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Light vector bosons and the weak mixing angle in the light of new reactor-based CE ν NS experiments

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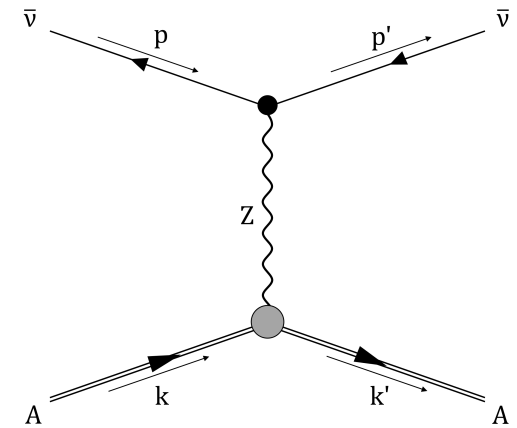
Light Dark World 2023 Karlsruhe
19 to 21 September 2023



Coherent elastic neutrino-nucleus scattering (CEvNS)

CEvNS chronology:

- Daniel Freedman (1974): weak (SM) NC, flavor-blind, threshold-free!
- First detection with π -DAR ν 's: COHERENT CsI (2017) & LAr (2021)
- Reactor experiments (2019 - ...):
CONNIE (Si), CONUS (Ge), NCC-1701 (Ge), ν GEN (Ge)
- Additional running and future experiments:
CCM (Ar), Miver (Ge/Si), NEON (NaI[Tl]), ν -cleus (CaWO_4 , Al_2O_3),
RED100 (Xe), Ricochet (Ge/Zn), Texono (Ge), CONUS+ (Ge)



The Channel:

- Coherence = enhancement $\sim N^2$

Upper limit on neutrino energy:

$$E_\nu \leq \frac{1}{2R_A} \approx \frac{197}{2.5\sqrt[3]{A}} [\text{MeV}]$$

- Observable = nuclear recoil energy T_A

$$\frac{d\sigma}{dT_A} = \frac{G_F^2 m_A}{4\pi} \left[(1 - 4\sin^2\theta_W)Z - N \right]^2 \left(1 - \frac{m_A T_A}{2E_\nu^2} \right) F^2(T_A)$$

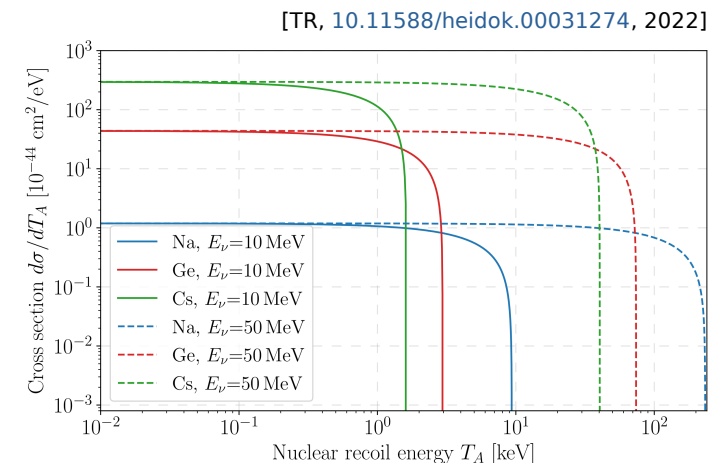
$\rightarrow$ detector material $\rightarrow$ neutrino source
 $\rightarrow$ energy threshold / neutrino source

- Very low energy threshold needed:

- $T_A \sim N^{-1}$

- Quenching: $T_A \rightarrow$ “detectable” energy

Cross section σ
vs.
nuclear recoil T_A



Brand-new results of the CONUS experiment!

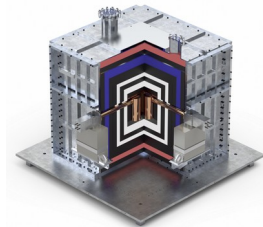
Reactor: Brokdorf nuclear power plant

- @17 m to 3.9 GW_{th}: inside reactor building!
- Neutrino flux: $2.3 \cdot 10^{13}/\text{cm}^2/\text{s}$
- Reactor information (fuel composition, thermal power) + **data-driven reactor spectra**

Background & shield

[Bonet et al., [10.1140/epjc/s10052-023-11240-4](https://arxiv.org/abs/10.1140/epjc/s10052-023-11240-4), 2021]

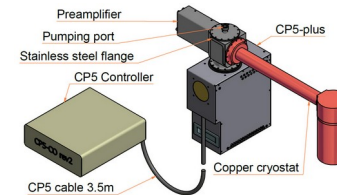
- Overburden: 10-45 m w.e.
- Active and passive components: bkg reduction $\times 10^3$ - 10^4
- Critical reactor-correlated bkg under control!
- Background level $O(10)/\text{keV}/\text{d}/\text{kg}$



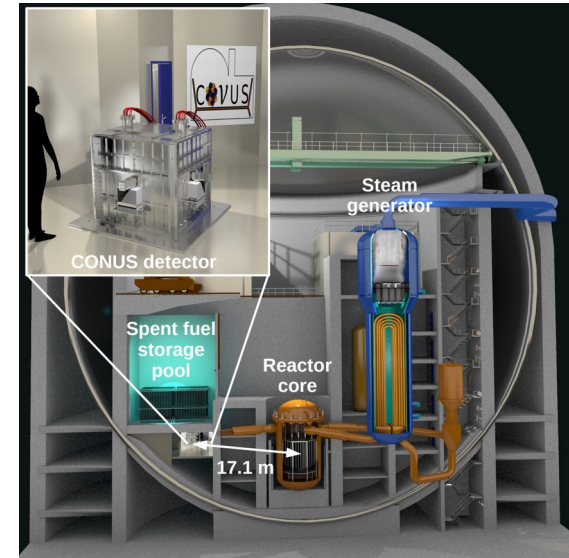
Detectors:

[Bonet et al., [10.1140/epjc/s10052-021-09038-3](https://arxiv.org/abs/10.1140/epjc/s10052-021-09038-3), 2021]

- HPGe PPC: 4*1kg
- $E^{\text{thr}} \sim 300\text{eV} \rightarrow \mathbf{210\text{ eV}}$
- Low Bkg design
- Electric cryocooling
- **New DAQ: Pulse shape discrimination**



[Bonet et al., [10.1007/JHEP05\(2022\)085](https://arxiv.org/abs/10.1007/JHEP05(2022)085), 2022]



New (and last) RUN-5 data sets:

Brand-new results of the CONUS experiment!

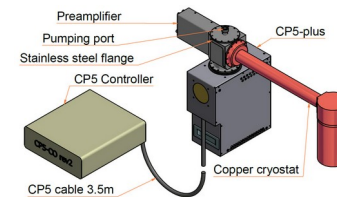
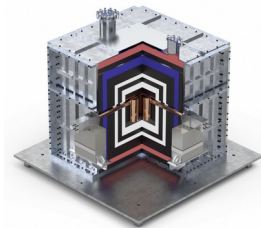
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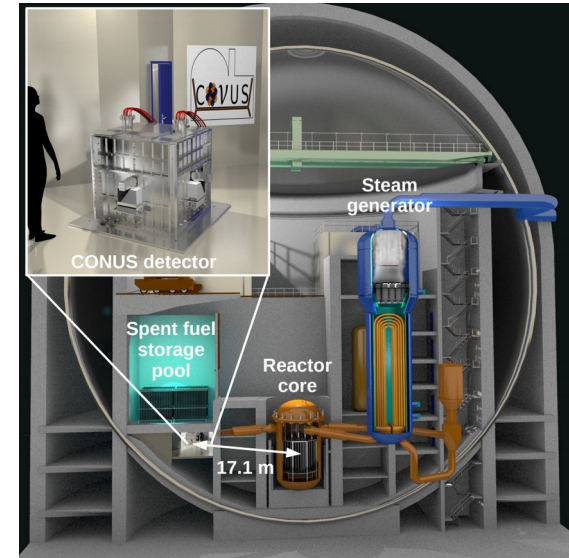


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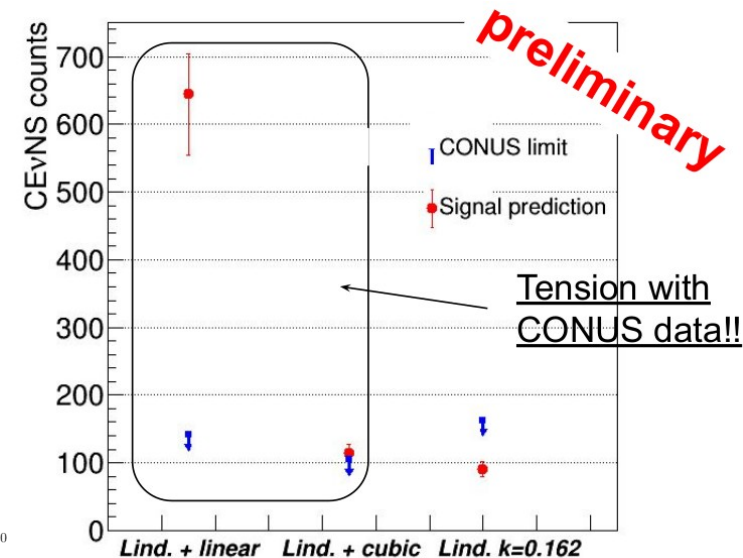
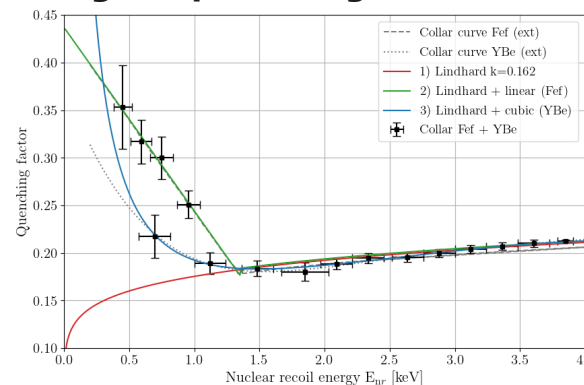


New (and last) RUN-5 data sets: ON~458 kg·d, OFF~293 kg·d

- CEvNS events < 163 events (90% C.L.)
- Factor <2 above SM prediction
- **More than one order of magnitude improvement** compared to previous analysis (2020)!
- Strongest limit on CEvNS from reactors!

Upcoming publication!

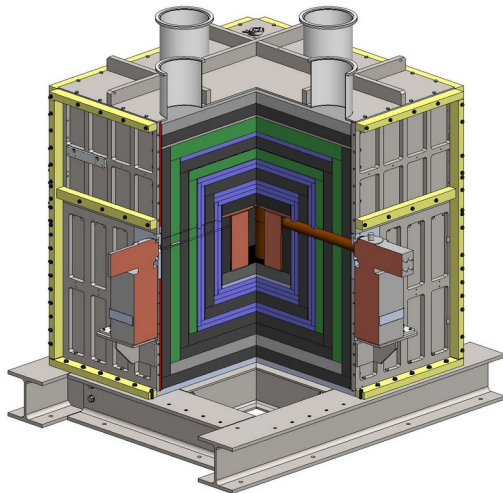
Signal quenching: $T \rightarrow E_{\text{ion}}$



New experiment at new site: CONUS+

Aim: First CEvNS detection @ reactor-site!

- Installation in Leibstadt (Switzerland) this summer
→ @ **20 m** distance to the 3.5 GW_{th} reactor core
- Refurbishment of existing COvUS Ge detectors
→ energy threshold < **200 eV**!
- Increased muon rejection efficiency via additional plastic scintillator layer
- **Full background characterization** at experimental site already performed
→ array of Bonner spheres + liquid scintillator cell



Future: Results from / upgrades of NCC-1701 (Ge), vGEN (Ge) and Texono (Ge)?

The weak mixing angle at low energy

- SM:** $SU(2)_L \times U(1)_Y$ + Higgs + renormalizability:

$$\sin^2 \theta_W^0 = e_0^2 / g_0^2 = 1 - m_W^2 / m_Z^2$$

→ respected by renorm. parameters (finite corrections)

- test SM @ quantum level / probe BSM physics

→ precision of $O(\pm 0.1\%)$ desired!

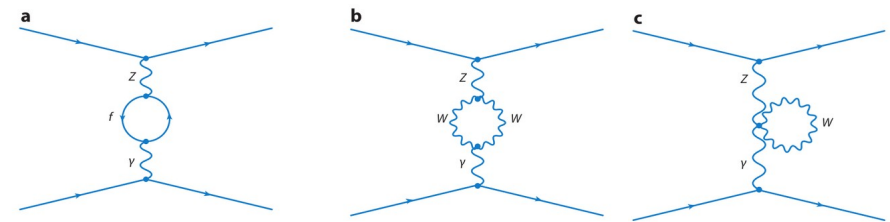
- At low energy γ -Z interference introduces small degree of parity violation

→ current experimental precision $O(\pm 1\%)$

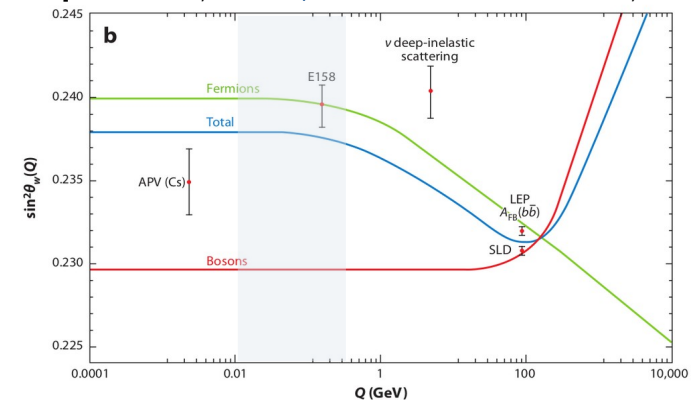
- Best precision achieved at Z pole: $O(\pm 0.1\%)$

- BUT** low- Q^2 measurements probe kinematic regions where Z pole measurements are insensitive to NP

→ **weakly-coupled light vector boson:**
dark parity violation / dark Z'



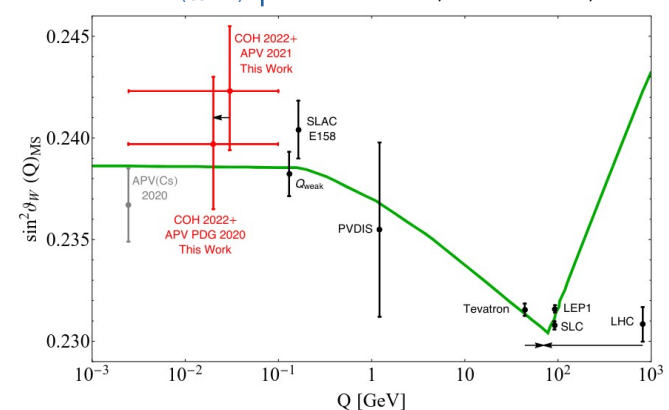
[Kumar et al., [10.1146/annurev-nucl-102212-170556](https://arxiv.org/abs/10.1146/annurev-nucl-102212-170556), 2013]



$$\Delta \sin^2 \theta_W(Q^2) \simeq -0.42 \varepsilon \delta \frac{m_Z}{m_{Z_d}} \left(\frac{m_{Z_d}^2}{Q^2 + m_{Z_d}^2} \right)$$

COHERENT
(Cs+Ar)

[Cadeddu et al., [10.1103/PhysRevD.102.015030](https://arxiv.org/abs/10.1103/PhysRevD.102.015030), 2020; Corona et al., [2303.09360](https://arxiv.org/abs/2303.09360), 2023]

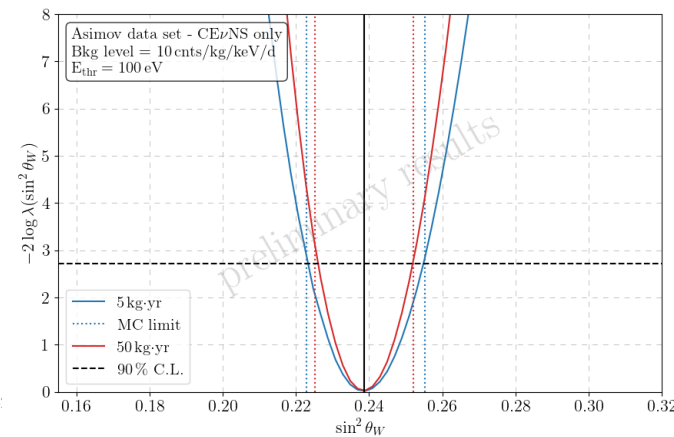
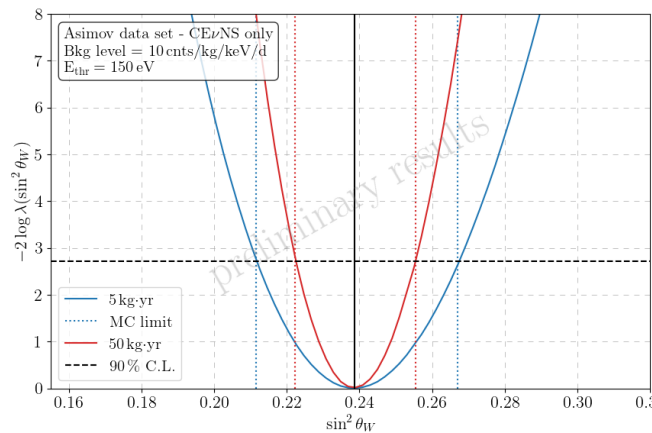
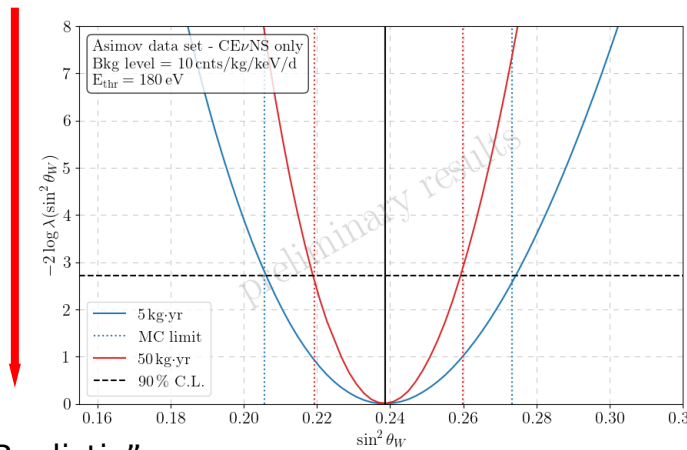


Future Weinberg angle sensitivity

$\Delta\chi^2$ contours for different mass and threshold configurations:

- EveS alone is generally not competitive
- Combined analysis dominated by contributions below 1keV $\rightarrow$ CE ν NS only
- Sensitivity limited by systematics: $\sim 5\%$ relative uncertainty

“Realistic” improvement: $\sim 3\text{-}8\%$ relative



“Realistic”
improvement:
 $\sim 1\text{-}6\%$ relative

Exposure	Rel. uncertainty
5 kg*yr	(-14; +15) %
50 kg*yr	(-8; +9) %

Exposure	Rel. uncertainty
5 kg*yr	(-11; +12) %
50 kg*yr	± 7 %

Exposure	Rel. uncertainty
5 kg*yr	(-6; +7) %
50 kg*yr	(-5; +6) %

Future light vector sensitivity

Simplified mediator model → contribution to CEνNS / EveS

$$\mathcal{L}_{Z'} = Z'_\mu \left(g_{Z'}^{\nu V} \bar{\nu}_L \gamma^\mu \nu_L + g_{Z'}^{e V} \bar{e} \gamma^\mu e + g_{Z'}^{q V} \bar{q} \gamma^\mu q \right) + \frac{1}{2} m_{Z'}^2 Z'_\mu Z'^\mu$$

- Assume universal coupling to quarks / electron / neutrinos: $(m_{Z'}, g_{Z'})$
- EveS more sensitive to lower mediator masses: m_e vs. m_{Ge}
- CONUS benchmark points (keV, GeV): EveS - $(5.2 \cdot 10^{-7}, 1.1 \cdot 10^{-2})$; CEνNS - $(2.7 \cdot 10^{-5}, 1.8 \cdot 10^{-3})$

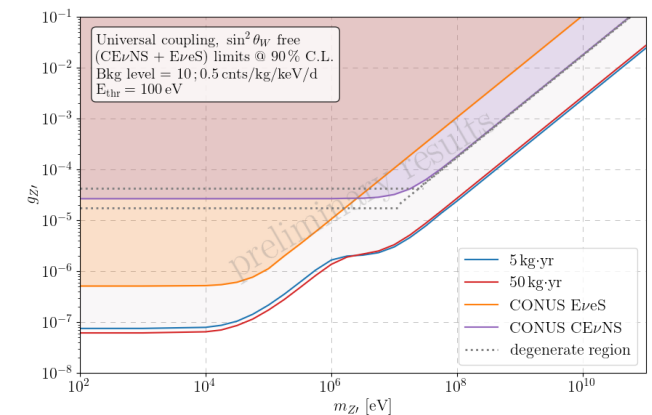
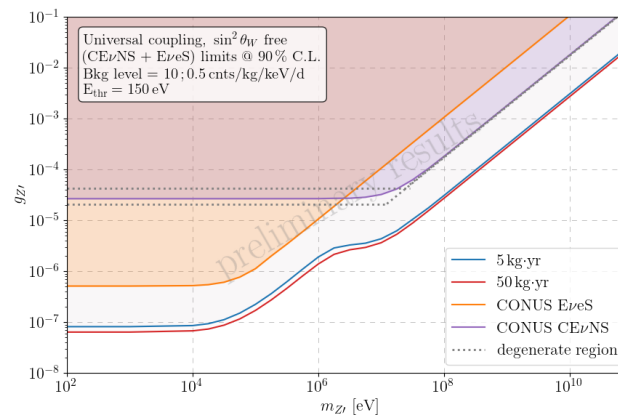
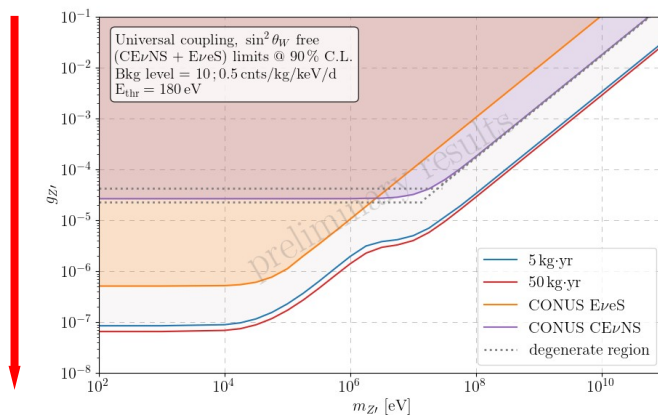
[Cerdeño et al., [10.1007/JHEP05\(2016\)118](https://arxiv.org/abs/10.1007/JHEP05(2016)118), 2016]

$$\frac{d\sigma}{dT} \propto \left[Q_{SM} + \frac{3g_{Z'}^2}{\sqrt{2}G_F} \frac{Z+N}{|q|^2 + m_{Z'}^2} \right]^2$$

[Cadeddu et al., [10.1007/JHEP01\(2021\)116](https://arxiv.org/abs/10.1007/JHEP01(2021)116), 2021]

[Bonet et al., [10.1007/JHEP05\(2022\)085](https://arxiv.org/abs/10.1007/JHEP05(2022)085), 2022]

“Realistic” improvements: CEνNS~7-29%, EveS~3-12%



“Real.” improv.:
EveS~18-23%
CEνNS~15%

Exposure	Limit @ keV	Limit @ GeV
5 kg*yr	$8.6 \cdot 10^{-8}$	$3.4 \cdot 10^{-4}$
50 kg*yr	$6.6 \cdot 10^{-8}$	$2.9 \cdot 10^{-4}$

Exposure	Limit @ keV	Limit @ GeV
5 kg*yr	$8.2 \cdot 10^{-8}$	$3.1 \cdot 10^{-4}$
50 kg*yr	$6.4 \cdot 10^{-8}$	$2.7 \cdot 10^{-4}$

Exposure	Limit @ keV	Limit @ GeV
5 kg*yr	$7.6 \cdot 10^{-8}$	$2.4 \cdot 10^{-4}$
50 kg*yr	$6.2 \cdot 10^{-8}$	$2.7 \cdot 10^{-4}$

Reactor neutrinos for low masses ↔ π-DAR neutrinos for higher masses

Future light vector sensitivity

THE benchmark (mediator) model $\rightarrow$ $U(1)_{B-L}$

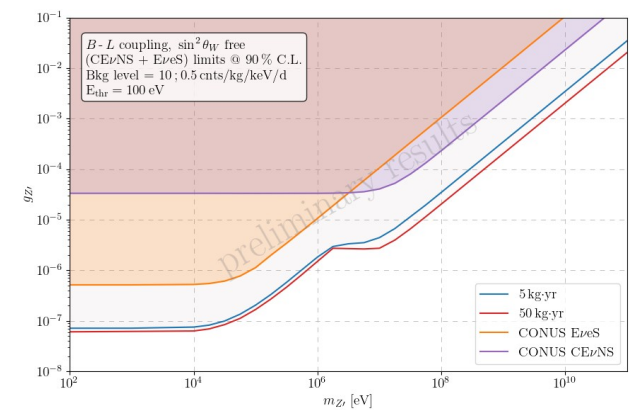
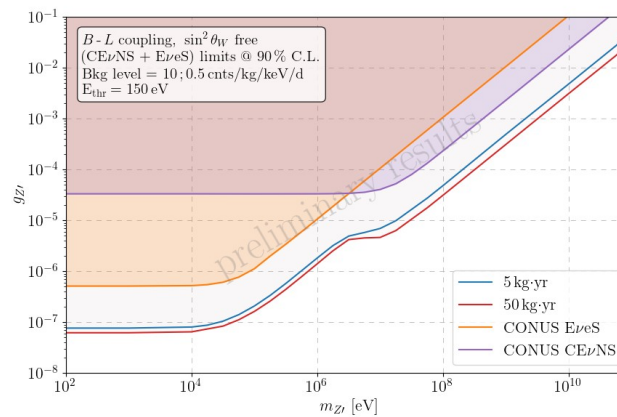
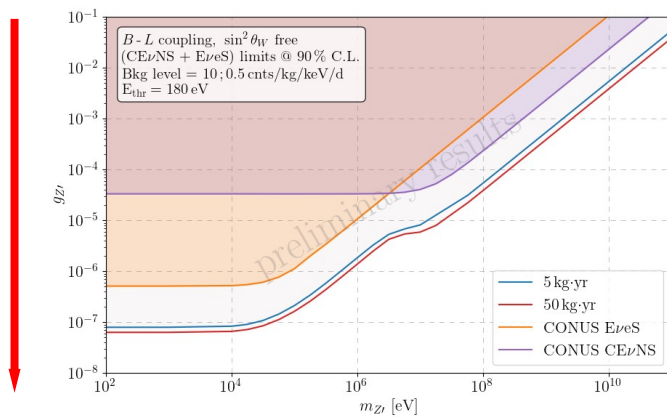
$$\mathcal{L}_{Z'} = Z'_\mu \left(g_{Z'}^{\nu V} \bar{\nu}_L \gamma^\mu \nu_L + g_{Z'}^{e V} \bar{e} \gamma^\mu e + g_{Z'}^{q V} \bar{q} \gamma^\mu q \right) + \frac{1}{2} m_{Z'}^2 Z'_\mu Z'^\mu$$

$$\frac{d\sigma}{dT} \propto \left[Q_{SM} - \frac{g_{Z'}^2}{\sqrt{2} G_F} \frac{Z + N}{|q|^2 + m_{Z'}^2} \right]^2$$

[Cadeddu et al., [10.1007/JHEP01\(2021\)116](https://arxiv.org/abs/1010077), 2021]

- Assume gauged B-L charge: $(m_{Z'}, g_{Z'})$
- EveS more sensitive to lower mediator masses: m_e vs. m_{Ge}
- CONUS benchmark points (keV, GeV): EveS - $(5.2 \cdot 10^{-7}, 1.1 \cdot 10^{-2})$; CEvNS - $(3.3 \cdot 10^{-5}, 2.3 \cdot 10^{-3})$

“Realistic” improvements: CEvNS~12-35%, EveS~2-10%



“Real.” improv.:
EveS~15-21%
CEvNS~27-43%

Exposure	Limit @ keV	Limit @ GeV
5 kg*yr	$8.0 \cdot 10^{-8}$	$5.4 \cdot 10^{-4}$
50 kg*yr	$6.3 \cdot 10^{-8}$	$3.9 \cdot 10^{-4}$

Exposure	Limit @ keV	Limit @ GeV
5 kg*yr	$7.7 \cdot 10^{-8}$	$4.8 \cdot 10^{-4}$
50 kg*yr	$6.2 \cdot 10^{-8}$	$3.1 \cdot 10^{-4}$

Exposure	Limit @ keV	Limit @ GeV
5 kg*yr	$7.2 \cdot 10^{-8}$	$3.5 \cdot 10^{-4}$
50 kg*yr	$6.1 \cdot 10^{-8}$	$2.0 \cdot 10^{-4}$

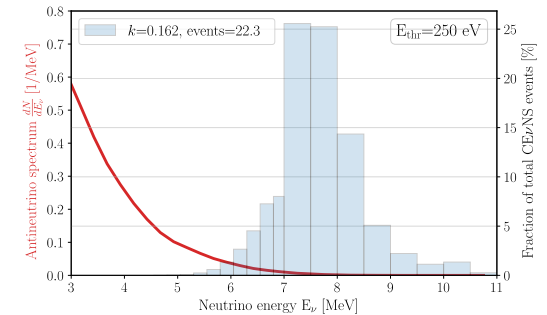
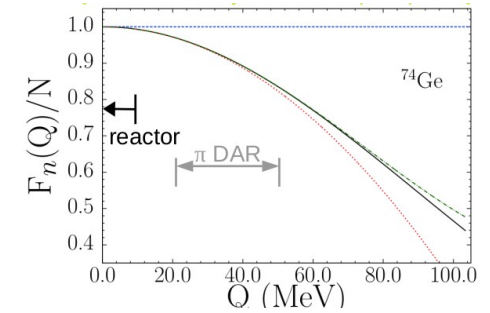
Reactor neutrinos for low masses $\leftrightarrow$ π -DAR neutrinos for higher masses

Opportunities for CEvNS experiments

SM investigations with enhanced precision:

- **Nuclear form factors:**
Model-independent extraction of neutron density distributions
→ Beam-reactor complementarity
- **Measuring reactor antineutrino spectrum:**
CEvNS sensitive to all flavors of high-E part where uncertainty is largest

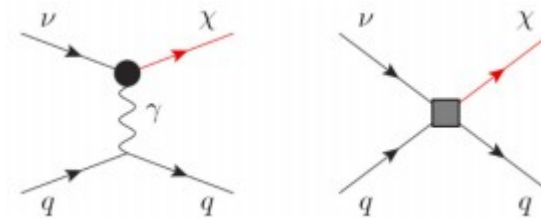
[Patton et al., [10.1103/PhysRevC.86.024612](https://arxiv.org/abs/10.1103/PhysRevC.86.024612), 2012]



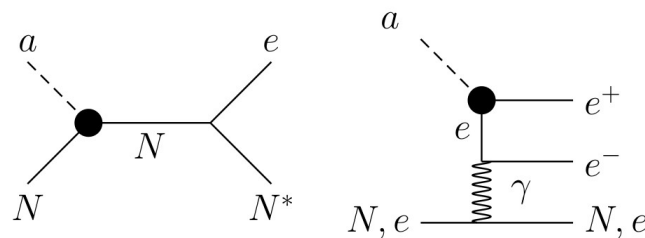
BSM investigations:

- **Light sterile neutrinos:**
Use CEvNS for ν flux measurements
- **New fermion searches:**
Test further ν interactions
→ ν mass, DM, ...
- **Probing portals:**
ALPs, dark photons, etc.

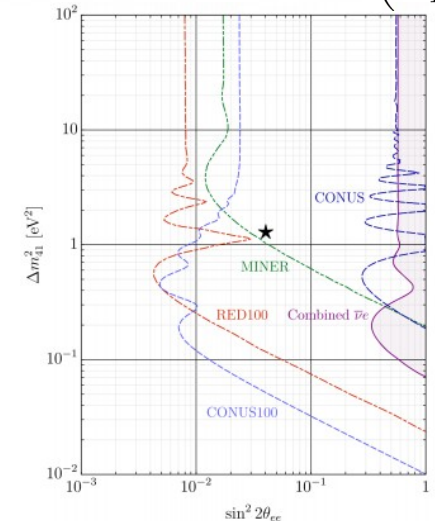
[Chang & Liao, [10.1103/PhysRevD.102.075004](https://arxiv.org/abs/10.1103/PhysRevD.102.075004), 2022]



[Aristizabal Sierra et al., [10.1007/JHEP03\(2021\)294](https://arxiv.org/abs/10.1007/JHEP03(2021)294), 2021]



$$P_{\nu_\alpha \rightarrow \nu_\alpha} = 1 - \sin^2 2\theta_{\alpha\alpha} \sin^2 \left(\frac{\Delta m_{41}^2 L}{4E_\nu} \right)$$



[Berryman, [10.1103/PhysRevD.100.023540](https://arxiv.org/abs/10.1103/PhysRevD.100.023540), 2019]

Conclusion

- **CEvNS opens new path to high-statistics neutrino physics**

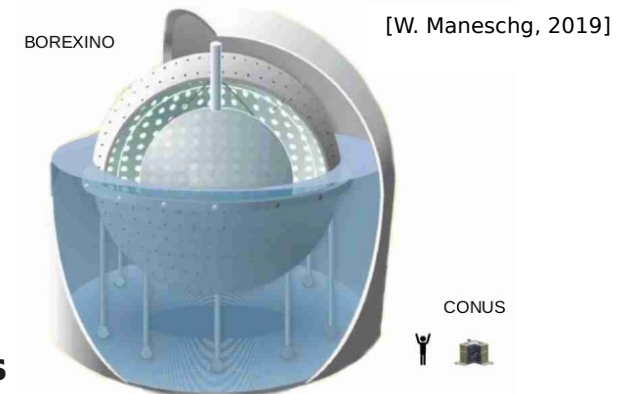
- beams and reactors go hand in hand!
- full spectrum of modern detection technologies
- “car-size” neutrino detectors!

- **Sensitivity for the next-gen. of Ge-based reactor experiments**

- Weinberg angle:
 - relative uncertainty $\lesssim 10\%$
 - systematics become crucial!
- Light vector bosons (universal coupling & B-L):
 - probing the next order of magnitude
 - improved threshold critical for CEvNS-dominated limit regions

- **CONUS RUN-5 SM results to be published soon**

→ **BSM analyses are coming!**



**Paper to
appear soon!**

Next-generation experiments (π -DAR, reactors & DM DD!) = huge playground for phenomenology

(Neutrino magnetic moment, neutrino charge radius, vector NSI, tensor NSI, light scalars, light vectors, sterile neutrinos, dark matter, weak mixing angle, nuclear form factors, neutrino-electron-scattering, supernova neutrinos, antineutrino reactor spectra, neutrino floor, etc.)

Thank You!

Backup

Assume a CONUS-like experiment

Reactor: commercial nuclear power plant

- Experimental site ~ 20 m to $3.5 \text{ GW}_{\text{th}}$
- ν flux: $\sim 1.5 \cdot 10^{13} / \text{cm}^2 / \text{s}$ (3% uncertainty)
- Typical PWR fuel composition:
56.1% U235, 7.6% U238, 30.7% Pu239, 5.6% Pu241

Background

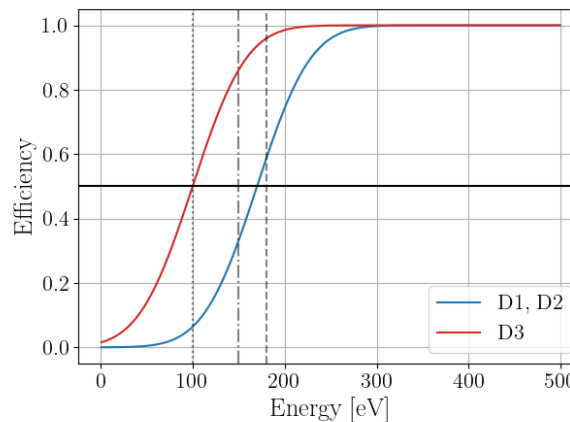
- No critical reactor-correlated bkg!
- Background levels:
 - 10 cnts/keV/d/kg for $E_{\text{ion}} \leq 1 \text{ keV}$
 - 0.5 cnts/keV/d/kg for $E_{\text{ion}} > 1 \text{ keV}$

Detectors:

- Data collection: $t_{\text{OFF}} = 0.5 \cdot t_{\text{ON}}$:

R1 - $\text{exp}_{\text{ON}} = 5 \text{ kg} \cdot \text{yr}$

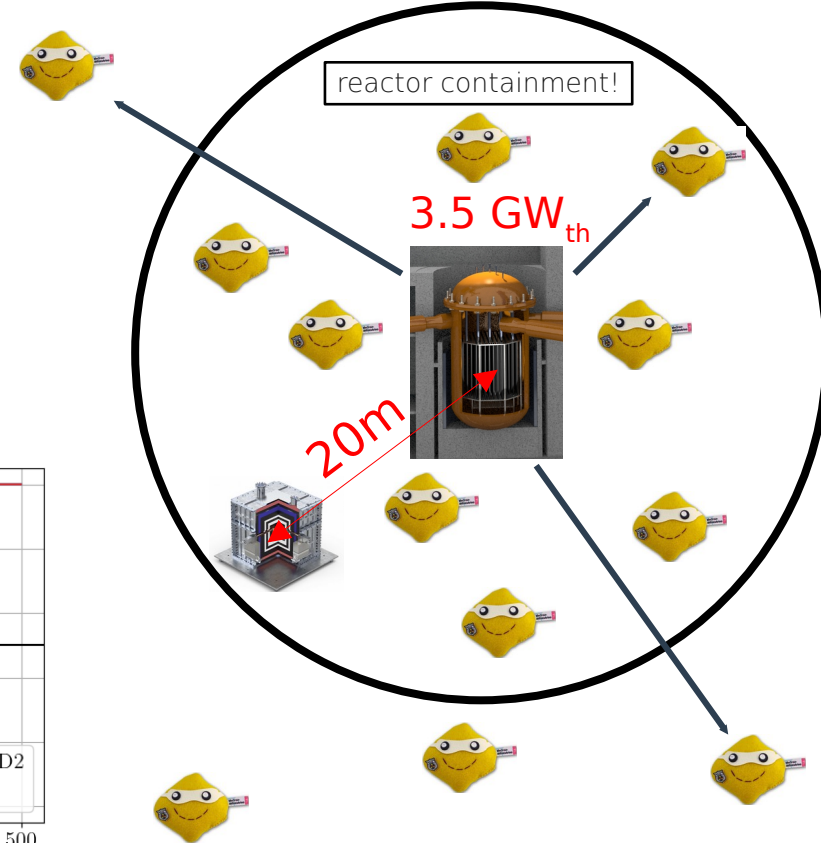
R2 - $\text{exp}_{\text{ON}} = 50 \text{ kg} \cdot \text{yr}$



[Bonhomme et al., [10.1140/epjc/s10052-022-10768-1](https://arxiv.org/abs/10.1140/epjc/s10052-022-10768-1), 2022]

- CONUS HPGe detectors with improved trigger efficiency:
 $E_{\text{thr}} \sim 3 \cdot \text{FWHM}_{\text{pulser}}$
- Quenching according to Lindhard model (1% uncertainty)

[Bonet et al., [10.1007/JHEP05\(2022\)085](https://arxiv.org/abs/10.1007/JHEP05(2022)085), 2022]



Detector	Threshold E_{thr}
D1 (conservative)	180 eV
D2 (expected)	150 eV
D3 (optimistic)	100 eV

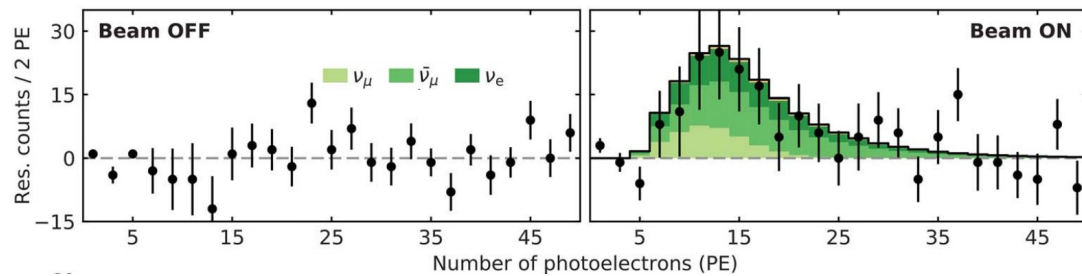
[particlezoo.net, 2023]

CEvNS observations and constraints

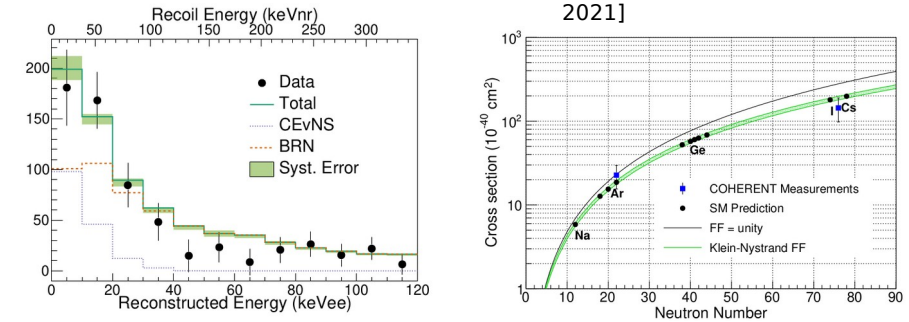
Observations:

COHERENT CsI[Na] (2017): $5 * 10^{20}$ POT
obs. 134 ± 22 , *pred.* 173 ± 48

[Akimov et al., [10.1126/science.aao0990](https://doi.org/10.1126/science.aao0990), 2017]

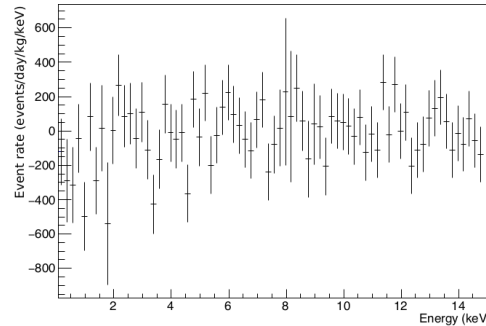


COHERENT LAr (2021): $13.7 * 10^{22}$ POT
obs. 159 ± 43 , *pred.* 128 ± 17
 [Akimov et al., [10.1103/PhysRevLett.126.012002](https://doi.org/10.1103/PhysRevLett.126.012002), 2021]

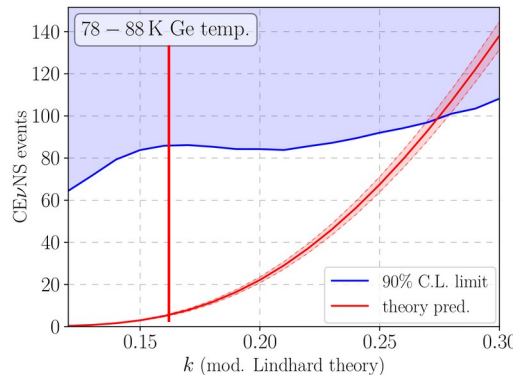


Limits:

CONNIE Si (2019):
 $R_{NP} < 40 * R_{SM}$ @ 95% C.L.
 2.1kg*d ON + 1.6kg*d OFF
 [Aguilar-Arevalo et al., [10.1103/PhysRevD.100.092005](https://doi.org/10.1103/PhysRevD.100.092005), 2019]



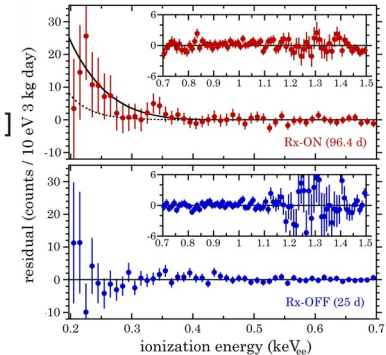
CONUS Ge (2020):
 $R_{SM} < 0.4/\text{kg/d}$ @ 90% C.L.
 (x17 lower than prediction)
 249kg*d ON + 59kg*d OFF
 [Bonet et al., [10.1103/PhysRevLett.126.041804](https://doi.org/10.1103/PhysRevLett.126.041804), 2020]



“Measurement [...] Reactor Antineutrinos”

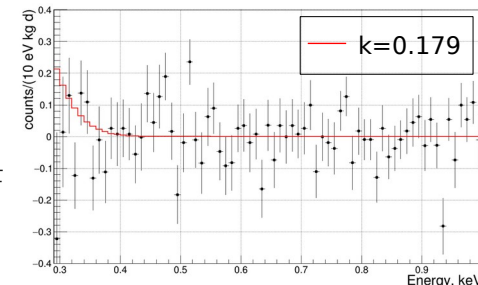
[Colaesi et al., [10.1103/PhysRevLett.129.211802](https://doi.org/10.1103/PhysRevLett.129.211802), 2022]

NCC-1701 Ge:
 289kg*d ON + 75kg*d OFF



ν GEN Ge (2022):
no excess,
 $k < 0.177$ @ 90% C.L.
 133kg*d ON + 66kg*d OFF

[Alekseev et al., [10.1103/PhysRevD.106.L051101](https://doi.org/10.1103/PhysRevD.106.L051101), 2022]

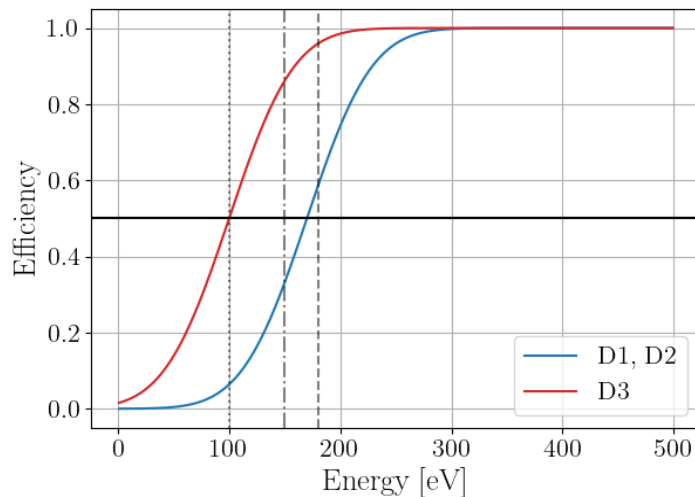


Outlook: Future Ge detectors at reactor sites

Progress for kg-size Ge detectors @ reactor sites

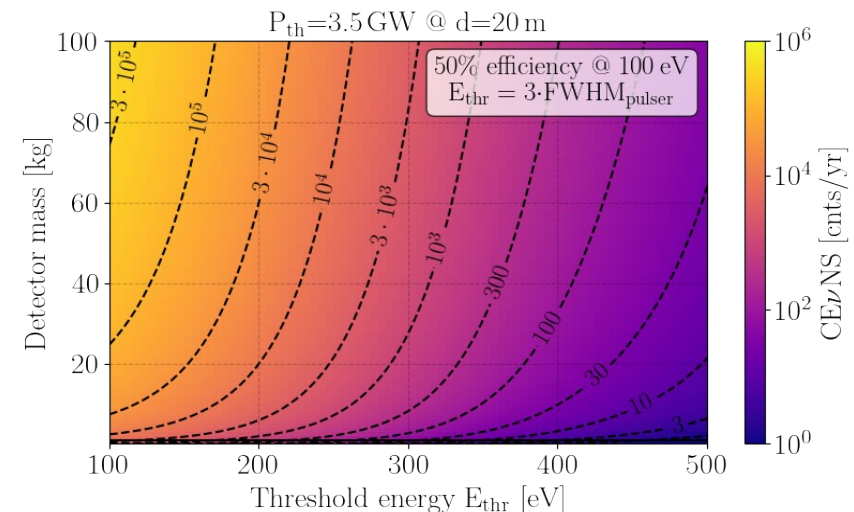
Experimental perspective (single diode):

- Improve pulser width $\text{FWHM}_{\text{pulser}}$
→ improves $E_{\text{thr}} \sim 3 \cdot \text{FWHM}_{\text{pulser}}$!
- Improve trigger efficiency
→ record more events in critical region!
- Background and environmental stability at experimental site
→ increases exposure



Theoretical perspective (whole set-up):

- Energy threshold and exposure are key for a strong CEvNS signal / BSM reach
- Trade-off between reactor flux and background
- Incorporate major uncertainties in SM and BSM studies!



HPPC Ge detectors offer a scalable technology also in critical environments

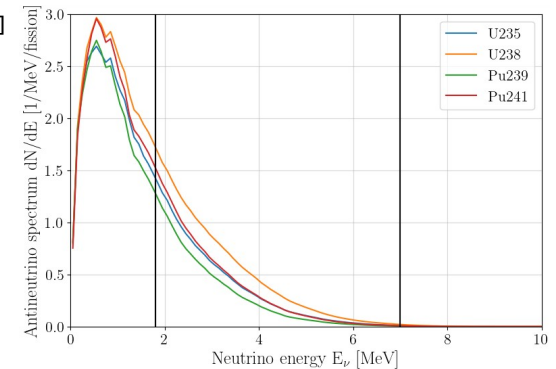
- energy threshold and trigger efficiency are crucial! exposure via multiple diodes?
- quenching (description) strongly affects CEvNS prediction

Reactor antineutrino spectrum: data-driven approach

[Estienne et al., [10.1103/PhysRevLett.123.022502](https://arxiv.org/abs/10.1103/PhysRevLett.123.022502), 2019]

1) Low energy region $E < 1.8$ MeV:

Summation spectra of Estienne et al. (2019)



2) Intermediate energy $1.8 \text{ MeV} \leq E < 7 \text{ MeV}$

Daya Bay (2021):

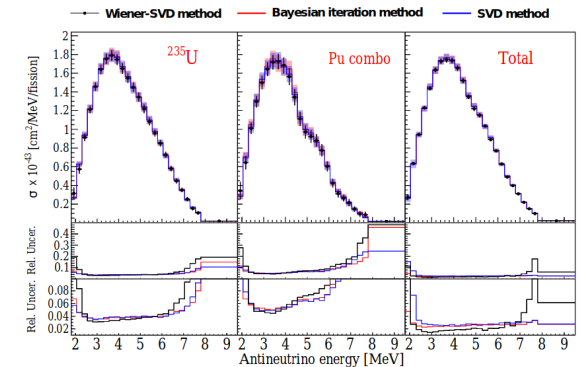
- unfolded IBD spectra of U235, (Pu239+ Pu241), total
- method to construct data-based predictions

$$S_{\text{pred}} = R \cdot \begin{pmatrix} S_{\text{total}} \\ S_{235} \\ S_{\text{combo}} \\ S_{238} \\ S_{241} \end{pmatrix} \quad R = \left(I_{25} \mid \Delta f_{235} I_{25} \mid \Delta f_{239} I_{25} \mid \Delta f_{238} I_{25} \mid (\Delta f_{241} - 0.183 \times \Delta f_{239}) I_{25} \right)$$

Difference in fission fractions

$\longrightarrow$ Mueller
 $\longrightarrow$ Huber

Extend with bin-to-bin
 uncorrelated uncertainties:
 10% Huber, 15% Mueller

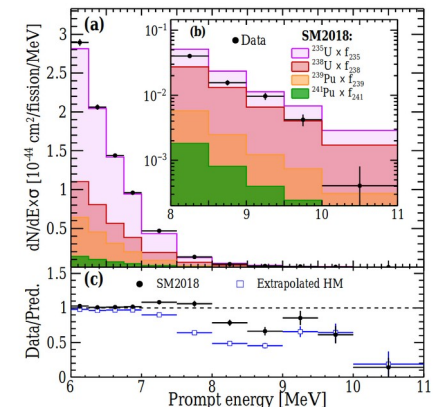


3) High energy $E \geq 7$ MeV:

Daya Bay measurement of combined high energy spectrum (2022)

→ Application of Daya Bay data-based method with Estienne et al. spectra

[An et al., [10.1088/1674-1137/abfc38](https://arxiv.org/abs/10.1088/1674-1137/abfc38), 2021]



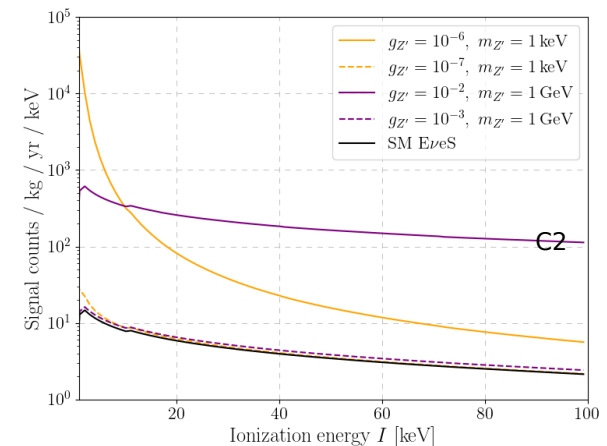
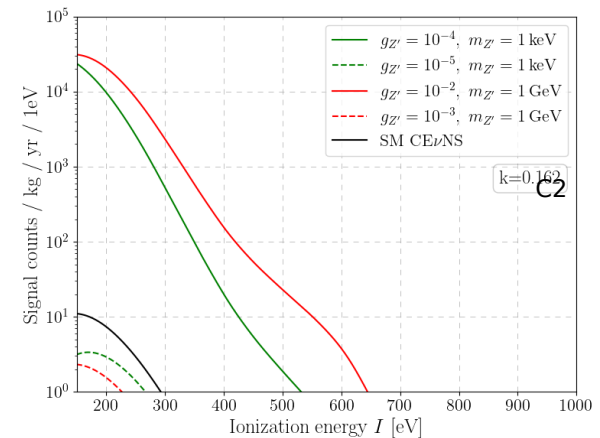
Spectrum also used in current CONUS analysis!

[An et al., [10.1103/PhysRevLett.129.041801](https://arxiv.org/abs/10.1103/PhysRevLett.129.041801), 2022]

Sensitivity determination

(Binned log-) Likelihood function:

- ON and OFF data are fitted together:
 $\log \mathcal{L} = \log \mathcal{L}_{\text{ON}} + \log \mathcal{L}_{\text{OFF}} + \text{pull term}$
 with $\log \mathcal{L}_{\text{ON}}(\sin^2 \theta_W, g_{Z'}, m_{Z'}; b_{\text{CE}\nu\text{NS}}, b_{\text{E}\nu\text{eS}}, \Phi_{\bar{\nu}}, k)$
 $\log \mathcal{L}_{\text{OFF}}(b_{\text{CE}\nu\text{NS}}, b_{\text{E}\nu\text{eS}})$
- Parameter list:
 - Weinberg angle $\sin^2 \theta_W$
 - Z' model parameters $g_{Z'}, m_{Z'}$
 - bkg normalizations $b_{\text{CE}\nu\text{NS}}, b_{\text{E}\nu\text{eS}}$
 - reactor neutrino flux ($\sim 3\%$) $\Phi_{\bar{\nu}}$
 - parameter of Lindhard model ($\sim 1\%$) k
- profile LH ratio for limits + χ^2 -distribution of test statistic
- Sensitivity estimates via two methods:
 - Asimov data set
 - MC sampling + median average of limits



Sensitivity to the Weinberg angle
 CEvNS - $E_{\text{ion}} < 1 \text{ keV}$

Sensitivity to a light vector boson:
 CEvNS - $E_{\text{ion}} < 1 \text{ keV}$
 EνeS - $E_{\text{ion}} < 100 \text{ keV}$

Experimental requirements at reactor site

Goal: Detecting CEvNS with high accuracy!

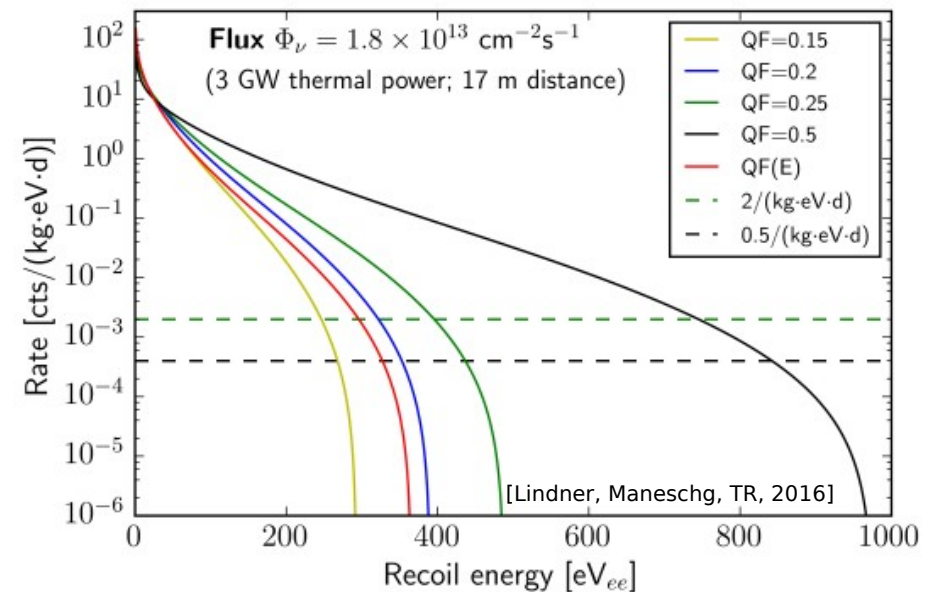
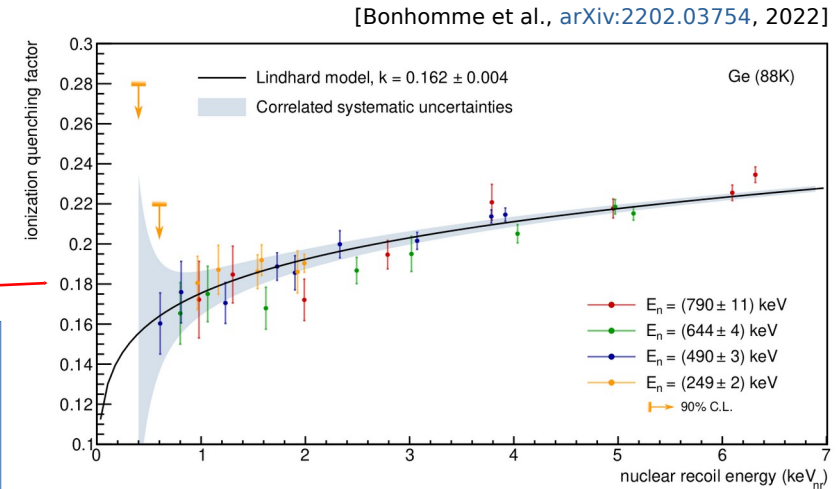
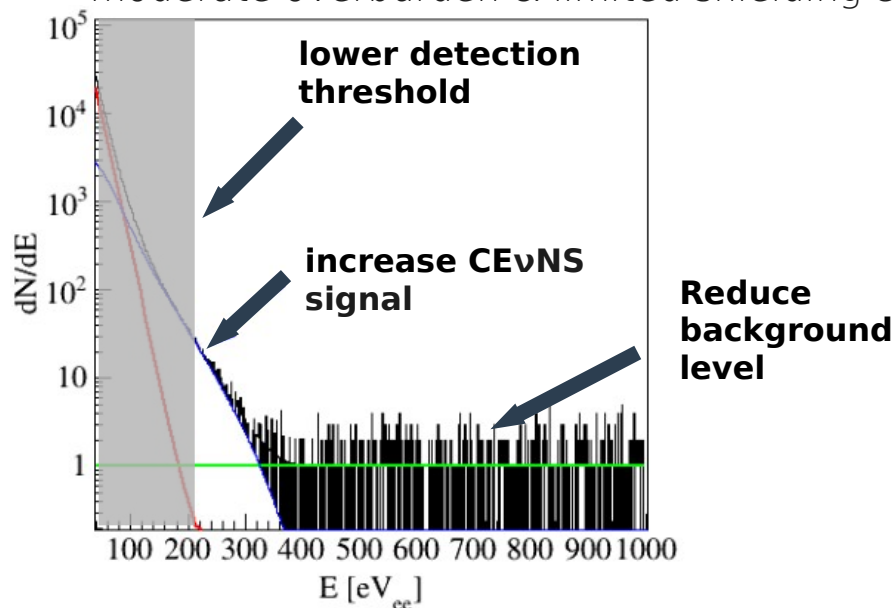
Several obstacles to overcome:

1) Beat $1/R^2$ factor
→ strong (= commercial) power plant,
close to reactor core

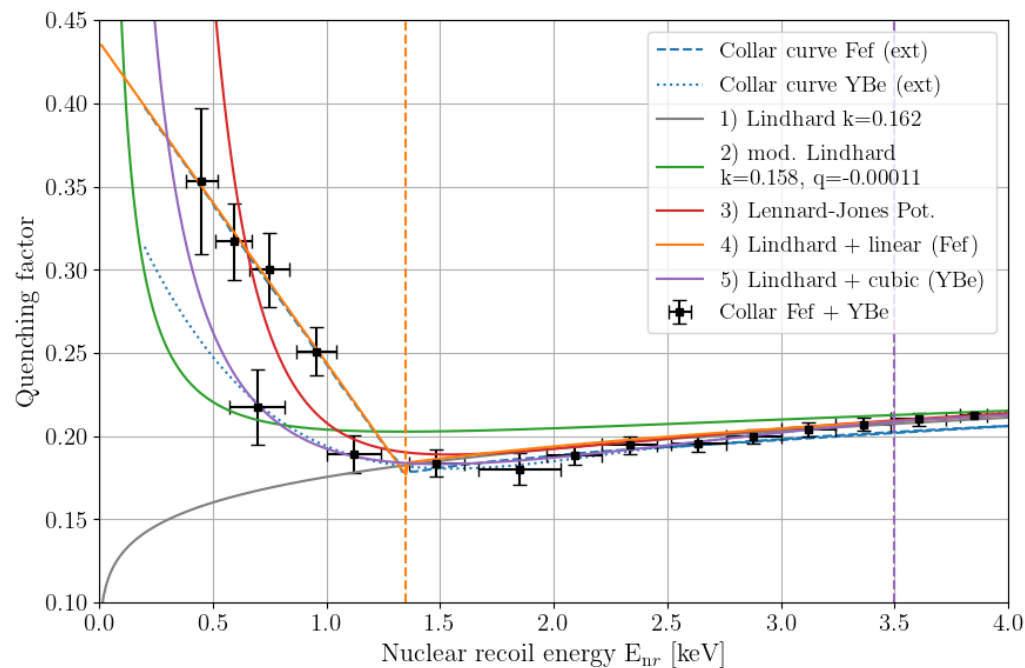
2) Compensate quenching ($E_{\text{recoil}} \rightarrow E_{\text{ion}}$)
→ lowest possible detection threshold

E_v : 10MeV
max E_{Recoil} : 3keV
max E_{ion} : ~600eV

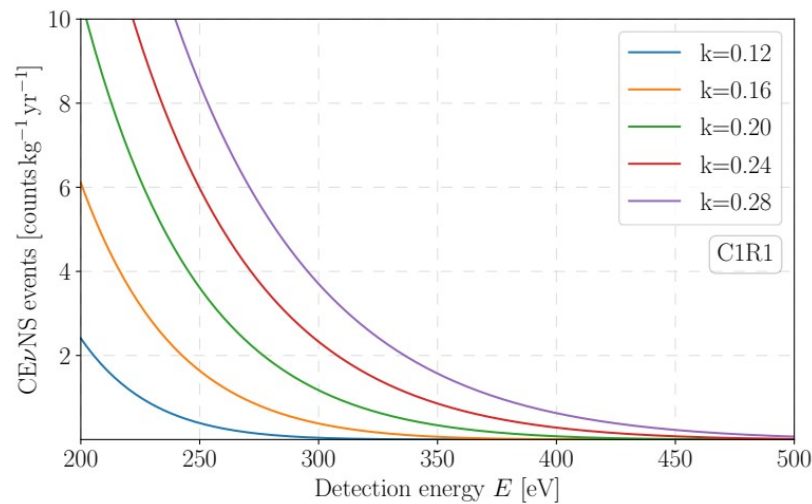
3) Low background outside lab condition
→ moderate overburden & limited shielding capacities



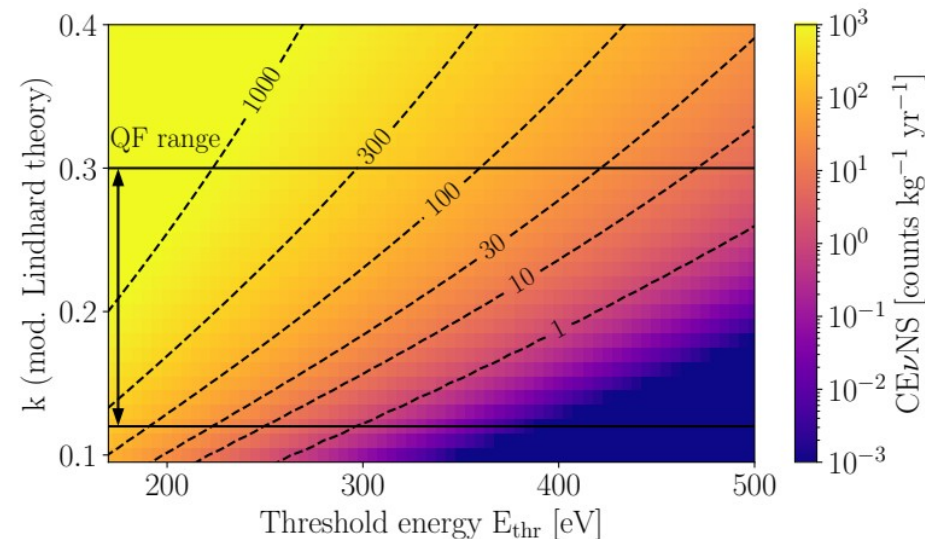
Impact of quenching at low energy



Quenching according to mod. Lindhard model:

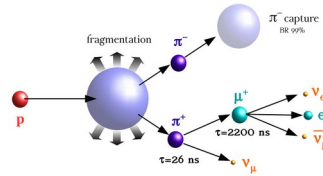


[TR, PhD thesis, 2022]



Two complementary approaches

π -decay-at-rest neutrinos:



- Pulsed GeV-proton beam hitting heavy target
→ multiple ν flavors
- Time correlation of events
→ background suppression $\times(10^3\text{-}10^4)$
- Higher ν energies
→ larger cross section, but reduced coherence

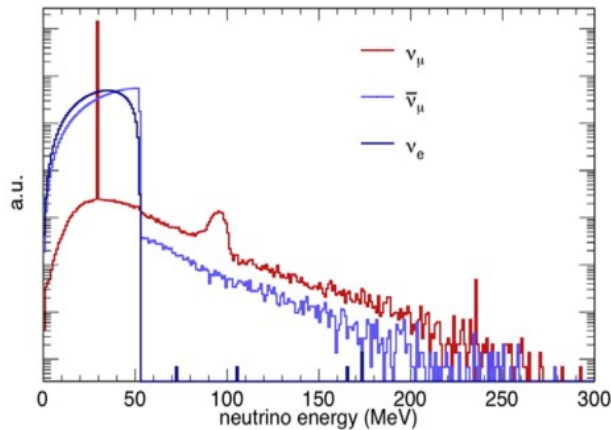
→ **COHERENT, CCM, “CE ν NS@ESS”...**

Reactor antineutrinos:

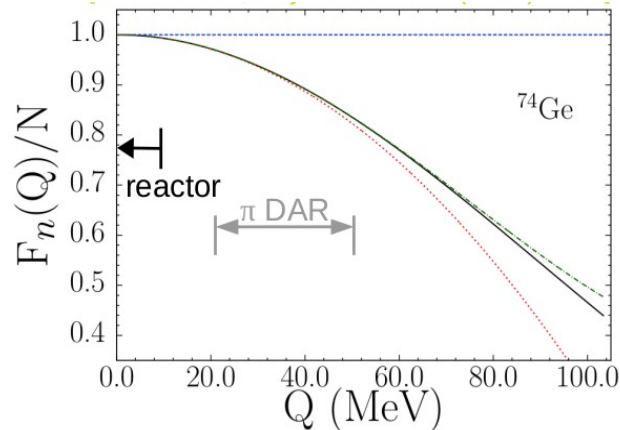
- β decays in nuclear reaction chains → only $\bar{\nu}_e$
- Strongest artificial ν source on earth:
 $\sim 10^{20}$ $\bar{\nu}_e$'s/GW/s
- ν energies up to 10 MeV → coherent regime!
- Close to reactor core: no lab conditions!
→ no cryogenic liquids, no remote control

→ **CONNIE, CONUS, NCC-1701, ν GEN ...**

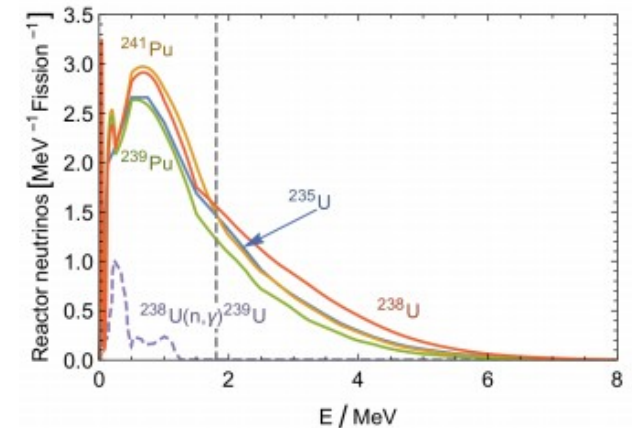
[Akimov et al., [arXiv:1509.08702](https://arxiv.org/abs/1509.08702), 2015]



[Patton et al., [10.1103/PhysRevC.86.024612](https://arxiv.org/abs/10.1103/PhysRevC.86.024612), 2012]



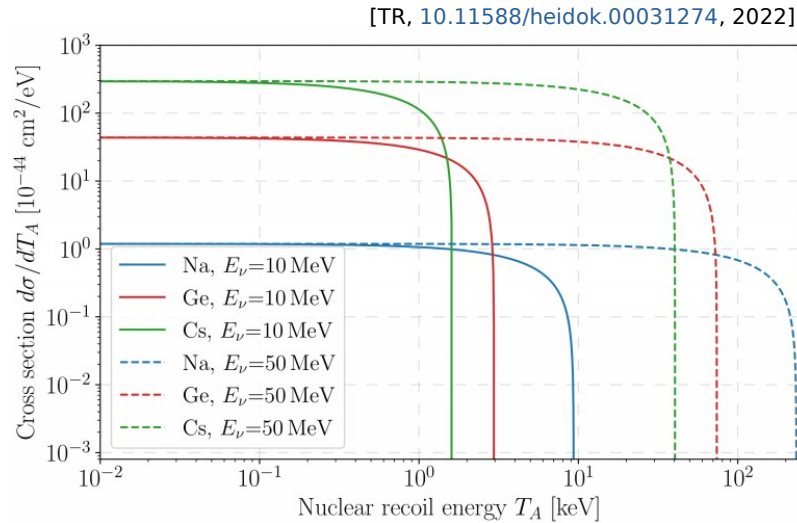
[Vitagliano et al., [10.1103/RevModPhys.92.045006](https://arxiv.org/abs/10.1103/RevModPhys.92.045006), 2020]



Beam-reactor complementarity:

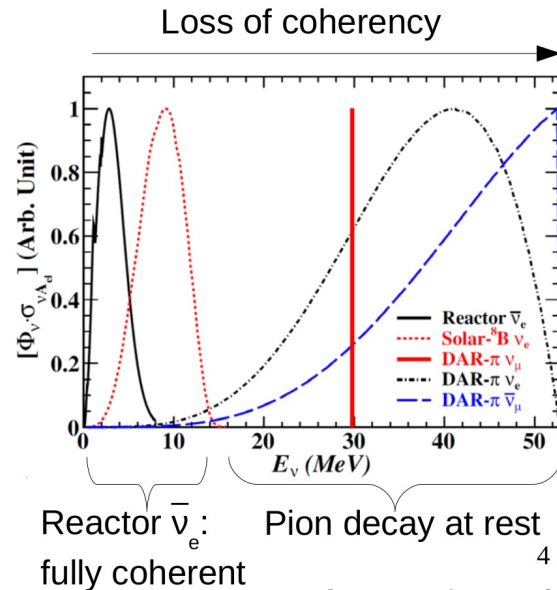
CE ν NS at reactor site as high statistic baseline for multi-target and multi-flavored beam investigations!

CEvNS with different sources and targets



Element	N	r_A [fm]	$E_\nu^{\max}$ [MeV]	$T_A^{\max}$ [keV]
Na	12	3.6	27.7	71.5
Si	14	3.8	25.9	51.3
Ar	22	4.4	23.1	28.5
Ge	38/40/42	5.2	18.9	10.5
I	74	6.3	15.7	4.16
Xe	75/77/78	6.4	15.5	3.93
Cs	78	6.4	15.4	3.85

[TR, 10.11588/heidok.00031274, 2022]



[W. Maneschg, 2017]

Neutrino source	Target	$T_A^{\max}$ [keV]	E (QF $\in \{0.1, 0.15, 0.2\}$) [keV]
Nuclear reactor (10 MeV)	Na	9.33	0.93 / 1.40 / 1.87
	Si	7.64	0.76 / 1.15 / 1.53
	Ar	5.37	0.54 / 0.81 / 1.07
	Ge	2.96	0.30 / 0.44 / 0.59
	I	1.69	0.17 / 0.25 / 0.34
	Xe	1.64	0.16 / 0.25 / 0.33
π -DAR source (50 MeV)	Cs	1.62	0.16 / 0.24 / 0.32
	Na	232.4	23.2 / 34.9 / 46.5
	Si	190.4	19.0 / 28.6 / 38.1
	Ar	134.0	13.4 / 20.1 / 26.8
	Ge	73.8	7.38 / 11.1 / 14.8
	I	42.3	4.23 / 6.34 / 8.45
	Xe	40.9	4.09 / 6.13 / 8.17
	Cs	40.4	4.04 / 6.03 / 8.07