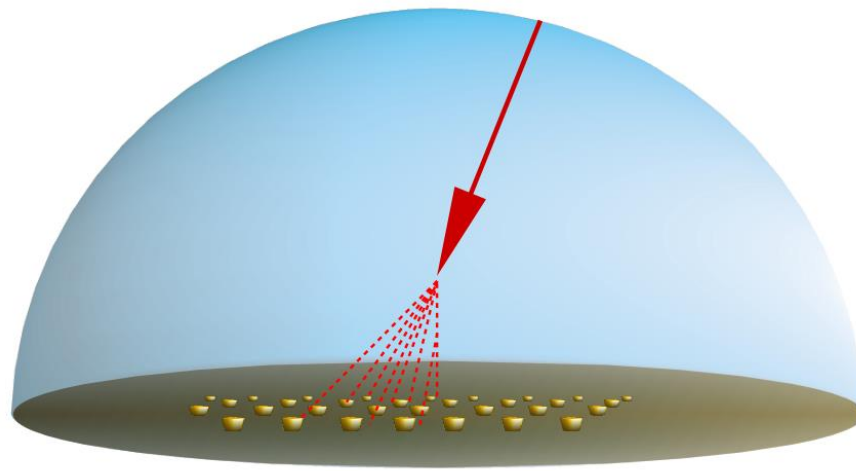




Exploring Deep Network Architectures to reconstruct Cosmic Ray Induced Air Showers at the Pierre-Auger-Observatory

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III. Physikalisches Institut A, RWTH Aachen



Bundesministerium
für Bildung
und Forschung

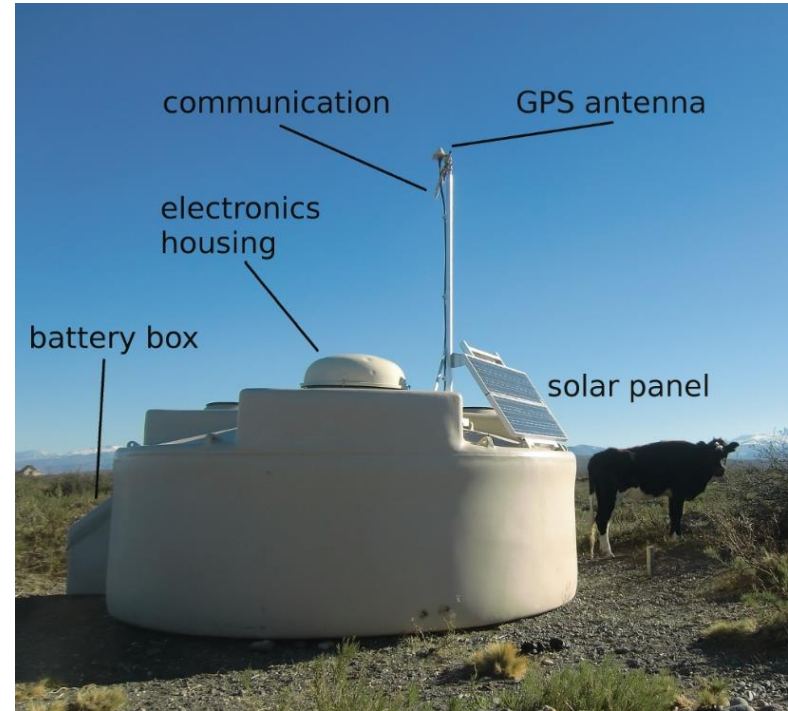
Jonas Glombitza – glombitza@physik.rwth-aachen.de

Aachen, 21.02.2017

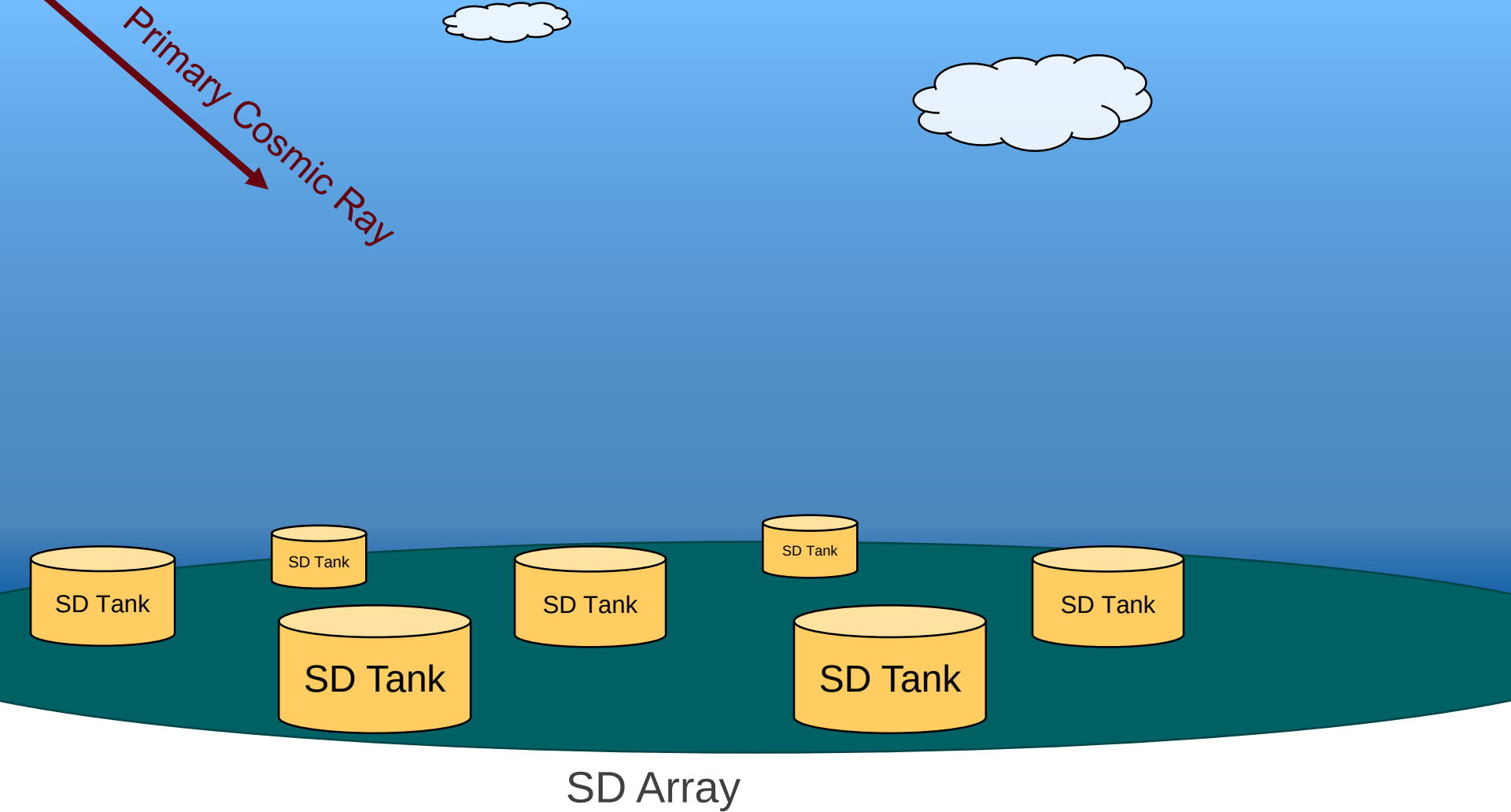
RWTHAACHEN
UNIVERSITY

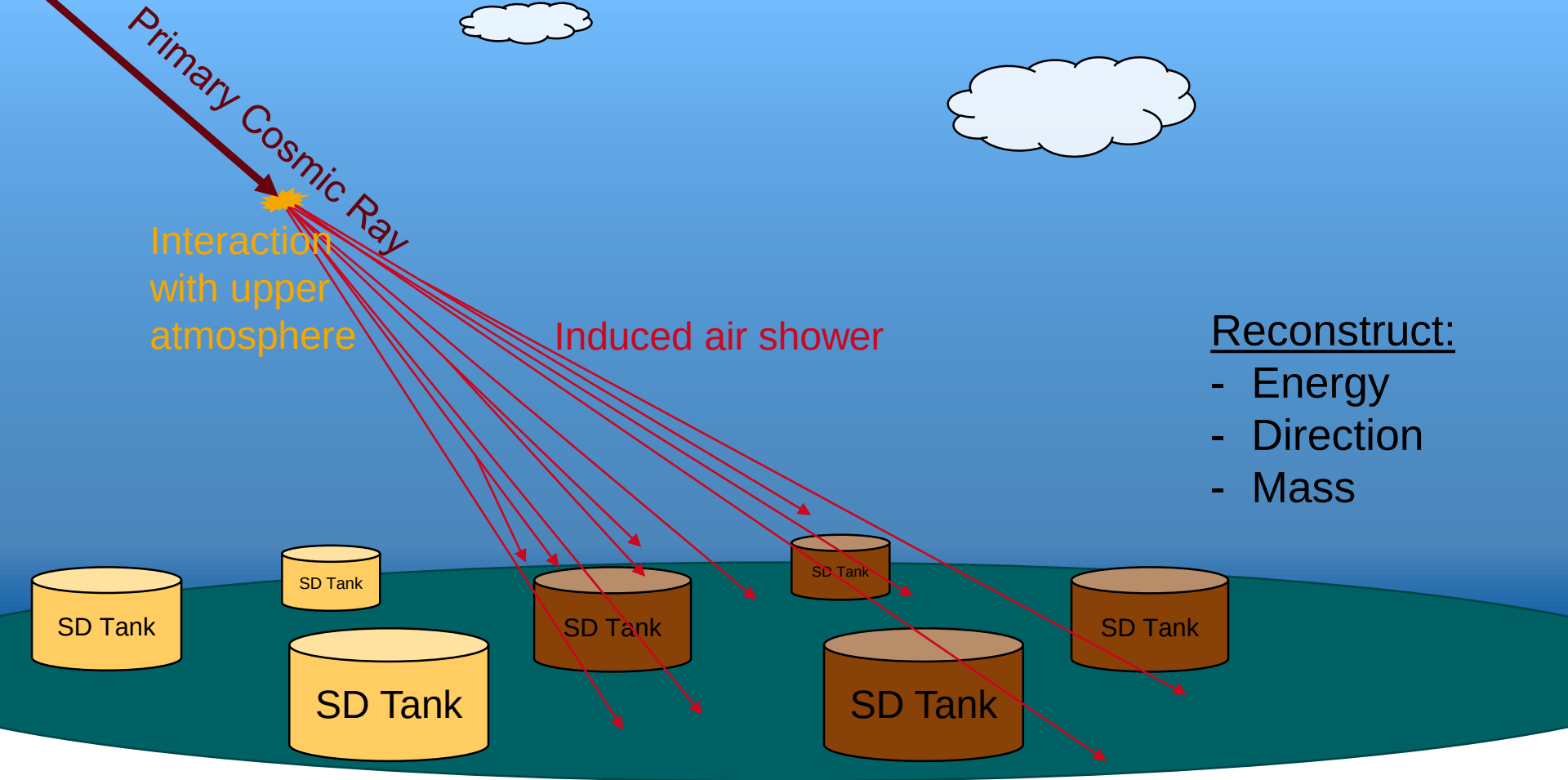
The Pierre-Auger-Observatory

- Cosmic ray observatory in Argentina
- Completed 2008
- Detection of UHECR
 - $E > 10^{17.5} \text{ eV}$
- Hybrid technique
 - Fluorescence telescopes
 - 1660 Surface detectors (SD)
- Array size $\sim 3000 \text{ km}^2$



*Surface detector of the
Pierre-Auger-Observatory*



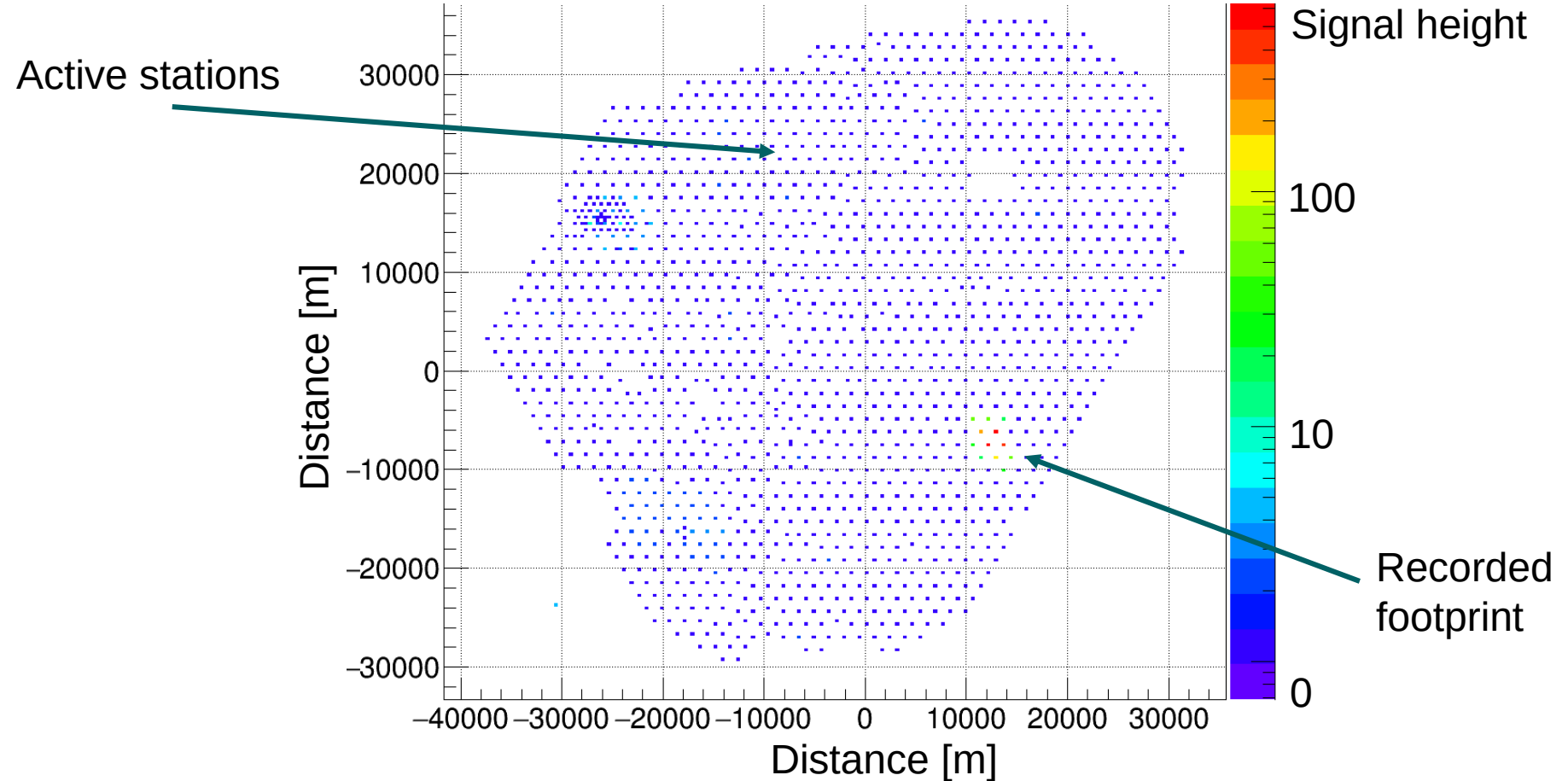


Reconstruct:

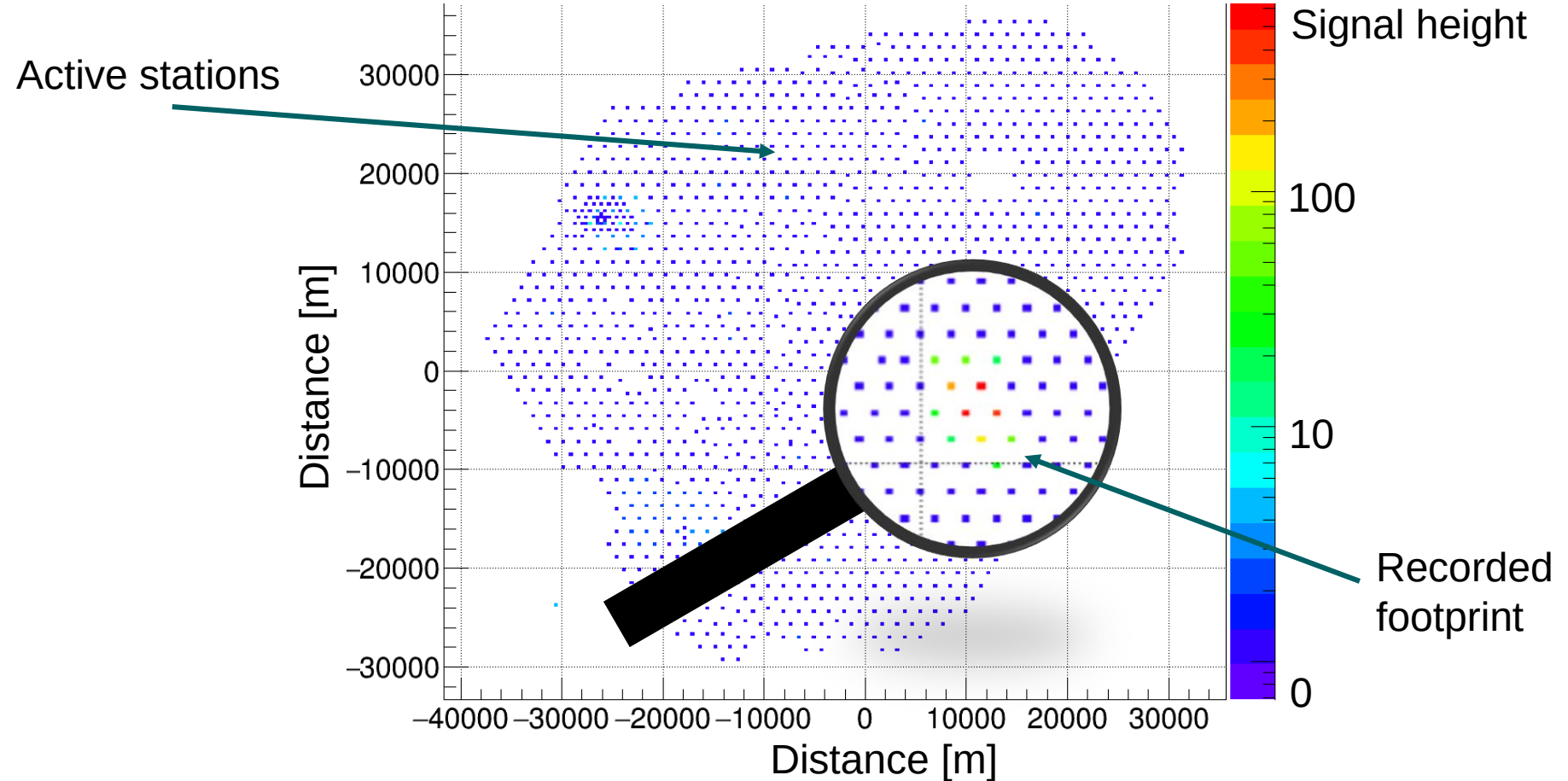
- Energy
- Direction
- Mass

SD array measures the footprint of the air shower

Surface Detector Array

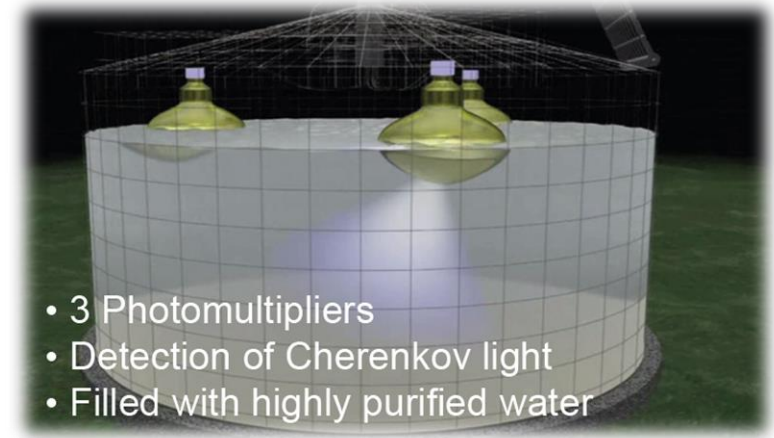
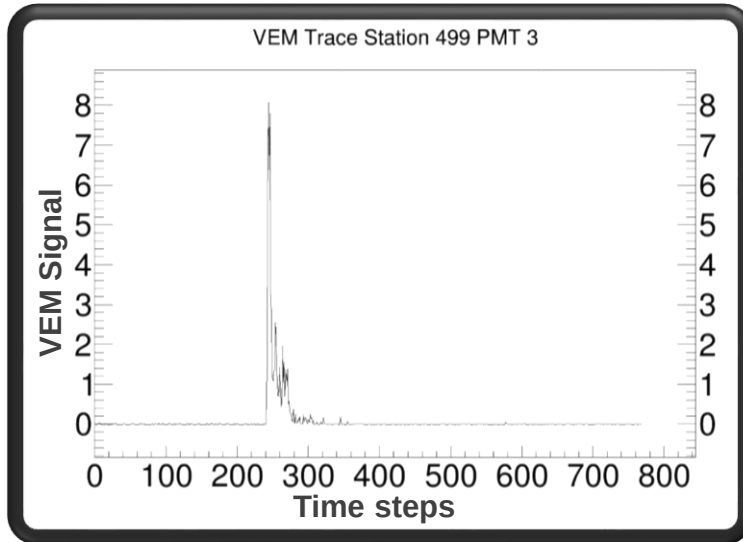


Surface Detector Array



Particle detection

- Particles travel through Detector tank filled with water
 - Production of Cherenkov light
 - Photomultipliers detect the light
 - Time trace



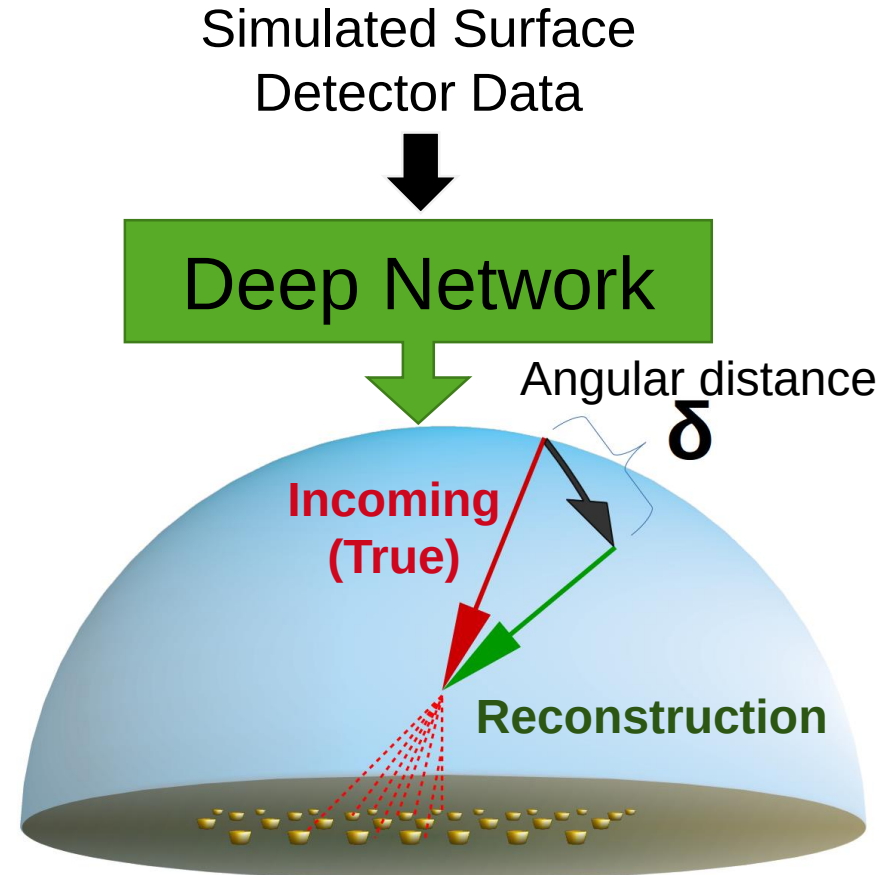
- 3 Photomultipliers
- Detection of Cherenkov light
- Filled with highly purified water

VEM = Vertical Equivalent Muons

Motivation – Image Recognition on Air showers

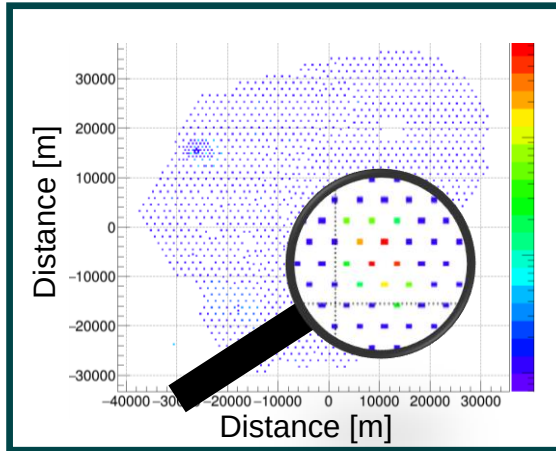
- Investigate if air shower reconstruction can be improved
 - Shower axis, energy, mass
- Test Case: Reconstruct shower axis
- Benchmark
 - Expert reconstruction including physics
 - 1° angular resolution
- Minimize Mean-Squared-Error between true and reconstructed shower axis

$$-L = (\vec{r} - \vec{r}')^2, \quad \vec{r} = \begin{pmatrix} X \\ Y \\ Z \end{pmatrix}$$

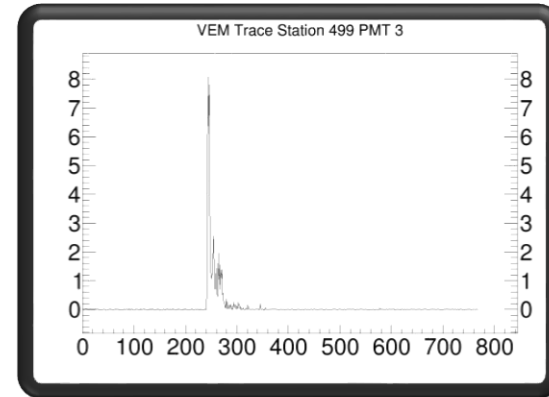


Data Format

- Use simulated Data from EPOS LHC ~ 29000 events
 - Energy Range $10^{19} - 10^{19.5}$ eV, Protons



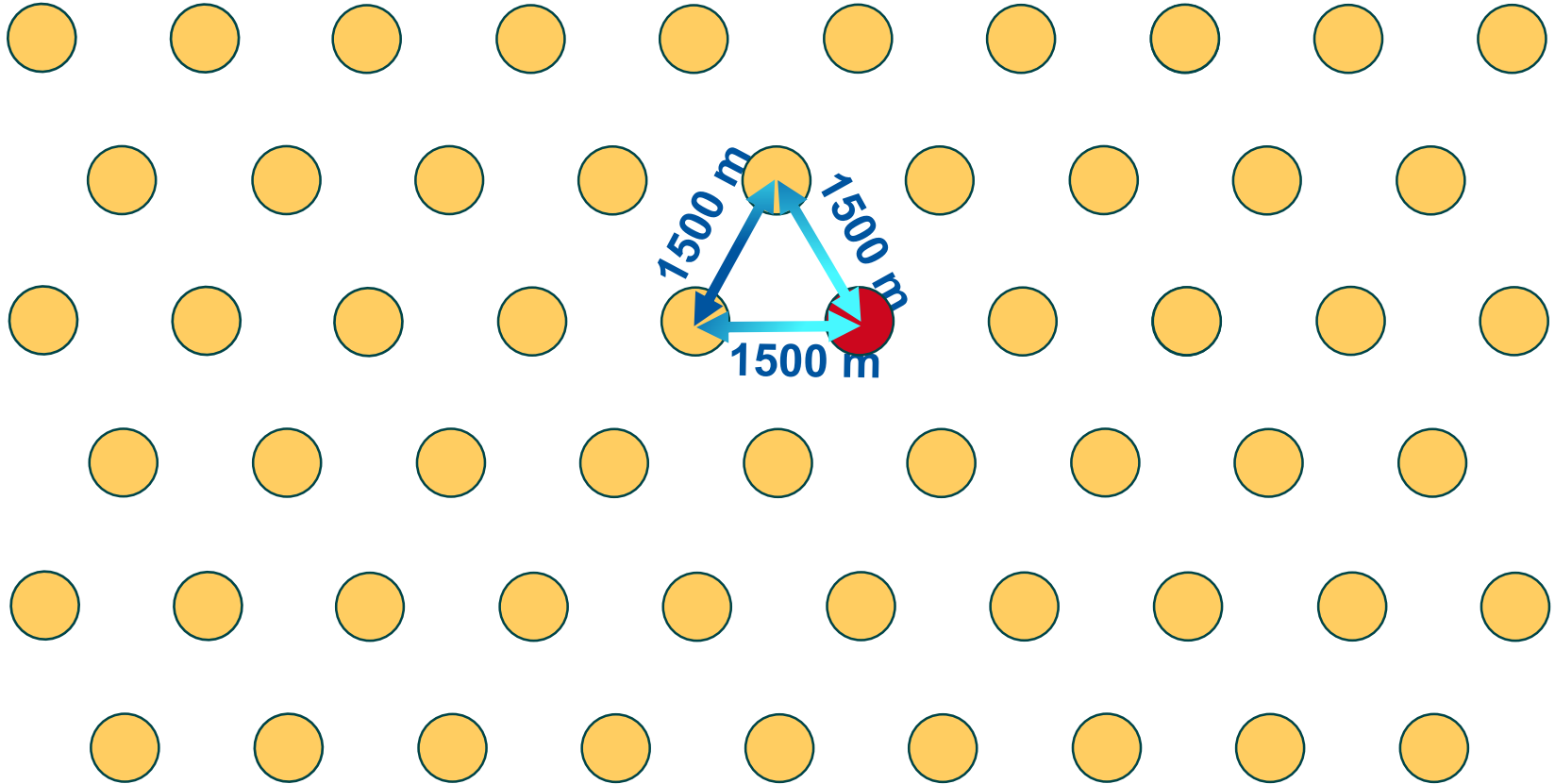
Shower footprint



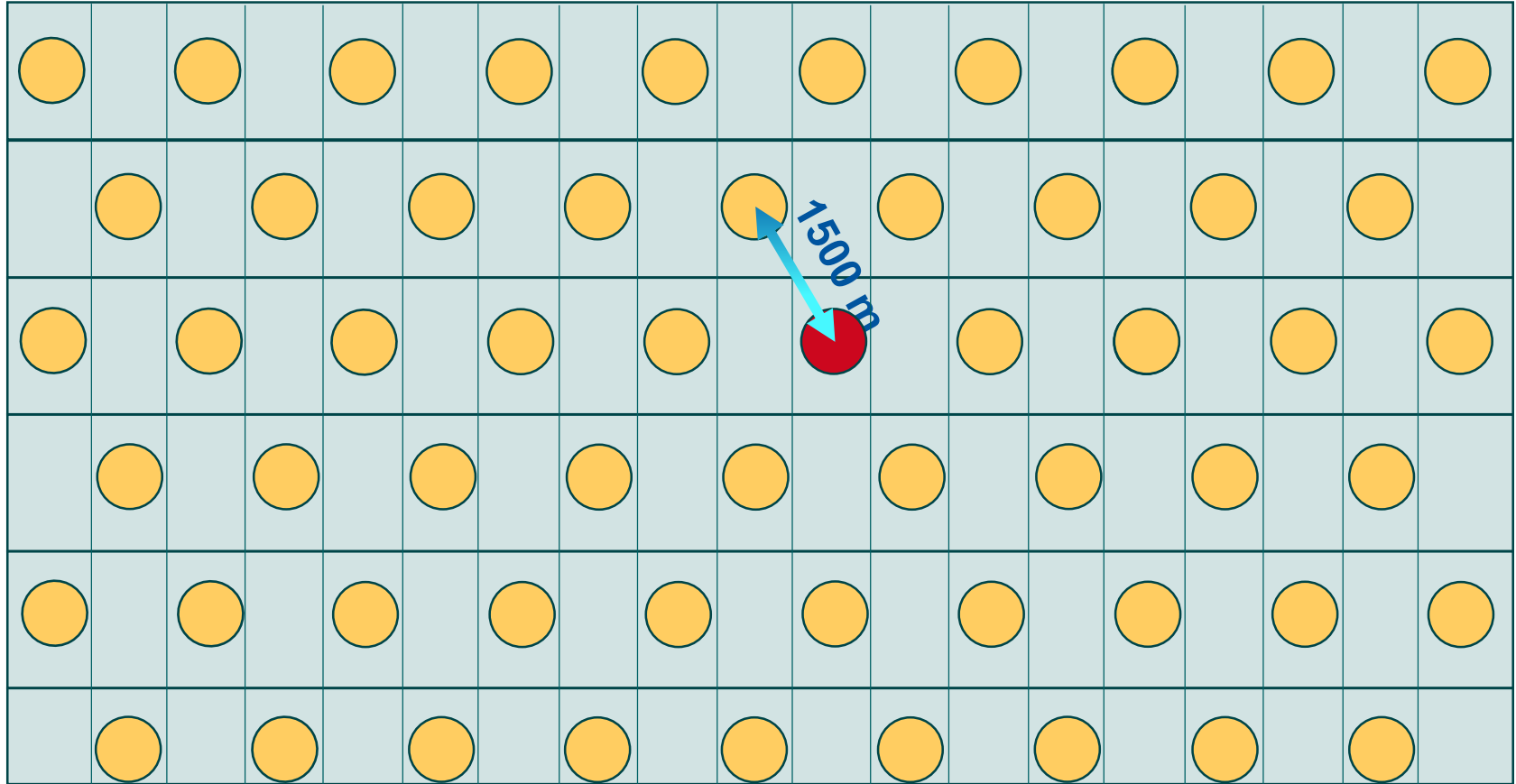
Photomultiplier time trace

- 2 spatial dimensions
- Need Cartesian grid for convolutional operations

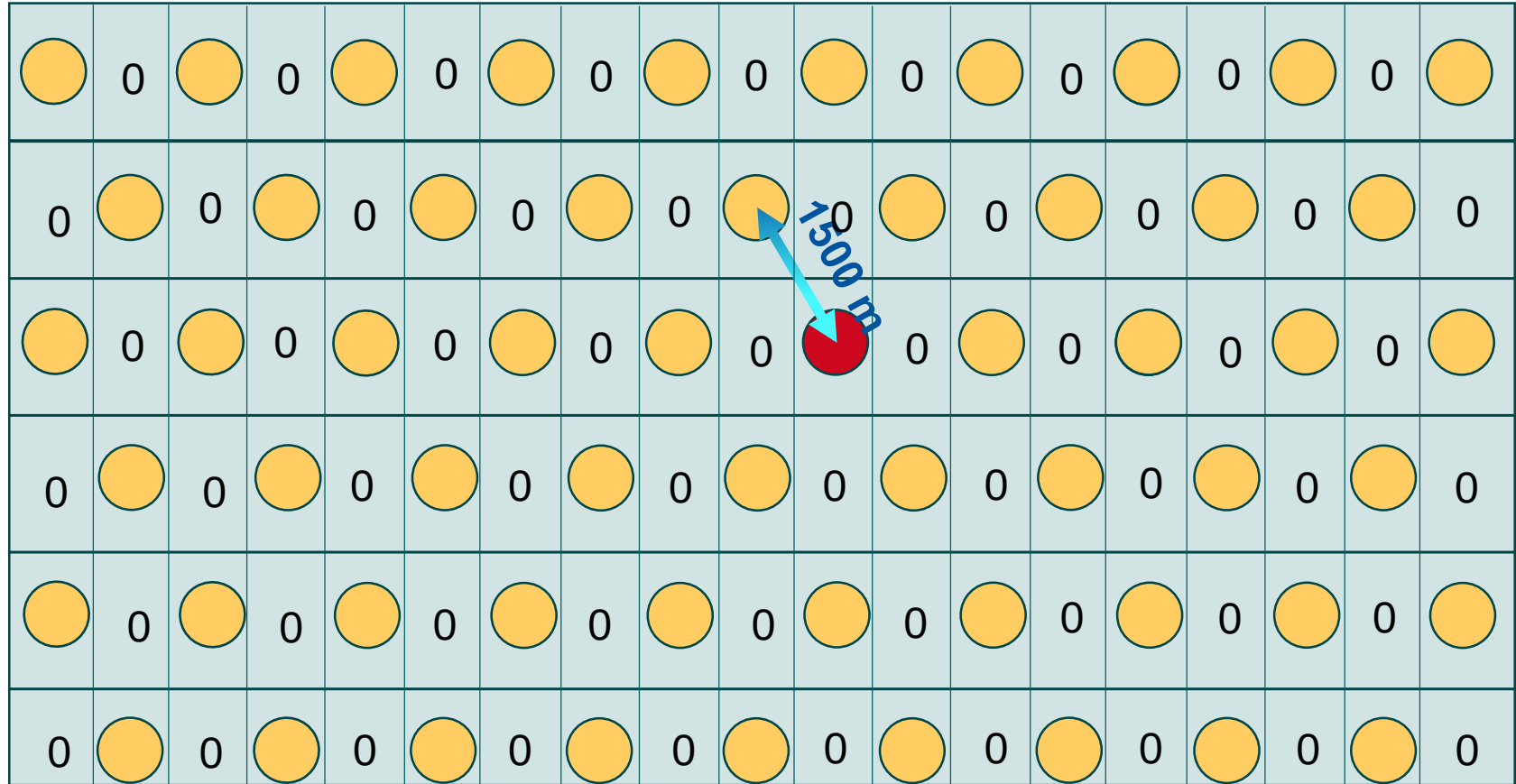
Processing the Footprint – The Surface Detector Grid



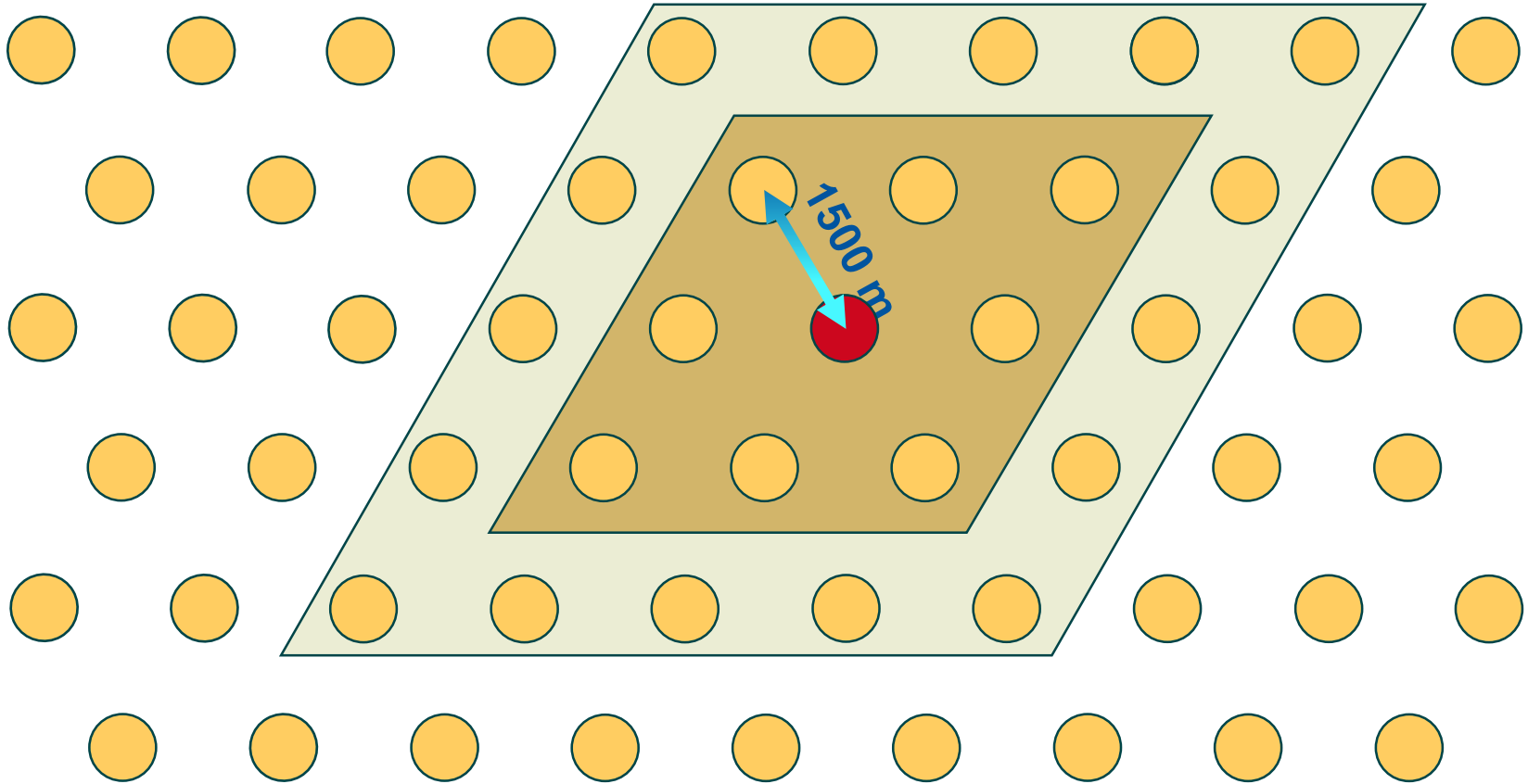
Checkerboard grid



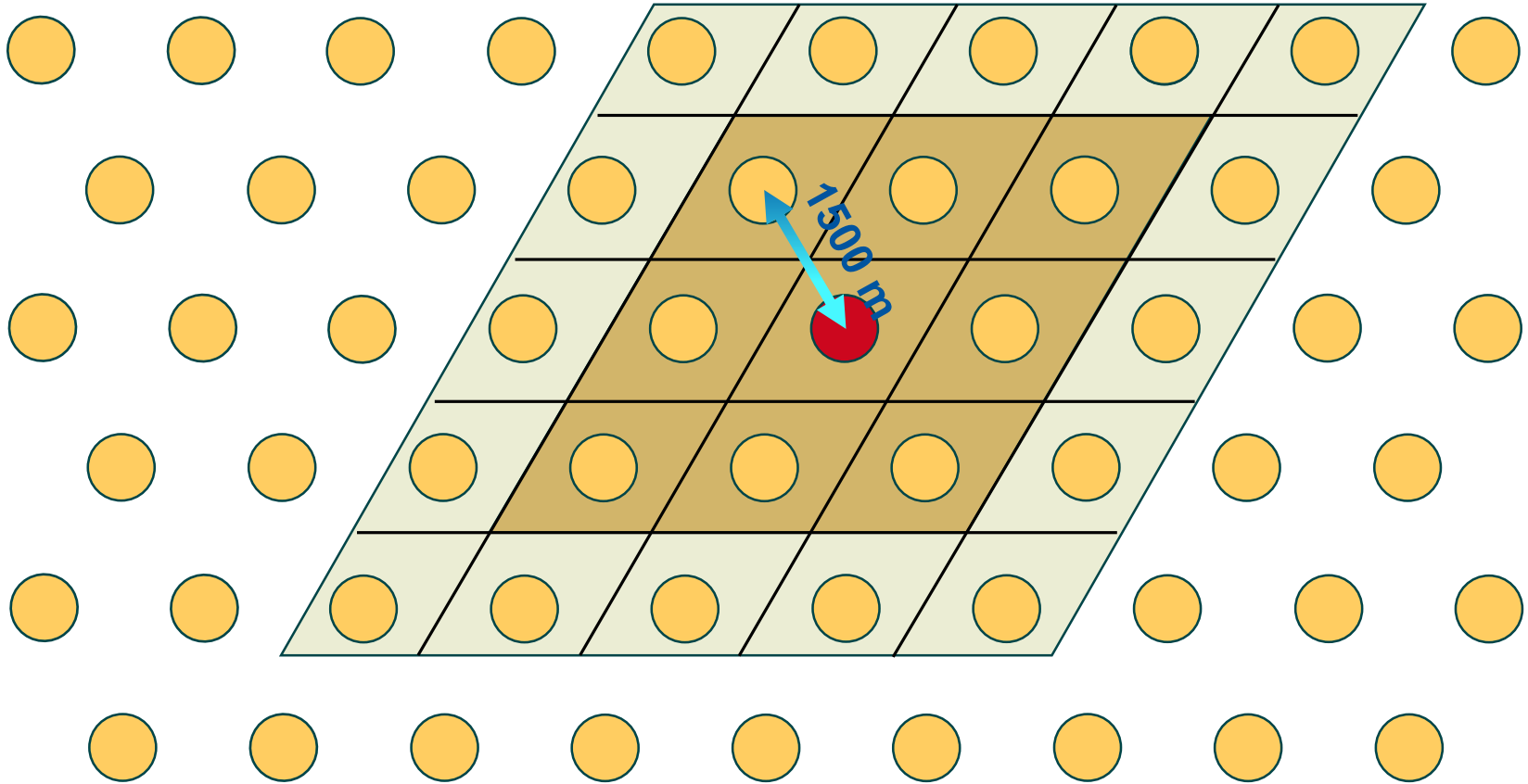
Checkerboard grid



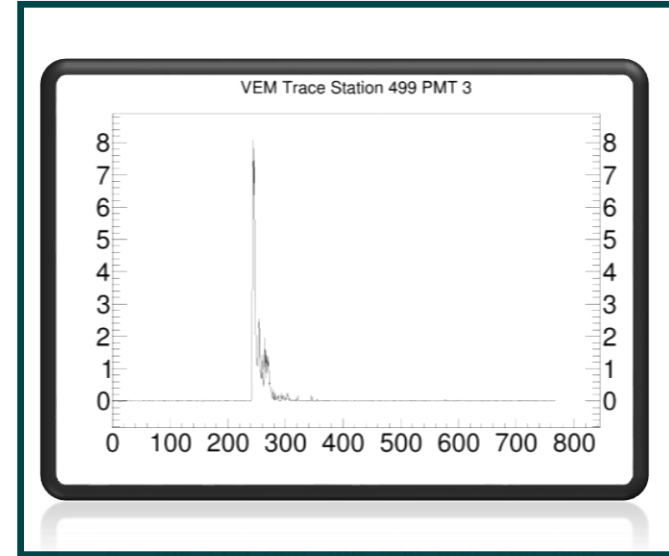
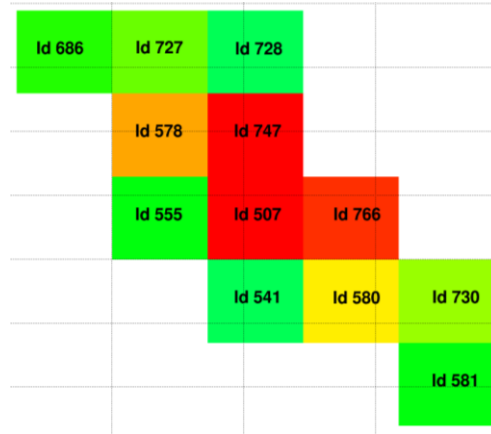
Rhombic grid



Rhombic grid



Data format



✓ Shower footprint

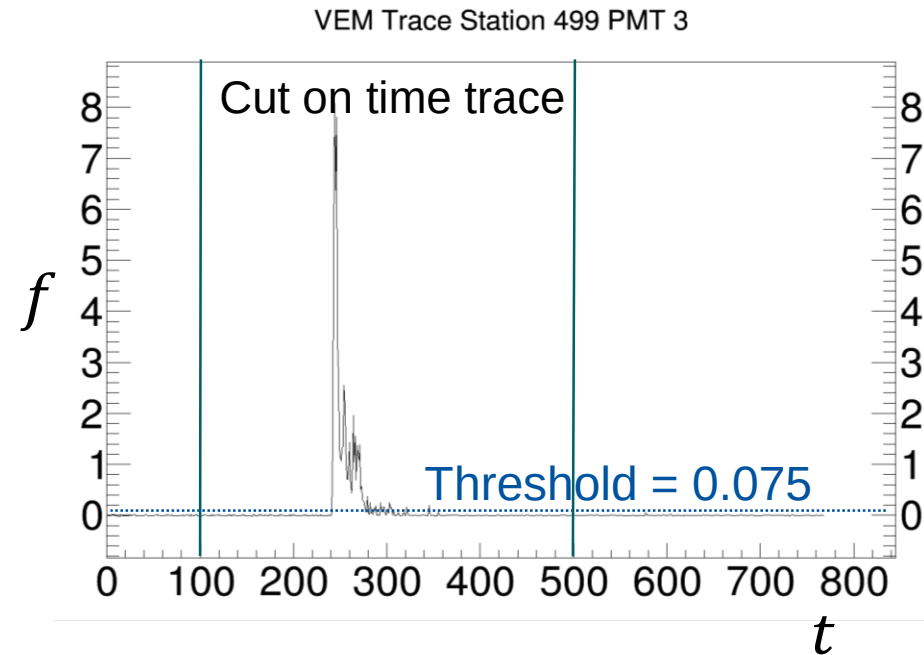
- 2 space dimensions
- Need Cartesian grid for convolutional operations

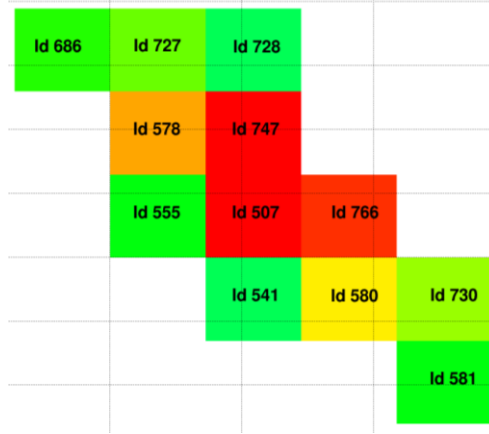
Photomultiplier time trace

- 1 time dimension
- Select signal regions

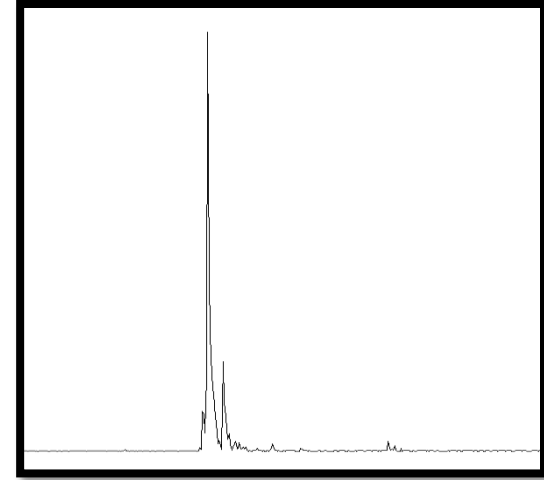
Preprocessing the Time Trace

- Limit time trace to interesting part
- Networks work best on normalized input
 - Set threshold to cut away noise
 - $f_{Normalized} = \log_{10}(10 * f(t) + 0.1)$
 - Background $\rightarrow -1$
 - Signal $\rightarrow >0$
- Critical for successfully training the investigated network architectures





✓ Shower footprint



✓ Time trace

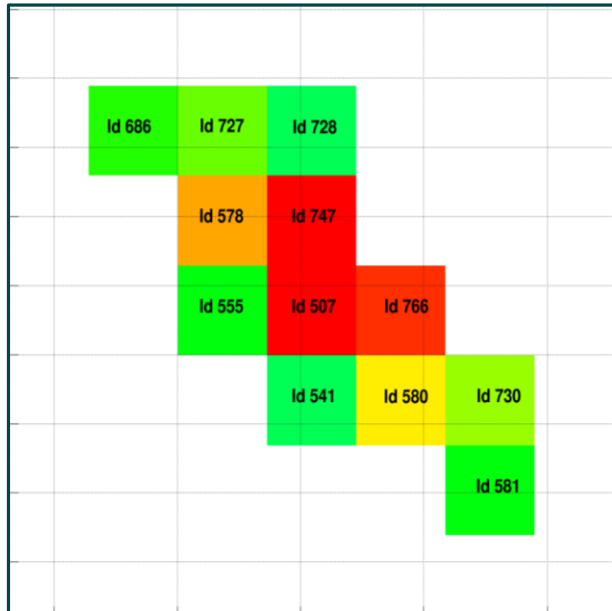


Test different architectures!

With keras – Tensorflow backend

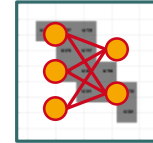
Network Architectures with 2D Input

- Sum over preprocessed time trace
→ 2D Map for every event

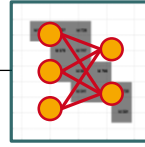


Architectures

- 5 Layer Neuronal Network (Fully Connected)
- 11 Layer Convolutional Network
- ~ each 1,000,000 parameters



Fully Connected Network

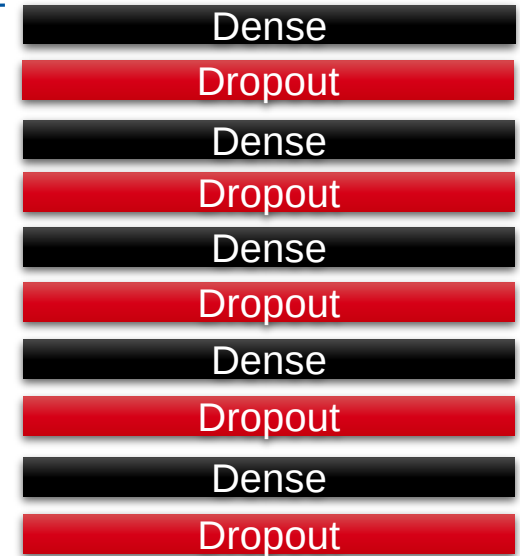


- 5 Blocks:
 - 1 Fully connected layer
 - ReLu activation
 - 1 Dropout layer

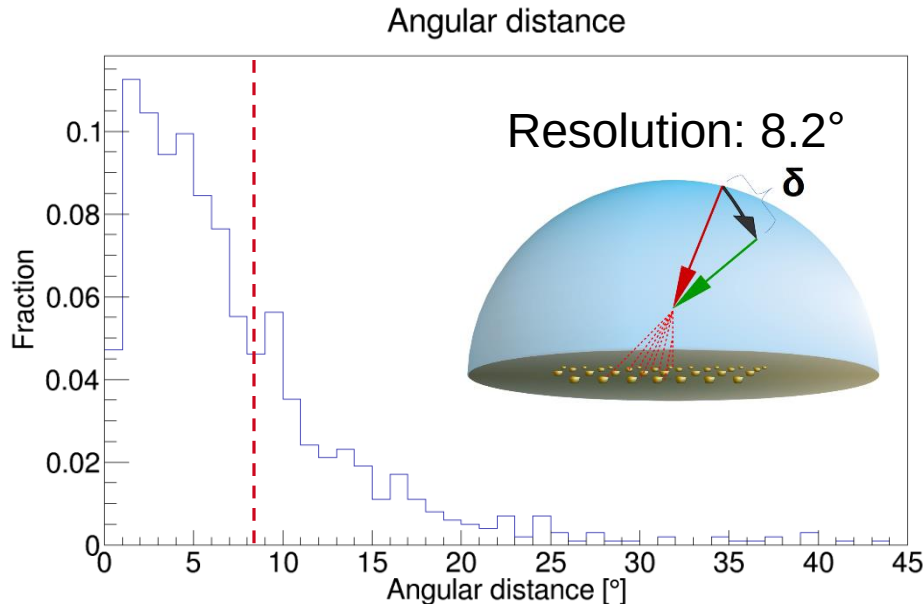
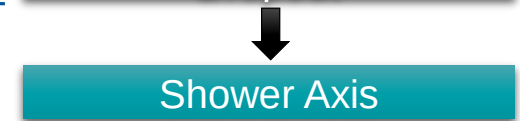
Input



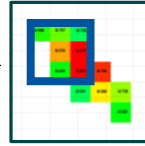
Hidden layer



Output

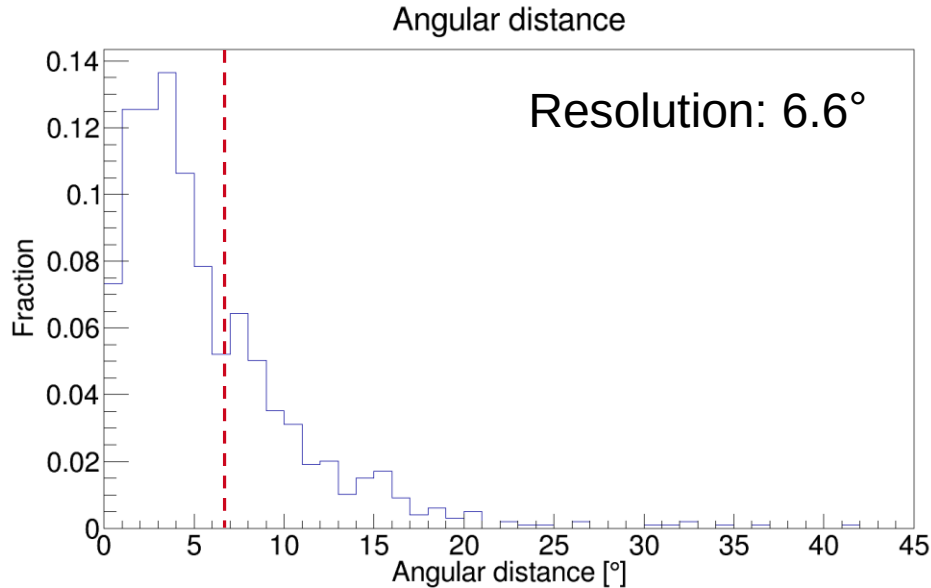


2D Convolutional Network



Input

- 11x 2D Convolutional layers with
 - ReLu
 - Batch Normalization
- 1 Fully connected layer
- Dropout



Hidden layer

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

2D Convolution

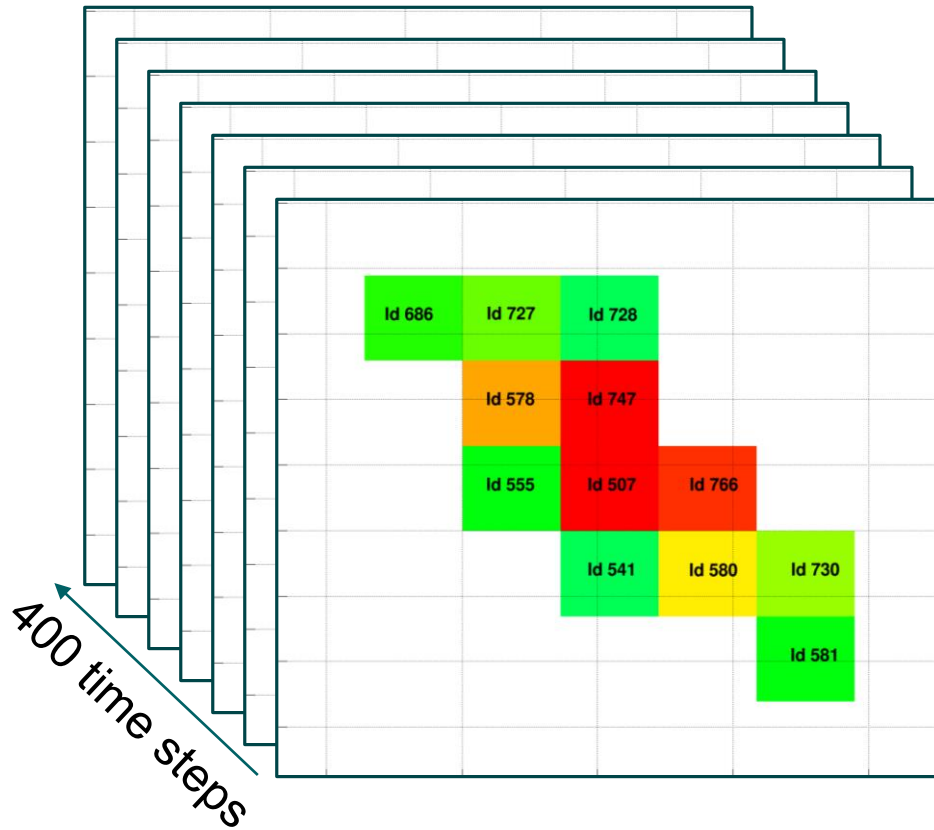
Dense

Dropout

Output

Shower Axis

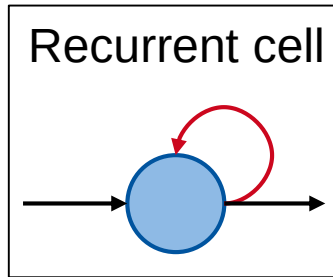
Spatiotemporal Input – Recurrent Networks



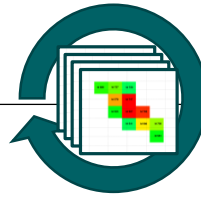
- Feed network a series of 2D images
- Per Event:
 - Stack of 400 shower maps

Architecture

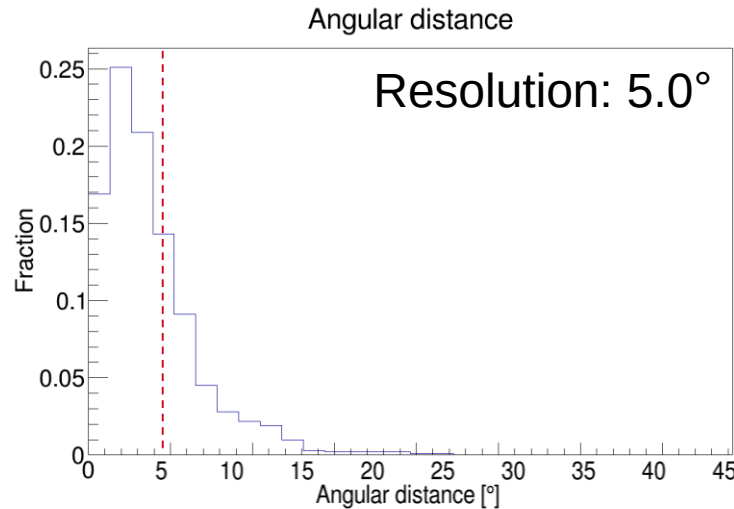
- LSTM + Convolutional Network



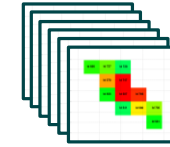
LSTM + Convolution



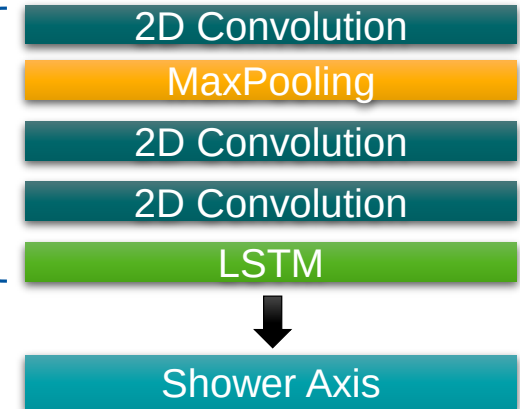
- 3 Convolutional layer
 - ReLu activation
- 1 Pooling layer
- Recurrent cell
 - Hard to train (long series)
- Only ~ 50,000 parameters



Input

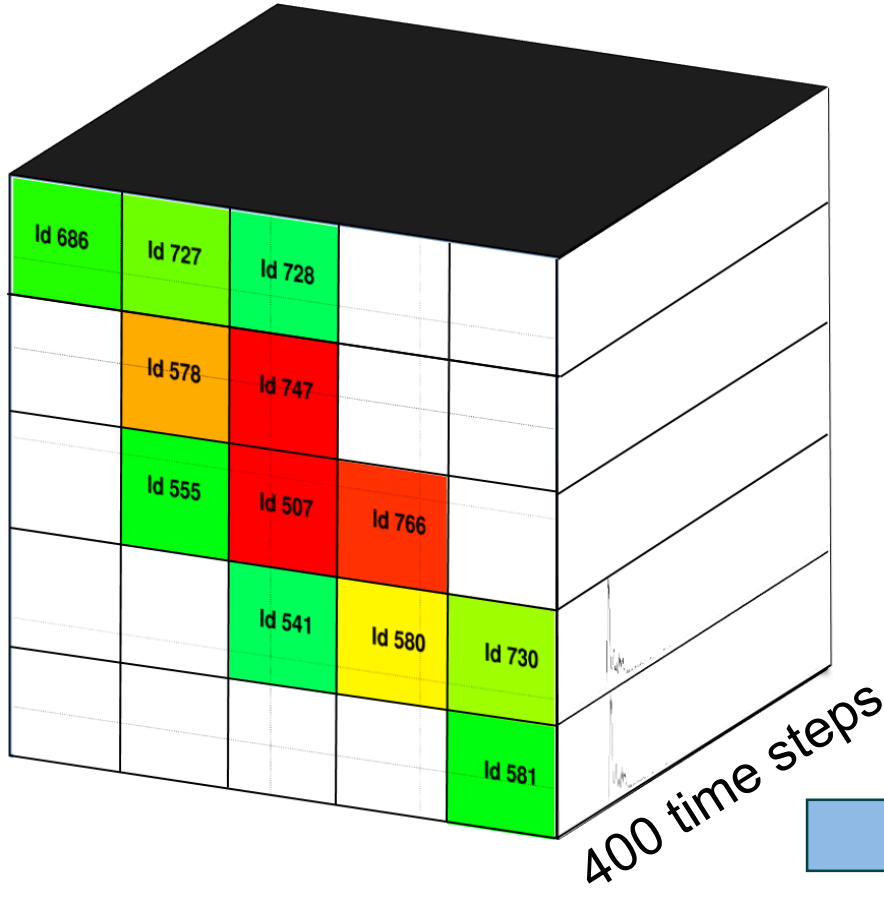


Hidden layer



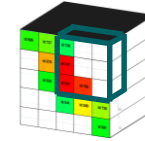
Output

Spatiotemporal Input



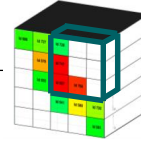
Architectures

- 11 Layer 3D Convolutional Network



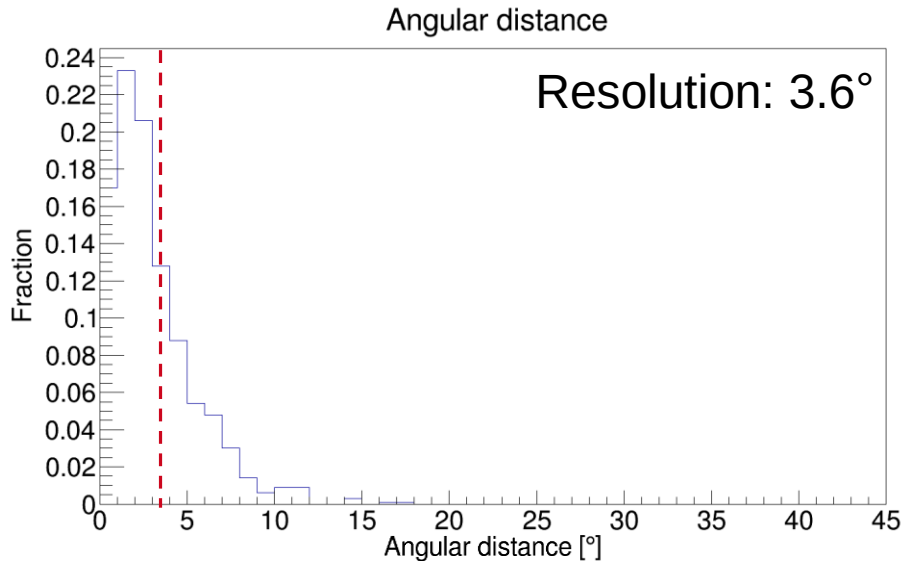
3D Footprint

3D Convolutional Network



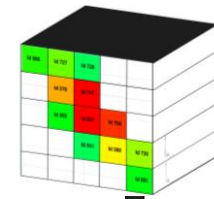
Input

- Same structure as 2D architecture
 - 11 Convolution layers
 - ReLu, Batch Normalization
 - Fully connected dense layer
 - Dropout



Hidden layer

Output



3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

3D Convolution

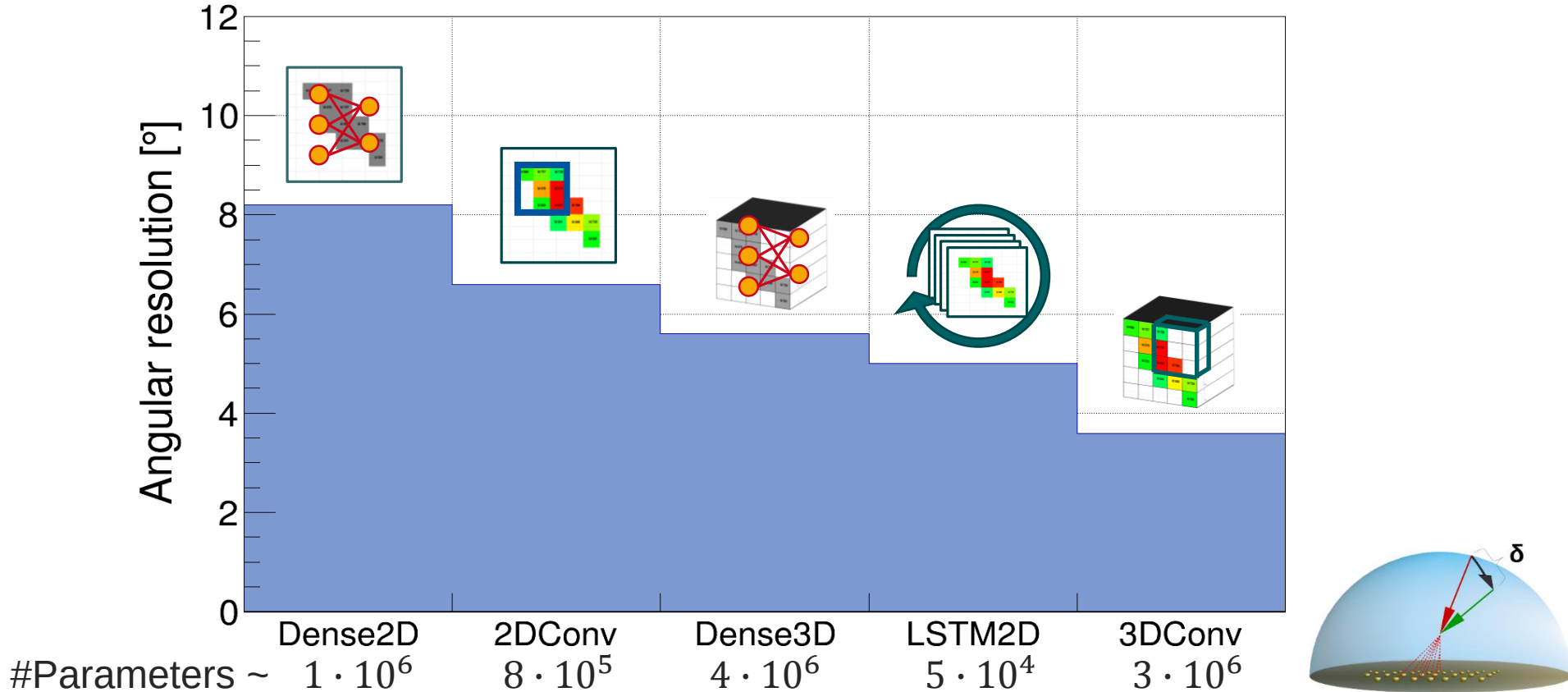
Dense

Dropout

Shower Axis

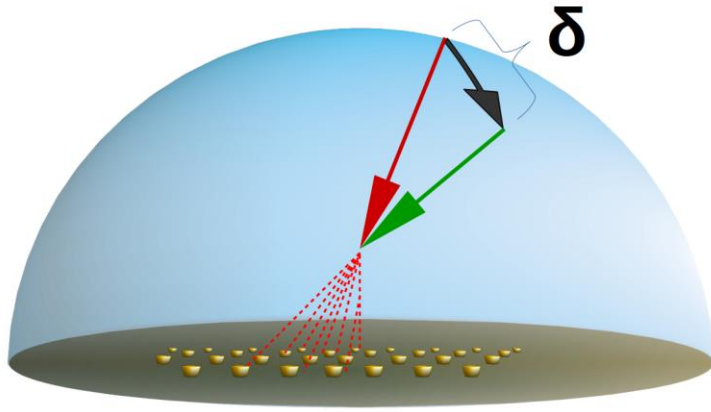
Different Network Architectures - Preprocessing

Angular resolution for different Network Architectures



Summary

- Successfully trained deep networks with up to 12 hidden layers
 - Reached resolution of 3.6°
- Based on row data
 - Without applying physical knowledge
- More complicated architectures perform better on spatiotemporal data
 - 3D Convolutions
 - LSTM powerful
 - But hard to train
- Scaling the input data can be most important



Backup

Exploring Deep Network Architectures to reconstruct Cosmic Ray Induced Air Showers at the Pierre-Auger-Observatory

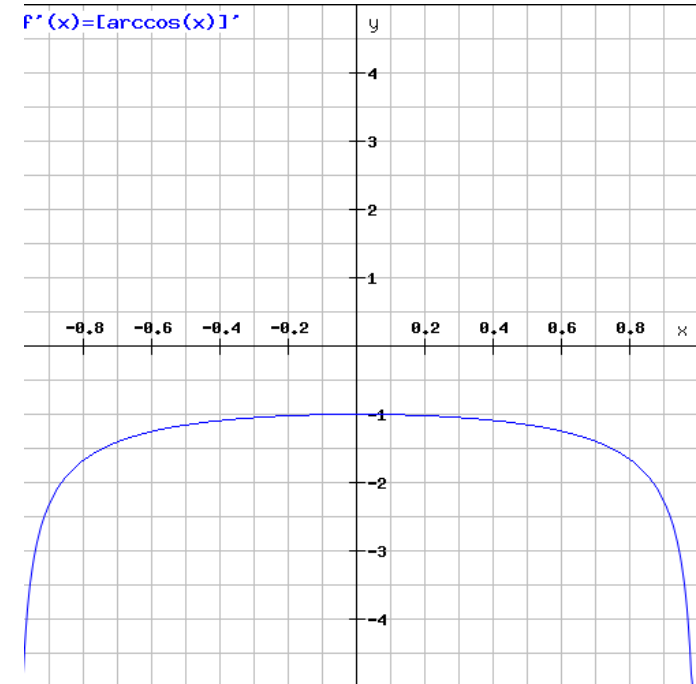
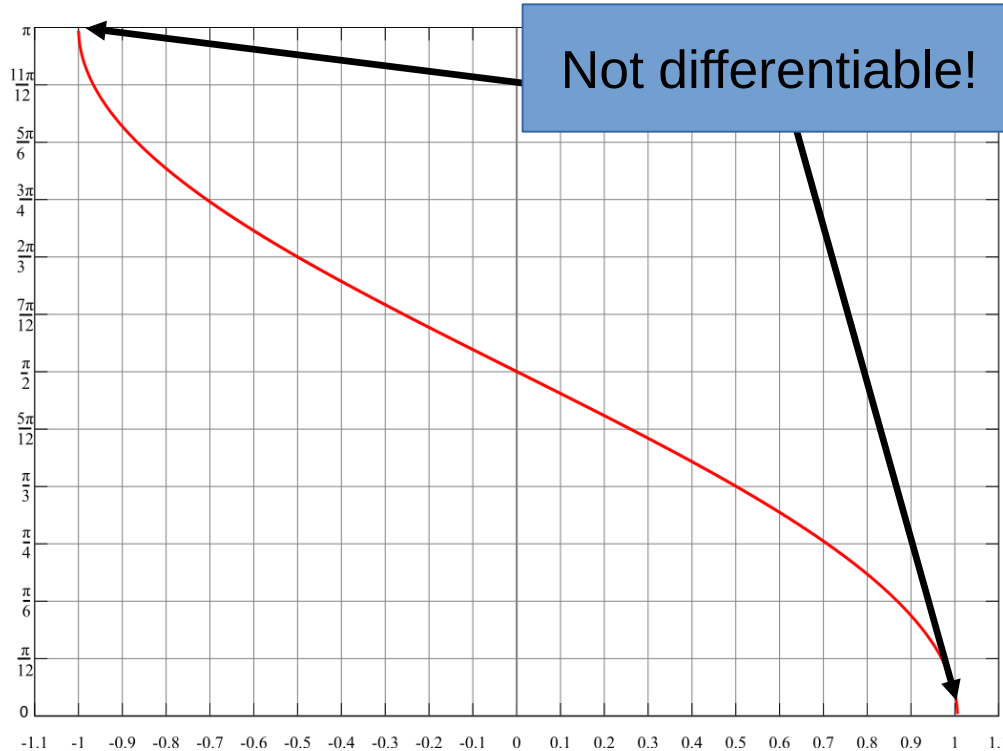
Jonas Glombitza, Martin Erdmann, David Walz, Marcus Wirtz

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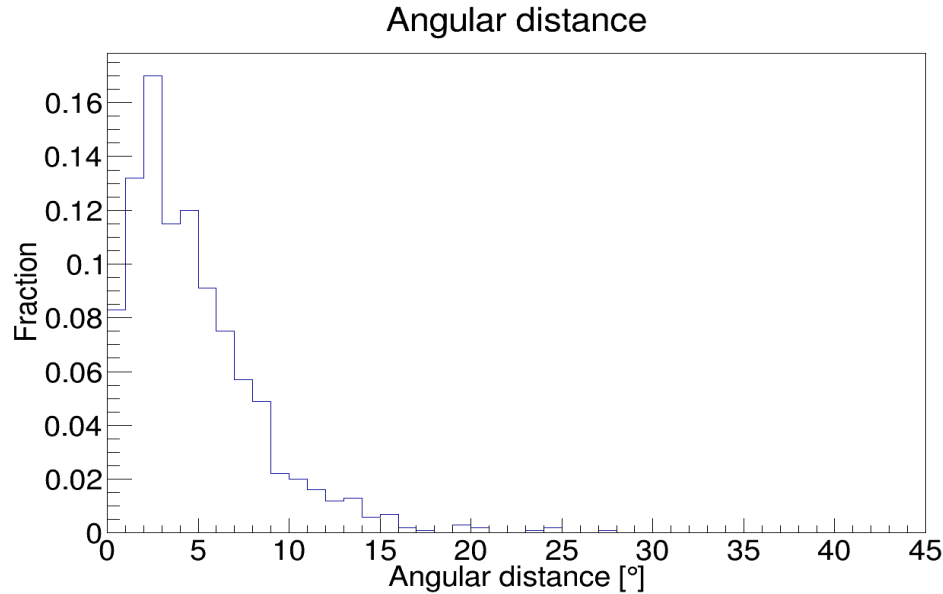
Network Performance – No Preprocessing

Use Angular Distance as Loss

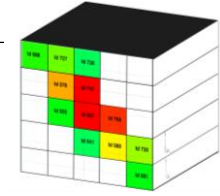
- Reconstruct Θ, Φ
- Angular distance as Loss $L = \arccos(\vec{x} \cdot \vec{y})$, $\vec{x}$ and $\vec{y}$ in spherical coordinates



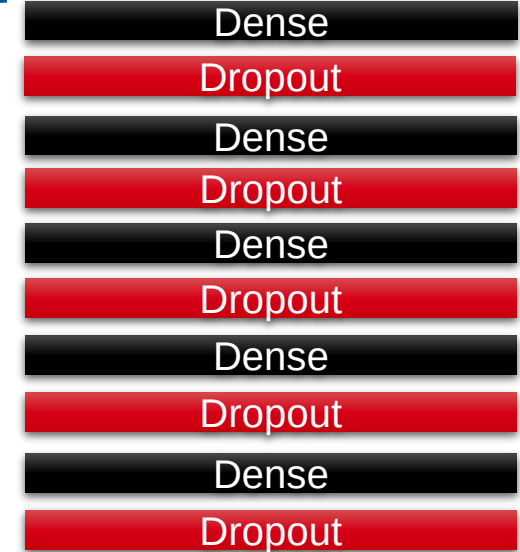
Fully connected 3D



Input



Hidden layer



Output

Shower Axis