# Transition radiation in cross-media showers

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CORSIKA 8 Air-Shower Simulation and Development Workshop,

Heidelberg 12 of July 2022





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# Outlook of the talk

- Transition radiation
- C8 implementation:
  1)Cross media shower
  2)Propagators implementation
- Results: very simple example to start testing the framework.

Note: all files mentioned in the presentation are available here.



### Transition radiation

- It is produced when a charge particle crosses the boundary between two dielectric media.
- Simplest example: Particle at constant speed from vacuum to metal (high conductivity). The field in the vacuum is then the sum of a charge q and an image charge -q.

Thus, for this observer, two particles would annihilate at the boundary producing transition radiation.



### Transition radiation

• At a microscopic level, transition radiation can be calculated as the sum of three contributions: direct, reflected and transmitted.

Comparison with theory:



C. W. James, H. Falcke, T. Huege, and M. Ludwig, Phys.Rev. E84, 056602 (2011).



P. Motloch, J. Álvarez-Muñiz, P. Privitera, and E. Zas, Phys. Rev. D 93, 043010 (2016).

### Implementation in Corsika 8

- As long as one knows the medium where the track is, we know which contributions to calculate.
- The special case of a track crossing the boundary is automatically handled by the particle propagation algorithm in Corsika 8 (Cascade.inl line 166).
  - The distance to the next medium is calculated in each Monte Carlo step.
  - If this distance is the shorter than all interaction lengths, that is the one chosen to be advance.
  - This way we will always get tracks fully contained in one medium.

### Shower from air to ice in C8

- For a shower that crosses several mediums we need to construct a layered environment in C8.
- This is done by adding one volume node to another. In this case if we want to have ice after the atmosphere, we add the ice sphere as a child node to the atmosphere sphere.



### Shower from air to ice in C8

Ice sphere

auto node5 = EnvType::createNode<Sphere>(center, EarthRadius+4\_km); ---> Sphere for layer 5 of atm std::make shared< node5->setModelProperties (medium5);  $\longrightarrow$  Add medium properties to layer 5 of the atm.  $auto const medium_ice = \longrightarrow Medium properties: refr. Index, type of medium, mag. Field, uniform density model$ std::make\_shared<UniformRefractiveIndex<MediumPropertyModel<UniformMagneticField<HomogeneousMedium<EnvironmentInterface>>>>> Setting up the environment taking into account the hierarchy between volumes (from the outermost to the innermost). universe->addChild(std::move(nodel));  $\longrightarrow$  Universe is the volume that englobes the "world" we create File: em shower.cpp @7f879ff5 Atm layer 4  $\blacktriangleleft$ Layer 5 is a child node of atm layer 4 Atm layer 5 Ice sphere is a child node of atm layer 5

### Details of cross media showers

- If we use UrQMD for low energy hadronic interactions we only have tabulated crosssections for nitrogen, oxygen and argon.
- This makes running showers in ice or water slow, could be added in the future since would only mean adding hydrogen to cover these two cases.
- Currently UrQMD needs to do many iterations to find the next interaction, failing to do so after reaching 50000 iterations, this makes us set very high hadronic thresholds.
- Electromagnetic interactions with PROPOSAL seems to be working without any issues from one medium to another.

### Dificulties in implementing TR

- Transmitted ray paths are already not analytical in a planar interface between two mediums with constant refractive index (calculated through iterative process).
- Straight reflected paths cannot be analytically calculated if the boundary surface is spherical (can be calculated through an iterative process).
- Fresnel coefficients are usually not analytical, only in some approximations one can arrive at an analytical expression.

### Specific media propagators in C8

• We will consider a few assumptions in order to get a simplified case:

- Using a spherical ice surface, we can approximate the air-ice interface as flat. The shower dimensions are small compared to the Earth radius (not always true if very inclined).

- We simplify our case to an homogenues ice sphere with constant refractive index.

- For the test antenna we are looking at, the air contribution is negligible: neglecting it will speed up considerably our simulations since the ray path calculation is not analytical.



# Specific media propagators in C8

auto receive\_ref\_{-k\_ref}; auto distance\_ref\_{(destination\_ref - source).getNorm()};

#### File: SimplePropagator.inl



# Specific media propagators in C8

auto const beta\_perp\_perp\_fresnel = beta\_perp\_perp\*fresnel\_perp; auto const beta\_perp\_para\_fresnel = beta\_perp\_para\*fresnel\_para; auto const beta\_perp\_final = beta\_perp\_perp\_fresnel +

beta\_perp\_para\_fresnel;

We return two propagators, one for each ray path. Then the radio process will iterate over them. return {SignalPath(time, averageRefractiveIndex\_, ri\_source, ri\_destination, emit\_, receive\_, beta\_perp, distance\_, points), SignalPath(time\_ref, ri\_destination, ri\_source, ri\_destination, emit\_ref\_, receive\_ref\_, beta\_perp\_final, distance\_ref\_, points)};

File: SimplePropagator.inl

New argument that contains the direction of the field taking into account the Fresnel coef.

### Fresnel coefficients in C8

- As the radio module is right now, they cannot be in principle implemented due to not having all the information in one place.
- In the radio process one is lacking the environment properties.
- In the section were the propagator is implemented you don't get the direction of the electric field.
- Needs discussion to evaluate best solution. Currently I implemented a version that modifies the propagator to also calculate the electric field direction in it taking into account the Fresnel coefficients.

### Results: Simulation setup

- e- 10 TeV shower with only electromagnetic interactions.
- $45^{\circ}$  shower to have a more coherent emission in ice for both reflected and direct contributions.
- Ice antenna at (-4 km, 0 km, Earth Radius 400 m), intersection of shower axis with the interface is at (0 km, 0 km, Earth Radius).
- Thresholds: 2 MeV for electron, positrons and photons.

Atm layer 4	
Atm layer 5	
Ice sphere	en e

### Results: Density profile



### Results: 10 TeV electron EM shower at $45^{\circ}$





- Corsika 8 allows an easy, all in one, implementation of cross media showers.
- These showers need to be compared to other setups, like Corsika 7 to Geant4 for example.
- The radio interface is working as expected for multiple propagators and producing reasonable pulses.
- The radio interface needs some small changes to apply Fresnel coefficients.



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Thank you for your attention !

# **Backup slides**

### Vertical em shower



#### 10000 steps in showerAxis



# Simulation setup for EM shower

