

# Introducing *The Future of Particle Physics* (KIT Edition)

Chris Quigg

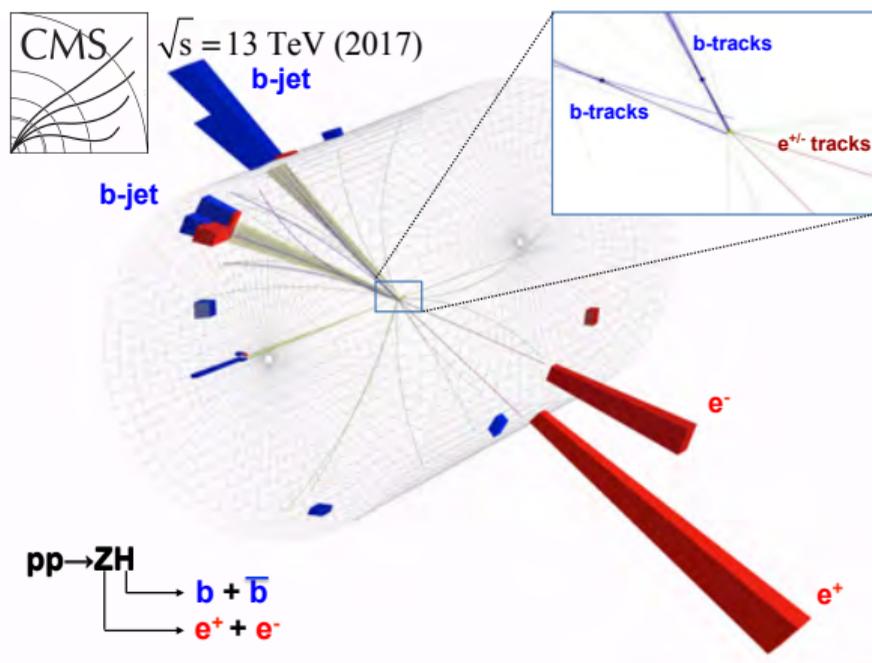
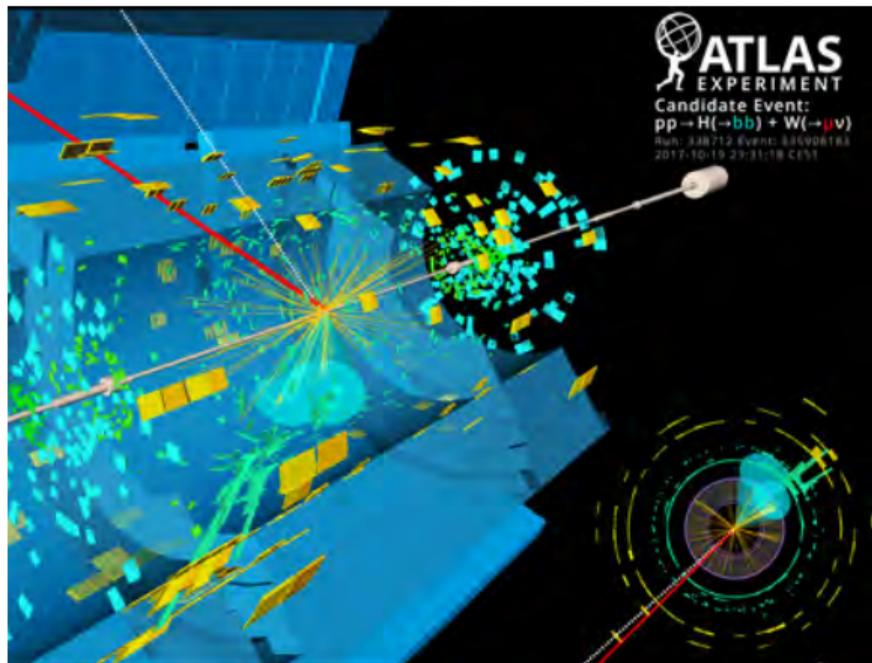
*Fermilab & CERN*



*The Future of Particle Physics: A Quest for Guiding Principles* · 10.2018

Supplemental reading: “Dream Machines,” arXiv:1808.06036

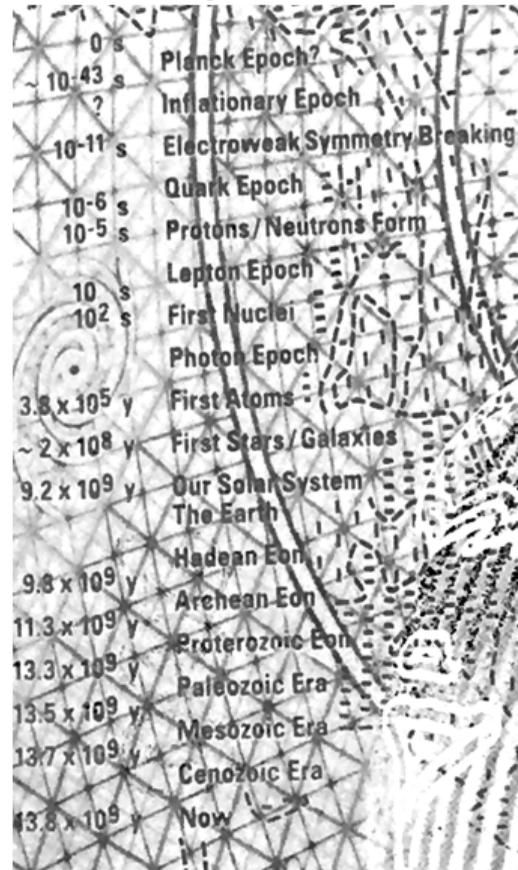
# We learn something every day: example of $H \rightarrow b\bar{b}$



# CHF200 Note (2018)



... many scales



# The importance of the 1-TeV scale

EW theory does not predict Higgs-boson mass

Thought experiment: conditional upper bound

$W^+W^-$ ,  $ZZ$ ,  $HH$ ,  $HZ$  satisfy  $s$ -wave unitarity,

provided  $M_H \lesssim (8\pi\sqrt{2}/3G_F)^{1/2} \approx 1 \text{ TeV}$

If bound is respected, perturbation theory is “everywhere” reliable

If not, weak interactions among  $W^\pm$ ,  $Z$ ,  $H$  become strong on 1-TeV scale

New phenomena (*H or something else*) are to be found around 1 TeV

# Where is the next important scale?

(Higher energies needed to measure  $HHH$ , verify that  $H$  regulates  $W_L W_L$ )

Planck scale  $\sim 10^{19}$  GeV (3 + 1-d spacetime)

Unification scale  $\sim 10^{15-16}$  GeV

$\Lambda_{\text{QCD}} \sim$  scale of confinement, chiral symmetry breaking

At what scale are charged-fermion masses set (Yukawa couplings)?

At what scale are neutrino masses set?

Will new physics appear at  $1\times, 10\times, 100\times, \dots$  EW scale?

# The Great Lesson of Twentieth-Century Science

The human scale of space and time is not privileged for understanding Nature, and may even be disadvantaged.

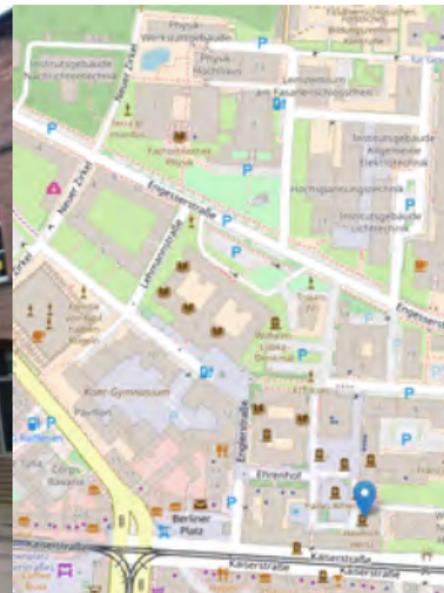
Renormalization group · Effective field theories

Resolution and extent in time and distance

Diversity and scale diversity in experimental undertakings

*The discovery that the human scale is not preferred is as important as the discoveries that the human location is not privileged (Copernicus) and that there is no preferred inertial frame (Einstein), and will prove to be as influential.*

# Heinrich Hertz on Maxwell's Equations



*One cannot study Maxwell's marvelous electromagnetic theory of light without sometimes having the feeling that these mathematical formulae have an independent existence and an intelligence of their own, that they are wiser than we are, wiser even than their inventor, that they give back to us more than was originally put into them. Über die Beziehungen zwischen Licht und Elektrizität*

# How to progress?

Explore the regions of the unknown,  
the unanswered questions

Try to divine where the secrets are hidden

Seek out soft spots in our current understanding,  
especially where the stories we tell are  
*unprincipled*  $\equiv$  not founded on sound principles

Supersymmetry: + *R*-parity +  $\mu$  problem + tame FCNC + ...

Big-Bang Cosmology: + inflation + dark matter + dark energy + ...

Particle content, even gauge groups of the Standard Model

# Guiding Principles

Symmetry (via Noether's Theorems) & Hidden Symmetry

Poincaré Invariance

Relativistic Quantum Field Theory

Unitarity, Causality

Working hypotheses:

Gauge Symmetry

Pointlike constituents

Minkowski spacetime (for most purposes)

...

# On-mass-shell accelerators

Large Hadron Collider Complex at CERN

Fermilab Main Injector

J-PARC Main Ring

BEPC II (IHEP-Beijing)

VEPP-2000 (BINP-Novosibirsk)

SuperKEKB (in commissioning)

Intensity improvement projects for  $\nu$  physics (Fermilab, J-PARC)

[Facility for Antiproton and Ion Research (Darmstadt)]

HL-LHC, promising  $3000 \text{ fb}^{-1}$  at  $\sqrt{s} \rightarrow 14 \text{ TeV}$

# Virtual accelerators

Japan: ILC,  $e^+e^-$  collisions initially at  $\sqrt{s} = 250$  GeV

HE-LHC (energy doubler for the LEP/LHC tunnel),  $pp$  at  $\sqrt{s} \approx 27$  TeV

CLIC-380,  $e^+e^-$  collisions initially up to  $\sqrt{s} = 380$  GeV

LHeC, to collide a 60-GeV  $e$  beam with the LHC  $p$  beam

Electron–Ion Collider, developed by the US nuclear physics community

CERN Future Circular Colliders: 100-km tunnel,  $hh$ ,  $ee$ ,  $eh$  studies

China: CEPC ( $e^+e^-$  Higgs factory) in large tunnel  $\rightsquigarrow$  SppC

# What LHC has taught us about the Higgs Boson

Evidence is developing as it would for a “standard-model” Higgs boson

Unstable neutral particle with  $M_H = 125.18 \pm 0.16$  GeV

Decays to  $W^+W^-$ ,  $ZZ$  implicate  $H$  as agent of EWSB

Decay to  $\gamma\gamma$  as expected (loop-level)      Indirect constraint on  $\Gamma_H$

Dominant spin-parity  $J^P = 0^+$

$Ht\bar{t}$  coupling from  $gg$  fusion,  $t\bar{t}H$  production link to fermion mass origin

$\tau^+\tau^-$  and  $b\bar{b}$  at expected rates

Only third-generation fermion couplings observed;  $\mu^+\mu^-$  constrained

Search-and-discovery phase  $\rightsquigarrow$  painstaking forensic investigation

# Questions about EWSB and the Higgs Sector

- 1 Is  $H(125)$  the only member of its clan? Might there be others—charged or neutral—at higher or lower masses?
- 2 Does  $H(125)$  fully account for electroweak symmetry breaking? Does it match standard-model branching fractions to gauge bosons? Are absolute couplings to  $W$  and  $Z$  as expected in the standard model?
- 3 Is the Higgs field the only source of fermion masses? Are the fermion couplings proportional to fermion masses?  $\mu^+\mu^-$  soon?  
How can we detect  $H \rightarrow c\bar{c}$ ?  $e^+e^-??$  basis of chemistry
- 4 What role does the Higgs field play in generating neutrino masses?
- 5 Are all production rates as expected?

# More questions about EWSB and the Higgs Sector

- 6 Can we establish or exclude decays to new particles? Does  $H(125)$  act as a portal to hidden sectors? When can we measure  $\Gamma_H$ ?
- 7 Can we find any sign of new strong dynamics or (partial) compositeness?
- 8 Can we establish the  $HHH$  trilinear self-coupling?
- 9 How well can we test the notion that  $H$  regulates Higgs–Goldstone scattering, i.e., tames the high-energy behavior of  $WW$  scattering?
- 10 What is the order of the electroweak phase transition?

## More new physics on the TeV scale and beyond?

Before LHC, much informed speculation—but no guarantees—about what might be found, beyond keys to EWSB.

Many eyes were on supersymmetry or Technicolor to enforce  
 $M_W \lll$  unification scale or Planck scale.

“WIMP miracle” pointed to the TeV scale for a dark matter candidate.

Some imagined that neutrino mass might be set on the TeV scale.

*No direct sign of physics beyond the standard model has come to light.*

Might first hints may come from precision measurements?

# Have we misconstrued naturalness and the hierarchy problem?

Did the existence of two once-and-done solutions to the hierarchy problem (supersymmetry and technicolor) lead us to view the discipline of naturalness too simplistically?



Available online at [www.sciencedirect.com](http://www.sciencedirect.com)



Nuclear Physics B (Proc. Suppl.) 140 (2005) 3–19

**NUCLEAR PHYSICS B**  
**PROCEEDINGS**  
**SUPPLEMENTS**

[www.elsevierphysics.com](http://www.elsevierphysics.com)

The final blunder was a claim that scalar elementary particles were unlikely to occur in elementary particle physics at currently measurable energies unless they were associated with some kind of broken symmetry [23]. The claim was that, otherwise, their masses were likely to be far higher than could be detected. The claim was that it would be unnatural for such particles to have masses small enough to be detectable soon. But this claim makes no sense when one becomes familiar with the history of physics. There have been a number of cases where numbers arose that were unexpectedly small or large. An early example was the very large distance to the nearest star as compared to the distance to the Sun, as needed by Copernicus, because otherwise the nearest stars would have exhibited measurable parallax as the Earth moved around the Sun. Within elementary particle physics, one has unexpectedly large ratios of masses, such as the large ratio of the muon mass to the electron mass. There is also the very small value of the weak coupling constant. In the time since my paper was written, another set of unexpectedly small masses was discovered: the neutrino masses. There is also the riddle of dark energy in cosmology, with its implication of possibly an extremely small value for the cosmological constant in Einstein's theory of general relativity.

## The Origins of Lattice Gauge Theory

K.G. Wilson

Smith Laboratory, Department of Physics, The Ohio State University, 174 W. 18th Ave., Columbus, OH 43210

# Questions about new physics on the TeV scale and beyond

- 11 Do quarks and leptons show signs of compositeness? Are they made of more elementary constituents?
- 12 Can we find evidence of a dark matter candidate?
- 13 Why is empty space so nearly massless? What is the resolution to the vacuum energy problem?
- 14 Will “missing energy” events signal the existence of spacetime dimensions beyond the familiar  $3 + 1$ ?
- 15 Can we probe dark energy in laboratory experiments?
- 16 Can we find clues to the origin of electroweak symmetry breaking? Is there a dynamical origin to the “Higgs potential?”

## More questions about new physics on the TeV scale and beyond

- 17 Might we find indirect evidence for a new family of strongly interacting particles, such as those that are present in supersymmetric extensions of the standard model, by seeing a change in the evolution of the strong coupling “constant,”  $1/\alpha_s$ , at the HE-LHC or a “100-TeV” collider?
- 18 How can we constrain—or provide evidence for—light dark-matter particles or other denizens of the dark in high-energy colliders or beam-dump experiments?
- 19 Does the gluon have heavy partners, indicating that QCD is part of a structure richer than  $SU(3)_c$ ?
- 20 How can technologies developed for accelerators advance the search for axions? How can we observe axions, dark photons, ...?

# Flavor: the problem of identity

We do not have a clear view of how to approach the diverse character of the constituents of matter

CKM paradigm: extraordinarily reliable framework in hadron sector

**BUT—many parameters:** no clue what determines them, nor at what energy scale they are set

Even if Higgs mechanism explains *how* masses and mixing angles arise, we do not know *why* they have the values we observe

Physics beyond the standard model!

# Flavor: the problem of identity (continued)

## Parameters of the Standard Model

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3	Coupling parameters, $\alpha_s$ , $\alpha_{em}$ , $\sin^2 \theta_W$
2	Parameters of the Higgs potential
1	Vacuum phase (QCD)
6	Quark masses
3	Quark mixing angles
1	CP-violating phase
3	Charged-lepton masses
3	Neutrino masses
3	Leptonic mixing angles
1	Leptonic CP-violating phase (+ Majorana phases?)
<hr/>	
26 <sup>+</sup>	Arbitrary parameters

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Will we see or diagnose a break in the SM? Contrast Landscape perspective

## Questions concerning the problem of identity

- 21 Can we find evidence of right-handed charged-current interactions?  
Is nature built on a fundamentally asymmetrical plan, or are the right-handed weak interactions simply too feeble for us to have observed until now, reflecting an underlying symmetry hidden by spontaneous symmetry breaking?
- 22 Are there additional electroweak gauge bosons, beyond  $W^\pm$  and  $Z$ ?
- 23 Is charged-current universality exact?  
What about lepton-flavor universality?

## More questions concerning the problem of identity

- 24 Where are flavor-changing neutral currents? In the standard model, these are absent at tree level and highly suppressed by the Glashow–Iliopoulos–Maiani mechanism. They arise generically in proposals for physics beyond the standard model, and need to be controlled.  
And yet we have made no sightings! Why not?
- 25 Can we find evidence for charged-lepton flavor violation?
- 26 Why are there three families of quarks and leptons? (Is it so?)

# Neutrinos ...

Neutrinos oscillate among the three known species,  $\nu_e, \nu_\mu, \nu_\tau$   
(discovered with neutrinos from natural sources)

Accelerator-based experiments  $\text{NO}\nu\text{A}$  and T2K

↪ DUNE and Hyper-Kamiokande + new short-baseline experiments

Tritium  $\beta$ -decay experiment KATRIN  
experiments that rely on reactors (JUNO)  
or natural sources (IceCube and KM3Net)

Puzzling results: LSND–MiniBooNE, “Reactor anomaly”

## Some outstanding questions in neutrino physics

- 27 What is the order of levels of the mass eigenstates  $\nu_1, \nu_2, \nu_3$ ? It is known that the  $\nu_e$ -rich  $\nu_1$  is the lighter of the “solar pair,” with the more massive  $\nu_2$ . Does the  $\nu_e$ -poor  $\nu_3$  lie above (“normal” mass hierarchy) or below (“inverted hierarchy”) the others?
- 28 What is the absolute scale of neutrino masses? **KATRIN vs. Cosmo?**
- 29 What is the flavor composition of  $\nu_3$ ? Is it richer in  $\nu_\mu$  or  $\nu_\tau$ ?
- 30 Is CP violated in neutrino oscillations? To what degree?
- 31 Are neutrinos Majorana particles? While this issue is primarily addressed by searches for neutrinoless double- $\beta$  decay, collider searches for same-sign lepton pairs also speak to it.
- 32 Do three light (left-handed) neutrinos suffice?

## Some outstanding questions in neutrino physics

- 33 Are there light sterile neutrinos? If so, how could they arise?
- 34 Do neutrinos have nonstandard interactions, beyond those mediated by  $W^\pm$  and  $Z$ ?
- 35 How can we detect the cosmic neutrino background?  
Each species, now:  $56 \text{ cm}^{-3}$      $T_\nu \approx 2 \text{ K} \approx 1.7 \times 10^{-4} \text{ eV}$
- 36 Are all the neutrinos stable?
- 37 Do neutrinos contribute appreciably to the dark matter of the Universe?
- 38 How is neutrino mass a sign of physics beyond the standard model?
- 39 Will neutrinos give us insight into the matter excess in the Universe (through leptogenesis)?

## Consider a neutrino factory

A Neutrino Factory based on a muon storage ring could provide a very strong second act for the coming generation of accelerator-based neutrino experiments.

Beyond its application to oscillation experiments as an intense source with known composition, an instrument that delivered  $10^{20}$   $\nu$  per year could be a highly valuable resource for on-campus experiments.

Neutrino interactions on thin targets, polarized targets, or active targets could complement the nucleon-structure programs carried out in electron scattering at Jefferson Lab and elsewhere.

# Don't forget the strong interactions!

QCD could be complete, up to  $M_{\text{Planck}}$  (modulo strong CP problem)  
... but that doesn't prove it must be

Prepare for surprises, such as

(Breakdown of factorization)

Free quarks / unconfined color

New kinds of colored matter

Quark compositeness

Larger color symmetry containing  $SU(3)_c$

# Questions pertaining to QCD

- 40 Are there new phenomena within QCD?  
Multiple production beyond diffraction + short-range order?  
Long-range correlations in  $y$  (or  $\eta$ )? Unusual event structures?
- 41 Will the expected high density of few-GeV partons lead to thermalization in  $pp$  collisions? What will be other consequences?
- 42 How will correlations among partons in a proton manifest themselves?
- 43 Can we distinguish spatial configurations of partons within protons?
- 44 What is the importance of intrinsic heavy flavors?
- 45 What body plans for hadrons can we identify beyond  $qqq$  and  $q\bar{q}$ ?

# Motivations for unified theories

Neutrality of atoms, balance of electron and proton charges

Quarks and leptons are spin- $\frac{1}{2}$  particles  
that come in matched sets  
as required by anomaly cancellation  
for a renormalizable  $SU(2)_L \otimes U(1)_Y$  theory

$SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$  gauge couplings tend to converge at high scales

## Questions for unified theories

- 46 What is the relationship of quarks to leptons?
- 47 Which quark doublet is matched with which lepton doublet?
- 48 Are there new gauge interactions linking quarks with leptons?
- 49 What is the (grand) unifying symmetry?
- 50 What determines the low-energy gauge symmetries?
- 51 What are the steps to unification? One more, or multiple?
- 52 Is perturbation theory a reliable guide to coupling-constant unification?
- 53 What sets the mass scale for the additional gauge bosons in a unified theory? ... for the additional Higgs bosons?

# Questions for unified theories

- 54 Is the proton unstable? How does it decay?
- 55 Is neutron–antineutron oscillation observable?
- 56 How can we incorporate gravity?
- 57 What is the nature of spacetime?  
Is it emergent?  
How many dimensions?

# A word about the astro/cosmo connection

We do not know what the Universe at large is made of

We do not know the complete thermal history of the universe

We have not accounted for the predominance of matter over antimatter in the observed universe

We do not know what provoked inflation (if it happened)

We do not know why the expansion of the universe is accelerating

*The First Three Minutes* is still a beautiful story, and so is  $\Lambda$ CDM!

# Tabletop precision experiments

Electric dipole moment  $d_e$ : CP/T violation

$$|d_e| < 8.7 \times 10^{-29} \text{ e cm}$$

ACME Collaboration, ThO

$$|d_e| < 1.3 \times 10^{-28} \text{ e cm}$$

NIST, trapped  $^{180}\text{Hf } ^{19}\text{F}^+$

$$(\text{SM phases: } d_e < 10^{-38} \text{ e cm})$$

(How) can we observe electric dipole moments of  $e, \mu, p$ ?

# “Tabletop” precision experiments

(Anti)proton magnetic moments: CPT test

$$\mu_{\bar{p}} = -2.792\,847\,344\,1(42) \mu_N$$

vs.

$$\mu_p = +2.792\,847\,344\,62(82) \mu_N$$

*BASE Collaboration @CERN Antiproton Decelerator*

# Exercise 1.

How should we respond if:

- (a) *The DAMA “seasonal variation” cannot be explained away?*
  - (b) *The LHC Higgs signal strength settles at  $\mu = 1.17 \pm 0.03$ ?  
Or if  $Ht\bar{t}$  remains high?*
  - (c) *The LHCb flavor anomalies persist?*
  - (d) *The  $(g - 2)_\mu$  anomaly strengthens?*
  - (e) *WIMP dark matter searches reach the neutrino floor?*
- ... (extra credit)

## Exercise 2.

*Sketch five “small-scale” (you define) experiments with the potential to change our thinking about particle physics or related fields.*

## Exercise 3.

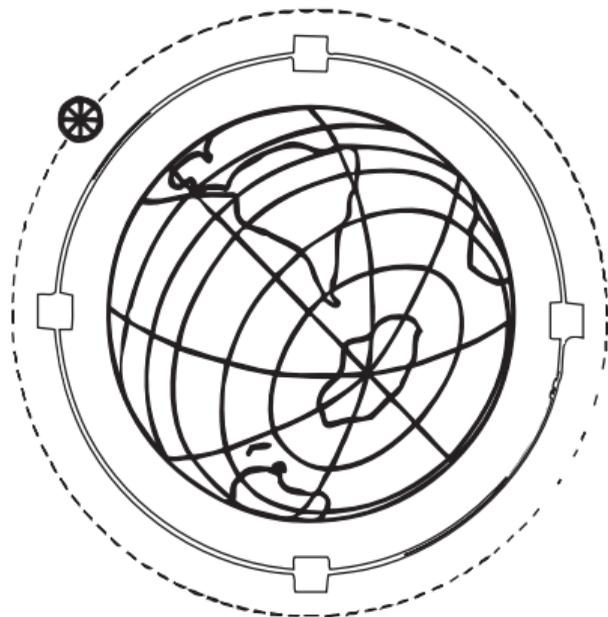
How would *you* assess the scientific potential (in view of cost and schedule) of

- (a) *The High-Luminosity LHC?*
- (b) *The High-Energy LHC?*
- (c) *A 100-TeV pp Collider (FCC-hh)?*
- (d) *A 250-GeV ILC?*
- (e) *A circular Higgs factory (FCC-ee or CEPC)?*
- (f) *A 380-GeV CLIC?*
- (g) *LHeC / FCC-eh? (or an electron-ion collider?)*
- (h) *A muon-storage-ring neutrino factory?*
- (i) *A multi-TeV muon collider?*
- (j) *The instrument of your dreams?*

# Big dreams

## Fermi's dream accelerator (1954)

$$E_{\text{beam}} = 5000 \text{ TeV}, \$1.7 \times 10^{11}$$



A THOUSAND TeV IN THE CENTER OF MASS:

INTRODUCTION TO HIGH ENERGY STORAGE RINGS<sup>†</sup>

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### I. INTRODUCTION

These lectures must begin with an apology. Normally at schools such as this, one expects the lecturer to be an acknowledged expert on the subject matter he is discussing. Here this is not the case. Design of high energy proton storage rings is not exactly my forte. Why am I doing this? There are several reasons, short of mental illness.\*

1. I want to learn this subject myself and there is no better way than trying to teach it. And Ferbel didn't stop me.

2. There needs to be a broader knowledge of accelerator physics in the elementary-particle community. Experimentalists at the storage rings find themselves especially closely coupled to their machine and its operation. And theorists can find interesting and challenging questions which lie at the frontier of the very active field of nonlinear mechanics.

3. Straightforward extrapolation of existing acceleration techniques would seem to lead to very large, expensive machines. While we may envision one, perhaps two generations of future accelerators using essentially existing techniques, the question of how to go beyond that is a difficult one. There seems to be a growing feeling that it is not too soon to start to face up to the problem. A look at the alternative--as we do here--can only provide stimulation.

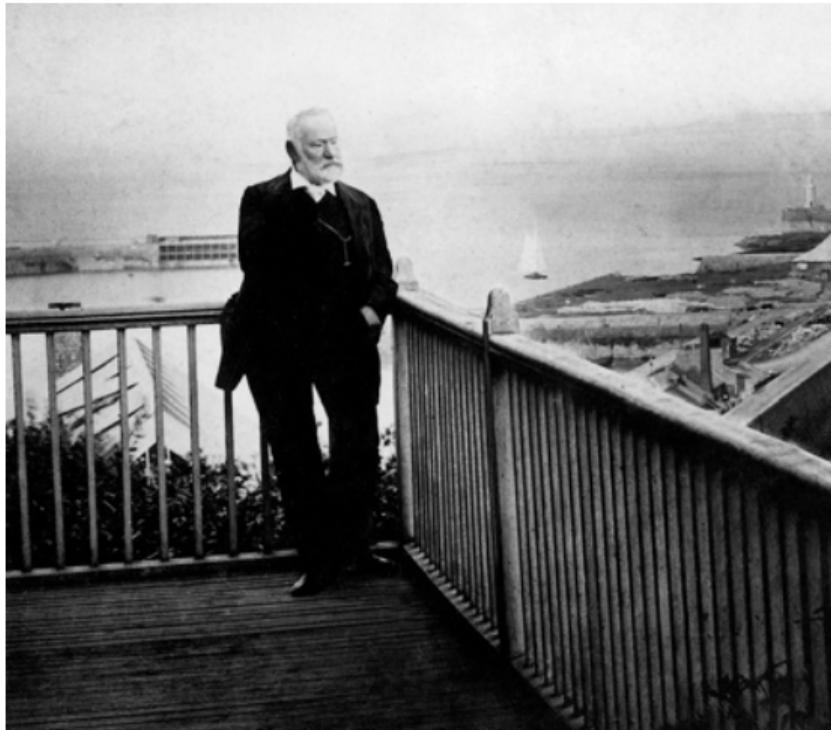
## Questions inspired by Big Dreams

- 58 Suppose we could reach gradients of many GeV—even a TeV—per meter. How would we first apply that bit of magic, and what characteristics other than gradient would be required?
- 59 If we *could* shrink multi-TeV accelerators, how might we shrink detectors that depend on particle interactions with matter?
- 60 What could we do with a low-emittance, high-intensity muon source?
- 61 What inventions would it take to accelerate beams of particles with picosecond lifetimes?
- 62 How can we imagine going far beyond current capabilities for steering beams? How might we apply high-transmissivity crystal channeling?
- 63 How would optimizations change if we could shape superconducting magnet coils out of biplanar graphene or an analogous material?

A final question (for now)!

How are we prisoners of conventional thinking?

How can we break out?



LA PLURALITÉ  
DES  
MONDES HABITÉS,

ÉTUDE

DE L'UN ESPÈCE LES CONDITIONS D'HABITABILITÉ DES TERRES CÉLESTES.  
INDICÉES AU POINT DE VUE DE L'ASTRONOMIE ET DE LA PHYSIOLOGIE;

CAMILLE FLAMMARION,

AUTEUR LITTÉRAIRE & CORRESPONDANT GÉNÉRAL DE PARIS, PASTEUR DE L'ÉCOLE DE  
SCIENCE DE NIMÈS, MÈMBRE DU SÉMINAIRE NIMÈS, ETC.

NUMÉRO DE LA COLLECTION  
ÉTUDE SUR LES MONDES HABITÉS, 1862  
LA TOUTE SÉRIÉUSEMENT GARDÉE ET DE LA DÉTERMINER

PARIS,

MALLET-BACHELIER, IMPRIMEUR-LIBRAIRE,  
DE MEGAN DES LIGNES, EN L'ÉTABLISSEMENT GÉNÉRAL DE PARIS,  
OPÉA DES AUGUSTINS, 35.

1862

Au delà du visible l'invisible, au delà de l'invisible l'inconnu.

—Victor Hugo, *Choses de l'infini* (1864)

Au delà de l'inconnu l'imaginaire!