

An aerial photograph of a radio telescope array. The array consists of a large number of dark, rectangular panels arranged in a grid pattern on a cleared, sandy area. The array is enclosed by a low metal fence. In the background, there is a line of trees and some residential buildings. The sky is clear and blue.

**Progress Report:  
LOFAR observation of  
NGC4631**

**Stefan Blex  
Astronomisches Institut  
Ruhr-Universität Bochum**



# Contents

- Motivation
- Observation data
- Processing steps:
  - Manual
  - After Prefactor
  - 'After' Factor
- Basic Analysis

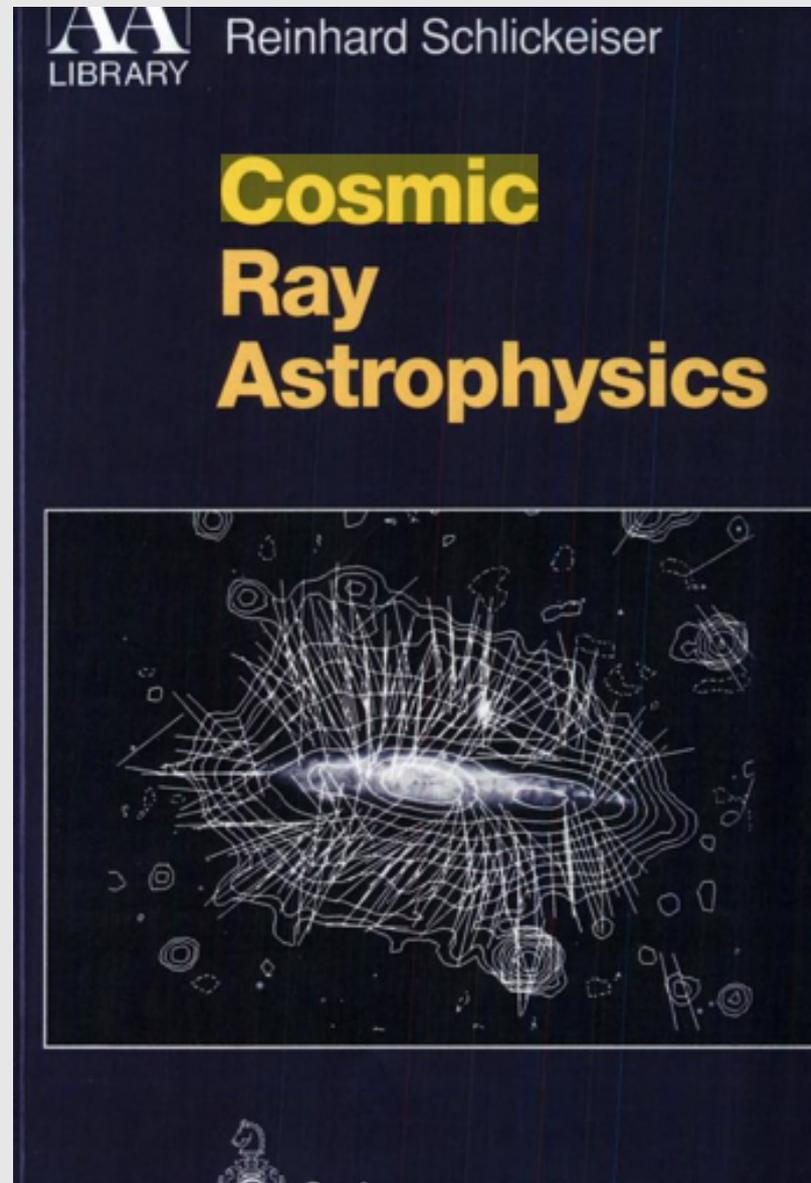


# Motivation

- large edge-on spiral galaxy
- distorted: interaction with NGC4627 and 4656/7
- central starburst
- galactic wind, X-ray halo



# Motivation



# Motivation



# Motivation

H $\alpha$

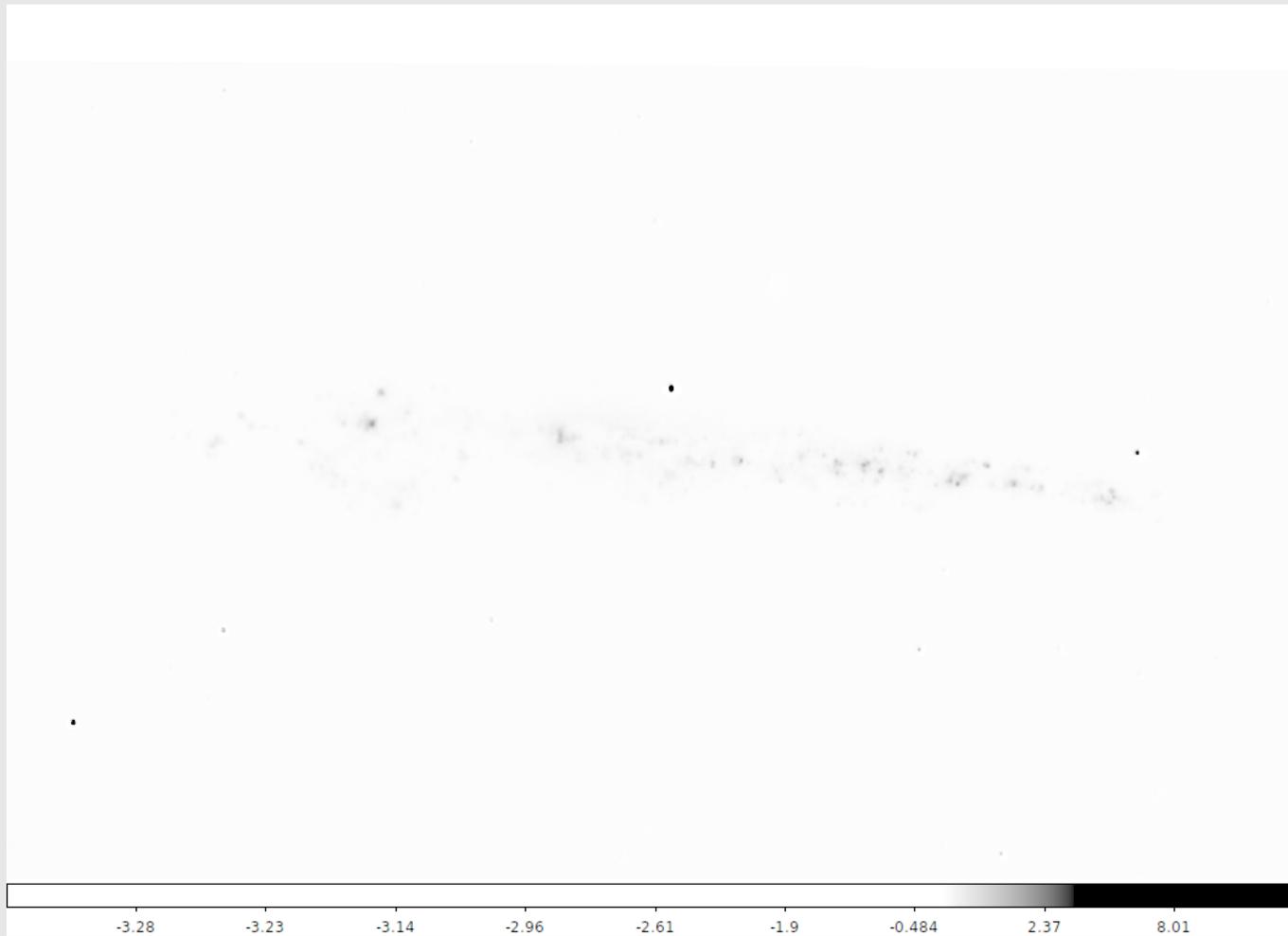


Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015

# Motivation

18cm

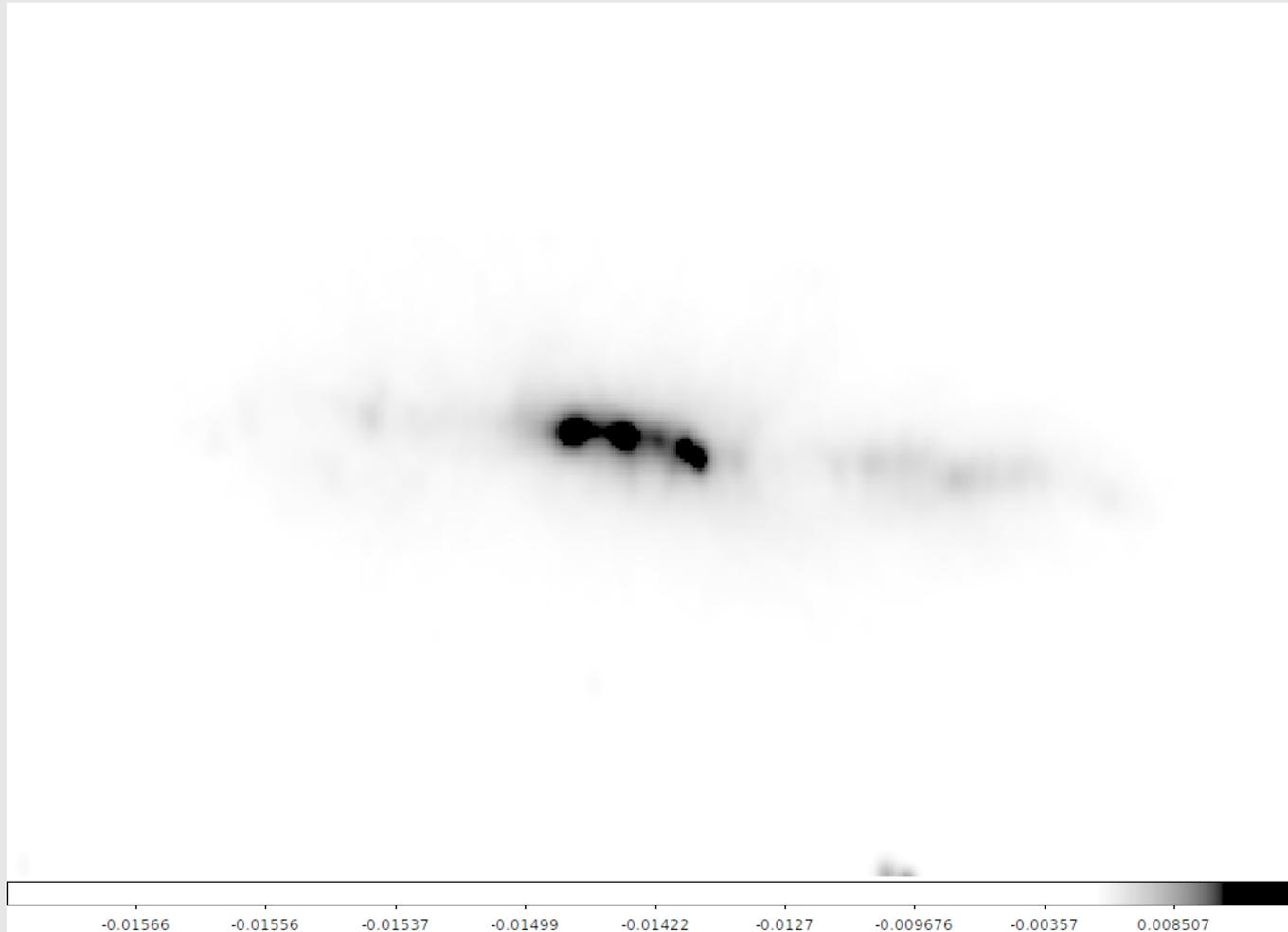
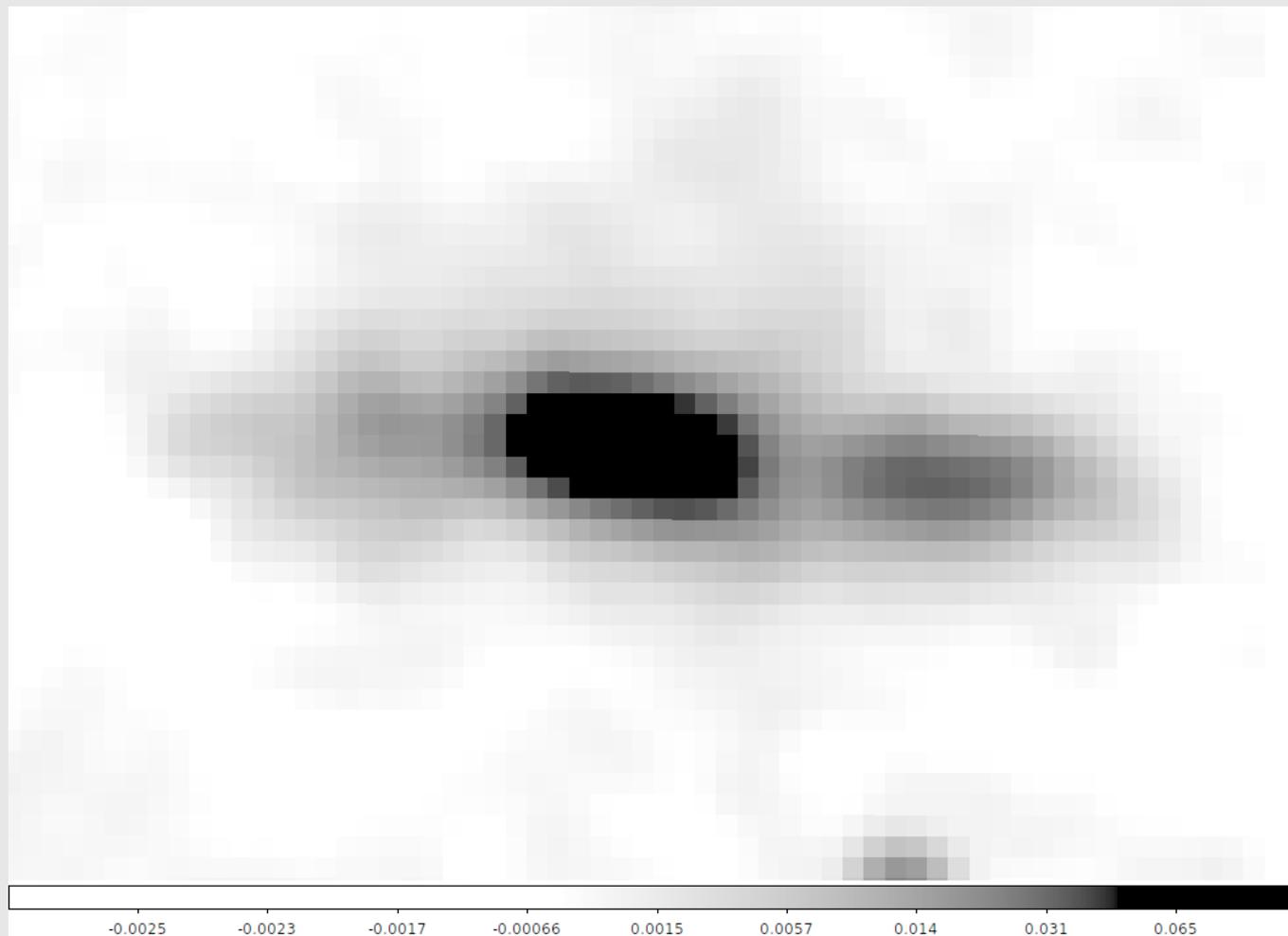


Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015

# Motivation

## NVSS



# Motivation

## VLSS

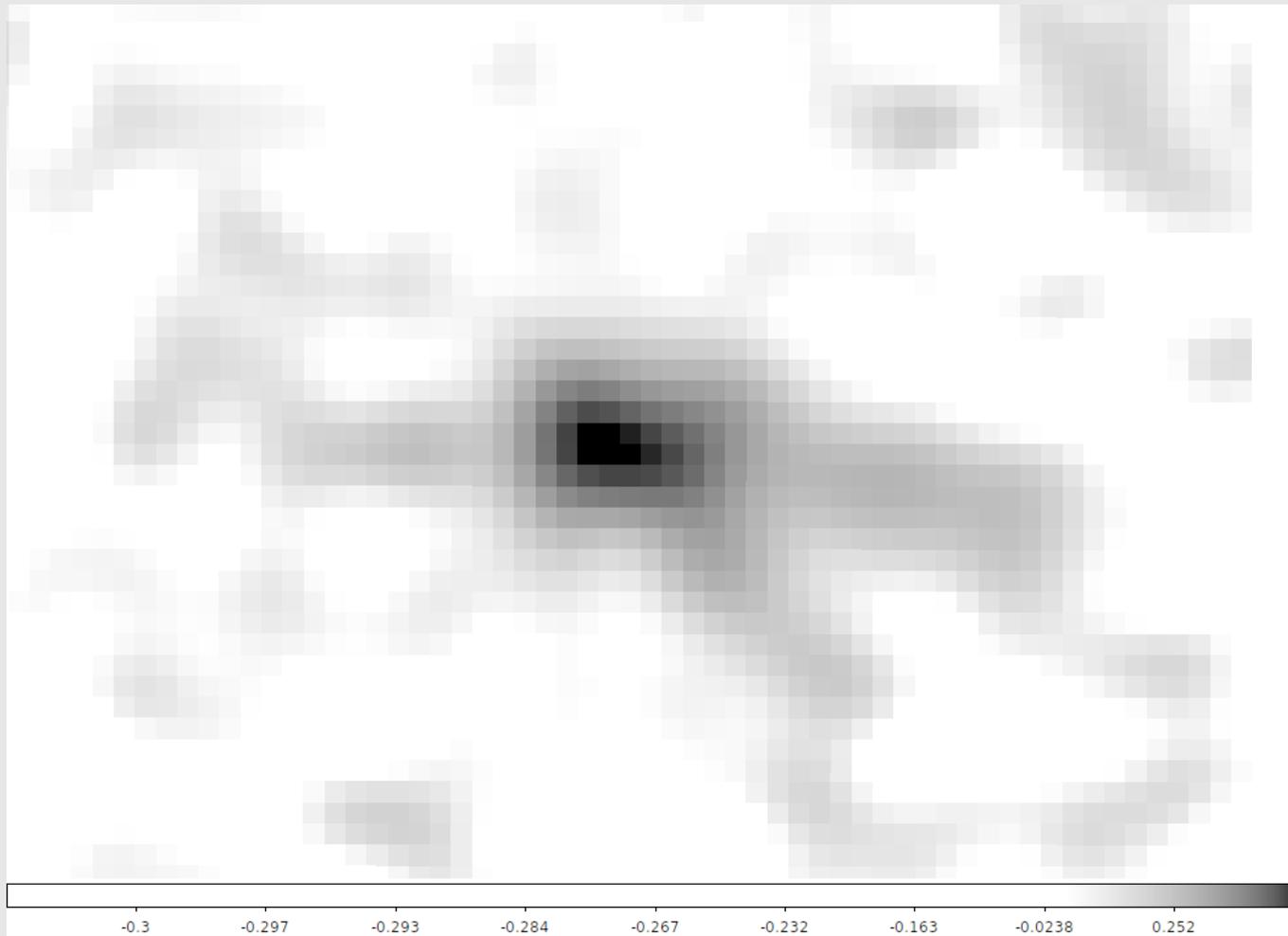


Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015



# The Observation

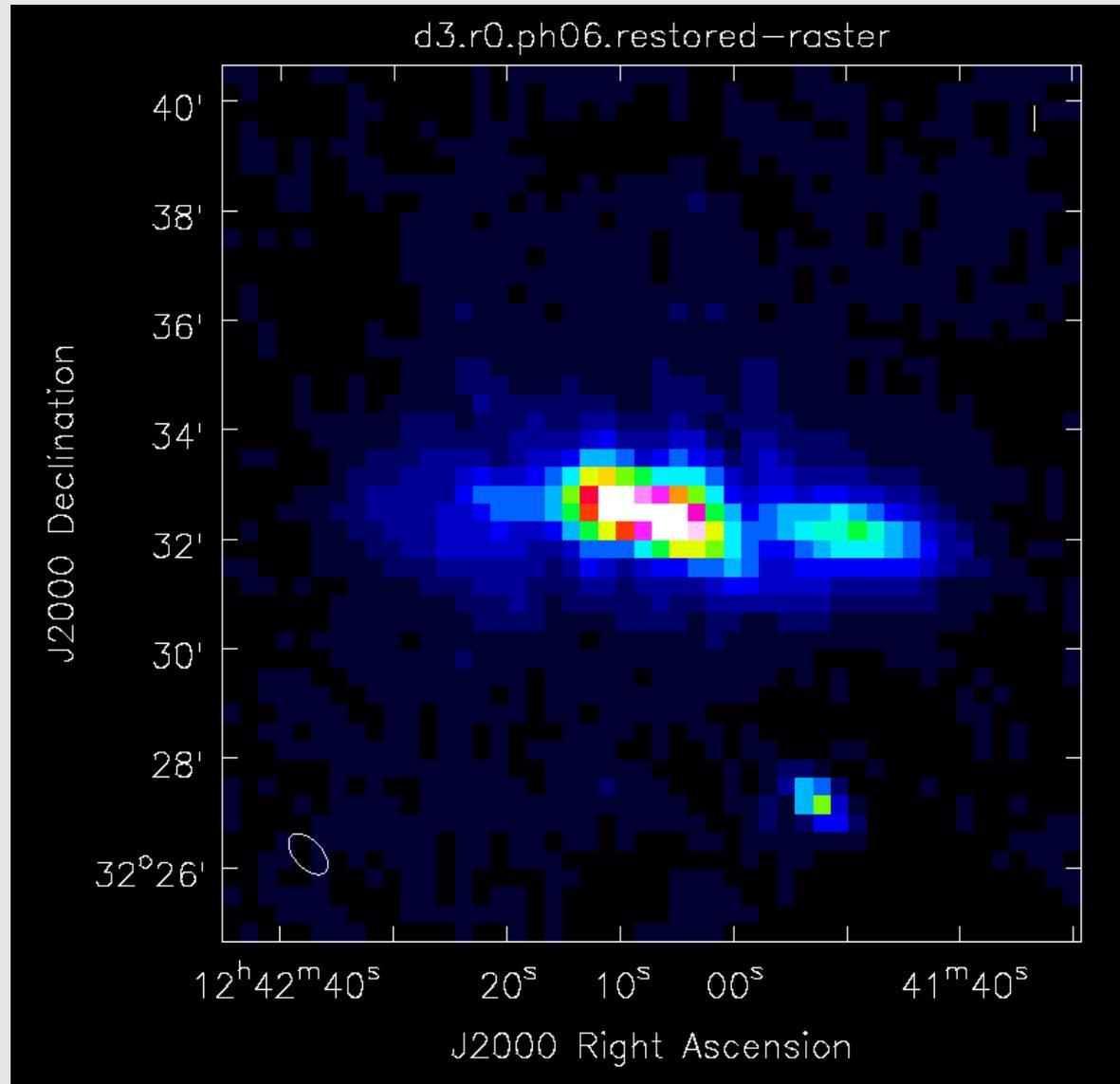
- NGC4631, Calibrator: 3C286
- Cycle 0 (18/19.04.13)
- Core & Remote Stations
- HBA, 324 Subbands
- Interlaced, 6.4h target, 1h cal.
- No preprocessing



# Manual calibration

- Demixing (trial and error)
- Flagging (AOFlagger, then by hand)
- Self-calibration
- *(Imaging)*

# Manual calibration

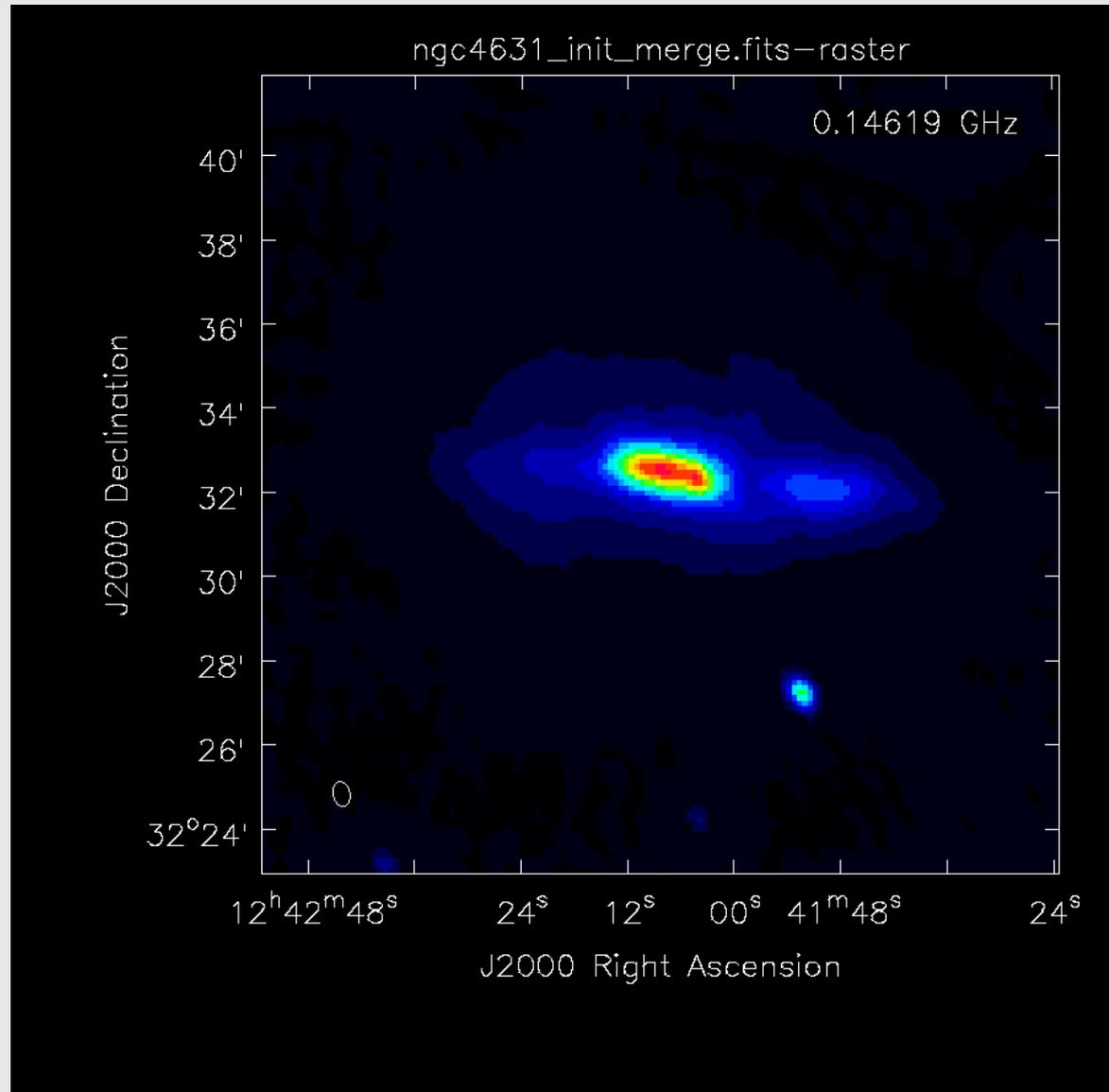




# Prefactor

- Long wait for a version that supports:
  - Data without preprocessing
  - Interlaced observations
- Initially, much fiddling required
- Current version runs a lot smoother
- Local bottleneck problem finally disappeared

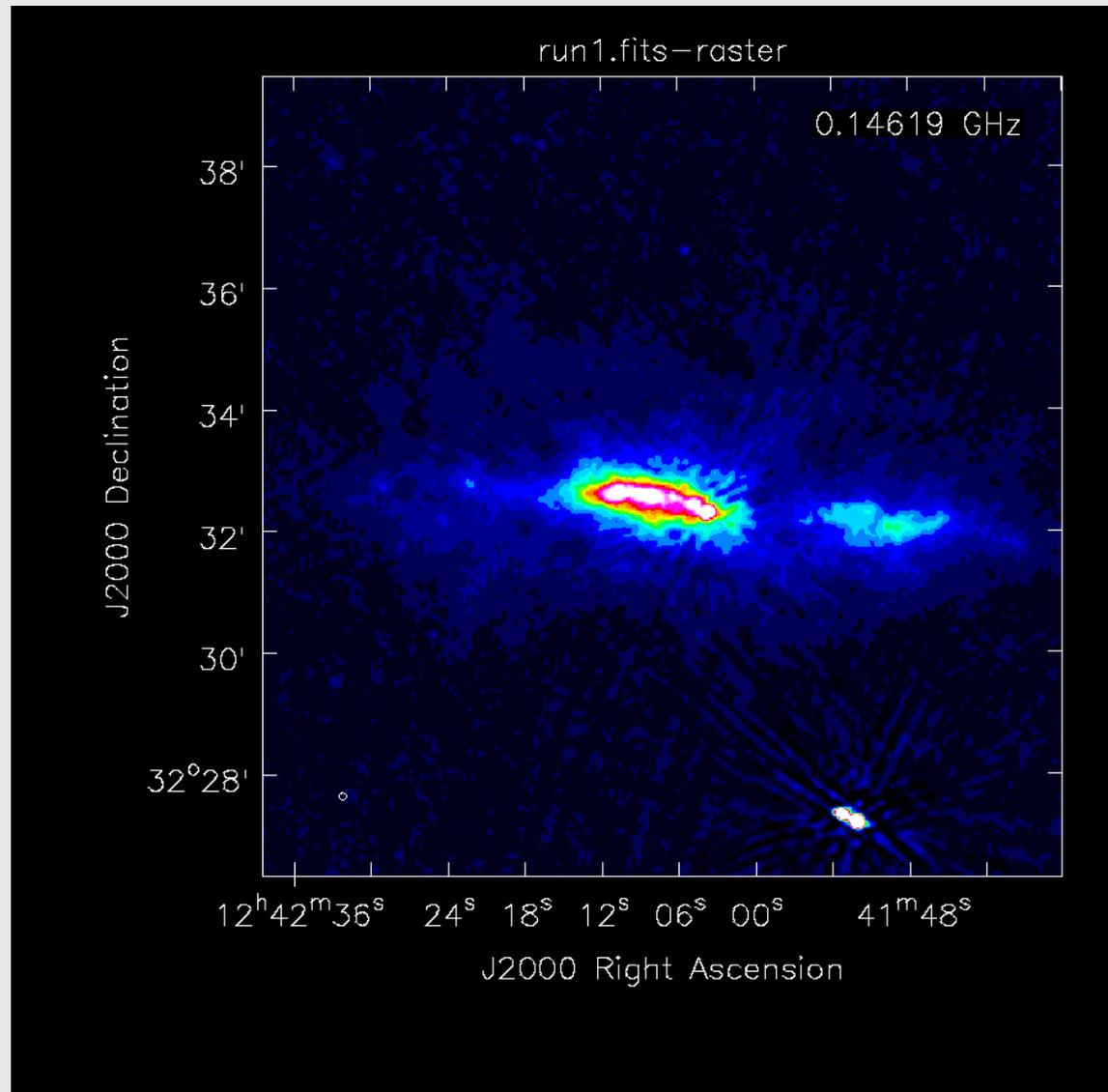
# Prefactor



# Factor

- Disregarding user error and some prerequisite installation trouble, runs smoothly on 1<sup>st</sup> try!
- Pretty harsh demands on hardware
- Settings are trial and error, at least for me (Andreas Horneffer helped a lot)
- I will need at least 3 full runs of Factor until satisfied with the result, 2<sup>nd</sup> run just finished

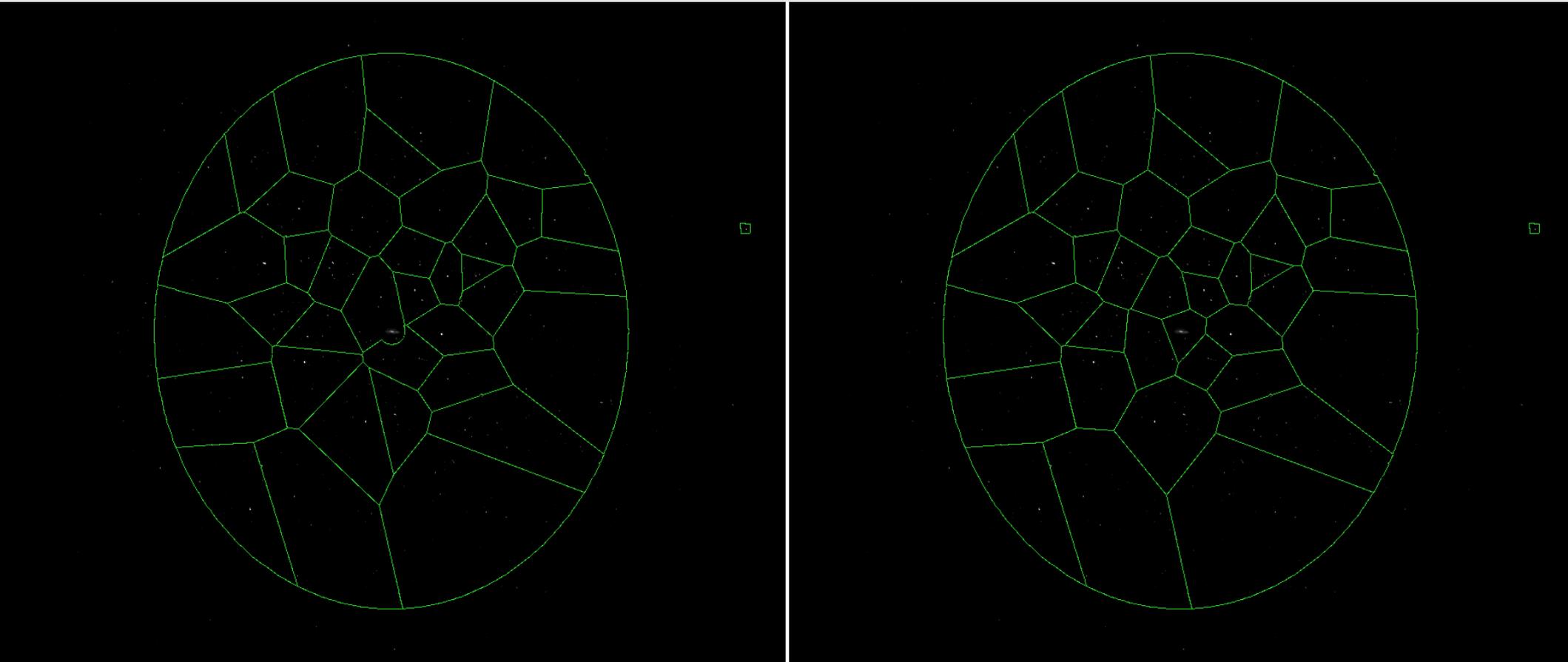
# Factor – 1<sup>st</sup> run



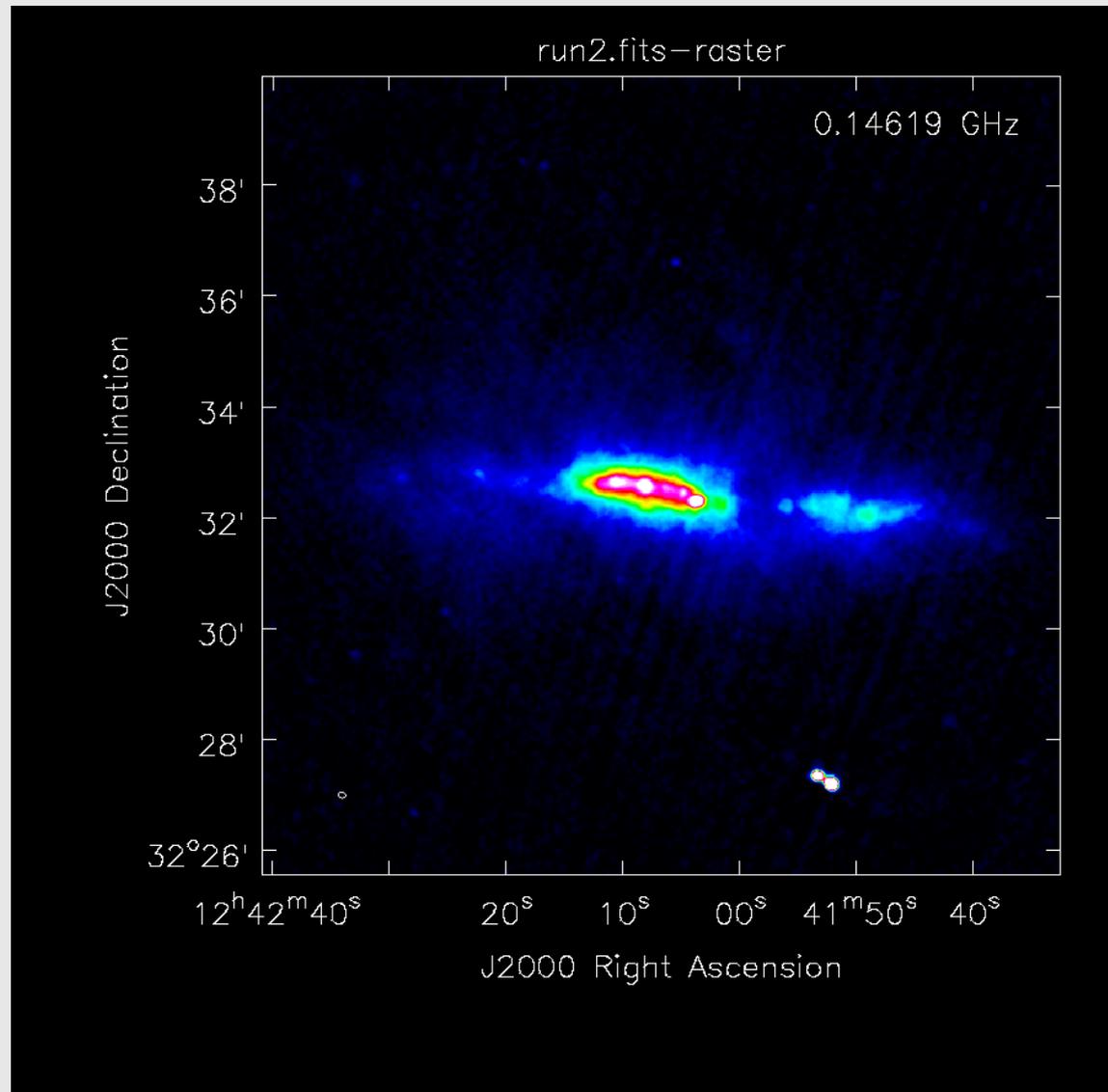


# Factor – facet adjustment

RUB



# Factor – 2<sup>nd</sup> run



# Analysis

## H $\alpha$

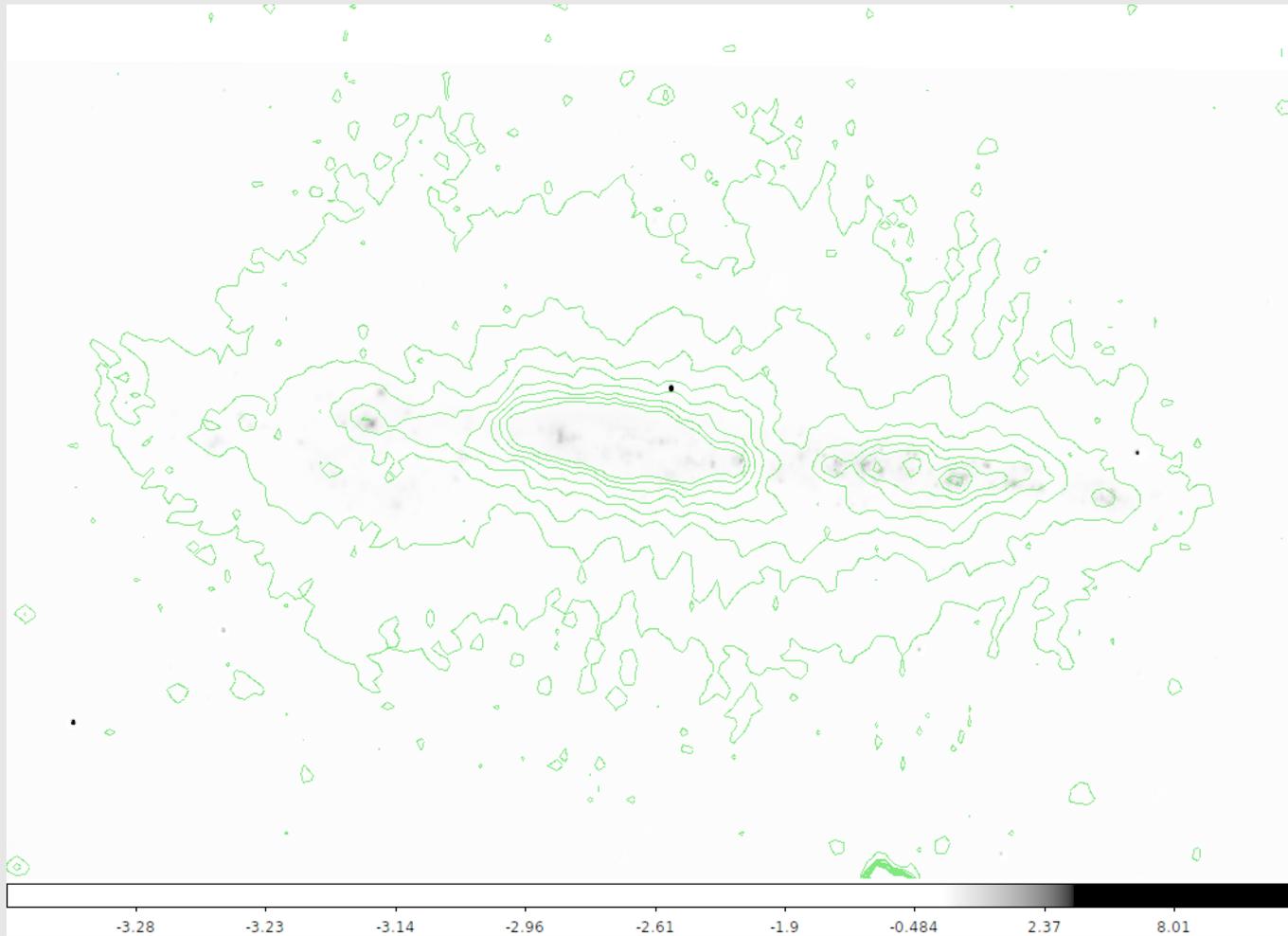


Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015

# Analysis

## 18cm

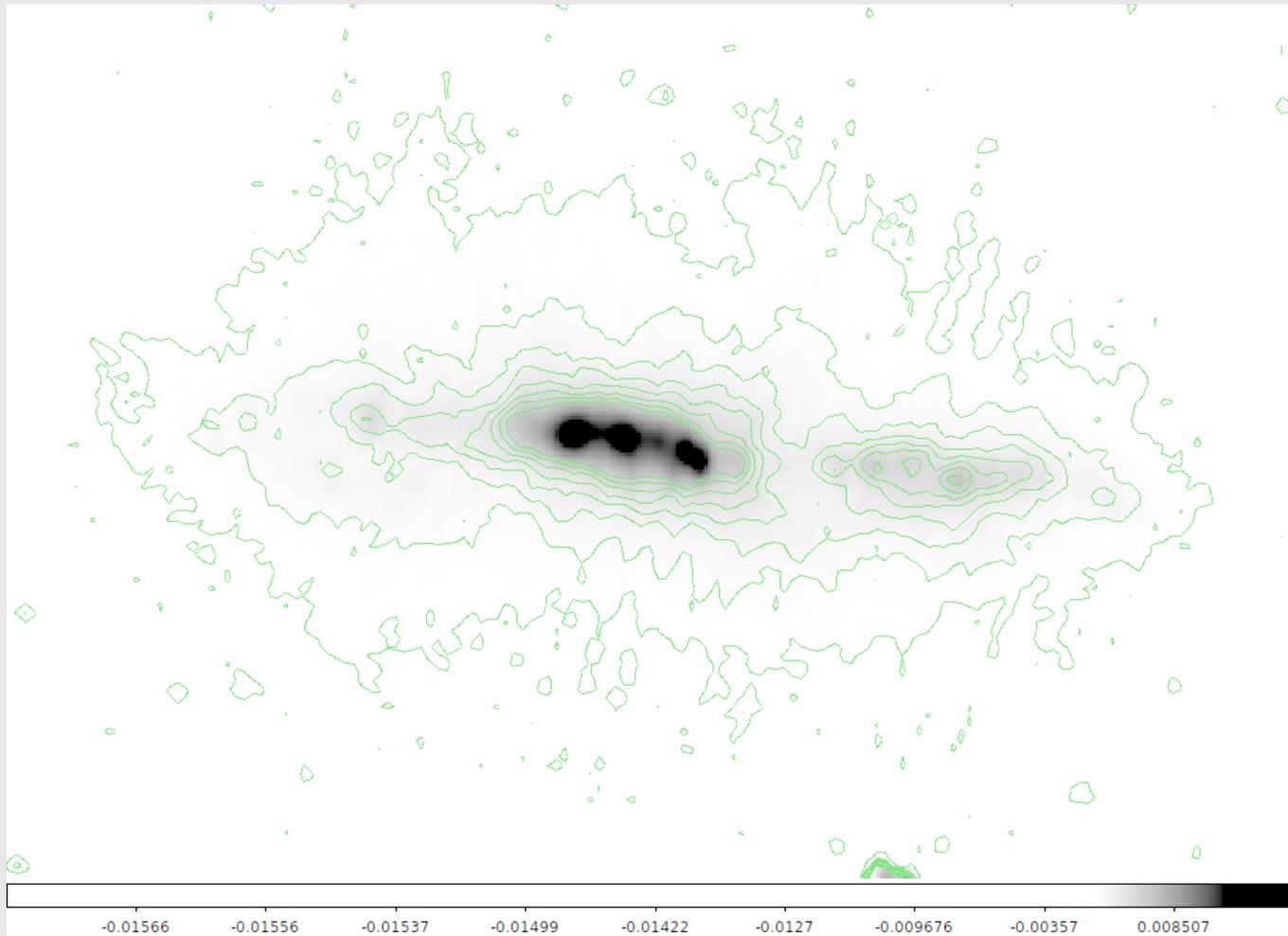


Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015

# Analysis

## NVSS

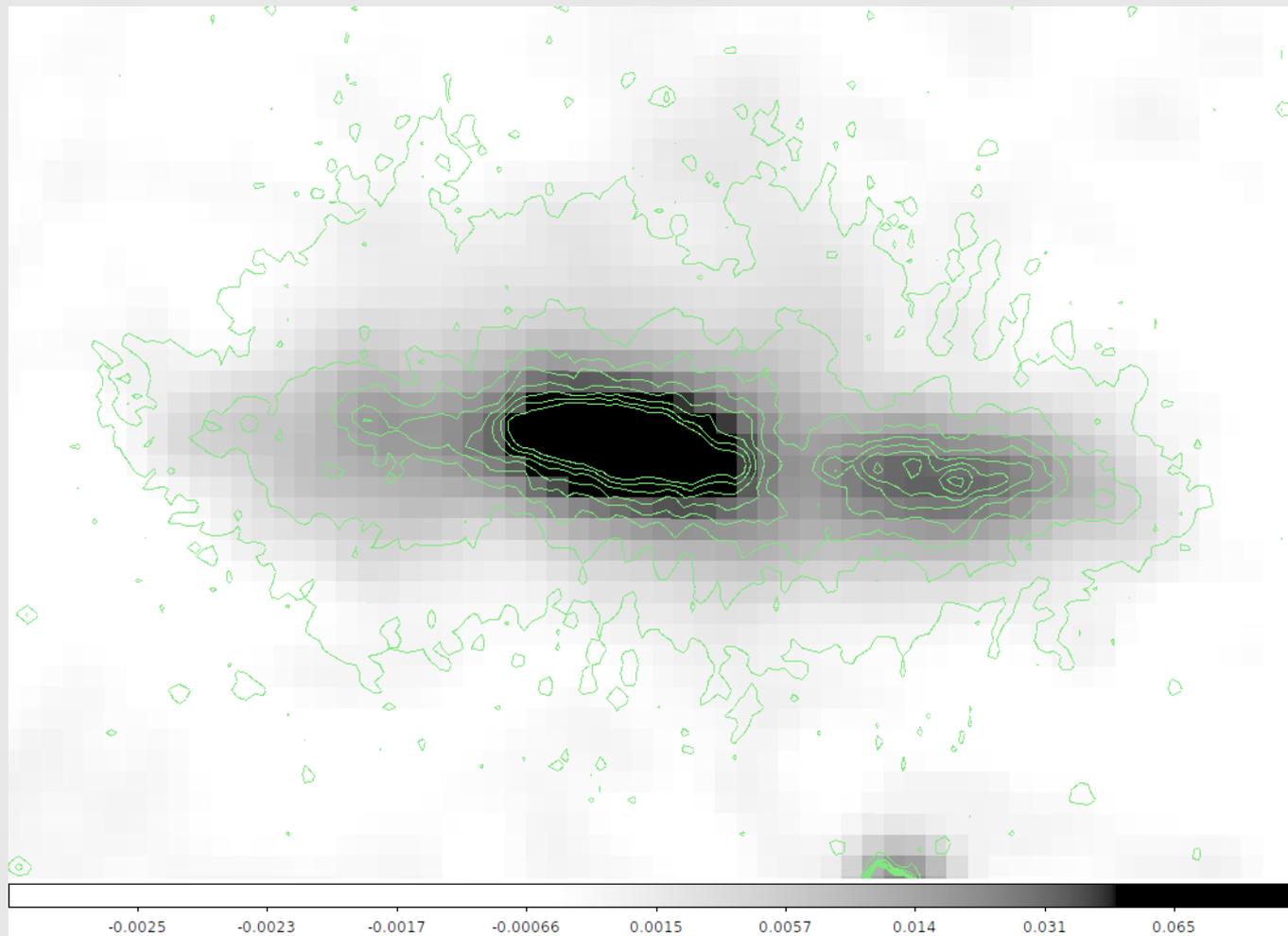


Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015

# Analysis

## VLSS

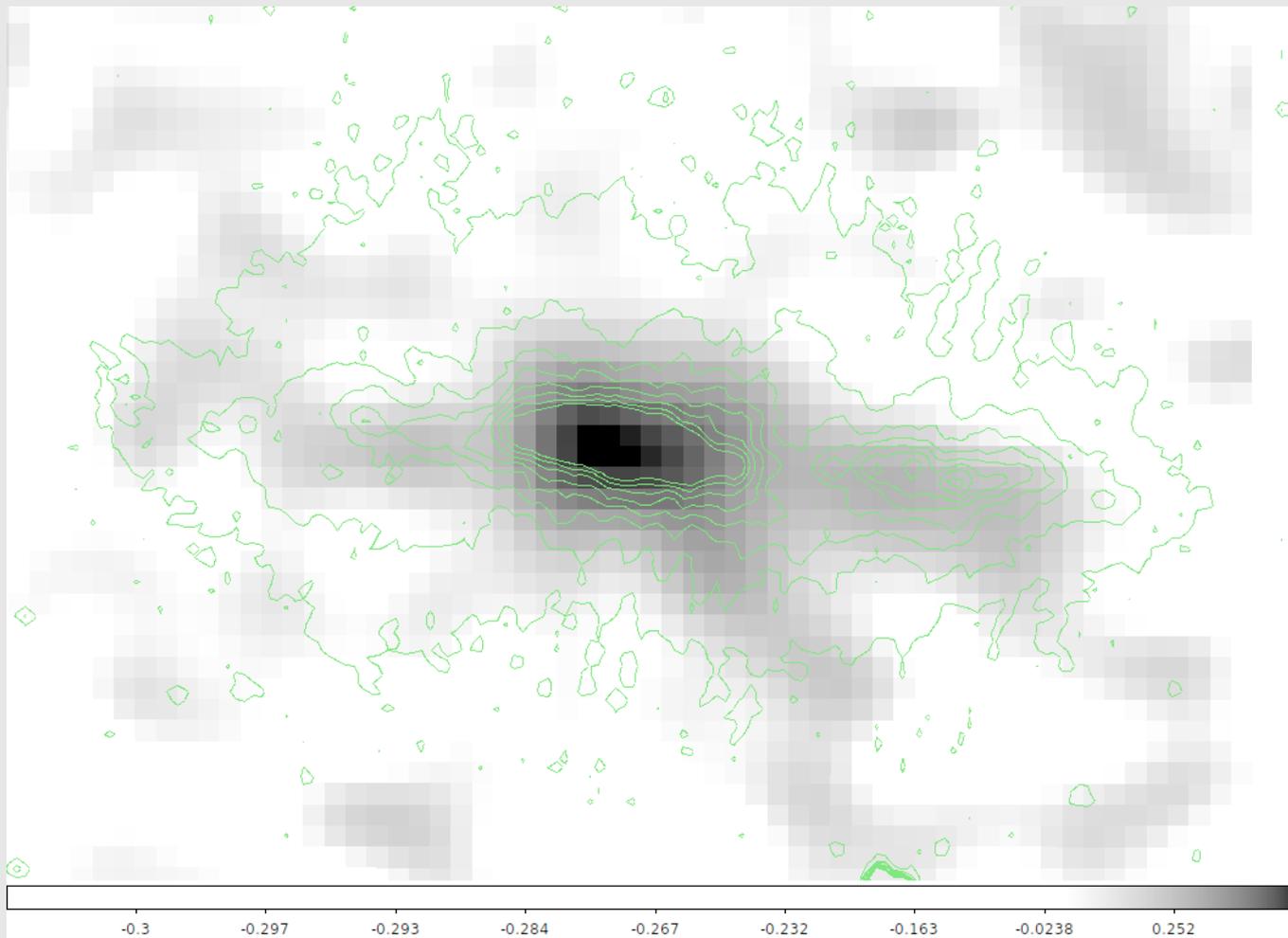
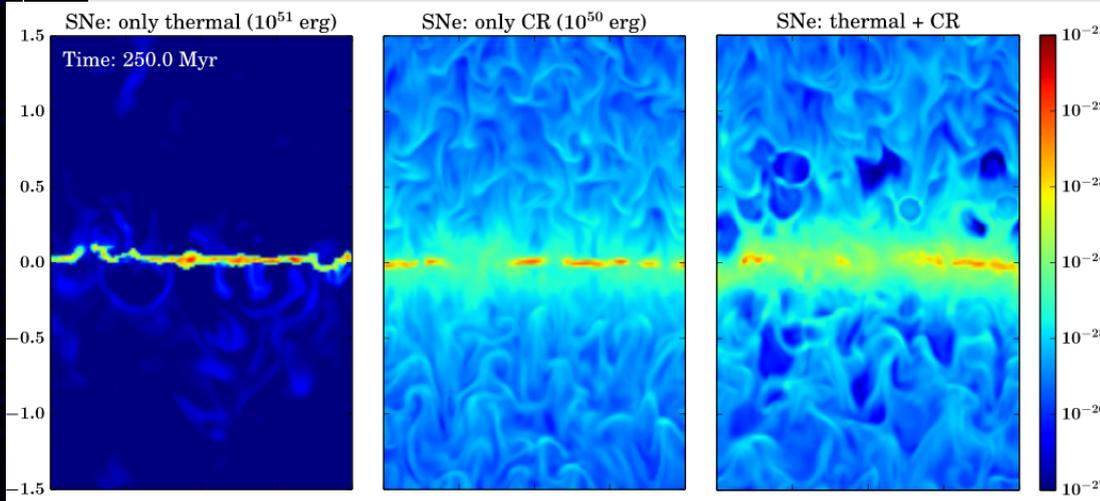
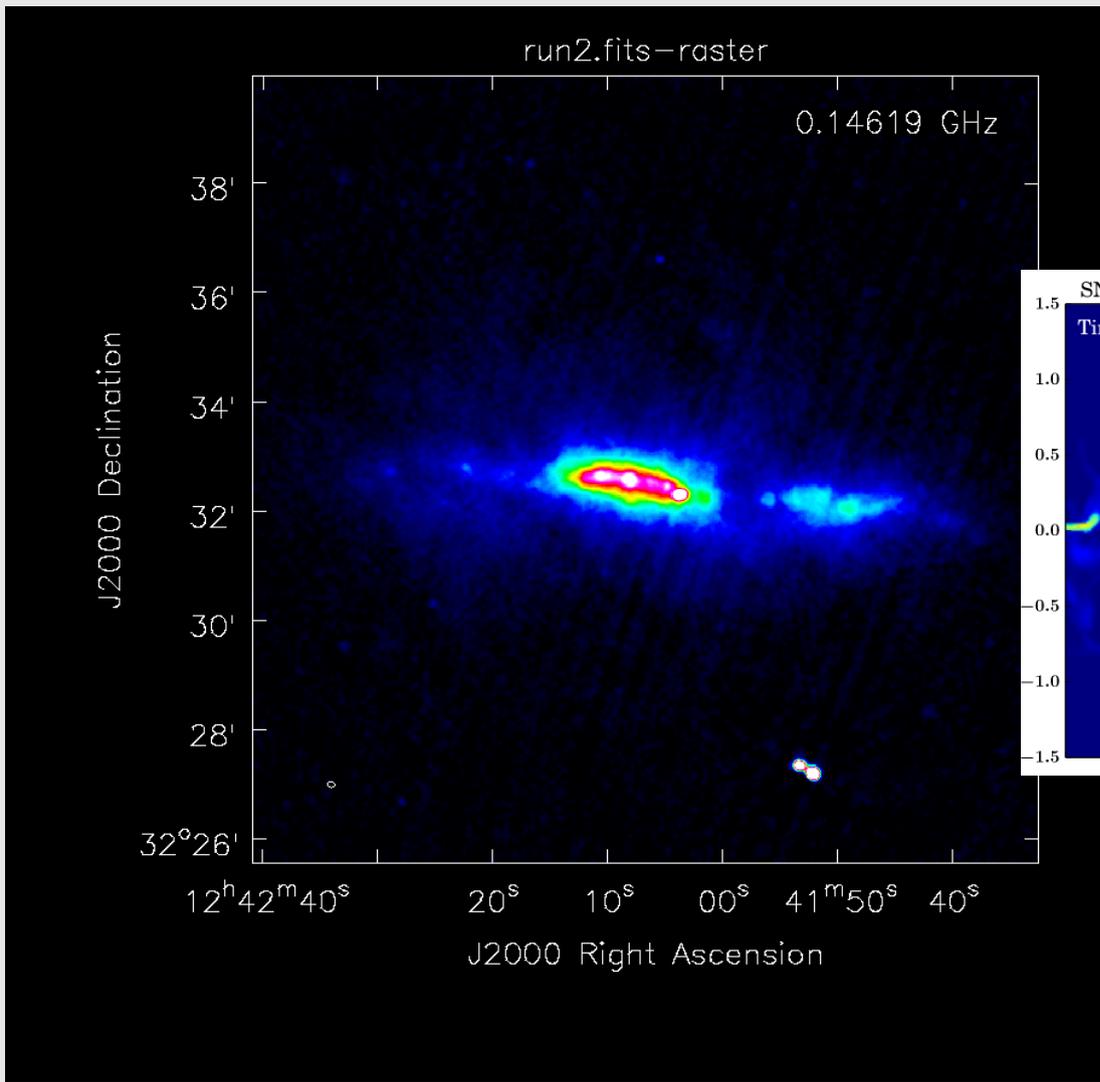


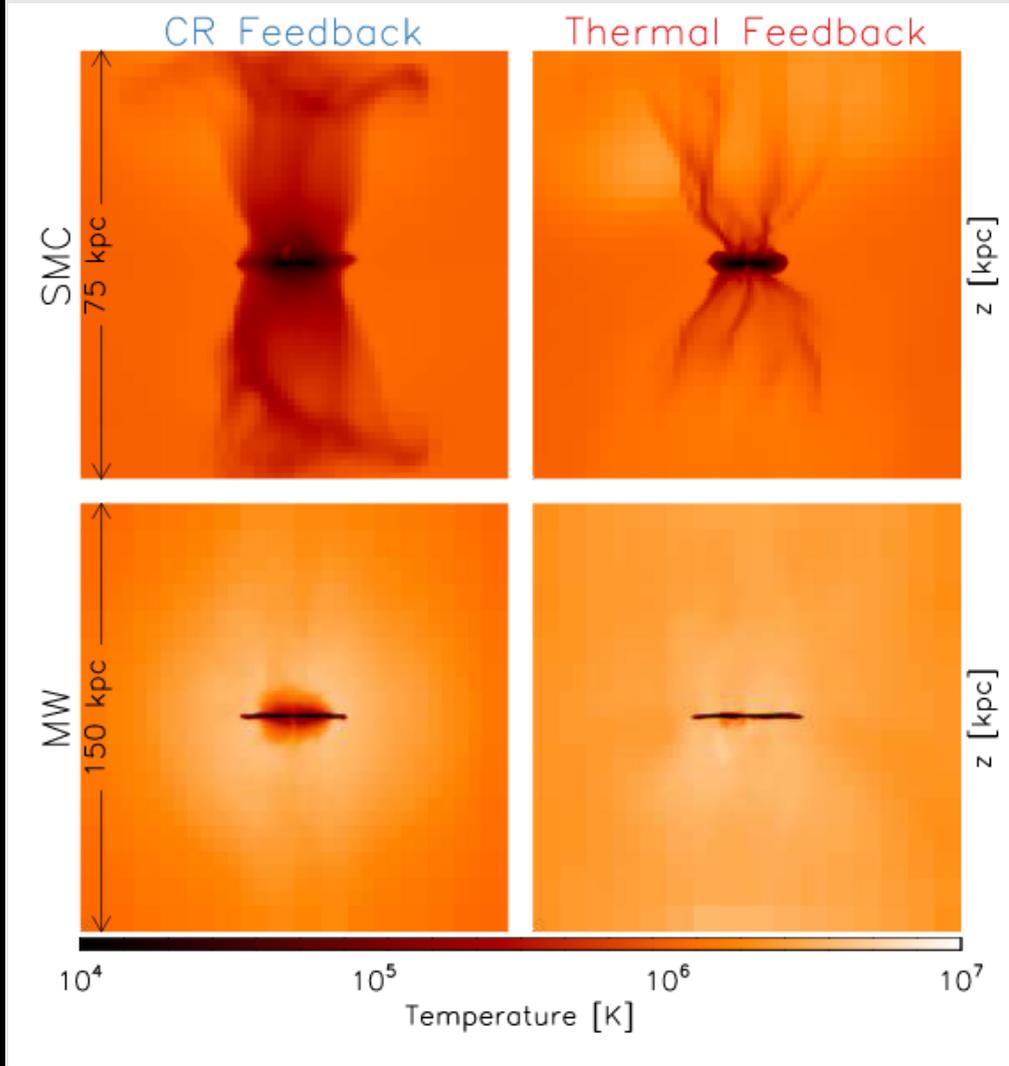
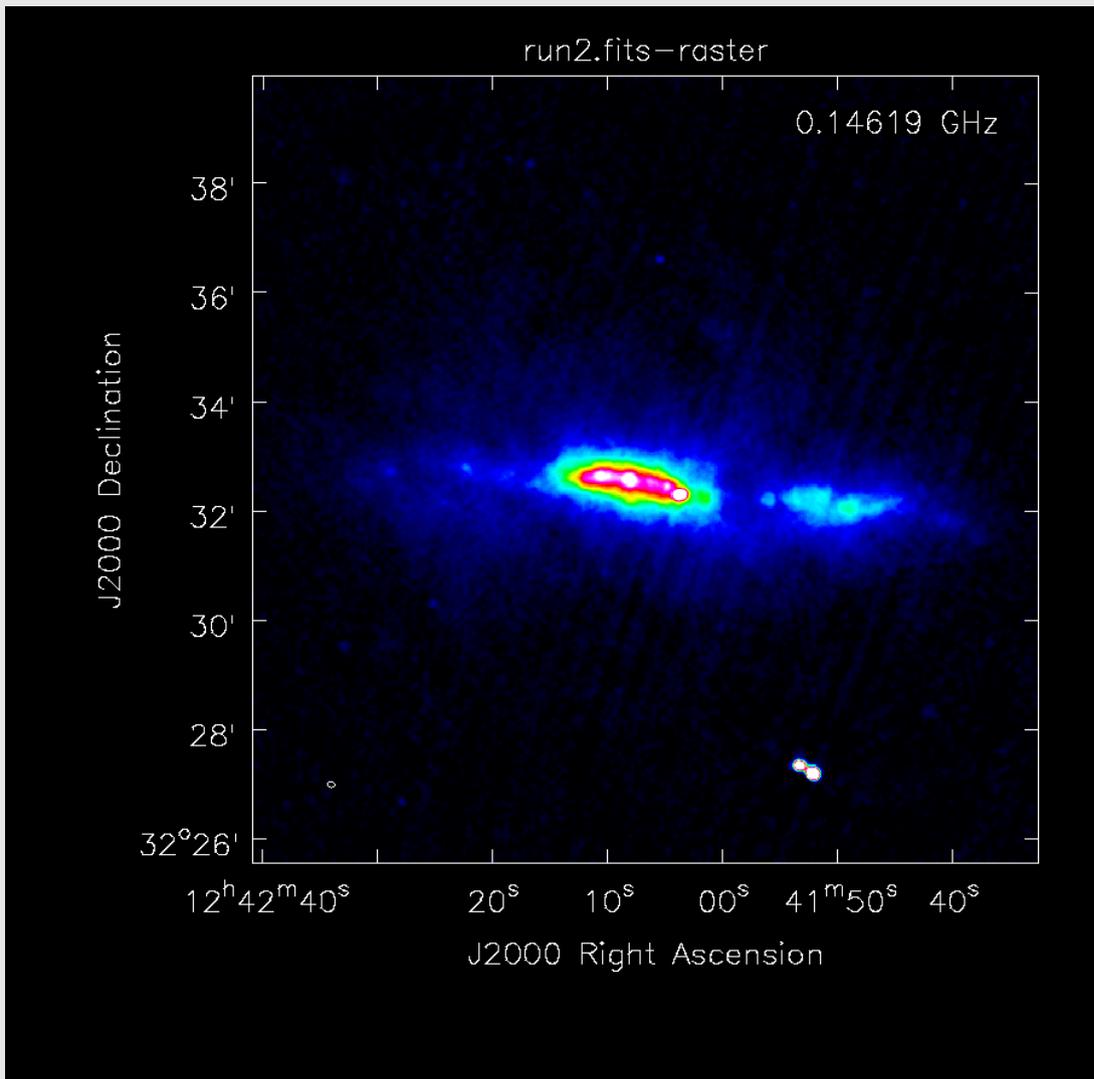
Image credit: Kennicutt et al. 2003  
reprocessed: D.J. Bomans 2015

# Analysis

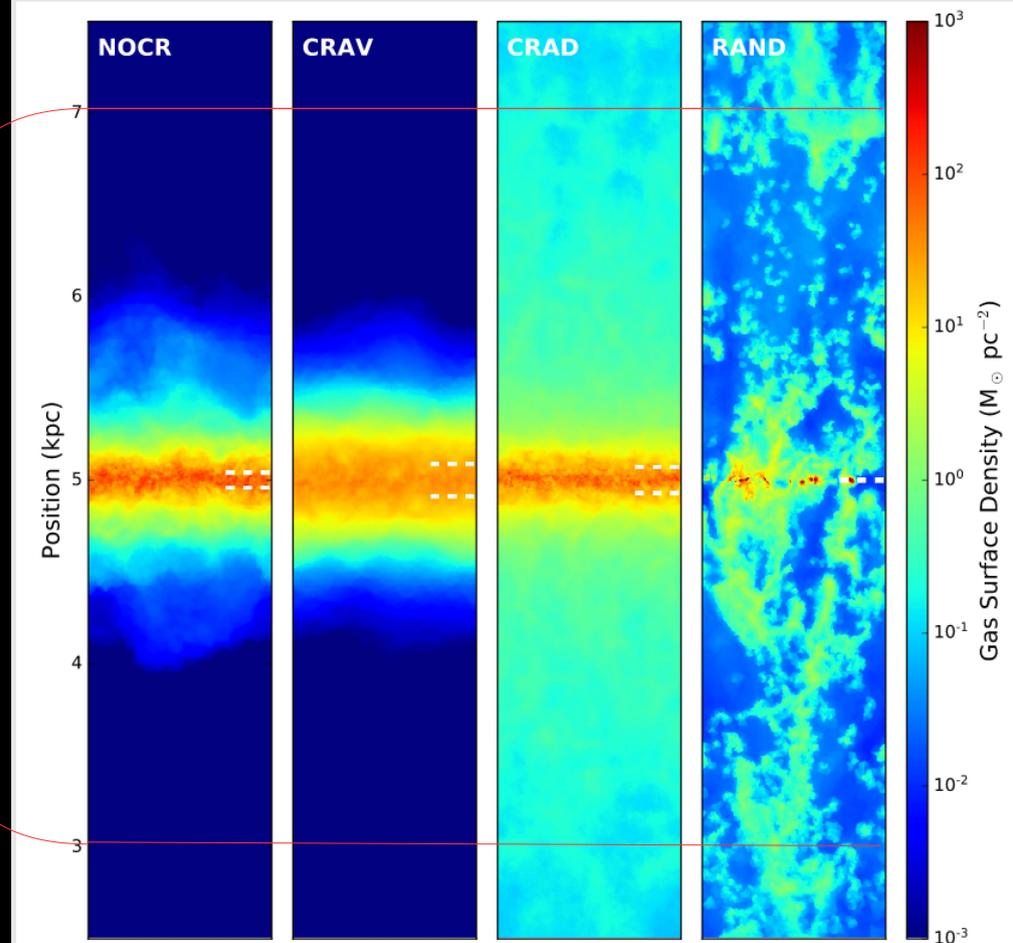
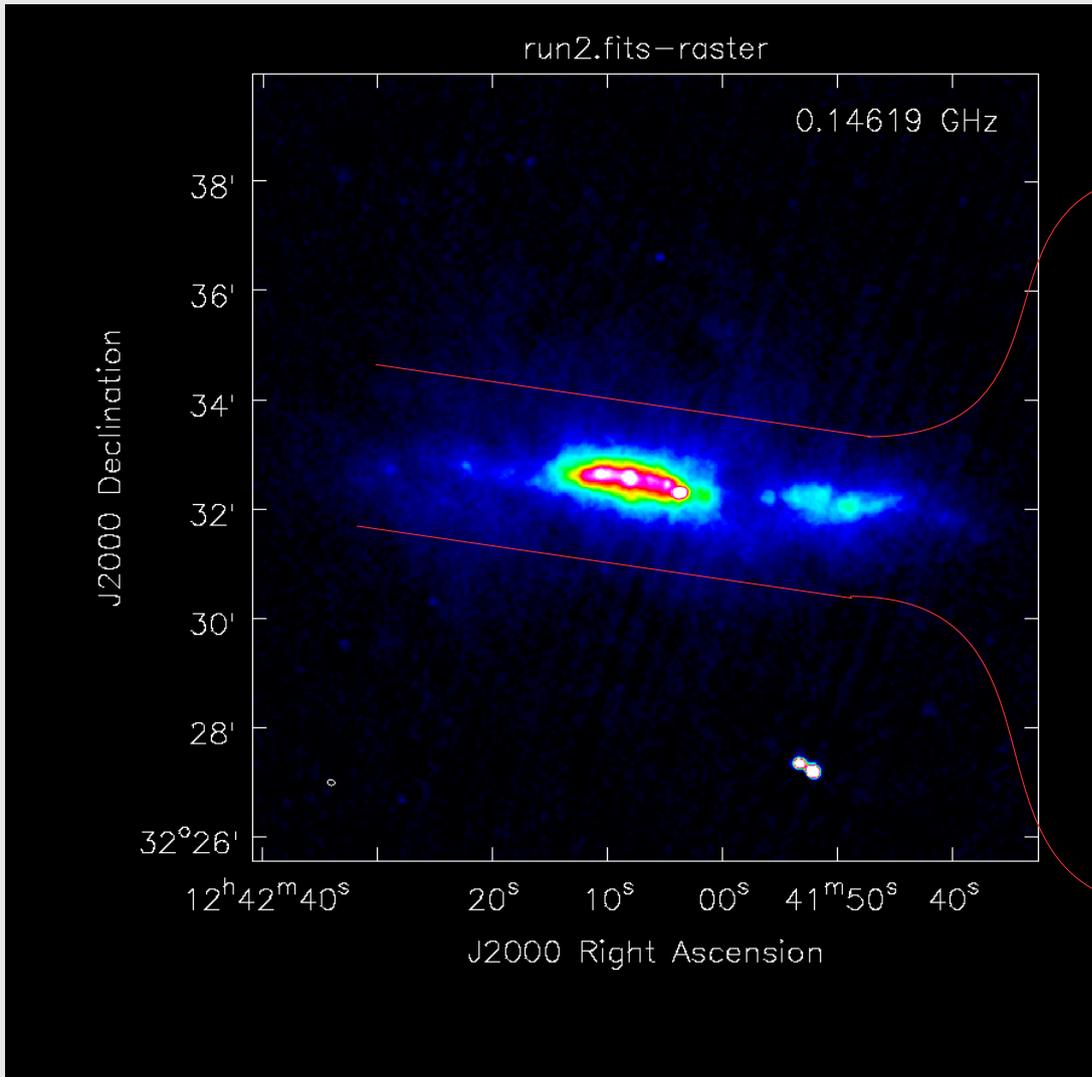


Girichidis et al., 2016

# Analysis

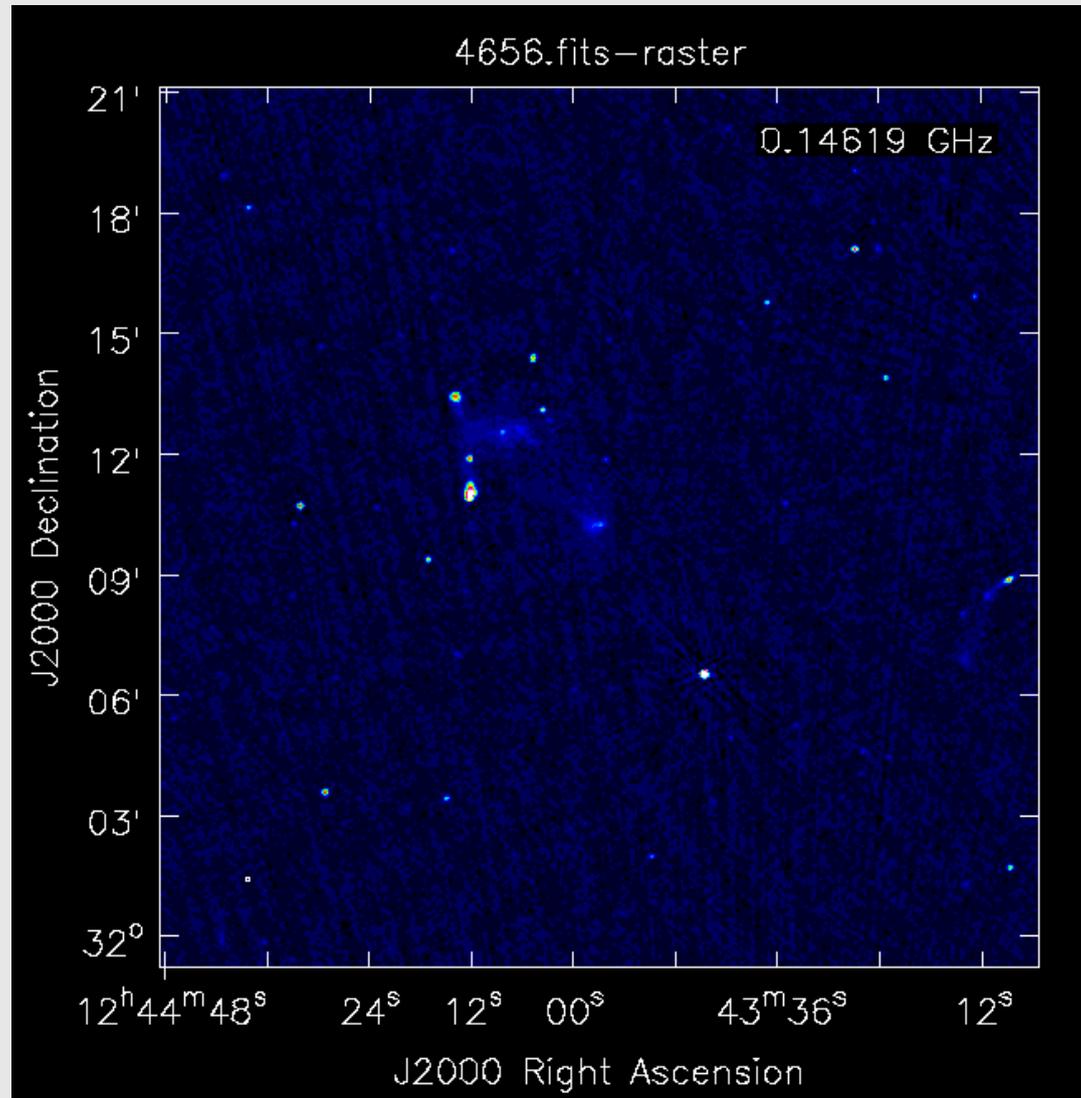


# Analysis



Simpson et al., 2016

# NGC 4656



# Summary

- Data reduction:
  - The current pipelines are great, manual labor is obsolete
  - Some fiddling still required
  - Factor is worth the effort
- Science:
  - Large outflows, consistent with other halo obs.
  - Filaments on north side up to 10 kpc
  - CR influence on halo clearly visible