm, G^0, h$
- * Φ_2, Φ_S (**dark particles** with **-1**): $H^\pm, h_1, h_2, h_3$ ($m_{h_1} < m_{h_2} < m_{h_3}$)

$\Rightarrow$ **DM**: *stable* particle dark matter candidate h_1

$\Rightarrow$ **explicit CPV**: introduced through **Im(A) $\neq 0$**

→ CPV after SSB, but vacuum is CP-symmetric $\Rightarrow$ CPV is **explicit**

→ solely in the dark sector - h_1, h_2, h_3 : states with mixed CP quantum number

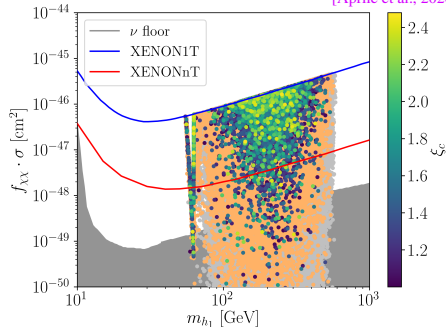
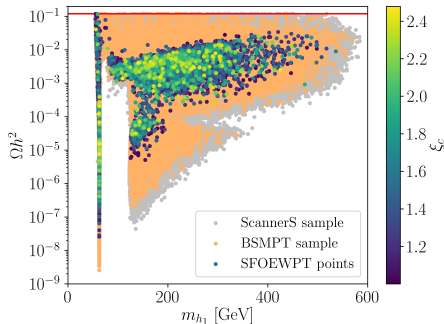
$\Rightarrow$ *not* constrained by EDM constraints

SFOEWPT points and DM Observables [LB, M. Mühlleitner, J. Müller, 2022]

$$\Omega_{\text{obs}} h^2 = 0.1200 \pm 0.0012$$

[Aghanim et al., 2018]

[Billard et al., 2013],
[Aprile et al., 2018],
[Aprile et al., 2020]



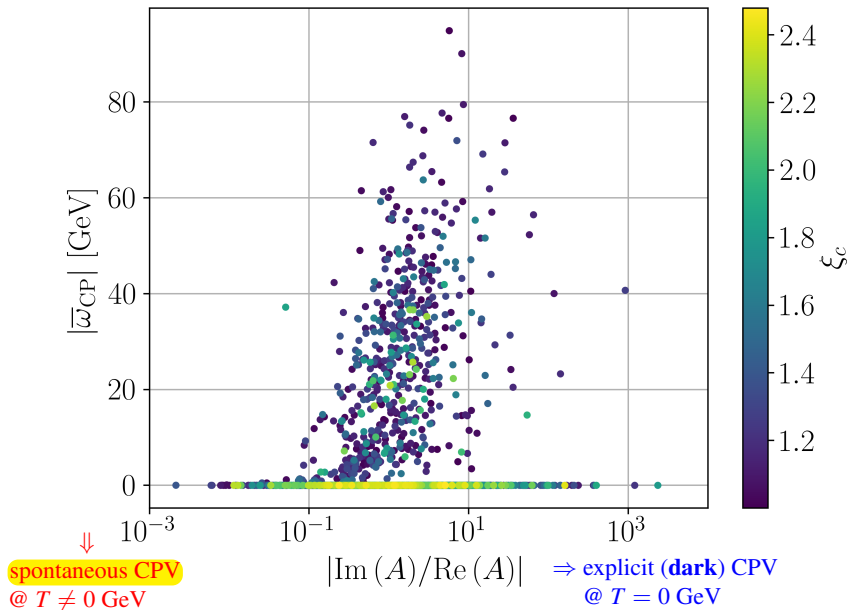
Viable SFOEWPT parameter points

- ⇒ compatible with *relic density* ($< \Omega h^2$)
- ⇒ above neutrino floor
- ⇒ testable at future *direct detection* experiments

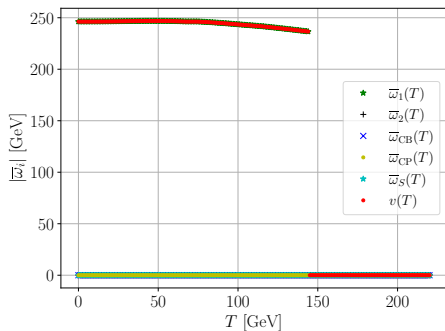
$$f_{\chi\chi} \cdot \sigma_{\text{SI, DM-nucl.}} \equiv \frac{\Omega_{\text{prod}} h^2}{\Omega_{\text{obs}} h^2} \cdot \sigma_{\text{SI, DM-nucl.}}$$

viable parameter points:

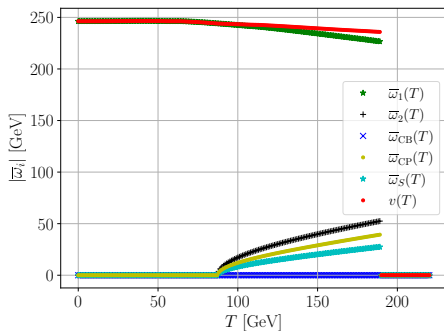
- * pass constraints imposed by: **ScannerS** [R. Coimbra et al., 2013] [M. Mühlleitner et al., 2020], **BSMPPT** [P. Basler, M. Mühlleitner, J. Müller, 2018/20]
- * $\text{BR}(h \rightarrow \text{inv.}) < 0.11$ [M. Aaboud et al., 2019]
- * $\mu_{h \rightarrow \gamma\gamma} = 1.12 \pm 0.09$ [A. Sirunyan et al., 2021]



- Two different VEV patterns in detail:



only ω_1 develops non-zero values



dark VEVs (except ω_{CB}) participate in EWPT
spontaneous CPV!

- @ $T \neq 0$: $\mathbb{Z}_2$ symmetry is broken → dark charge no longer conserved → dark sector **mixes** with SM-like particles
- ⇒ **additional non-standard CPV transferred to the SM-like couplings to fermions @ $T \neq 0$!**

Conclusion

- Dynamical generation of the baryon asymmetry of the universe (BAU) possible if *Sakharov conditions* fulfilled
 - *Electroweak Baryogenesis*: fulfill Sakharov conditions with
 - **BSM models**
 - **Non-standard *CP*-violation (CPV)**
 - *Strong first-order electroweak phase transition (SFOEWPT)*
 - **Extended Higgs sector models** as promising framework for EWBG:
 - **CP in the Dark**: viable **SFOEWPT** points that additionally show *spontaneous CPV* at finite temperature
- *Can these points successfully generate the BAU?*
- **Type-II R2HDM-EFT**: $|1 - \xi_c^{d4}|$ **minimised** by scalar dim-6 contributions
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Thanks for Your attention!

Baryon Asymmetry of the Universe and Sakharov Conditions

initial: *Big Bang* (**symmetric** universe) $\Leftrightarrow$ today: **BAU** (**asymmetric** universe)

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How can we generate a non-zero *baryon asymmetry of the universe*?

[Sakharov, 1967]: Dynamical generation of a BAU with an initially symmetric state possible if

Condition

Existence of B violating processes $\Rightarrow$ Sphaleron-mediated @ $T > T_{\text{EW}} = 100 \text{ GeV}$
[N. Manton, 1983], [F. Klinkhammer, N. Manton, 1984]

$\mathcal{C}$ and $\mathcal{CP}$ violation (CPV) $\Rightarrow$ *Cabibbo-Kobayashi-Maskawa* mechanism
[N. Cabibbo, 1963], [M. Kobayashi, T. Maskawa, 1973]

Departure from thermal equilibrium $\Rightarrow$ Electroweak phase transition (EWPT)
[D. Kirznits, 1972], [L. Dolan, R. Jackiw, 1974]

EWBG and an SFOEWPT

- EWBG takes place around $T \sim T_{EW}$
- EWPT happens and bubbles with non-zero vacuum expectation value (VEV) are created and expand
- Necessary departure from thermal equilibrium achieved through **strong first-order** EWPT (SFOEWPT)

→ ‘**first-order**’: discontinuity in VEV v at T_c :

$$V(v = 0, T_c) = V(v \neq 0, T_c)$$

- How do we see this in the potential?

→ global minimum jumps from symmetric to broken minimum @ T_c

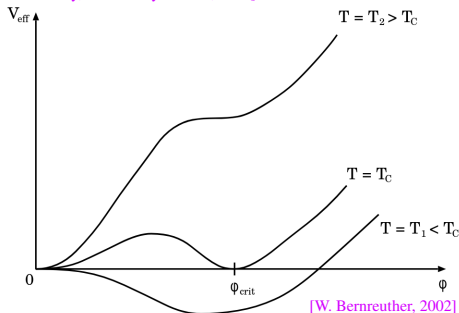
→ ‘**strong**’: conservation of BAU through sufficient suppression of the sphaleron rate inside the bubbles

$$\Gamma_{\cancel{B}+L}^{\text{sph}} \propto \exp - \frac{E_{\text{sph}}(T)}{T} \quad \Rightarrow \quad \xi_c \equiv \frac{v_c}{T_c} \gtrsim 1 \quad \text{baryon-wash-out condition}^* \quad [\text{M. Quiros, 1994}]$$

⇒ EWPT in SM only *smooth cross-over* [K. Kajantie et al., 1996]

⇒ Need BSM models that enable an **SFOEWPT*** + *non-standard CPV*

[D. Morrissey, M. Ramsey-Musolf, 2012]



Minimizing the One-Loop Corrected Effective Potential @ $T \neq 0$

- true vacuum state @ finite temperature (FT) including radiative corrections = global minimum of the **effective potential** @ $T \neq 0$
- general one-loop effective potential @ $T \neq 0$ splits into temperature-dependent and independent part [L. Dolan, R. Jackiw, 1974]

↖ optional finite shift

$$V^{(1)}(\omega, T) = \underbrace{V^{(0)}(\omega)}_{\text{tree-level}} + \underbrace{V^{\text{CW}}(\omega)}_{\substack{T\text{-indep.} \\ \text{Coleman-Weinberg} \\ \text{potential} \\ \text{renormalized in } \overline{MS}\text{-scheme}}} + \underbrace{V^T(\omega, T)}_{\substack{T\text{-dep.} \\ \text{UV finite} \\ \text{IR finite after resummation} \\ m^2 \rightarrow m^2 + \Pi^{(1)}(0)}} + \underbrace{V^{\text{CT}}(\omega)}_{\substack{\text{finite shift of} \\ \text{scalar masses} \\ \text{and mixing angles}}}$$

[S. Coleman, E. Weinberg, 1973] [M. Carrington, 1992],
[R. Parwani, 1992],
[P. Arnold, O. Espinosa, 1993]

[P. Basler et al., 2017]

- V^{CT} absorbs NLO scalar mass and angle shift [P. Basler et al., 2017]

$$0 = \partial_{\phi_i} (V^{\text{CW}} + V^{\text{CT}}|_{\vec{\omega}=\vec{\omega}_{\text{tree}}})$$

$$0 = \partial_{\phi_i} \partial_{\phi_j} (V^{\text{CW}} + V^{\text{CT}}|_{\vec{\omega}=\vec{\omega}_{\text{tree}}})$$