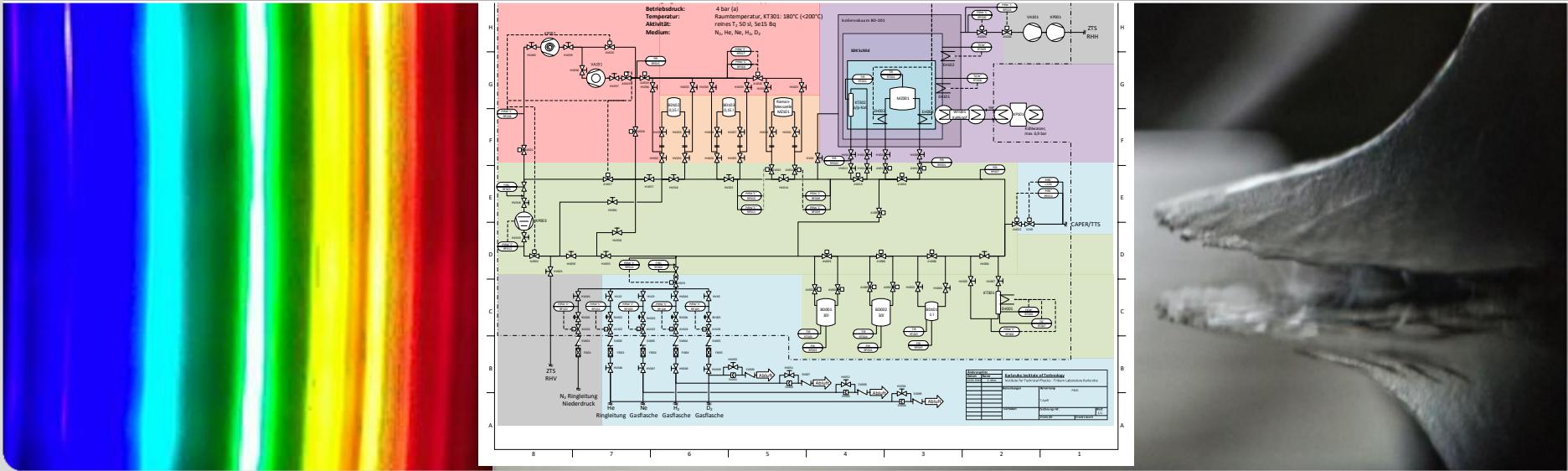
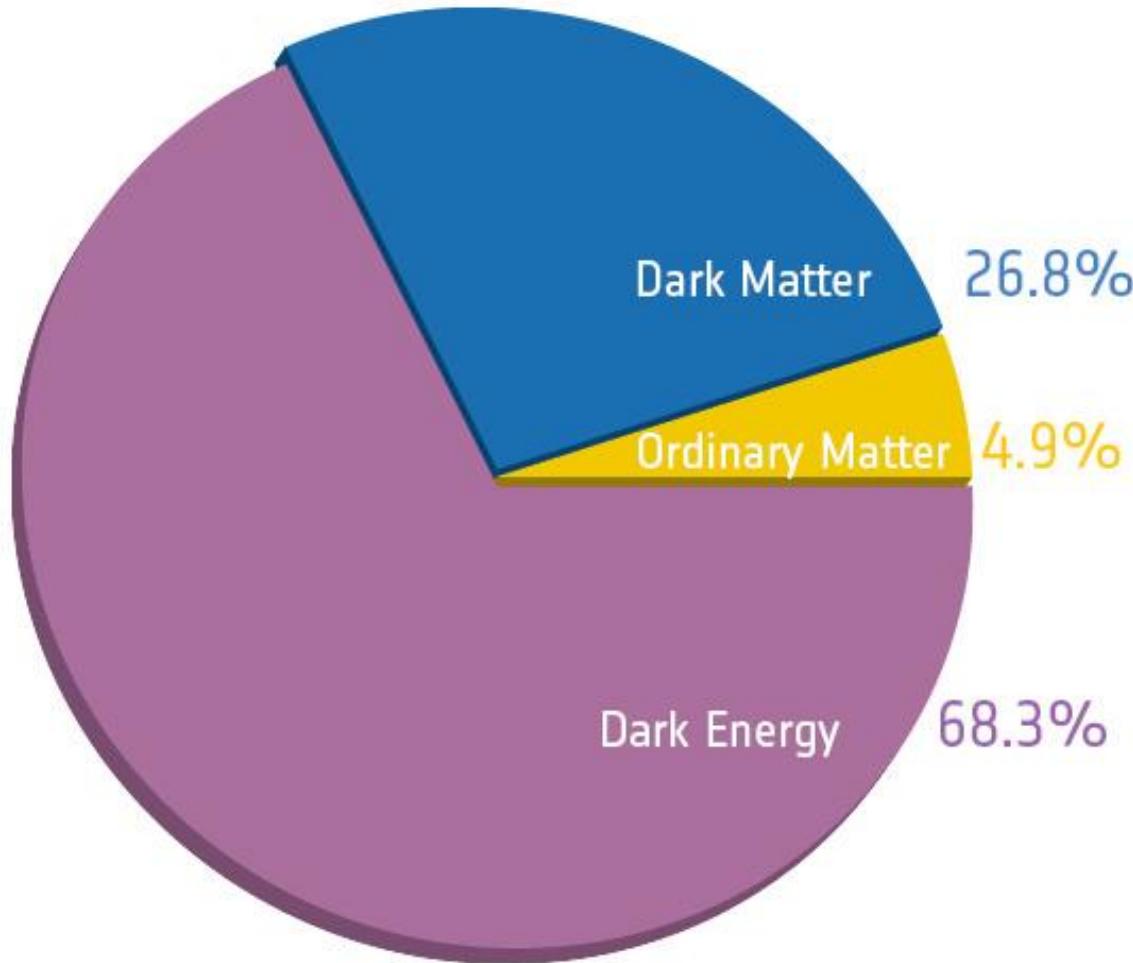


# Spectroscopic Investigation of Molecular Effects in Cold Hydrogen Isotopologues

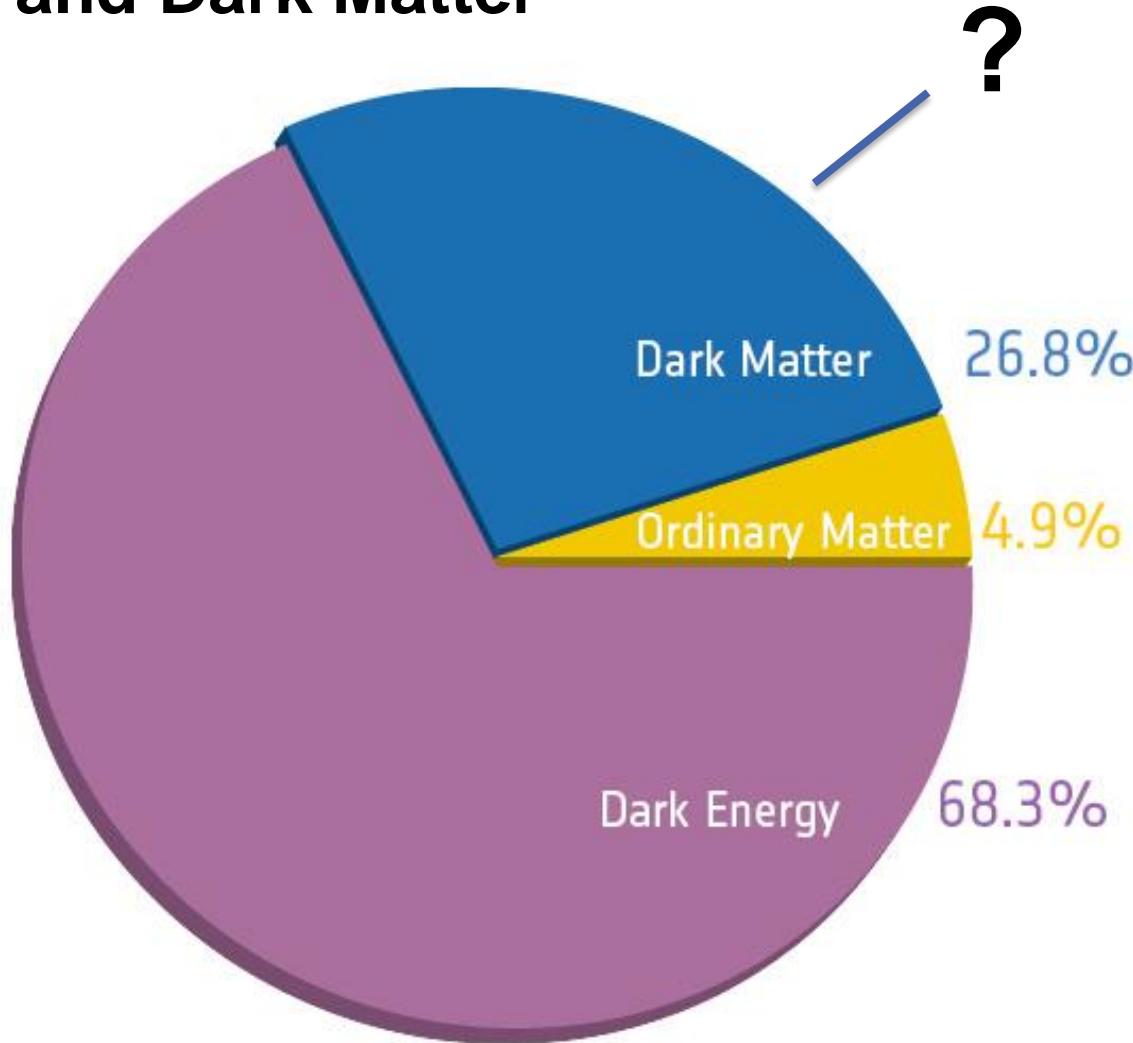
Sebastian Mirz – Tritium Laboratory Karlsruhe



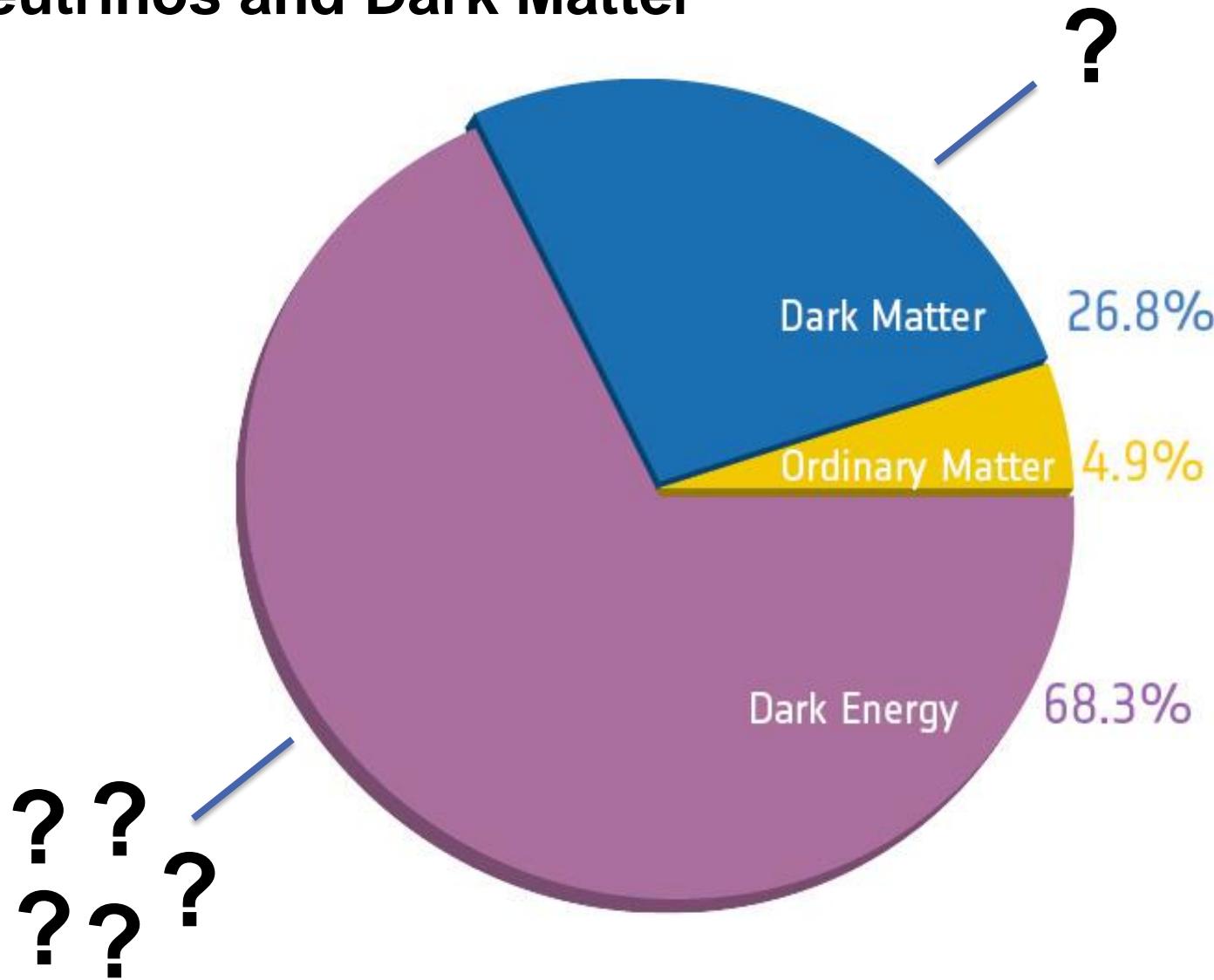
# Neutrinos and Dark Matter



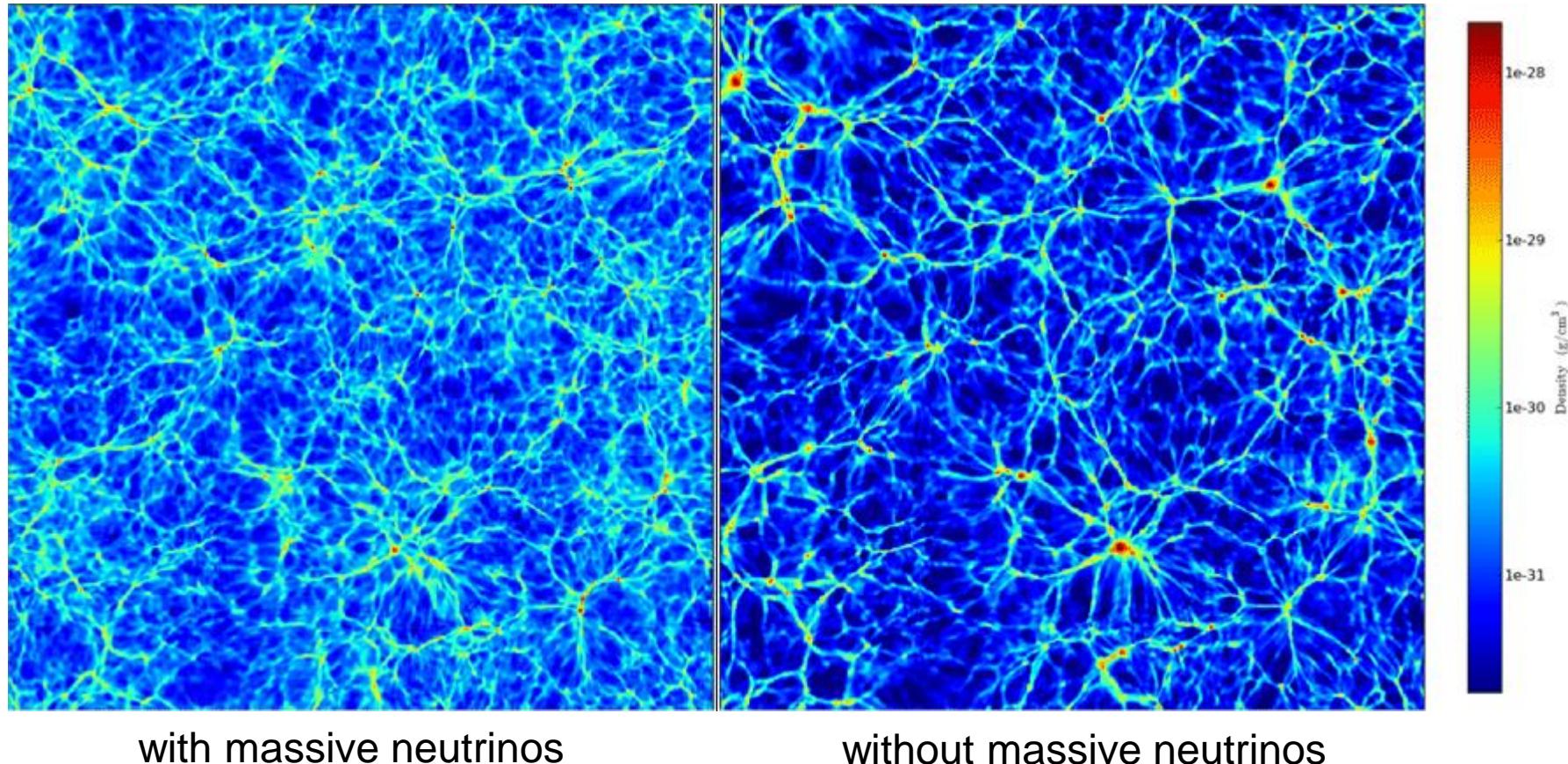
# Neutrinos and Dark Matter



# Neutrinos and Dark Matter



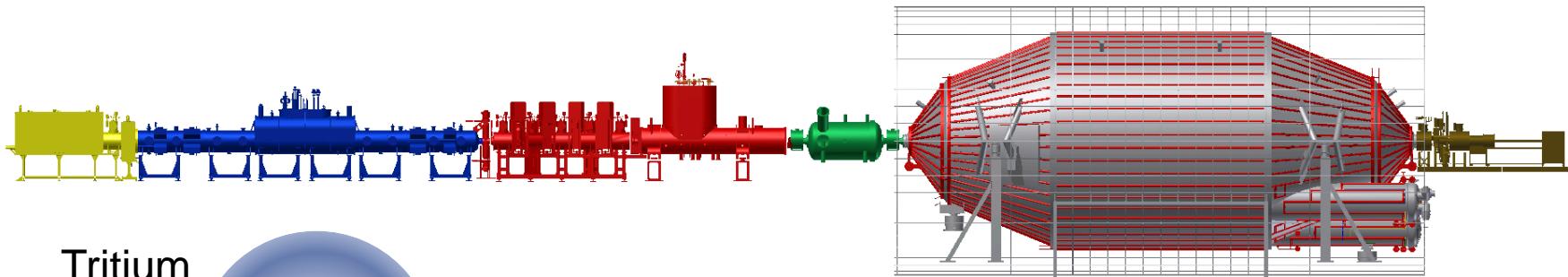
# Structure formation



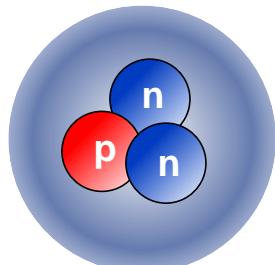
# How to weigh a neutrino?



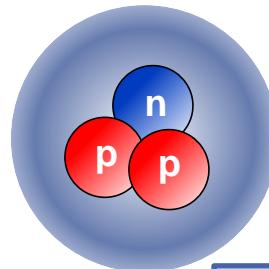
# The KArlsruhe TRItium Neutrino Experiment



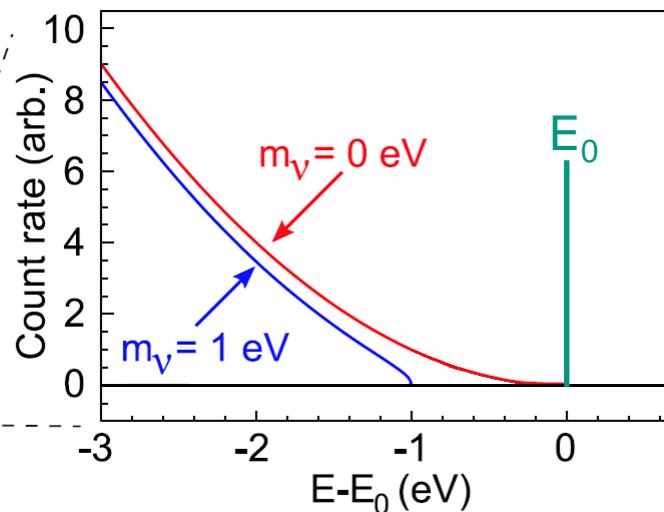
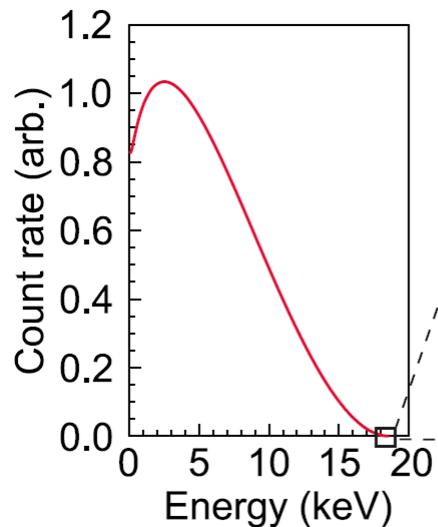
Tritium



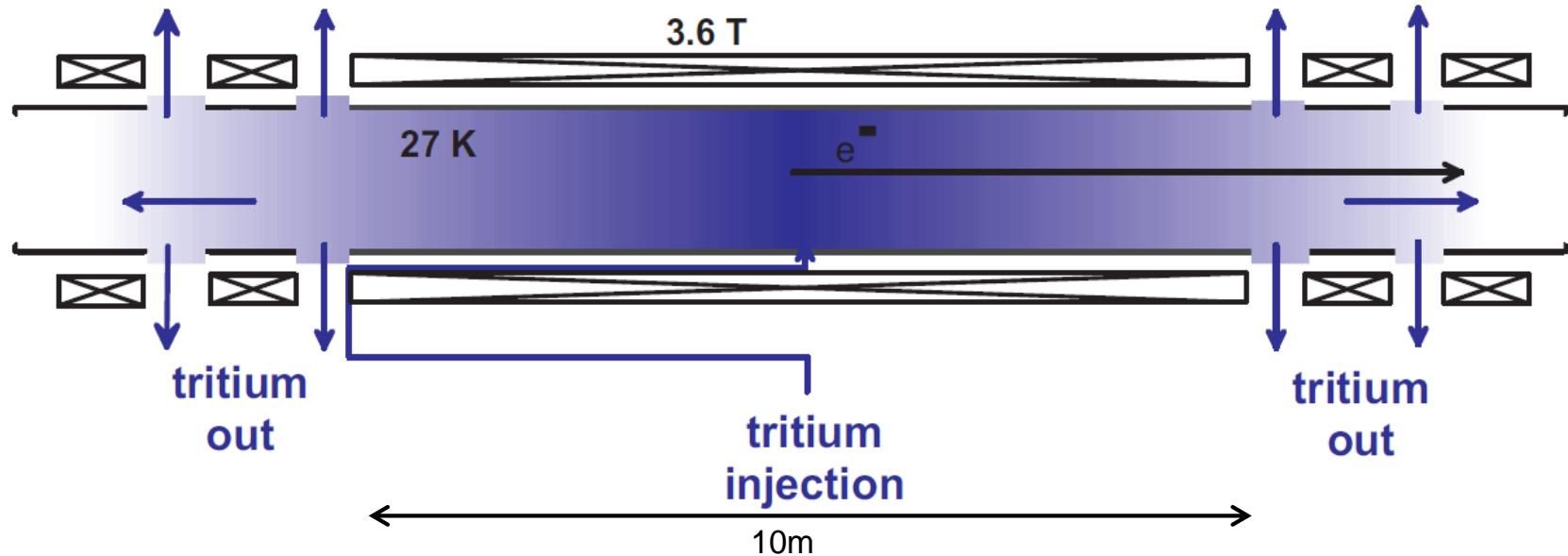
Helium-3



$$E_0 = 18.6 \text{ keV}$$

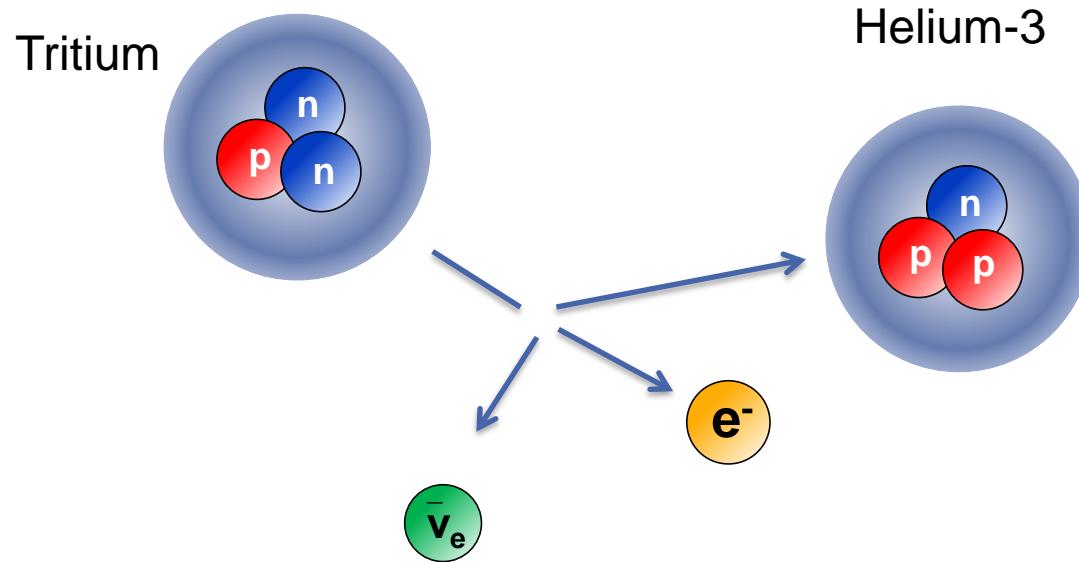


# KATRIN WGTS

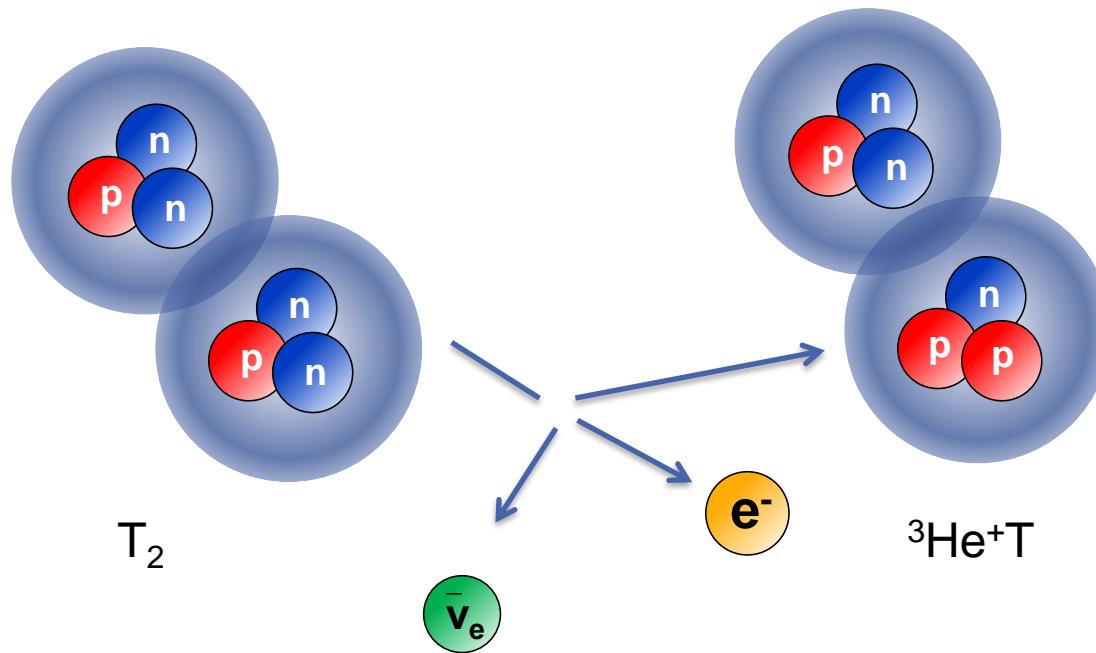


- ~27-30K
- ~ $10^{-3}$  mbar inlet pressure
- Tritium purity: 95 %

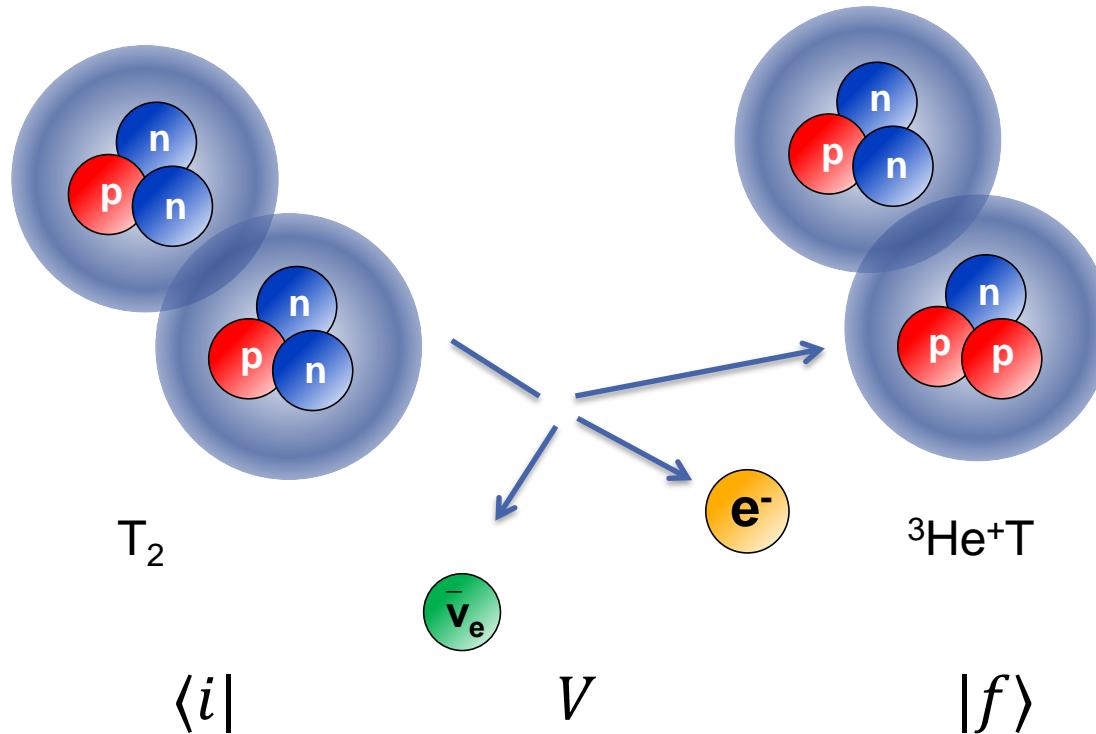
# Tritium decay



# Molecular beta decay



# Molecular beta decay



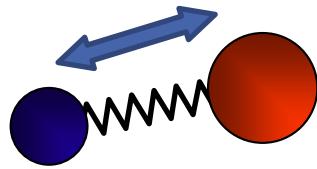
Fermi's golden rule:

$$W_{i \rightarrow f} = \frac{2\pi}{\hbar} |\langle i | V | f \rangle|^2 \rho(E_f)$$

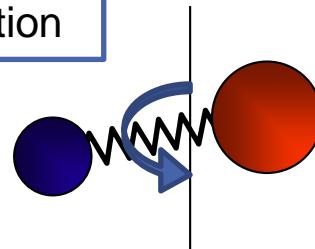
Beta spectrum depends on initial and final state distribution

# Initial state distribution

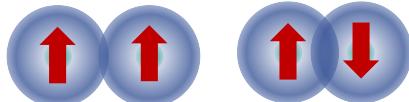
Vibration



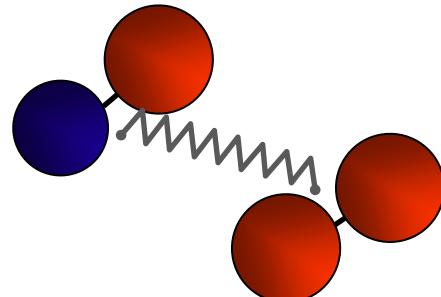
Rotation



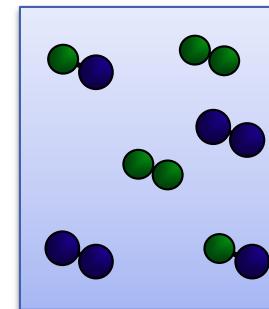
ortho/para  $T_2$ ,  $D_2$ ,  $H_2$



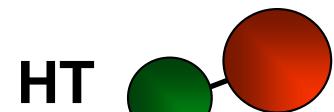
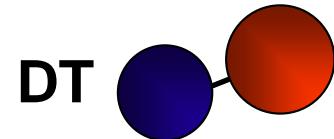
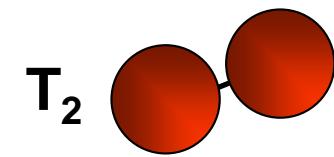
Van-der-Waals clusters



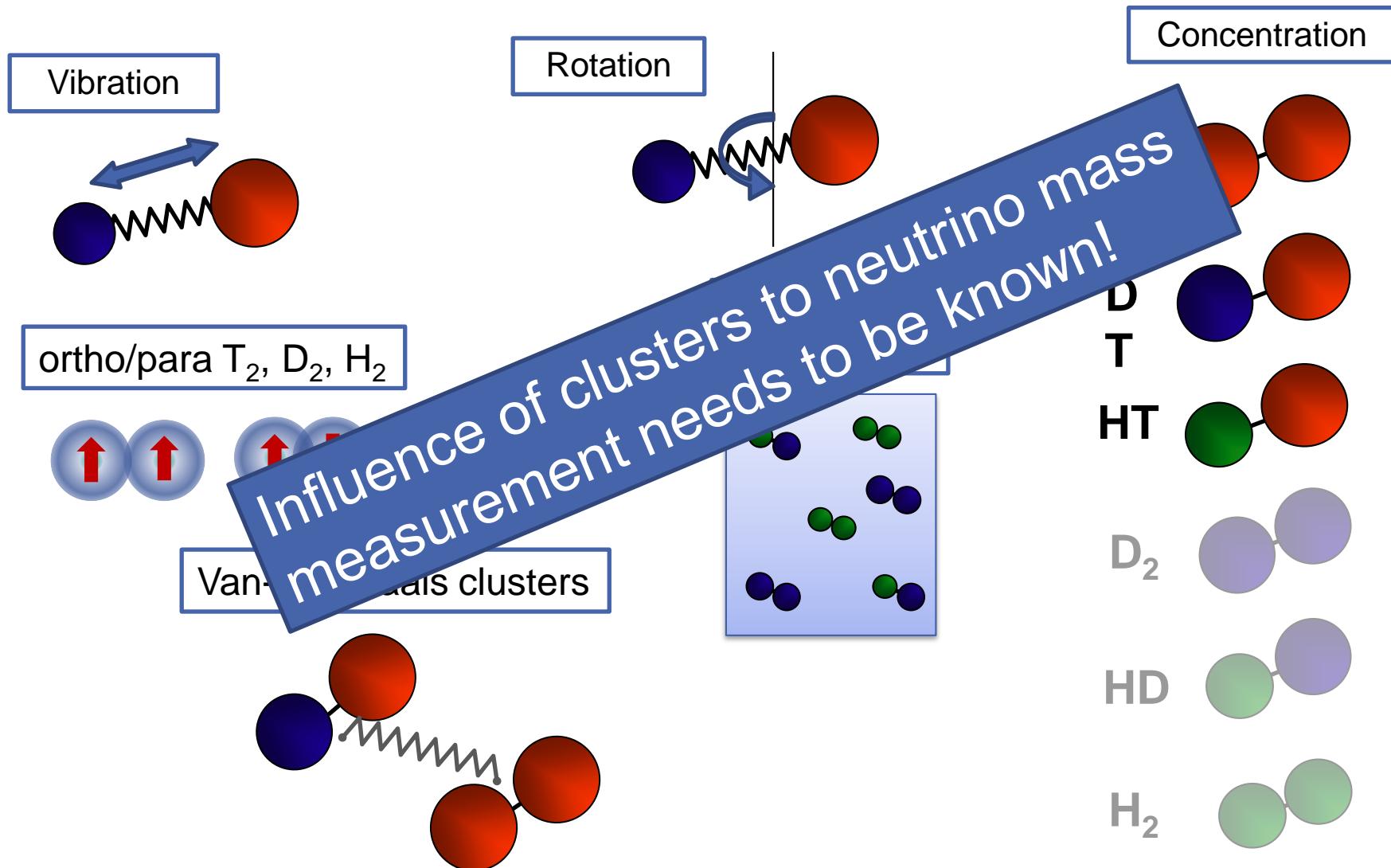
Temperature + pressure



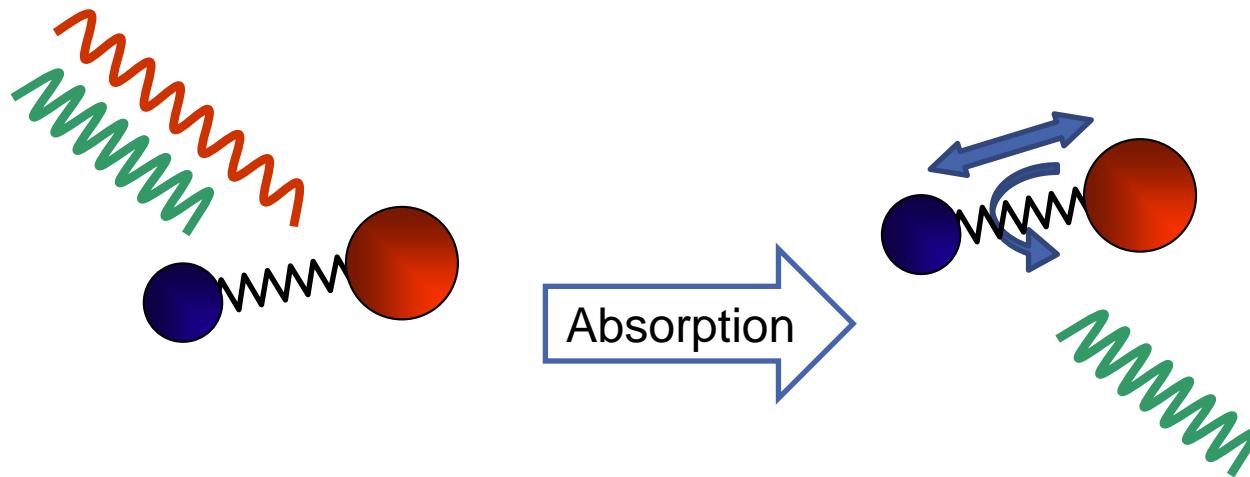
Concentration



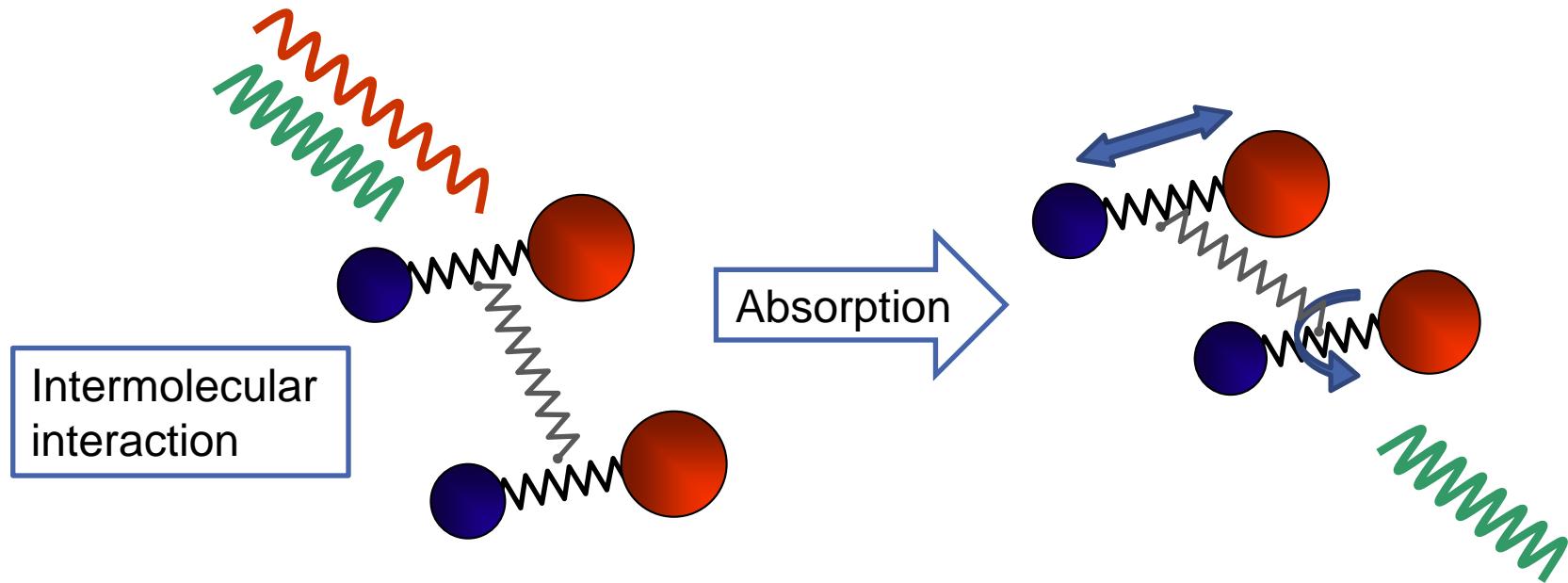
# Initial state distribution



# Experimental method: IR absorption spectroscopy



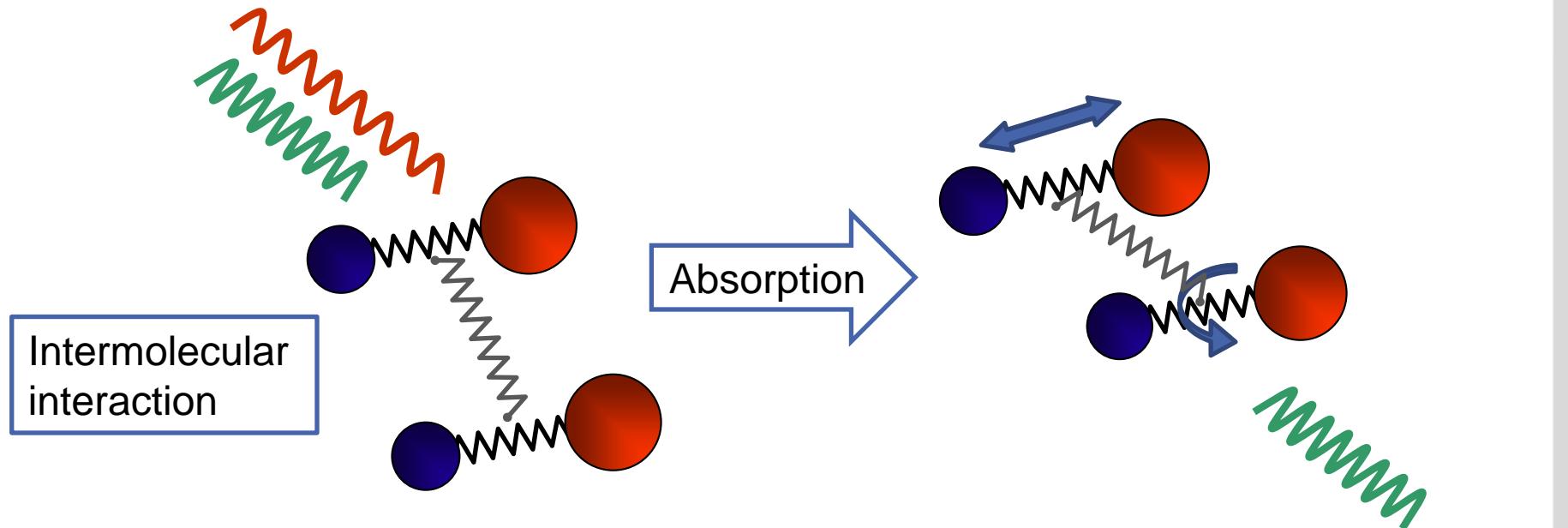
# Experimental method: IR absorption spectroscopy



Sensitive to:

- Intermolecular interaction  
→ Van-der-Waals complex
- Composition
- Vibration / rotation
- Ortho/para ratio
- Temperature

# Experimental method: IR absorption spectroscopy



Sensitive to:

- Intermolecular interaction  
→ Van-der-Waals complex
- Composition
- Vibration / rotation
- Ortho/para ratio
- Temperature

Ideal method: IR spectra only possible with intermolecular interaction

# Strategy

## ■ Demonstation in the liquid phase

- High cluster density
- High signal expected

With existing  
TApIR experiment

## ■ Gas phase without tritium

- Temperature and pressure dependency studies

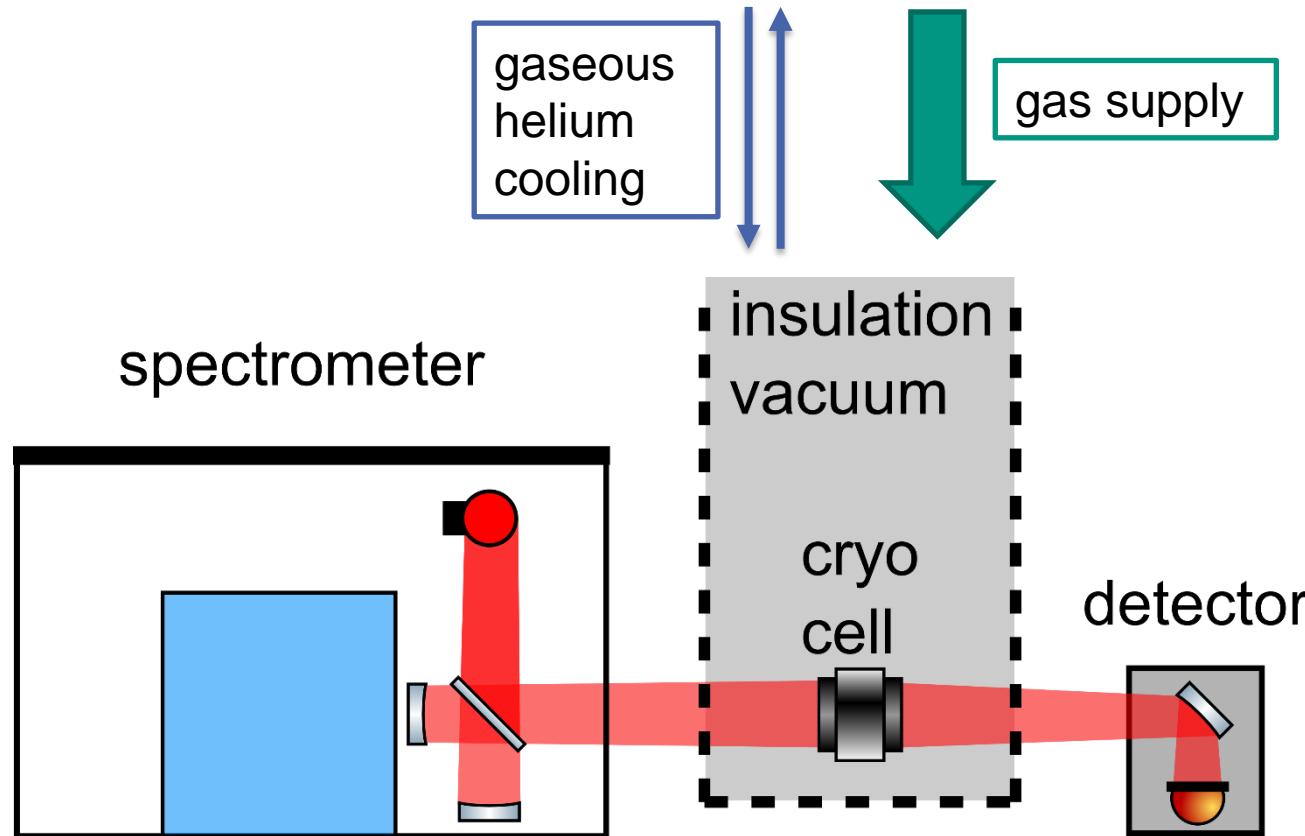
## ■ Gas phase with tritium

- New  $T_2$ ApIR experiment

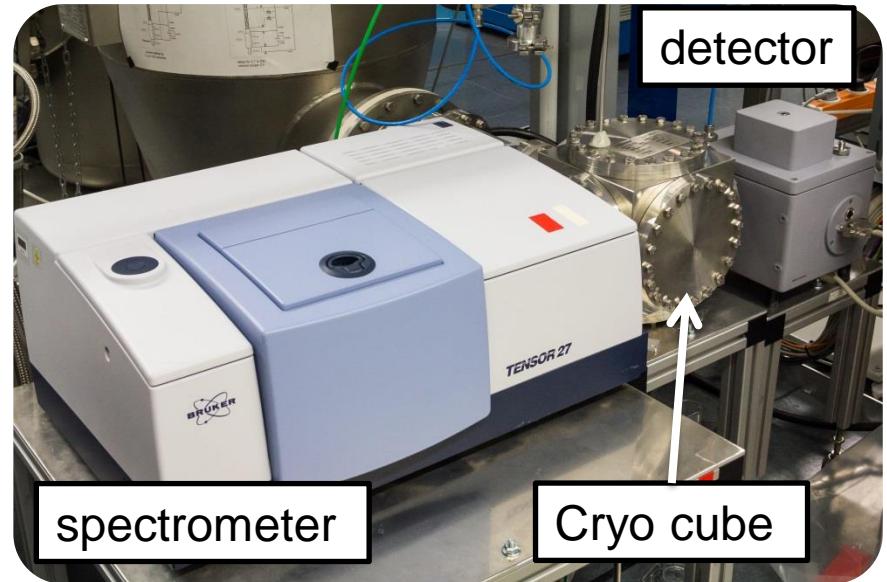
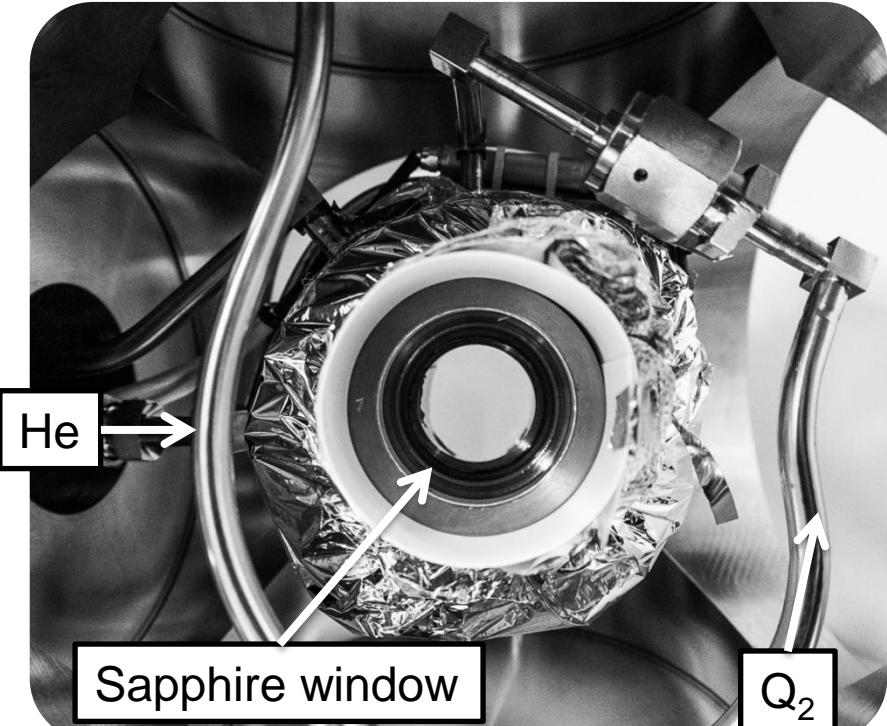


# TApIR Setup

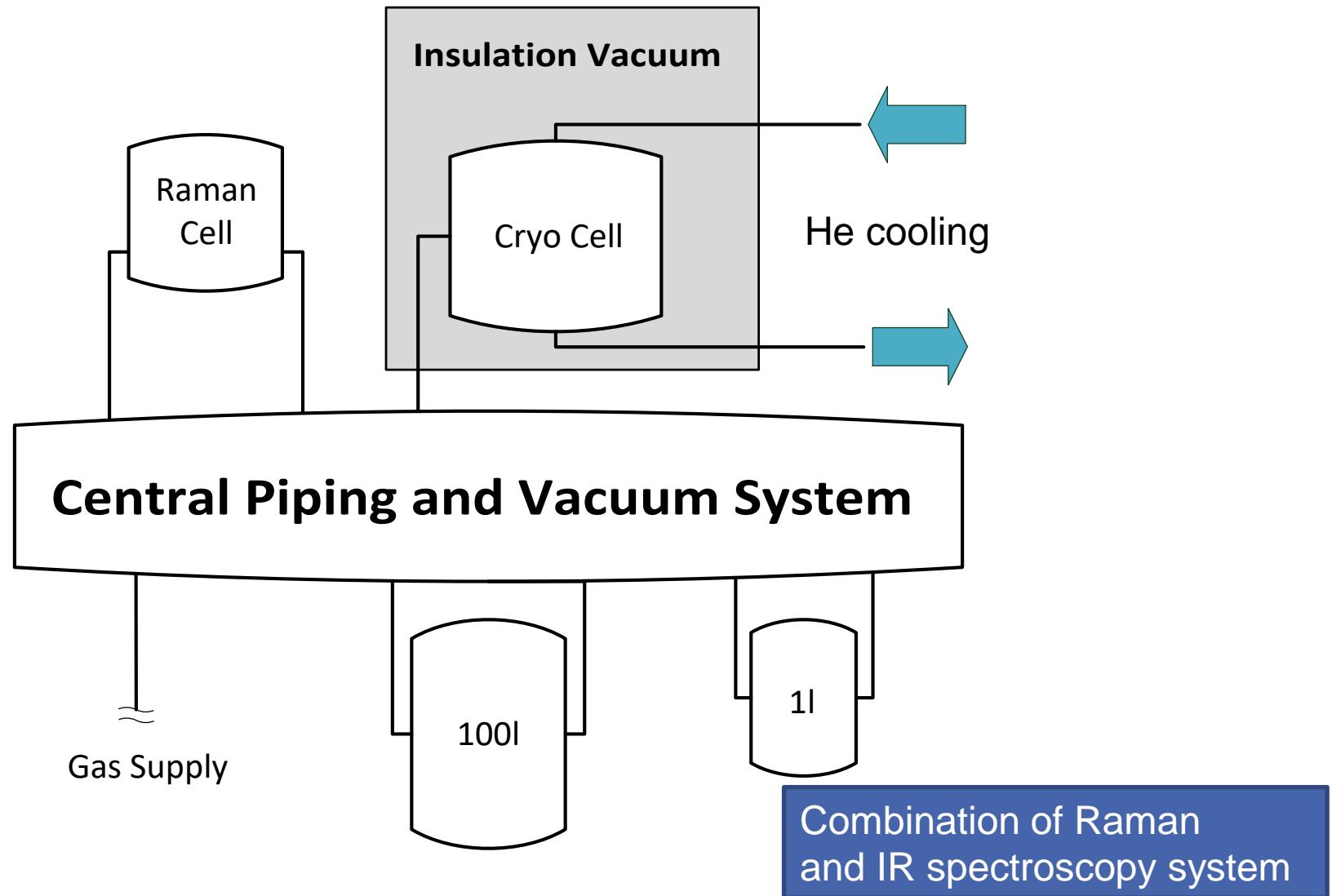
- Temperature: 18 K to 35 K
- Only H<sub>2</sub> HD D<sub>2</sub> mixtures



# Experimental Setup



# Simplified Flow Chart



# Strategy

## ■ Demonstation in the liquid phase

- High cluster density
- High signal expected

With existing  
TApIR experiment

## ■ Gas phase without tritium

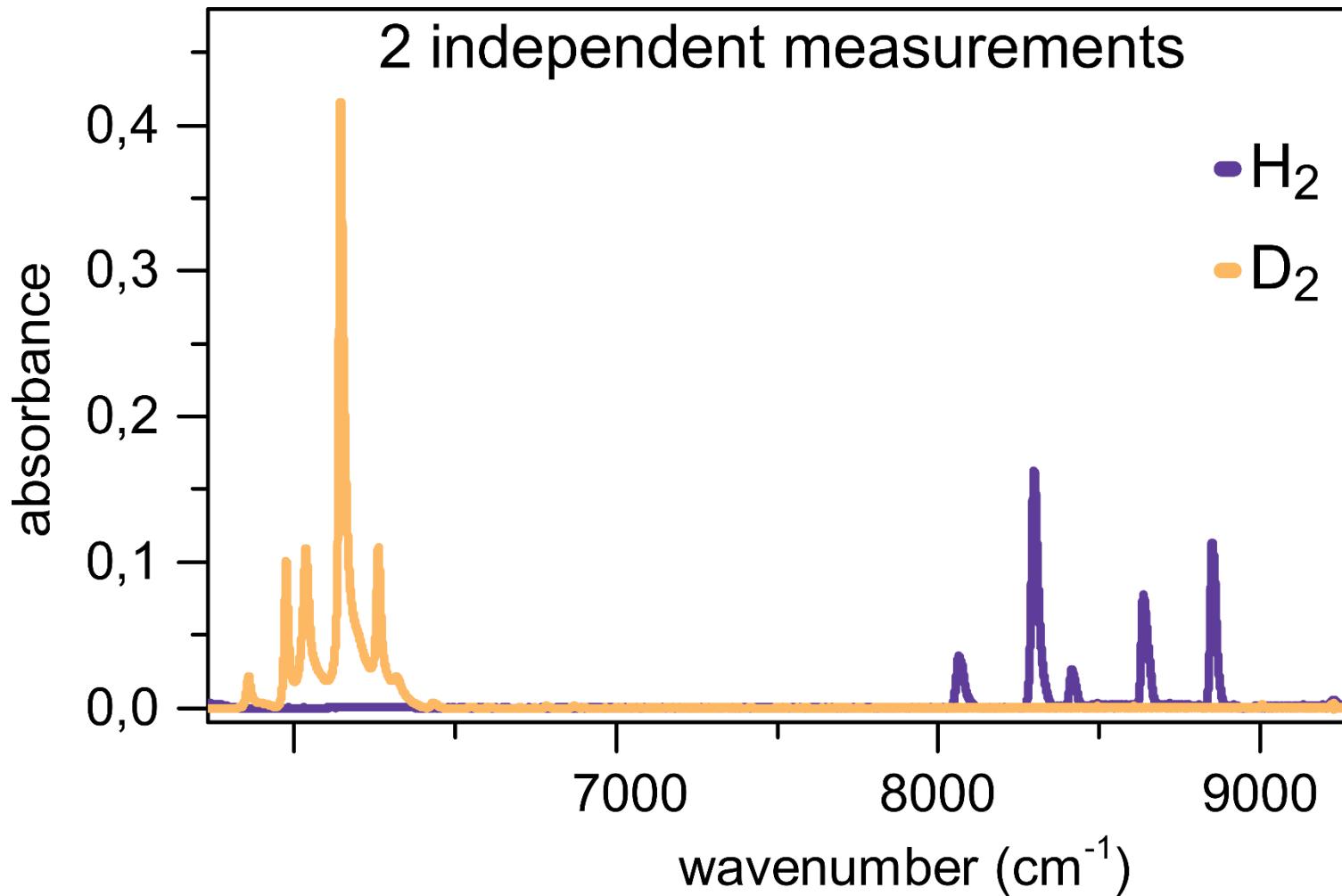
- Temperature and pressure dependency studies

## ■ Gas phase with tritium

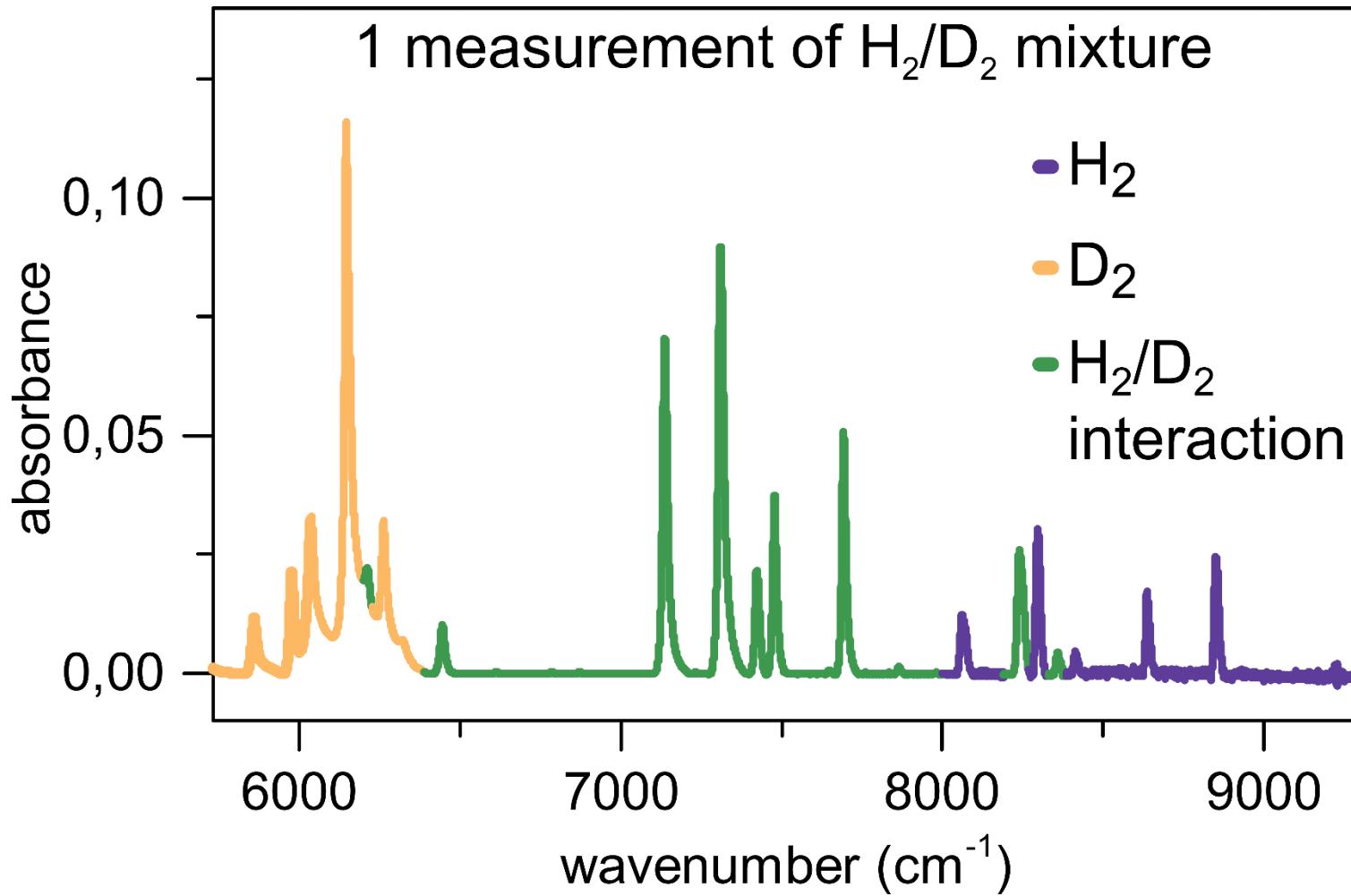
- New  $T_2$ ApIR experiment



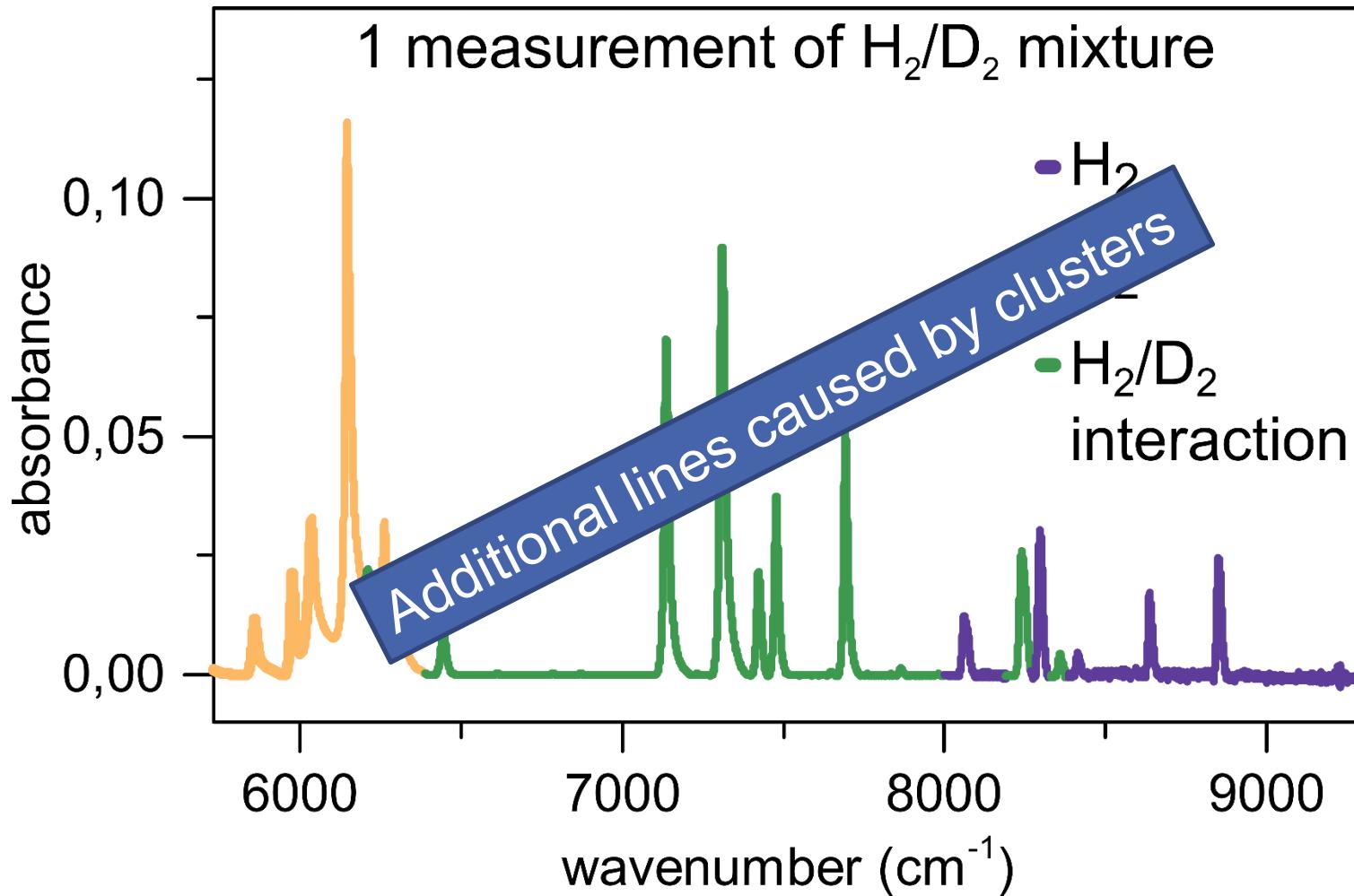
# Liquid phase, ~20 K



# Liquid phase, ~20 K



# Liquid phase, ~20 K



# Strategy

## ■ Demonstation in the liquid phase

- High cluster density
- High signal expected



With existing  
TApIR experiment

## ■ Gas phase without tritium

- Temperature and pressure dependency studies

## ■ Gas phase with tritium

- New  $T_2$ ApIR experiment



# Strategy

## ■ Demonstation in the liquid phase

- High cluster density
- High signal expected



With existing  
TApIR experiment

## ■ Gas phase without tritium

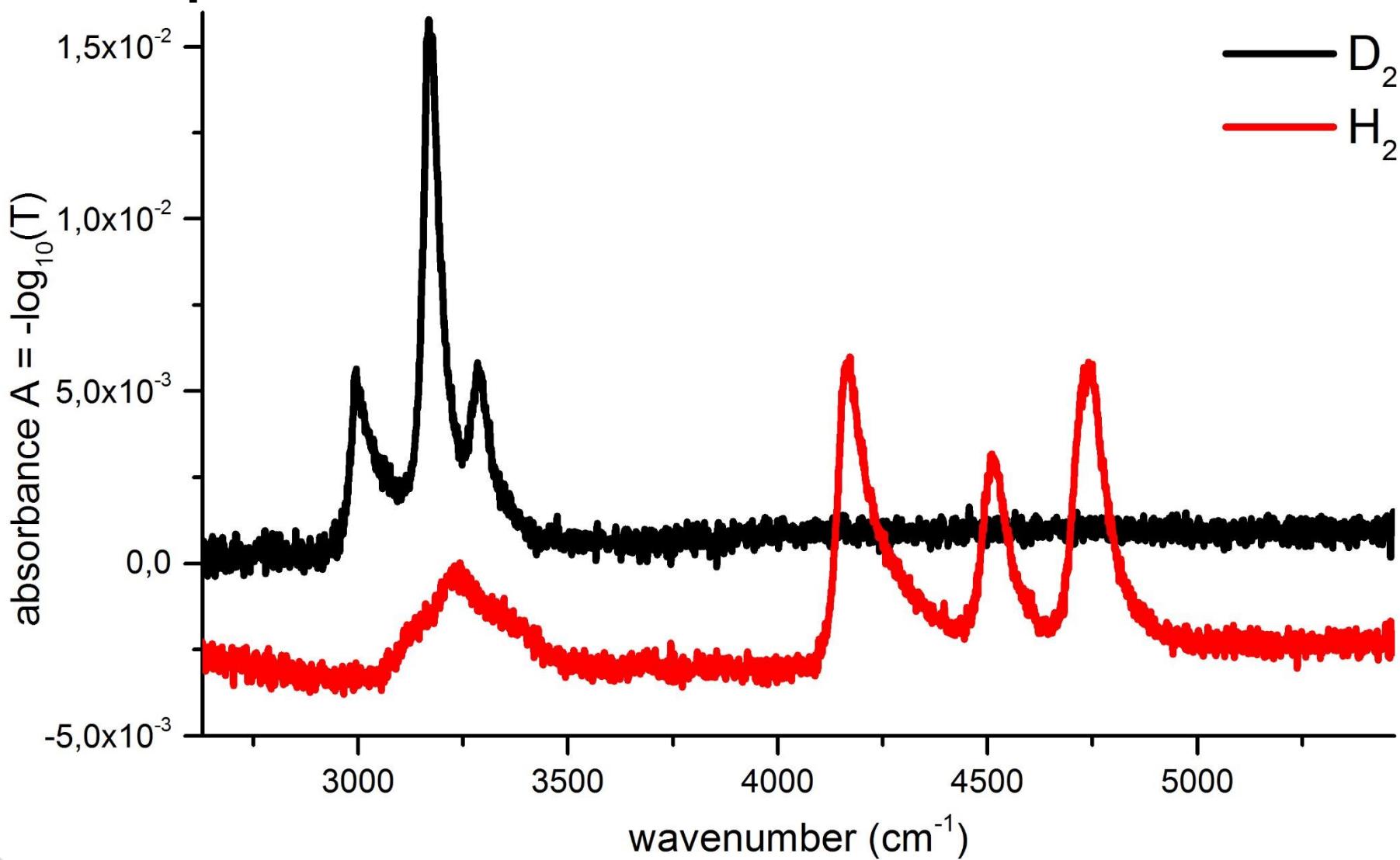
- Temperature and pressure dependency studies

## ■ Gas phase with tritium

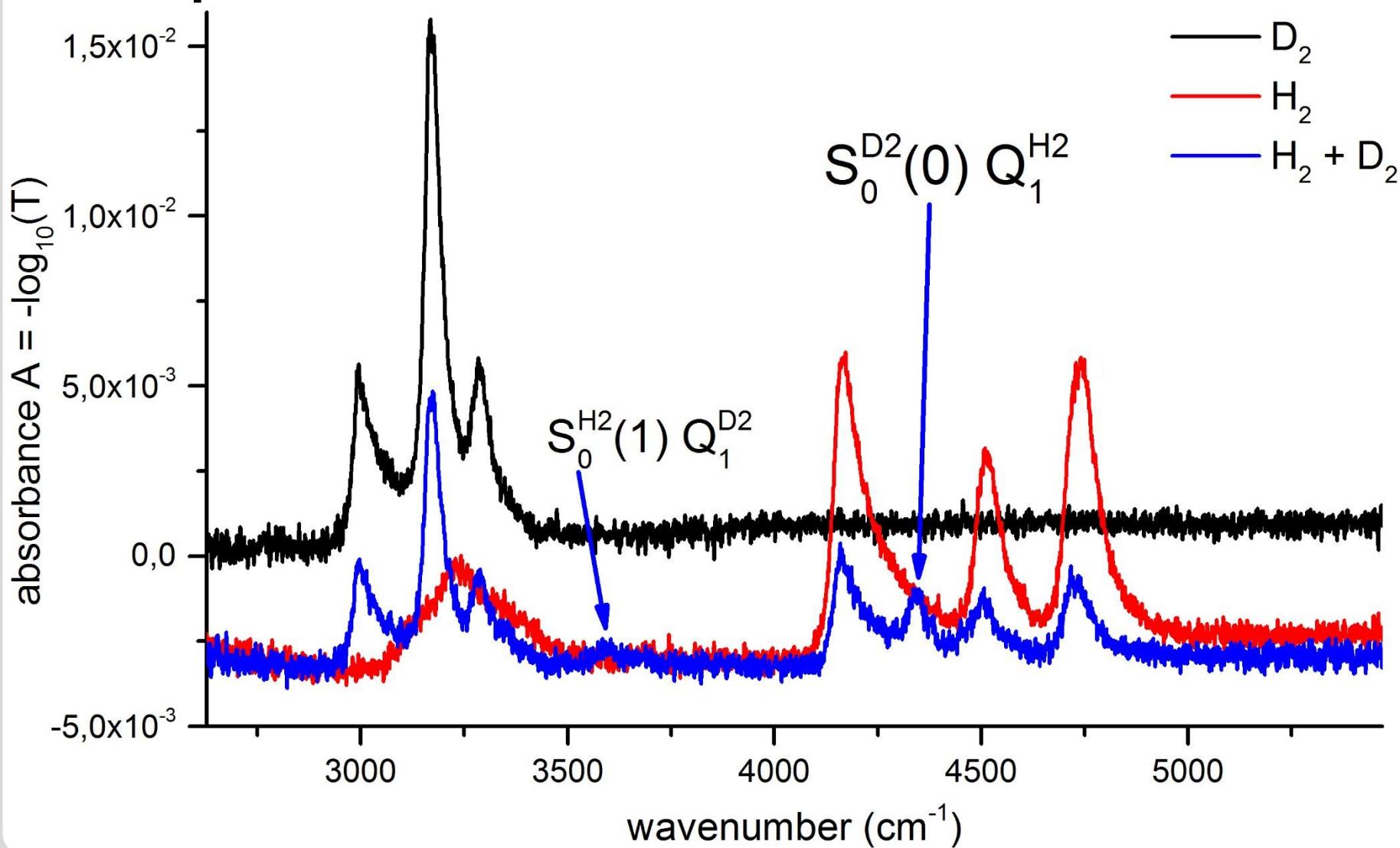
- New  $T_2$ ApIR experiment



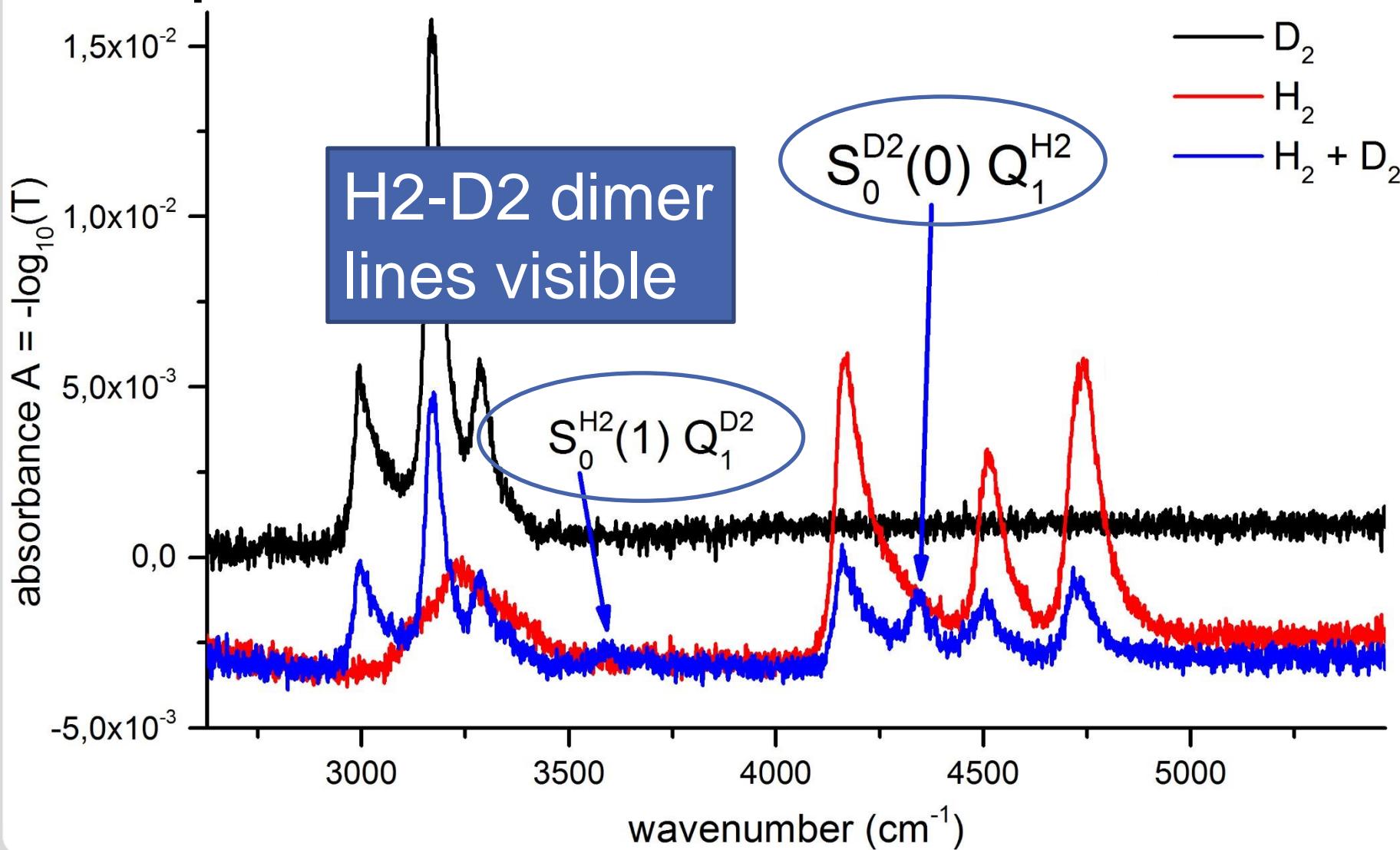
# $\text{H}_2\text{-D}_2$ -Dimers in the gaseous phase at ~30K, 2 bar pressure



# $\text{H}_2\text{-D}_2$ -Dimers in the gaseous phase at $\sim 30\text{K}$ , 2 bar pressure



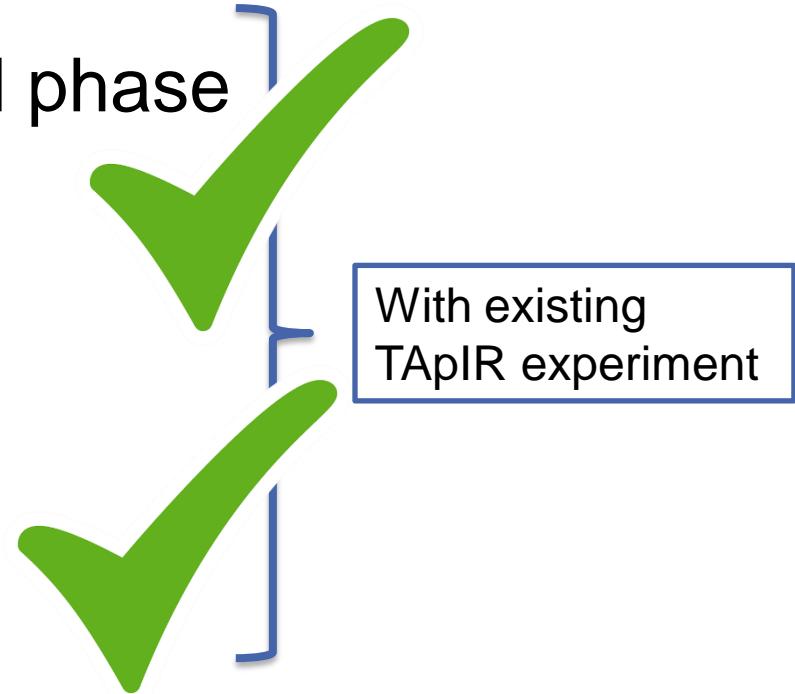
# $\text{H}_2\text{-D}_2$ -Dimers in the gaseous phase at $\sim 30\text{K}$ , 2 bar pressure



# Summary

## ■ Demonstation in the liquid phase

- High cluster density
- High signal expected



## ■ Gas phase without tritium

- Temperature and pressure dependency studies running at the moment

## ■ Gas phase with tritium

- New  $T_2$ ApIR experiment



# Outlook

- Gas phase with tritium
  - New T<sub>2</sub>ApIR experiment  
→ commissioning 2017



## ■ Impact on KATRIN

- Temperature and pressure influence on clusters, WGTS between ~27 and 33 K
- Cluster concentration?
  - systematic influence on neutrino mass can be simulated



# THANK YOU FOR YOUR ATTENTION!

# Spectroscopic Notation

$\Delta J_{\Delta\nu}(J_i)$

|                |     |     |     |     |
|----------------|-----|-----|-----|-----|
| $\Delta J = 0$ | $P$ | $Q$ | $R$ | $S$ |
| -2             | -1  | 0   | 1   | 2   |