#### Pattern Recognition in KM3NeT: A multi-dimensional challenge



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for the KM3NeT Collaboration

#### 21 February 2017







# KM3NeT Collaboration

www.km3net.org

Cyprus



ARCA Site

Morocco

Toulon

21 Feb 2017

#### **The ORCA Detector**



#### The ORCA Detector



#### **Measuring Neutrinos**



#### Motivation

- KM3NeT detectors form 3D images (6D maybe?)
- Our goal is to extract physics from these images
- Can we train a neural network to do the job for us?



## Inspired by:

- Convolutional Neural Network work developed by NOvA
- Paper: <u>arXiv:1604.01444</u>
- The following slides were taken from talks below

Cincinnati

UNIVERSITY OF



#### CHEP 2016

Using Modern Deep Learning Techniques to Categorize Neutrino Interactions

Adam Aurisano University of Cincinnati

IML LHC Machine Learning WG 14 April 2016

# The Convolutional Visual Network for Identification of NOvA Events.

An implementation of Convolutional Neural networks and its applications on neutrino interaction events.

#### **Network Layers**

#### Kernel Renormalization:

Kernels evolve as the training progresses through renormalization. This process uses non saturating functions.





(a) Standard Neural Net



(b) After applying dropout.

#### **Dropout:**

Randomly reset weights, effectively removing whole nodes at each step.

**Encourages complex dependence** and discourages overtraining

**CVN Neutrino Identification** 

#### CHEP - October 2016

#### **NOvA**



**Neutrino Event CVN:** Siamese network architecture based on GoogLeNet.



- Inspired by siamese architectures to allow the network to learn from features on each 2D view of the event.
- Using the caffe framework <a href="http://caffe.berkeleyvision.org/">http://caffe.berkeleyvision.org/</a>
- We train on Fermilab's Wilson cluster GPUs (2 к40s)
- Trained on 4.7 million simulated events of all neutrino interaction types plus cosmic rays



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#### **Initial Training**



- The initial training consisted of ~10 passes over 3.4 million training images.
  - With two k40 GPUs, this took ~1 week.
- We can judge how well our model works by producing a confusion matrix.
  - This shows the relationship between the true event category and what the PID thought was the most likely event category.
- The matrix is mostly diagonal events are mostly correctly identified.
- Mis-identified events mostly fall within blocks – while the interaction type is wrong, the selected neutrino flavor is still correct.

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#### What about ORCA?



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#### Our Detector

- Cylindrical volume with hexagonal string symmetry
- Extra dimensions from PMT orientation and hit time
- A 2-point model: Light production and detection points



#### **Detection Point**

- No need for true detection, just encode time and direction information
- Extrapolate light back in the direction of the PMT for a time  $t_{hit}$   $t_0$  (event start time)
- Ignores particle that produced the light



## Image Inputs

- Define four (x2) views (2D projections of the detector)
- Plot charge in detection and "production" positions



## Why Projections?

#### Why does deep and cheap learning work so well?

Henry W. Lin (Harvard), Max Tegmark (MIT)

(Submitted on 29 Aug 2016 (v1), last revised 28 Sep 2016 (this version, v2))

- Symmetry, locality, ... -> neural nets are good with images and drawings
- Hierarchical data -> deep neural network more efficient



#### Multi-view Convolutional Neural Networks for 3D Shape Recognition

Hang Su, Subhransu Maji, Evangelos Kalogerakis, Erik Learned-Miller (Submitted on 5 May 2015 (v1), last revised 27 Sep 2015 (this version, v3))

• Multi-view 2D images can outperform networks with truly 3D shape inputs



Figure 2. Precision-recall curves for various methods for 3D shape retrieval on the ModelNet40 dataset. Our method significantly outperforms the state-of-the-art on this task achieving 80.2% mAP.

#### What's next?

- **Deep learning** is being successfully implemented in HEP
- Can we use it to **reconstruct and/or classify** KM3NeT events?
- We have an initial concept, based on the NOvA experience
- NOvA used the **Caffe framework**, but should we?
- Challenging 6D input space. Is the best solution projections?
- Are **Convolutional NNs** a good approach?
- Other sugestions?

# Thank you!



#### **Backup Slides**

# Objectives

- ORCA: Determine the Neutrino Mass Hierarchy (NMH)
- ARCA: Discover/Observe high-energy neutrino sources in the universe





# The Challenge

- Measure neutrino direction and energy
- Search for oscillation patterns from matter effects
- Sensitive to difference in patterns between NH and IH
- Requires large statistics and good energy and direction res.



 $\theta_{\mathsf{Z}}$ 

# Trigger

- Optical background mostly from <sup>40</sup>K decays in the water
- Measured: 8 kHz uncorr., 340 Hz level-two coinc. / PMT [Eur. Phys. J. C 74, 3056 (2014)]
- Look for coincidences in time and PMT direction to reduce trigger rate.
- Causality further restricts space and time correlations for extra power.
- Final trigger rate ~59 Hz, with 70% of events containing a cosmic ray muon.



# **Optical Noise**

- Optical background in full detector ~500 MHz.
- Neutrino events ~40 hits in a ~500 ns window
- Expect **250 noise hits** (~14% purity)
- Trigger approach ~5 ns time residuals
- Calibrated using 2-fold coincidences
- Can achieve ~3 noise hits per trigger (>90% purity)





#### Limitations

- Hadronic showers can generate multiple cherenkov rings, but reconstructing them is limited by photon statistics
- Without hadronic info, reconstruction is limited by intrinsic kinematics
- Extracting some information on particle multiplicity would be a significant achievement for a deep learning technique



# **Trigger Performance**

- Input a conservative noise rate of 10 kHz uncorr. (500Hz level-two coinc.)
- Achieve a total triggered rate of 59 Hz
- About 70% of events contain a muon (41 Hz)
- High efficiency for  $\nu_{\mu}$  and  $\nu_{e}$  above 4 GeV
- Slightly more efficient for up-going neutrinos (Larger PMT coverage)



**Neutrino Rate:** 

1 v / 10 min

#### Reconstruction

- 1) Start with a track or shower hypothesis
- 2) Use causality to perform a robust hit selection
- 3) Find vertex and direction that best match hit pattern
- 4) Estimate track range for computing track energy (0.24 GeV / m)
- 5) Estimate **Shower energy** and direction from hit distribution after initial fit to the vertex position and time



# Shower Hypothesis

#### **Reco Performance**

- Energy resolution: ~25% (Close to limit arXiv:1612.05621)
- Angular resolution: Better than 10 degrees at relevant energies





## **Event Selection**



- Events are classified through a Random Decision Forest (RDF)
- At 10 GeV:
  - 90% of  $v_e$ -CC are shower-like
  - 70% of  $v_{\mu}$ -CC are track-like
- Most atmospheric muons are removed by containment cuts



#### **Deep Learning**



developer.nvidia.com/deep-learning-courses

- Deep learning is a new paradigm that has caused a renaissance in the machine learning community.
- Use sparsely connected neurons to allow for many hidden layers.
- Deep structure extracts increasingly complex features from the input data instead of needing engineered features.

14 April 2016

Adam Aurisano

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#### **CVN Classifier**



4.7 million, minimally preselected simulated events, pushed into LevelDB databases: 80% for training and 20% for testing.

Rescale calibrated energy depositions to go from 0 to 255 and truncate to chars for dramatically reduced file size at no loss of information

Fine tuned with 5 million cosmic data events taken from an out of beam time minimal bias trigger.

The architecture attempts to categorize events as  $\{V_{\mu}, V_{e}, V_{\tau}\} \times \{QE, RES, DIS\},$ NC, or Cosmogenic.

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#### Ongoing Work CVN and Reconstruction

#### Using the existing reconstruction.

Classify clusters by particle ID





Original CVN network modified to take 4 views (event + prong)

Define clusters

Trained on 50% purity prongs from all events no preselection

Room for improvement in classification and network optimization

**CVN Neutrino Identification** 

#### CHEP - October 2016

#### Fernanda Psihas

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