

Constraining Flavored Dark Matter: A Systematic Study in the DMFV Framework

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J. Heisig, M. Krämer and F. Wilsch

22.07.2025

Evidence for Dark Matter

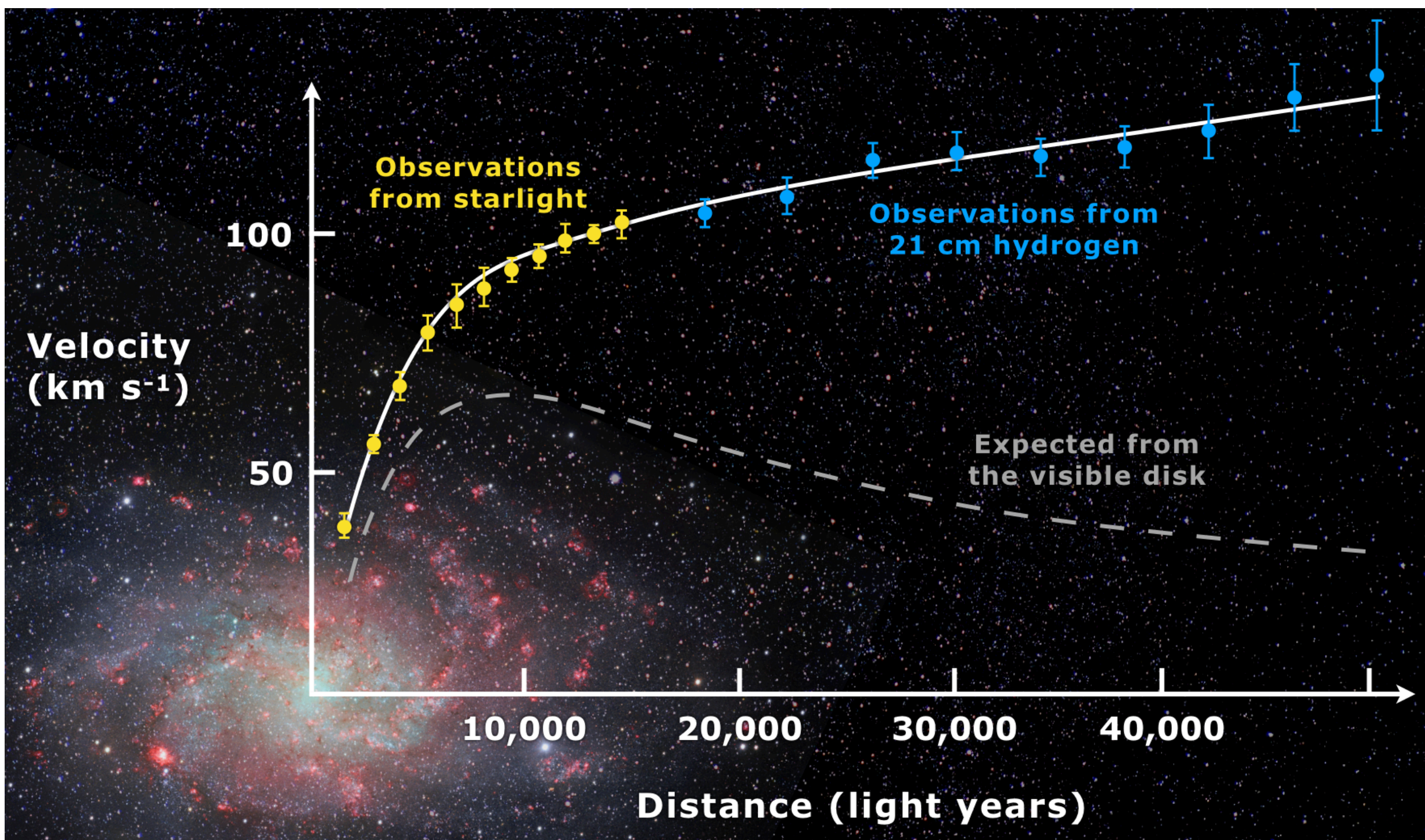
[Zwicky 1933]

[PLANCK collaboration 2018]

Evidence for Dark Matter

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Galaxy rotation curves

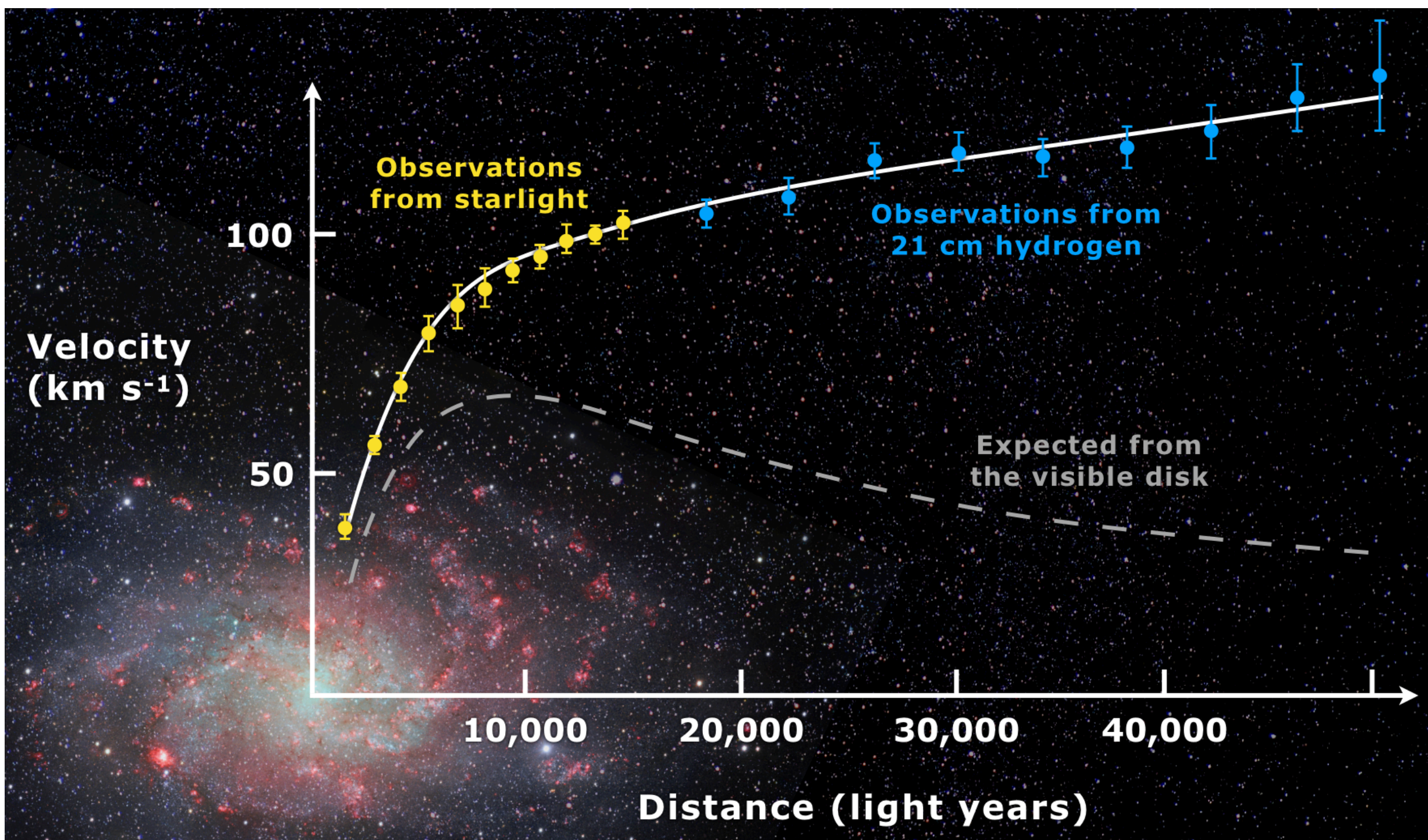


[Wikipedia]

Evidence for Dark Matter

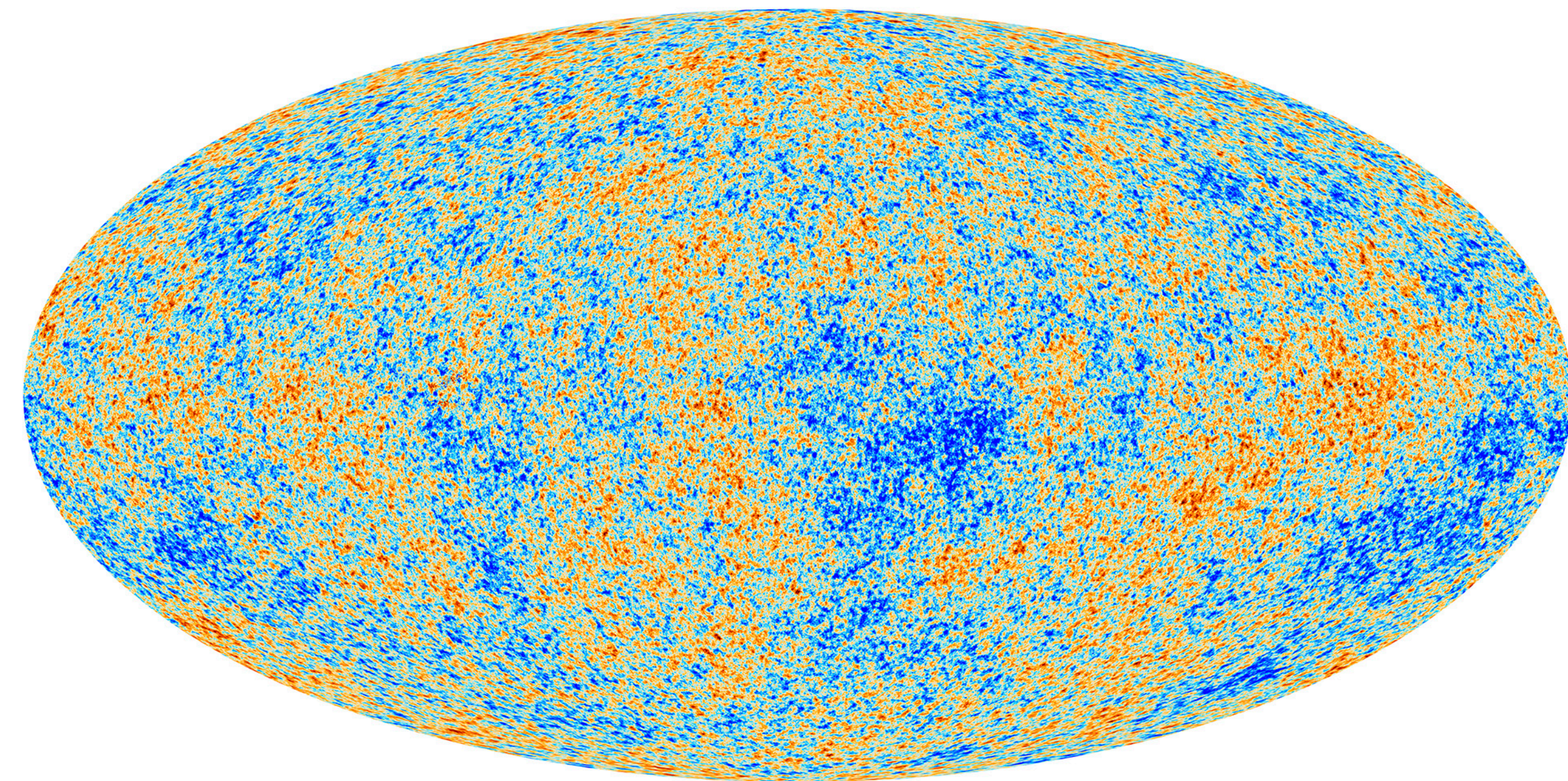
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Galaxy rotation curves



[Wikipedia]

CMB anisotropies



[ESA]

Evidence for Dark Matter

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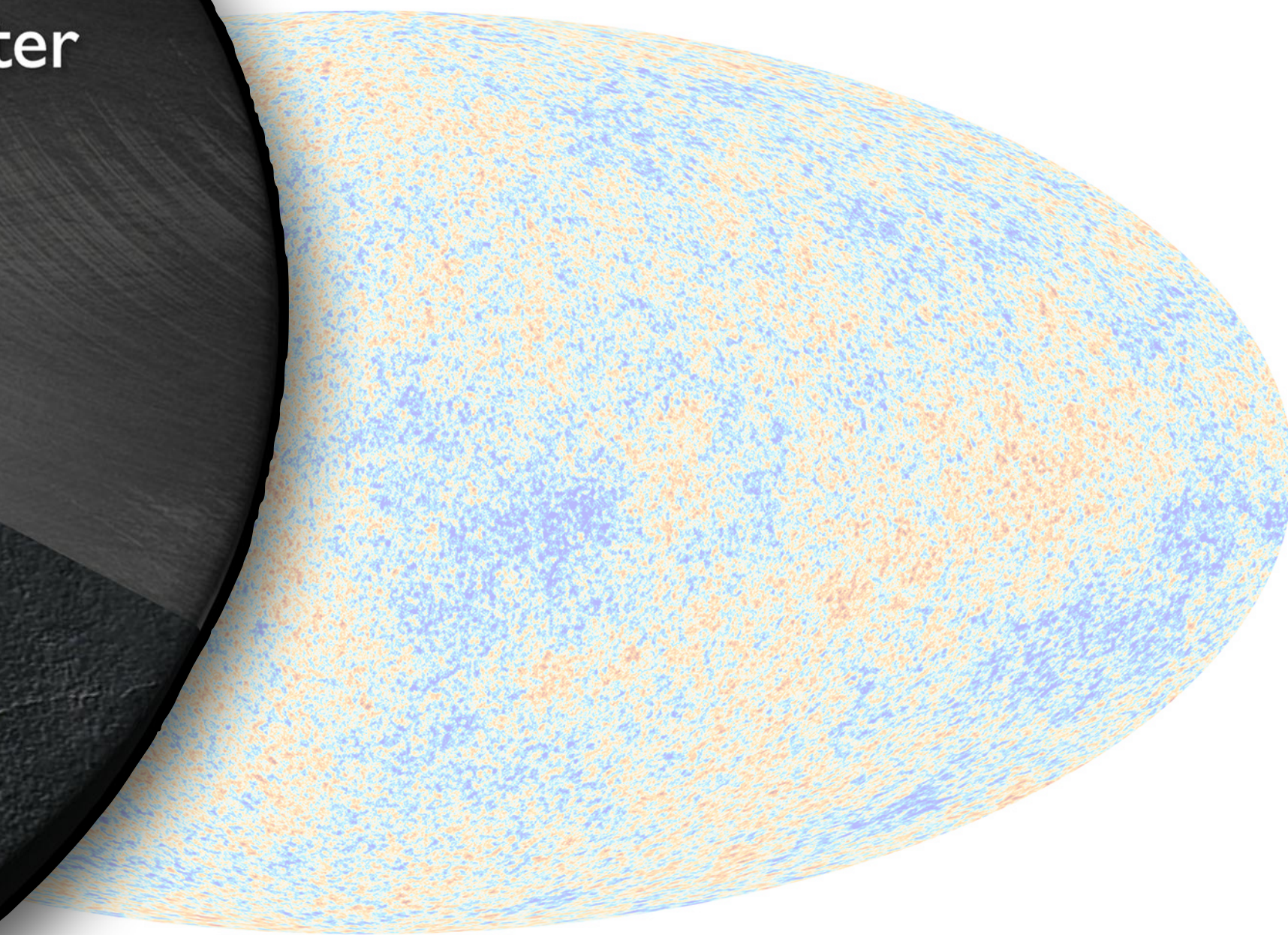
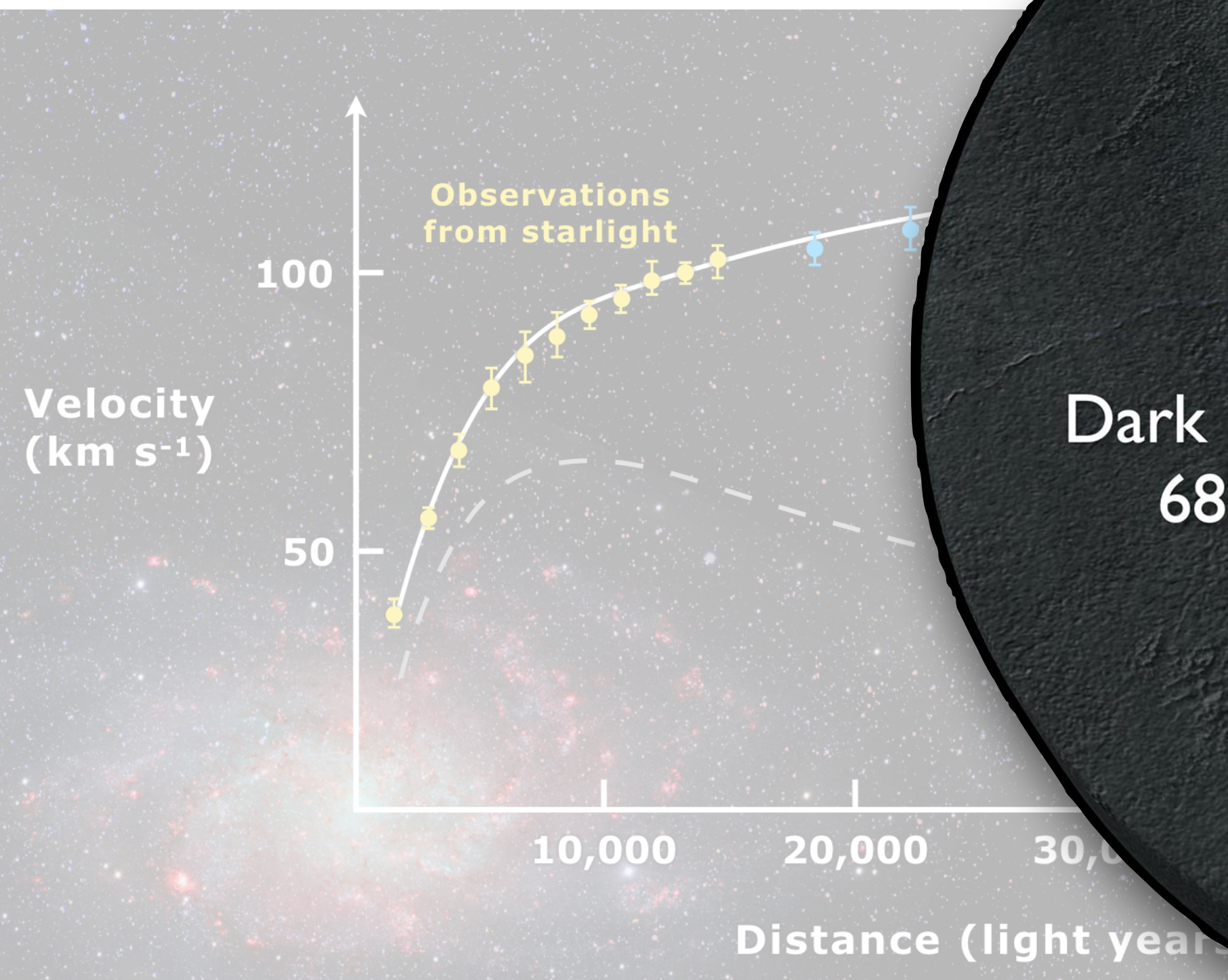
Ordinary Matter
4.9%

CMB anisotropies

Dark Matter
26.8%

Dark Energy
68.3%

Observations
from starlight



[ESA]

What is Dark Matter (DM)?



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We have no idea

What is Dark Matter (DM)?



We have no idea

**→ Study different
models for DM, here
flavoured DM**

Why Flavoured Dark Matter?

- Traditional weakly interacting massive particle (WIMP) models under severe pressure
→ Assume **DM has non-trivial flavour structure** like SM
- New source of flavour and CP violation
→ Different phenomenology
- Introduce **dark minimal flavor violation (DMFV)** concept to keep parameters minimal

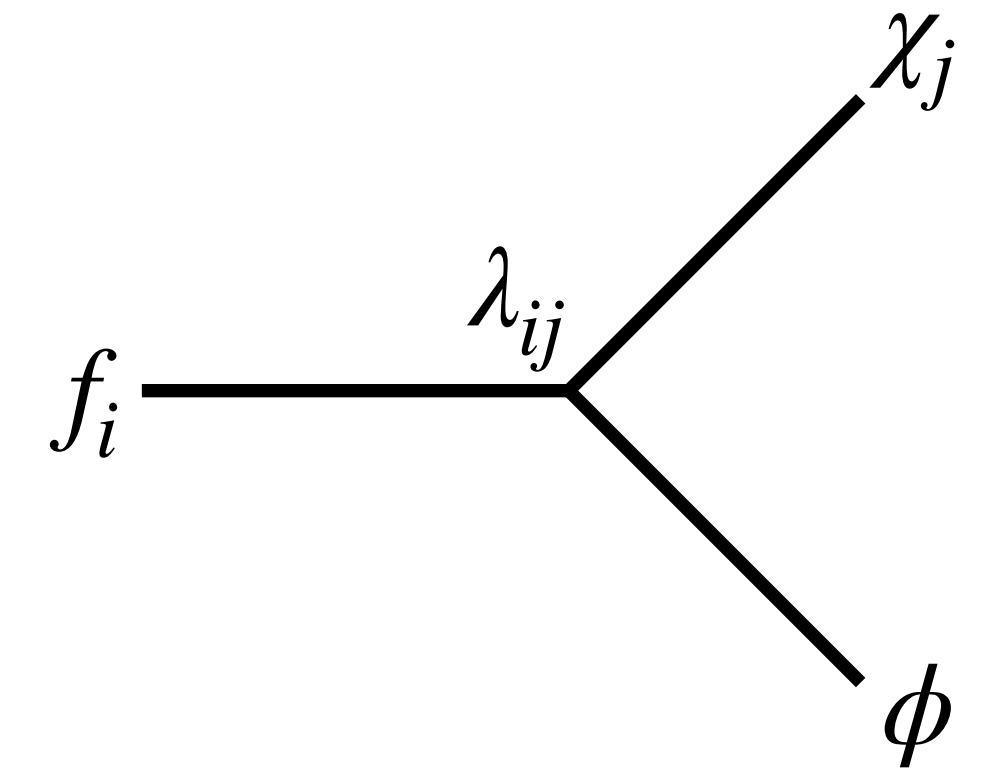


three generations of matter (fermions)			
	I	II	III
mass	$\approx 2.16 \text{ MeV}/c^2$	$\approx 1.273 \text{ GeV}/c^2$	$\approx 172.57 \text{ GeV}/c^2$
	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$
charge	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$
	u up	c charm	t top
spin	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$
QUARKS	d down	s strange	b bottom
	$\approx 4.7 \text{ MeV}/c^2$	$\approx 93.5 \text{ MeV}/c^2$	$\approx 4.183 \text{ GeV}/c^2$
LEPTONS	$\approx 0.511 \text{ MeV}/c^2$	$\approx 105.66 \text{ MeV}/c^2$	$\approx 1.77693 \text{ GeV}/c^2$
	-1	-1	-1
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$
	e electron	μ muon	τ tau
	$< 0.8 \text{ eV}/c^2$	$< 0.17 \text{ MeV}/c^2$	$< 18.2 \text{ MeV}/c^2$
	0	0	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino

[Wikipedia]

Dark Minimal Flavour Violation [Agrawal, Blanke, Gemmler 2014]

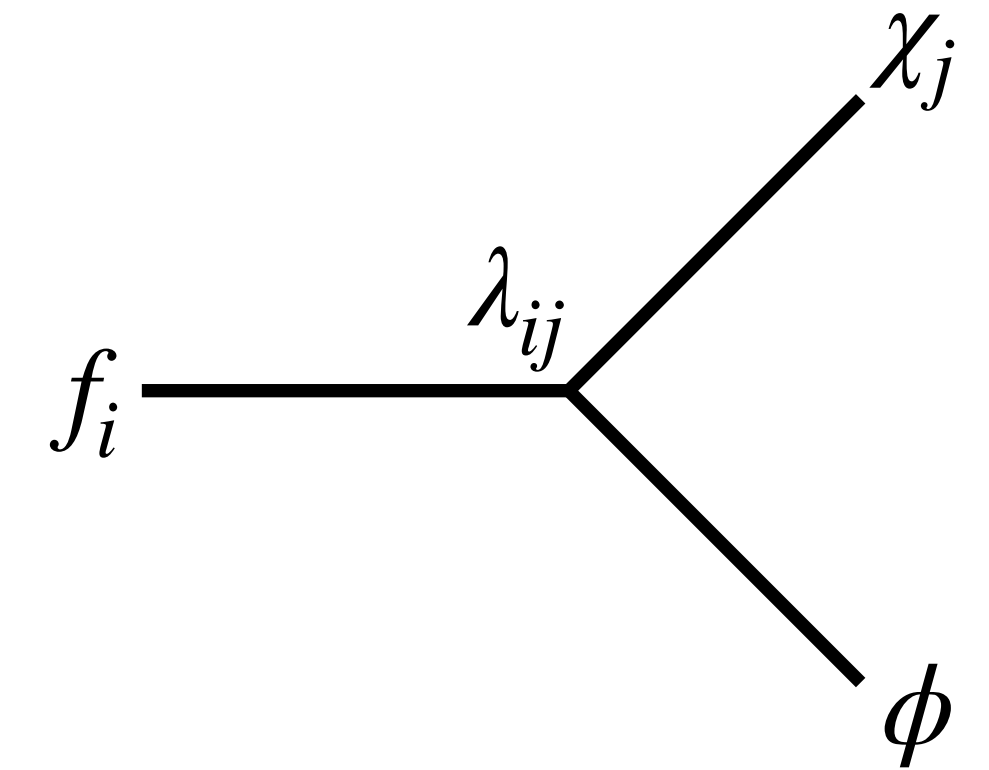
- Extend SM by **single new flavour symmetry and new fields** (DM flavour triplet $X = (X_1, X_2, X_3)^T$ and mediator Y)
- **Couple to SM fermions f** through complex 3×3 matrix λ
- Mass matrix M_X expanded in powers of λ



Dark Minimal Flavour Violation

[Agrawal, Blanke, Gemmler 2014]

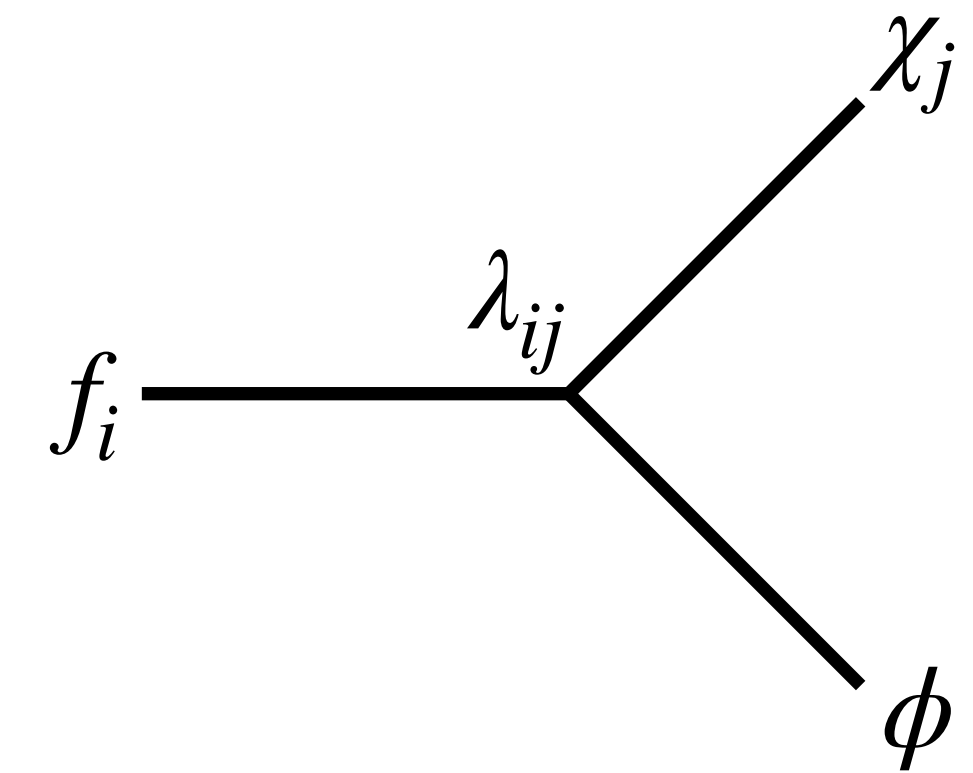
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X_j	Y	f_i

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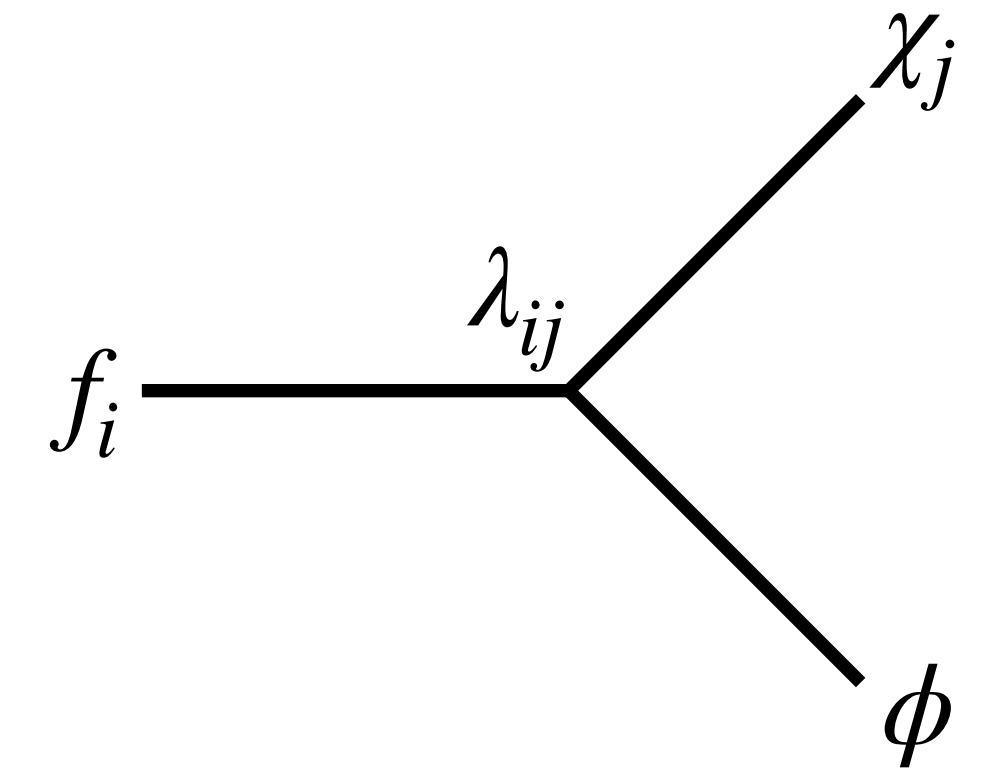
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X_j	Y	f_i
Spin 0 • Real scalar • Complex scalar	• Vector-like fermion	

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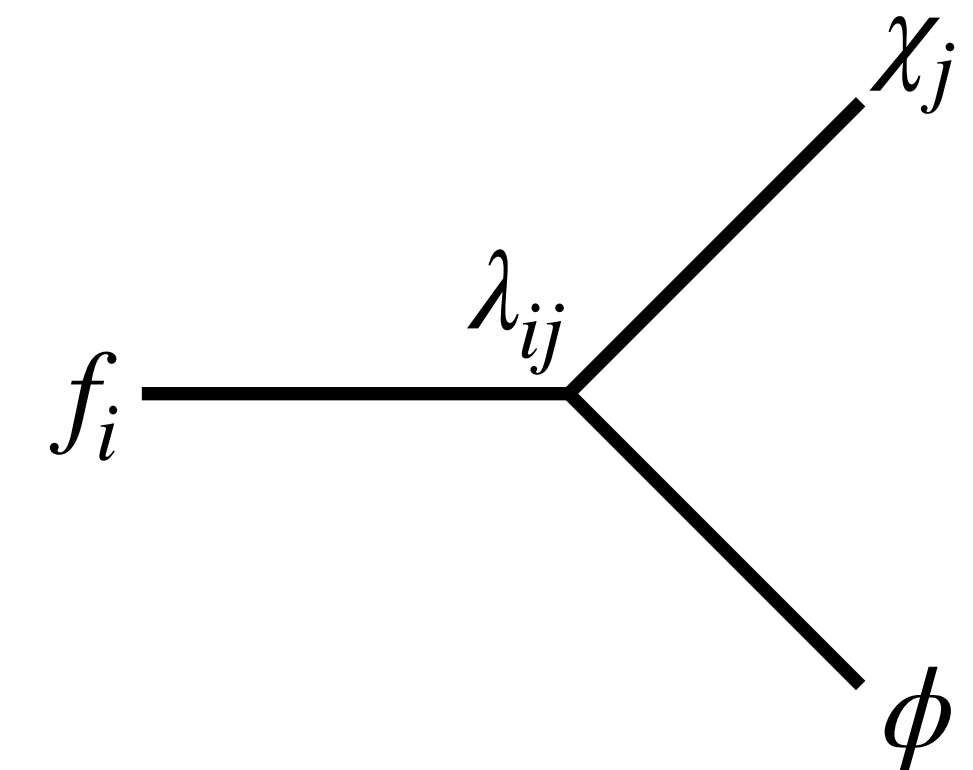
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Spin 1/2 • Majorana fermion • Dirac fermion	• Complex scalar	

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X_j	Y	f_i
Spin 0 • Real scalar • Complex scalar	• Vector-like fermion	• Up-type quarks u_R • Down-type quarks d_R • Charged leptons e_R
Spin 1/2 • Majorana fermion • Dirac fermion	• Complex scalar	• Quark doublet q_L • Lepton doublet ℓ_L

Previously Studied Models

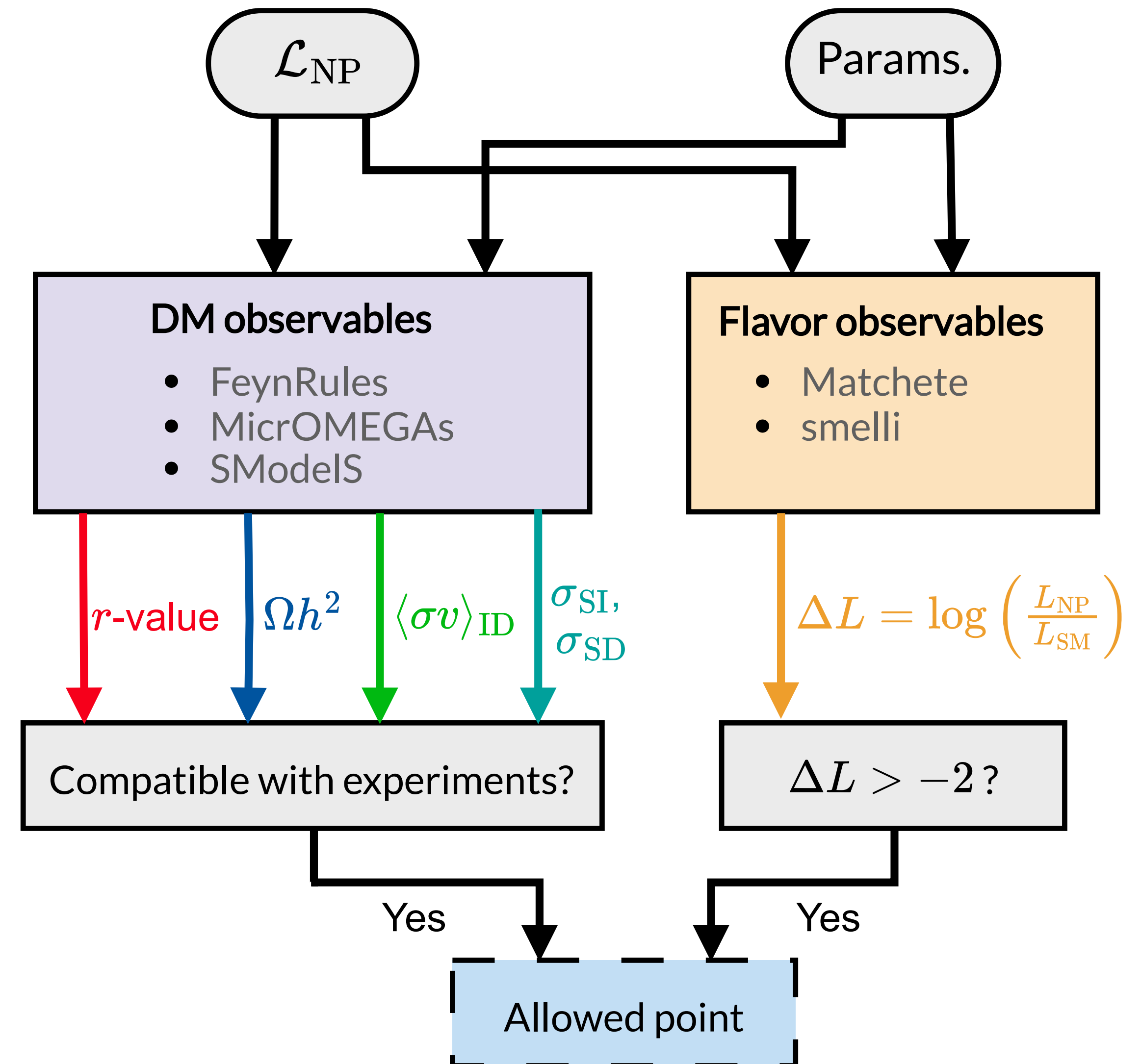
	u_R	d_R	q_L	e_R	ℓ_L
Real scalar DM	×	×	×	×	×
Complex scalar DM	×	×	×	[Acaroglu, Agrawal, Blanke 2022] [Acaroglu, Blanke, Tabet 2022] [Acaroglu, Agrawal, Blanke 2023]	[Acaroglu, Agrawal, Blanke 2022]
Dirac fermion DM	[Blanke, Kast 2017] [Jubb, Kirk, Lenz 2017] [Blanke et al. 2021]	[Agrawal, Blanke, Gemmler 2022] [Bensalem, Stolarski 2022]	[Blanke, Das, Kast 2018] [Blanke et al. 2021]	[Chen, Huang, Takhistov 2016]	×
Majorana fermion DM	[Acaroglu, Blanke 2022] [Acaroglu, LR et al. 2024]	×	×	×	×

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	u_R	d_R	q_L	e_R	ℓ_L
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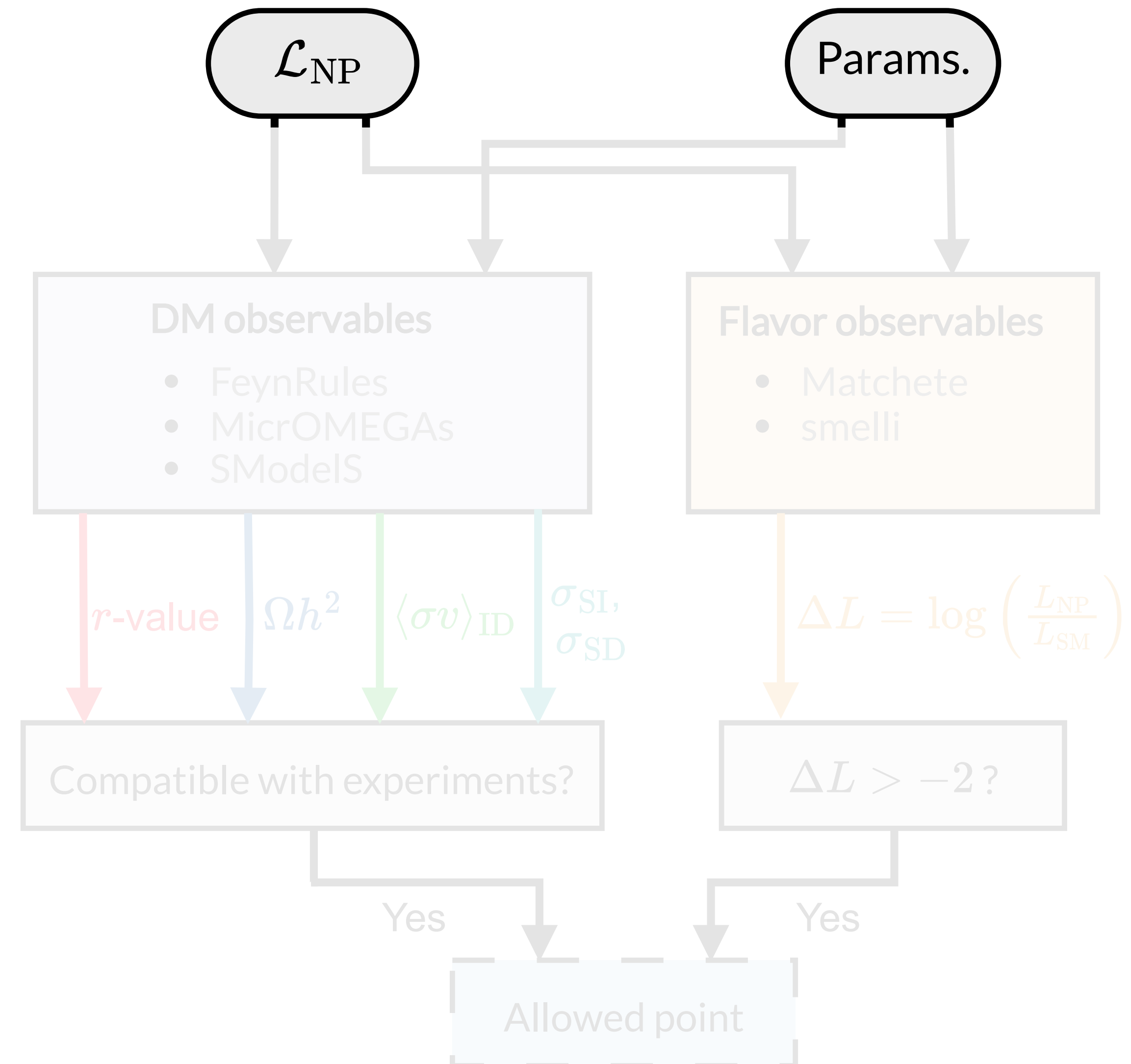
→ Develop framework to study phenomenology of all possible models

Tool-Chain



Tool-Chain - Input

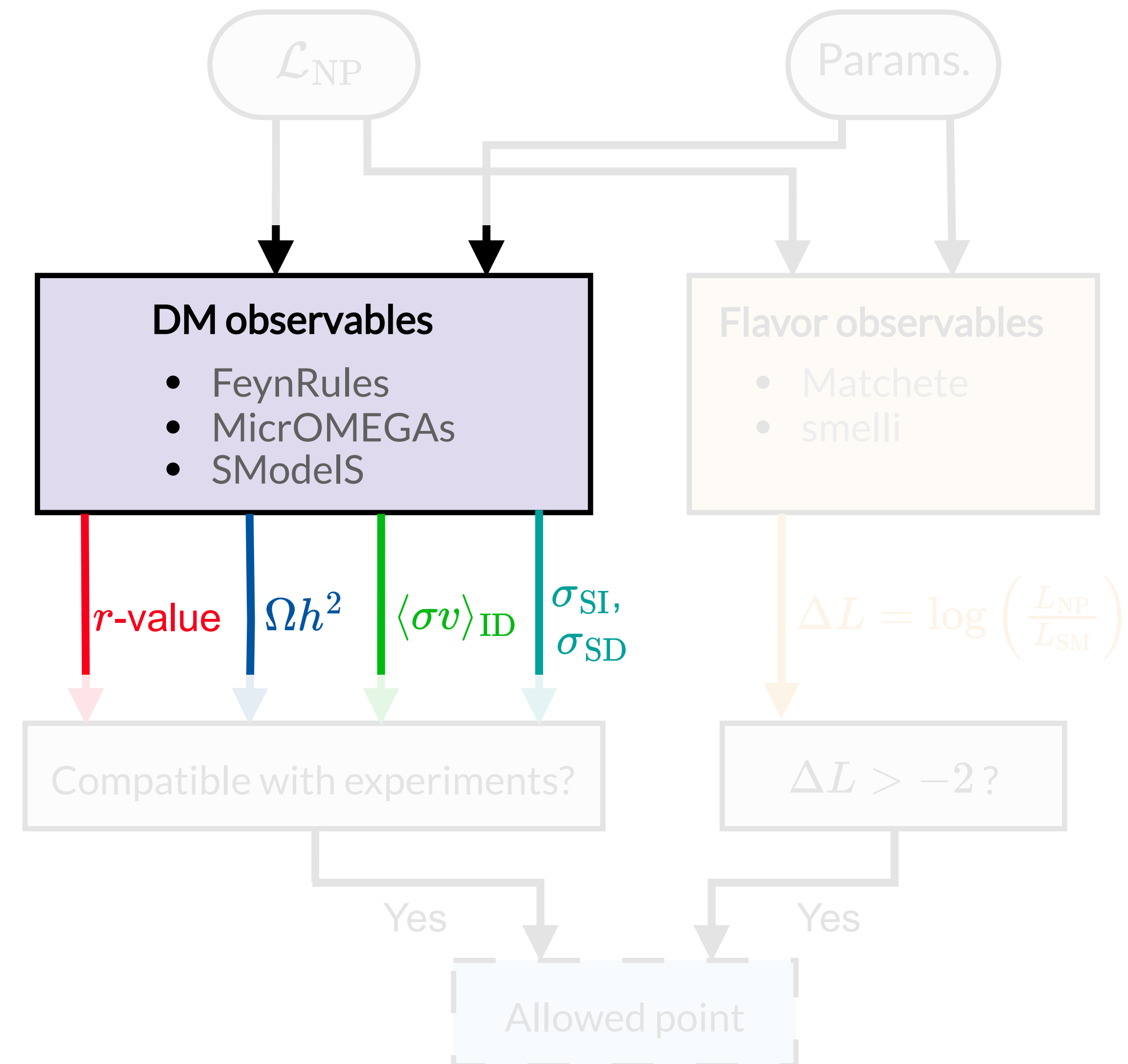
- Lagrangian $\mathcal{L}_{\text{NP}}$ describing all new physics (NP) interactions of DM model
- Values for masses M_{X_i} and M_Y
- Values for couplings to SM particles



Tool-Chain - DM Observables

[Alloul et al. 2014]

- Implemented all models in FEYNRULES
- All FeynRules, CalcHEP and UFO model files available on GitHub in [lena-ra/Flavored-Dark-Matter](https://github.com/lena-ra/Flavored-Dark-Matter)
- CalcHEP files used for implementation in MICRAMEGAS

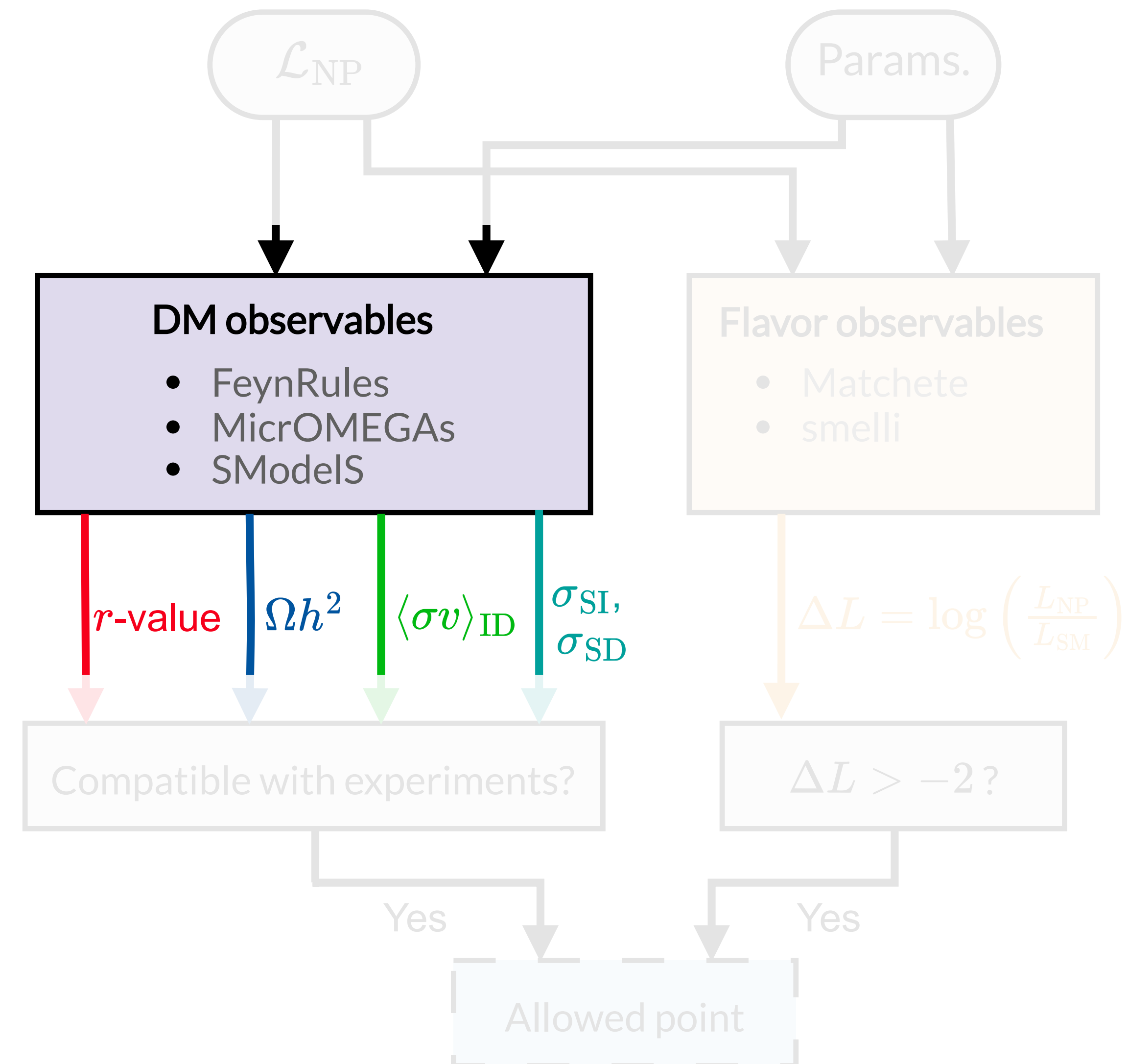


Tool-Chain - DM Observables

[Bélanger et al. 2018]

[Altakach et al. 2024]

- MICROMEGAS computes **DM properties**:
 - Relic abundance Ωh^2
 - Annihilation cross section $\langle \sigma v \rangle$ for every annihilation channel for indirect detection
 - Spin (in)dependent DM-nucleon cross section $\sigma_{\text{SI/SD}}$



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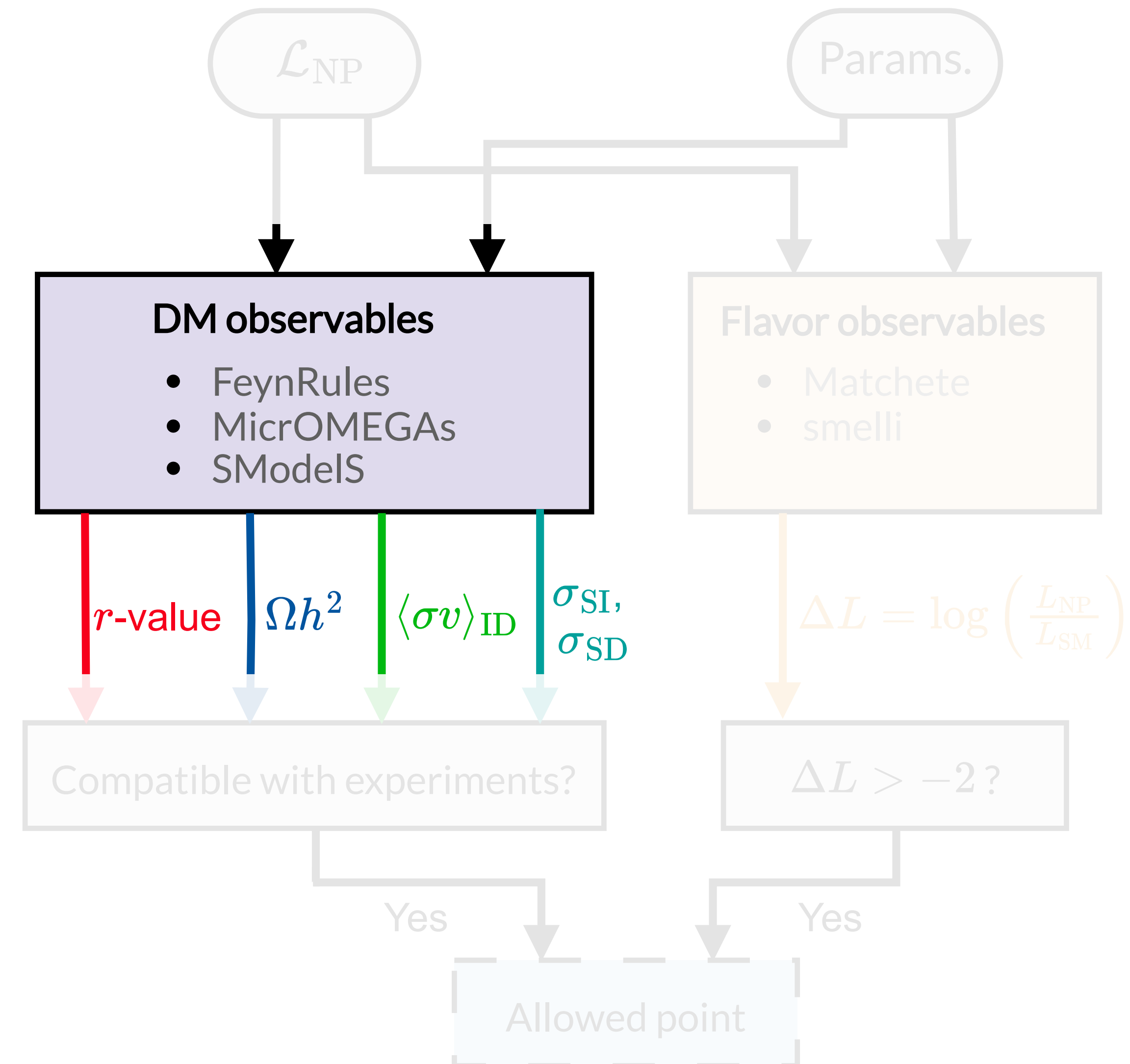
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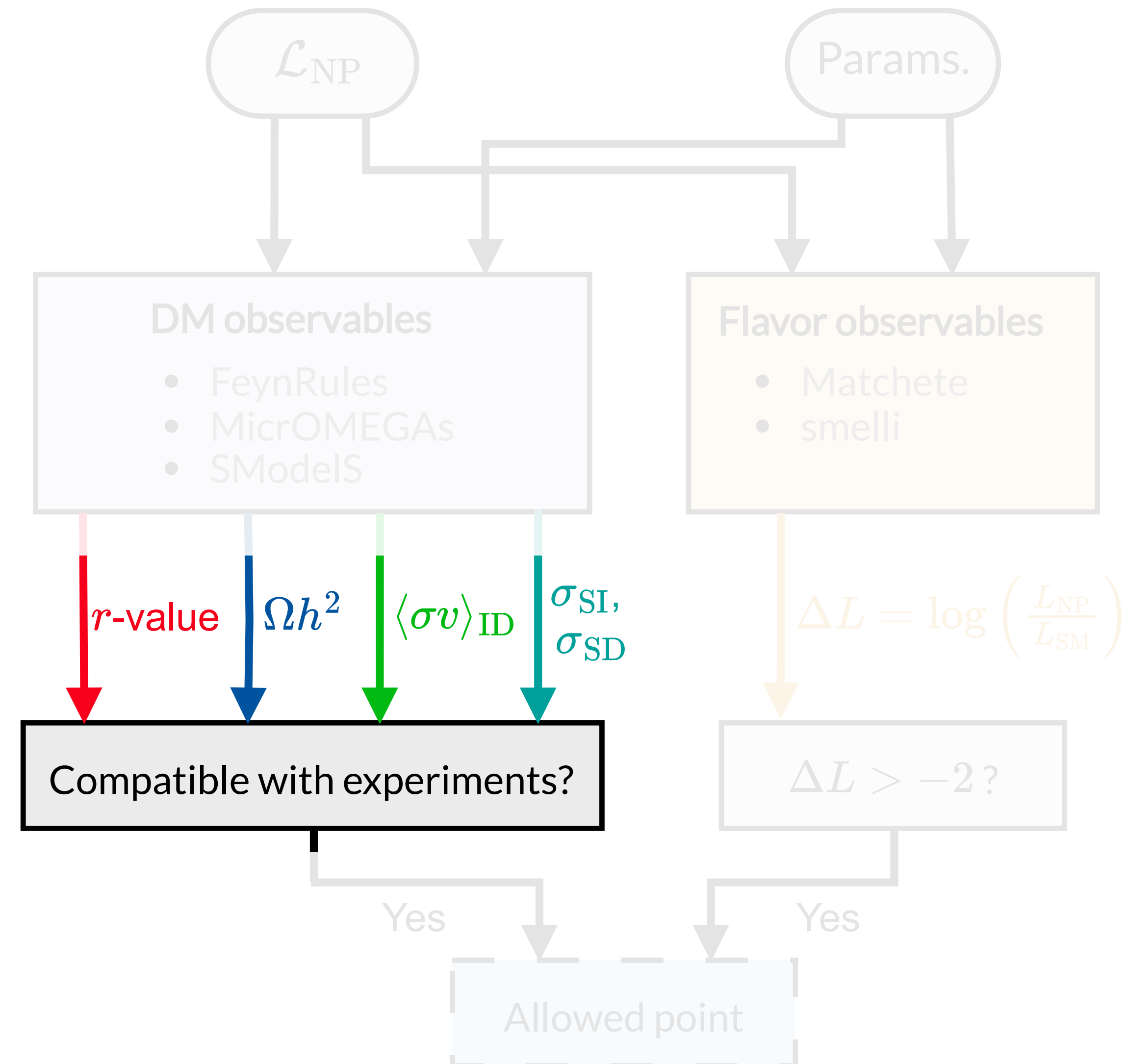
- SModelS computes constraints from colliders, returns

$$r\text{-value} = \frac{\text{signal cross section}}{\text{experimental upper limit}}$$



Tool-Chain - Check Constraints

- Exclude parameter point if one of the following is true:
 - $r\text{-value} \geq 1$
 - Ωh^2 not within $[0.11, 0.13]$
 - $\langle \sigma v \rangle_{\text{ID}} > \text{experimental limit from } \gamma\text{-ray, } \bar{p}, e^+ \text{ and } \nu \text{ data}$
 - σ_{SI} or $\sigma_{\text{SD}} > \text{experimental limits from CRESST-III, LZ, XENON1T, DarkSide-50, PICO-60}$



Tool-Chain - Flavor Observables

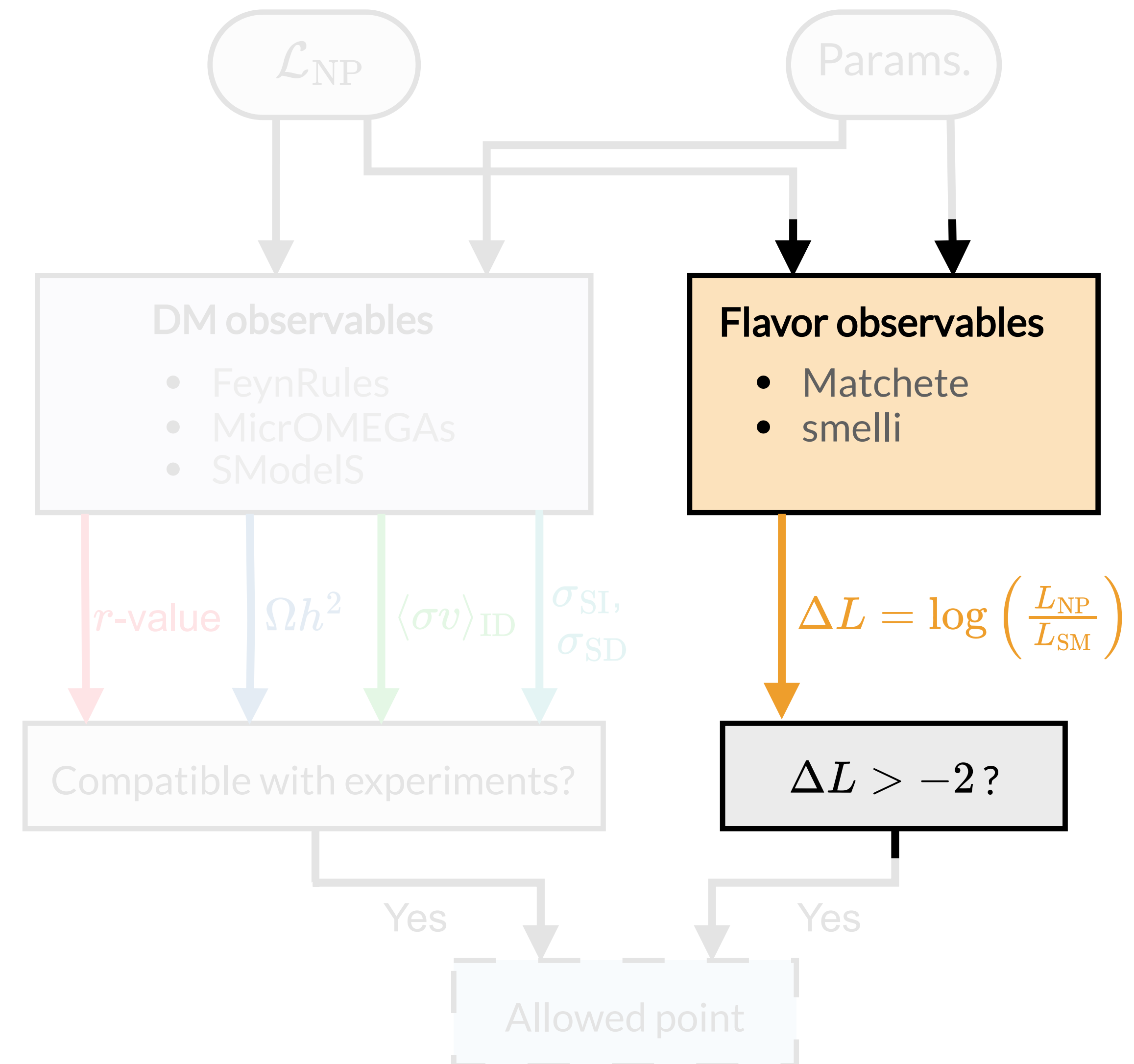
[Fuentes-Martín et al. 2023]

[Aebischer et al. 2019]

- MATCHETE computes **matching of DM model** to SM Effective Field Theory
- SMELLI **evolves SMEFT Wilson coefficients** from high energy scale to low energy scale where flavor observables are measured
- **Computes flavor observables**
- Provides **log-likelihood ratio**

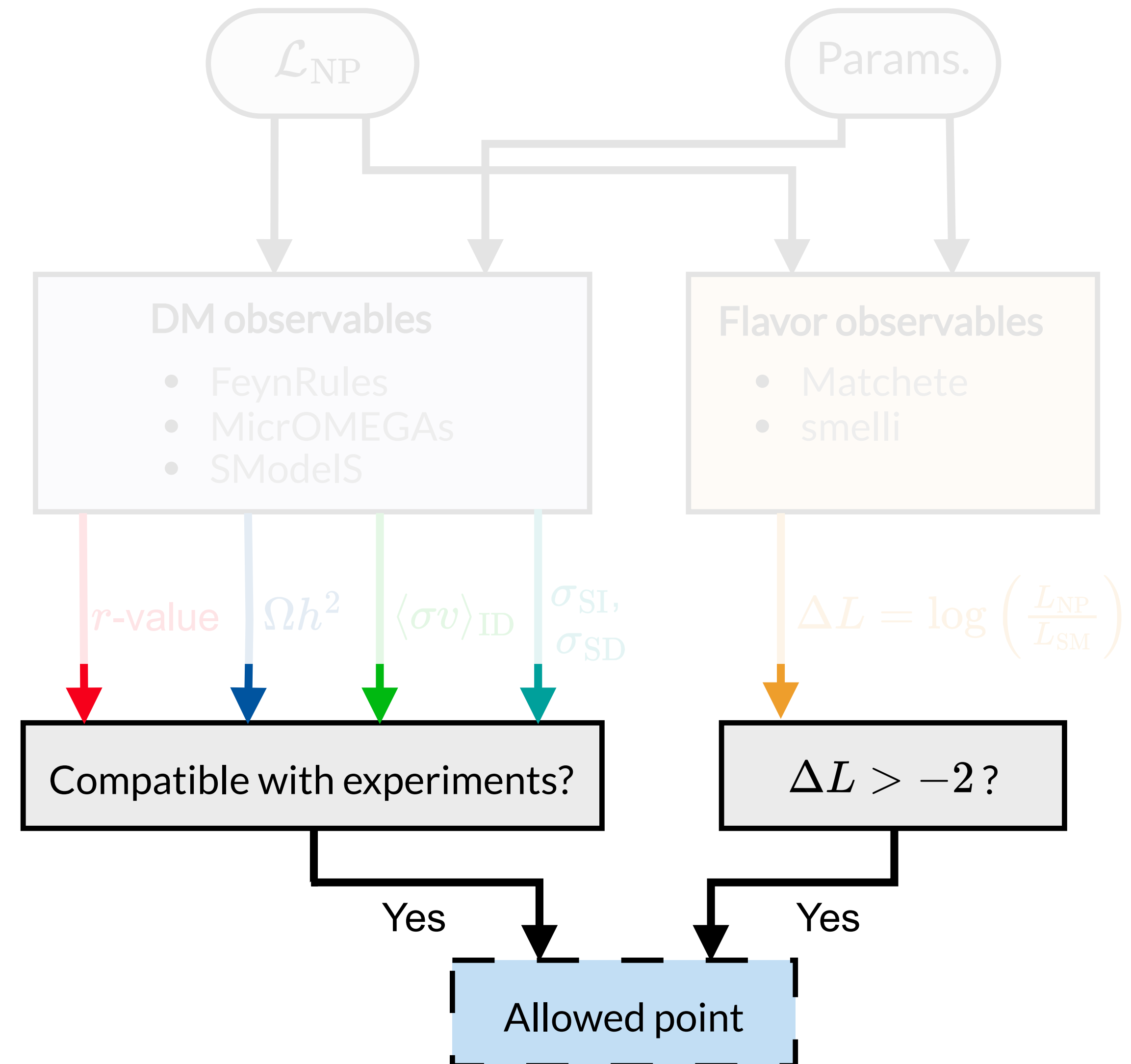
$$\Delta L = \log \left(\frac{L_{\text{NP}}}{L_{\text{SM}}} \right)$$

- Keep point if $\Delta L > -2$



Tool-Chain - Allowed Parameters

- Perform **random parameter scan**
- Keep point if not excluded by any observable
- Obtain **viable parameter space** for any of the flavored DM models



Previously Studied Models

	u_R	d_R	q_L	e_R	ℓ_L
Real scalar DM	×	×	×	×	×
Complex scalar DM	×	×	×	[Acaroglu, Agrawal, Blanke 2022] [Acaroglu, Blanke, Tabet 2022] [Acaroglu, Agrawal, Blanke 2023]	[Acaroglu, Agrawal, Blanke 2022]
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Majorana fermion DM	[Acaroglu, Blanke 2022] [Acaroglu, LR et al. 2024]	×	×	×	×

- Dark matter flavor triplet $\tilde{\chi}$, mediator φ

Majorana DM coupling to e_R

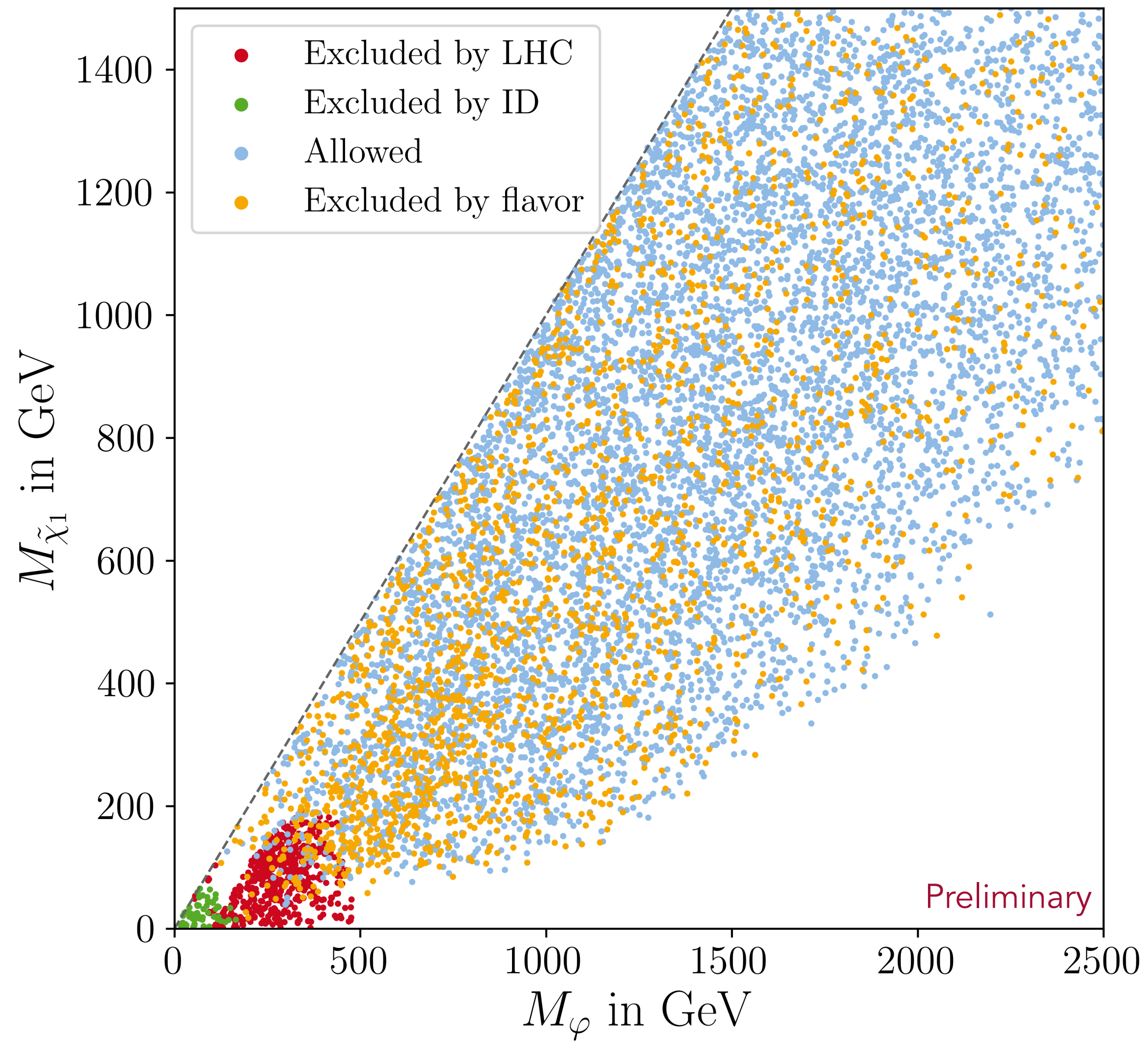
- Dark matter flavor triplet $\tilde{\chi}$, mediator φ
- New physics contribution with flavour and CP violating interaction:

$$\mathcal{L}_{\text{NP}} \supset \frac{1}{2} \bar{\tilde{\chi}} (i \not{\partial} - M_{\tilde{\chi}}) \tilde{\chi} + (D_\mu \varphi)^\dagger (D^\mu \varphi) - m_\varphi^2 \varphi^\dagger \varphi + (\lambda_{ij} \bar{e}_{R,i} \tilde{\chi}_j \varphi + \text{h.c.}) \\ + \lambda_{\varphi\varphi} (\varphi^\dagger \varphi)^2 + \lambda_{\varphi H} (\varphi^\dagger \varphi) (H^\dagger H)$$

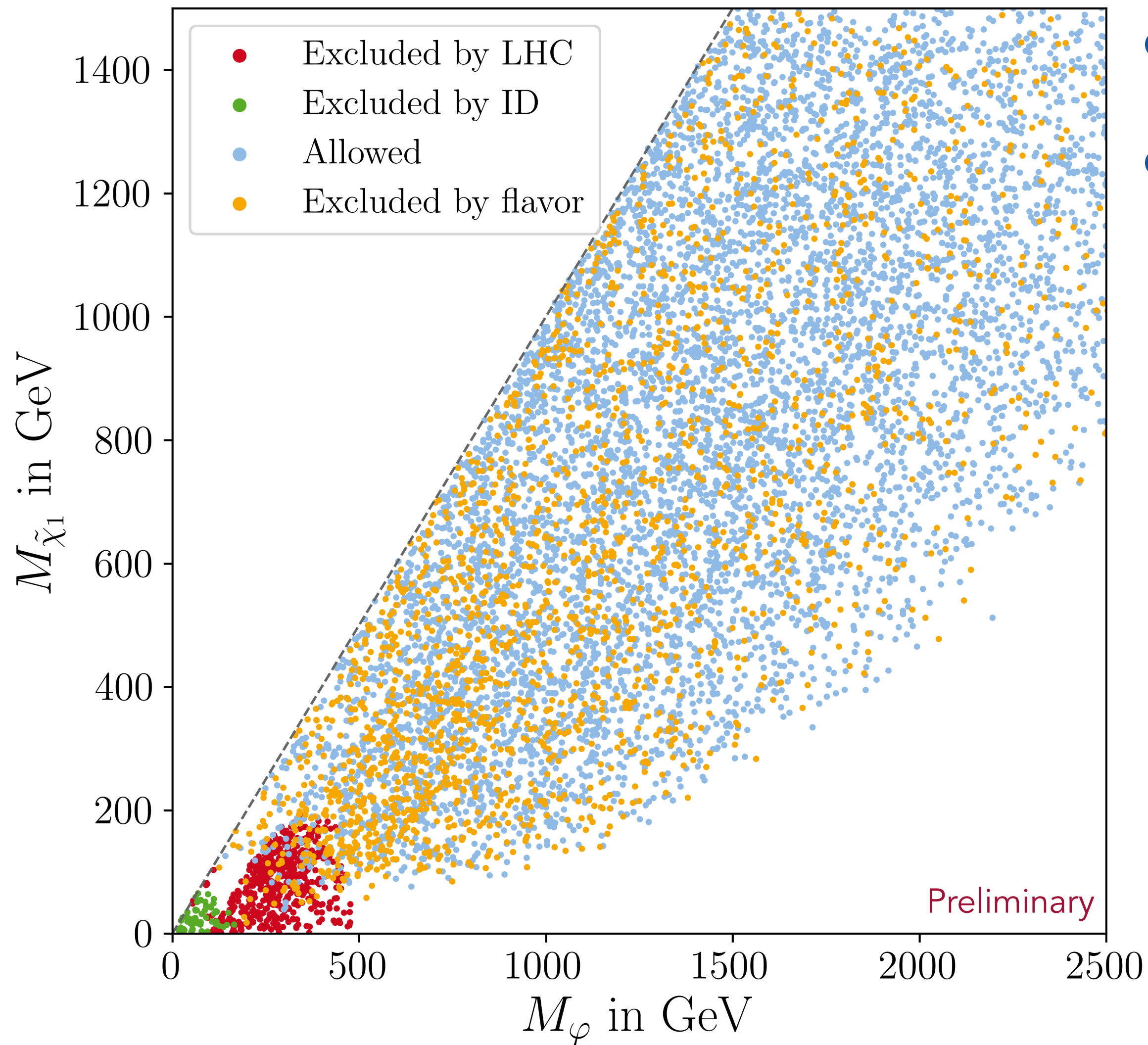
- Mass matrix cannot be generic $\rightarrow$ expand mass matrix in powers of λ :

$$M_{\tilde{\chi}} = m_{\tilde{\chi}} \left[\mathbf{1} + \frac{\eta}{2} (\lambda^\dagger \lambda + \lambda^T \lambda^*) + \mathcal{O}(\lambda^4) \right] \xrightarrow{\text{Diagonalize}} M_{\tilde{\chi}}^D = \text{diag}(M_{\tilde{\chi}_1}, M_{\tilde{\chi}_2}, M_{\tilde{\chi}_3})$$

Majorana DM coupling to e_R

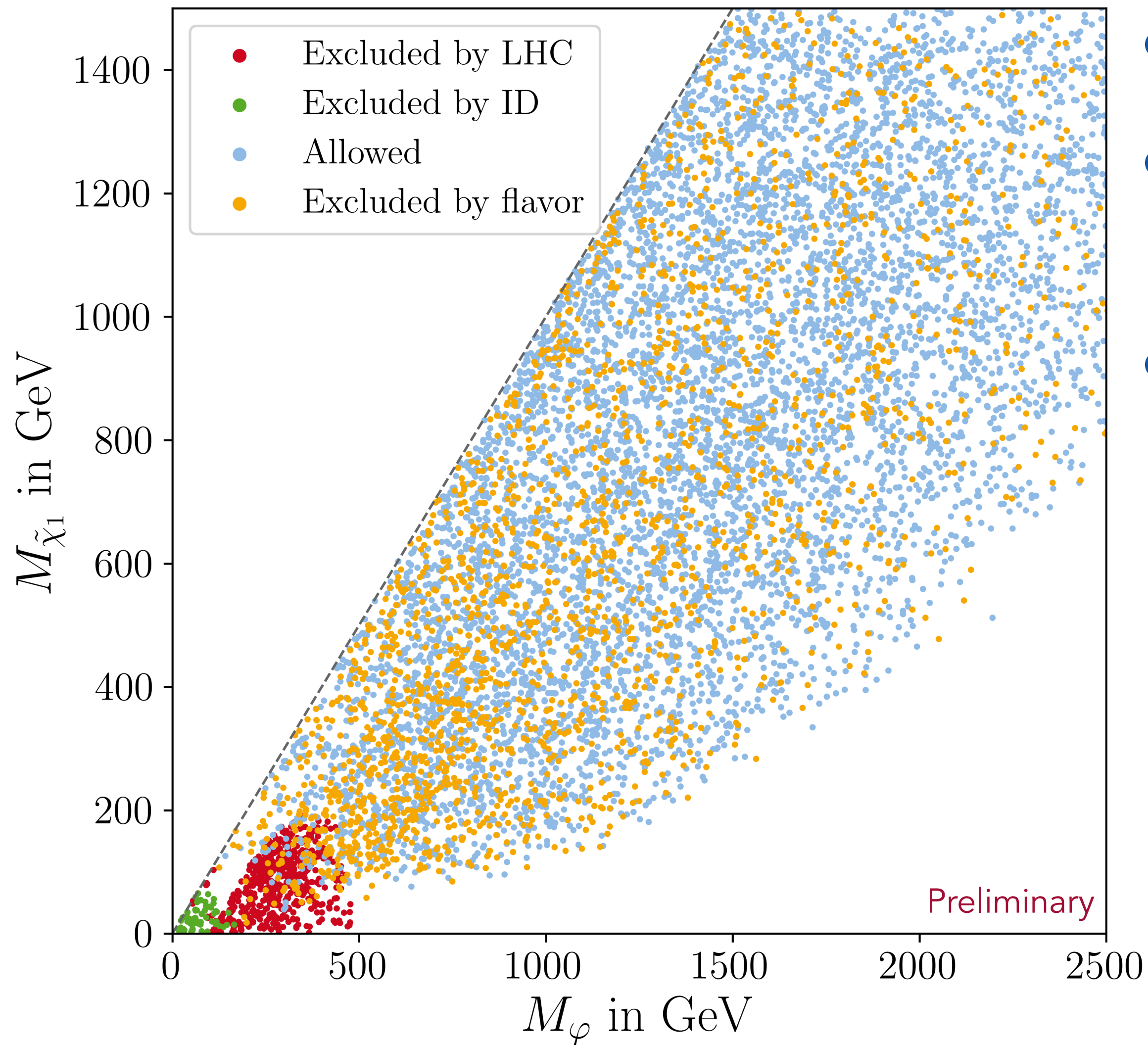


Majorana DM coupling to e_R



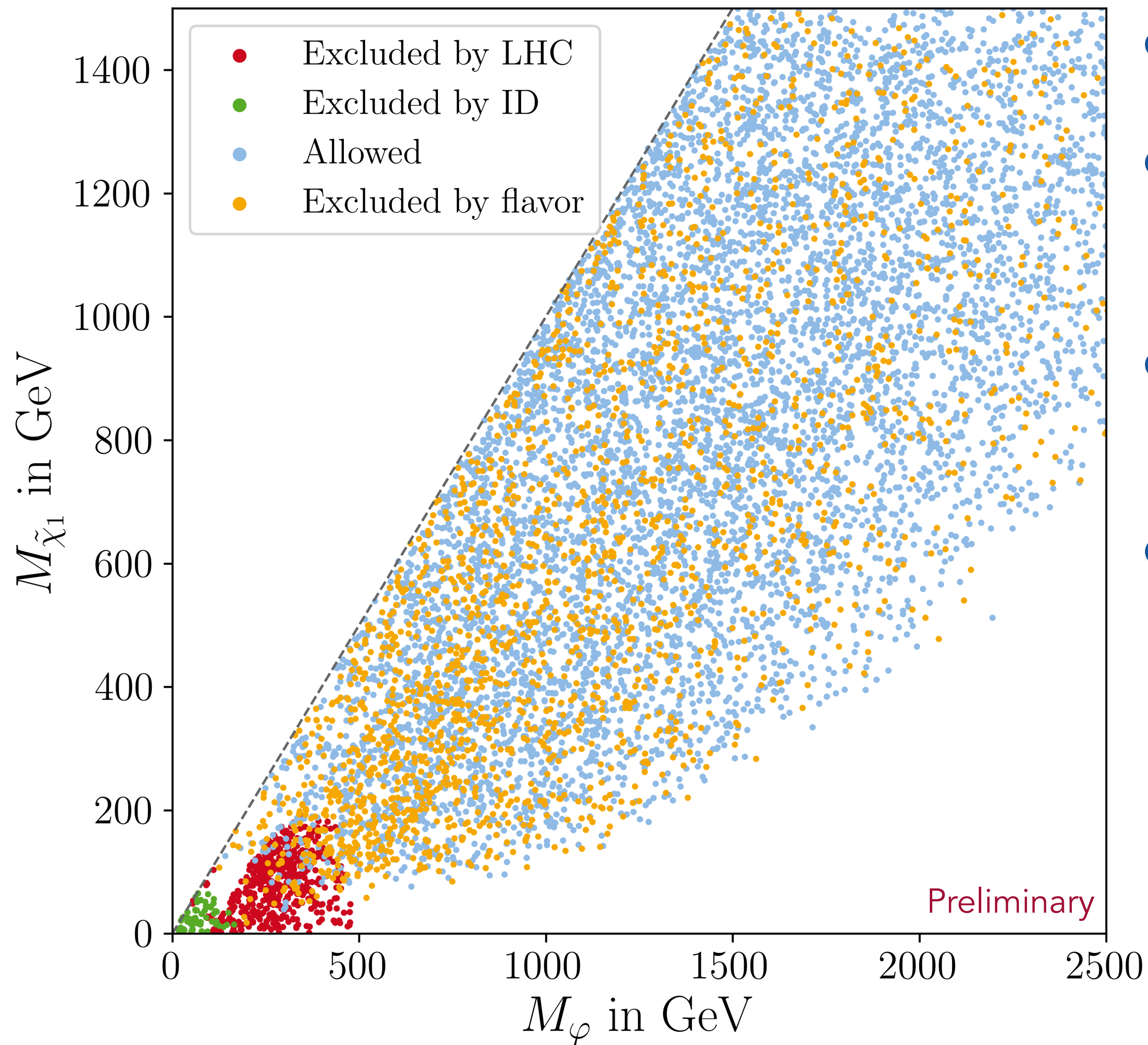
- Relic abundance most constraining
- Lower bound on DM mass because of upper limit on coupling

Majorana DM coupling to e_R

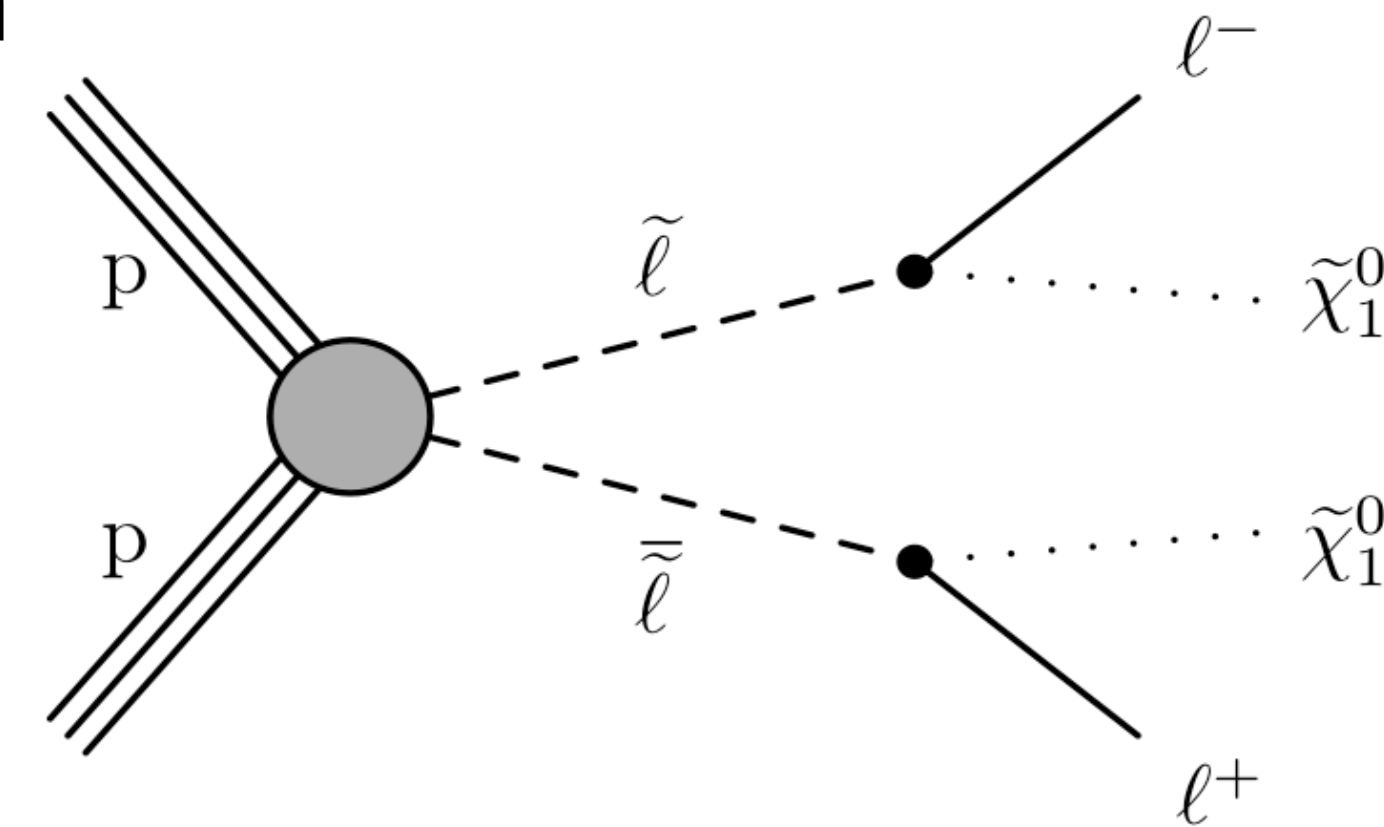


- Relic abundance most constraining
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- ID constrains weak because of p -wave suppression of annihilation

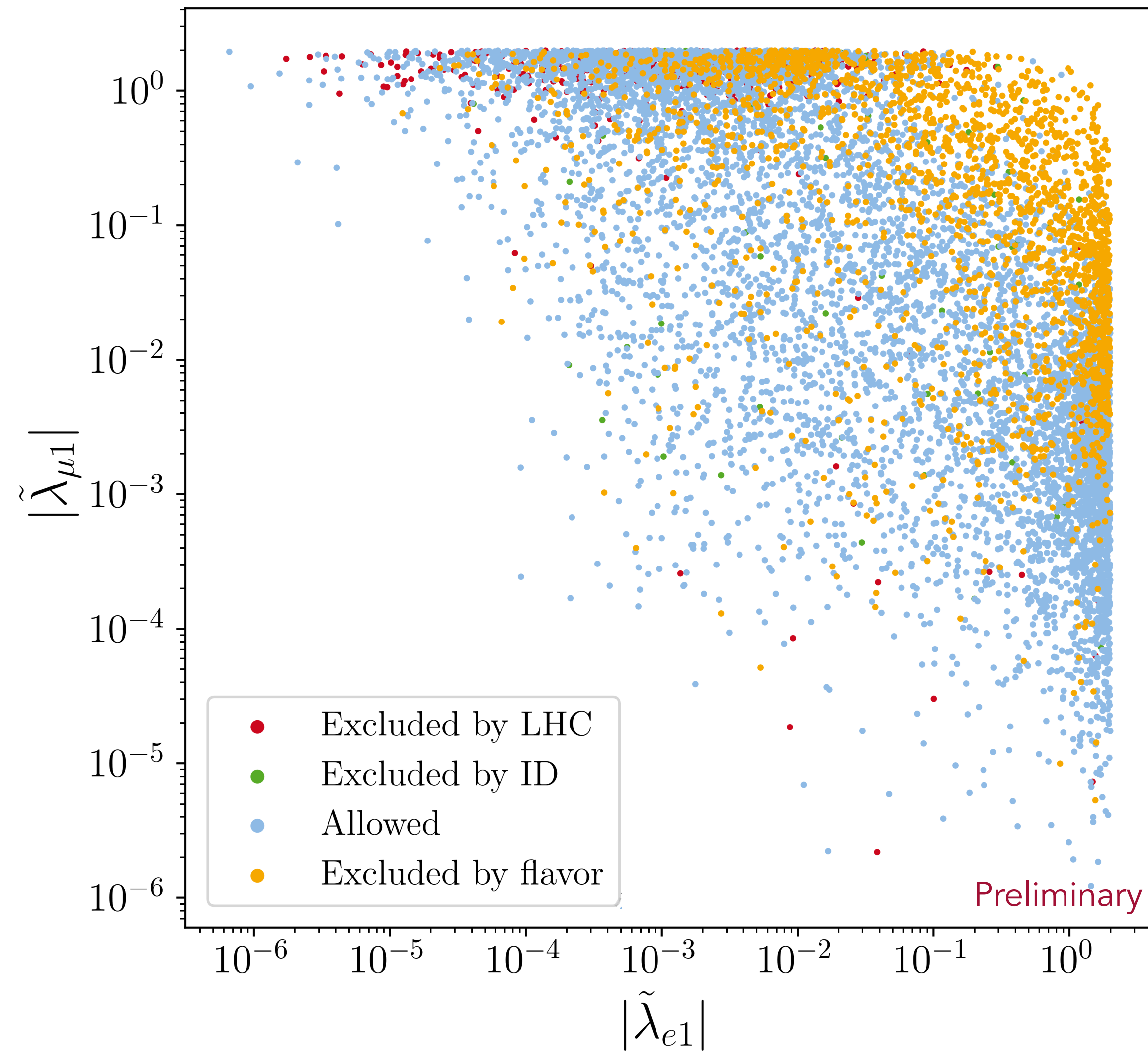
Majorana DM coupling to e_R



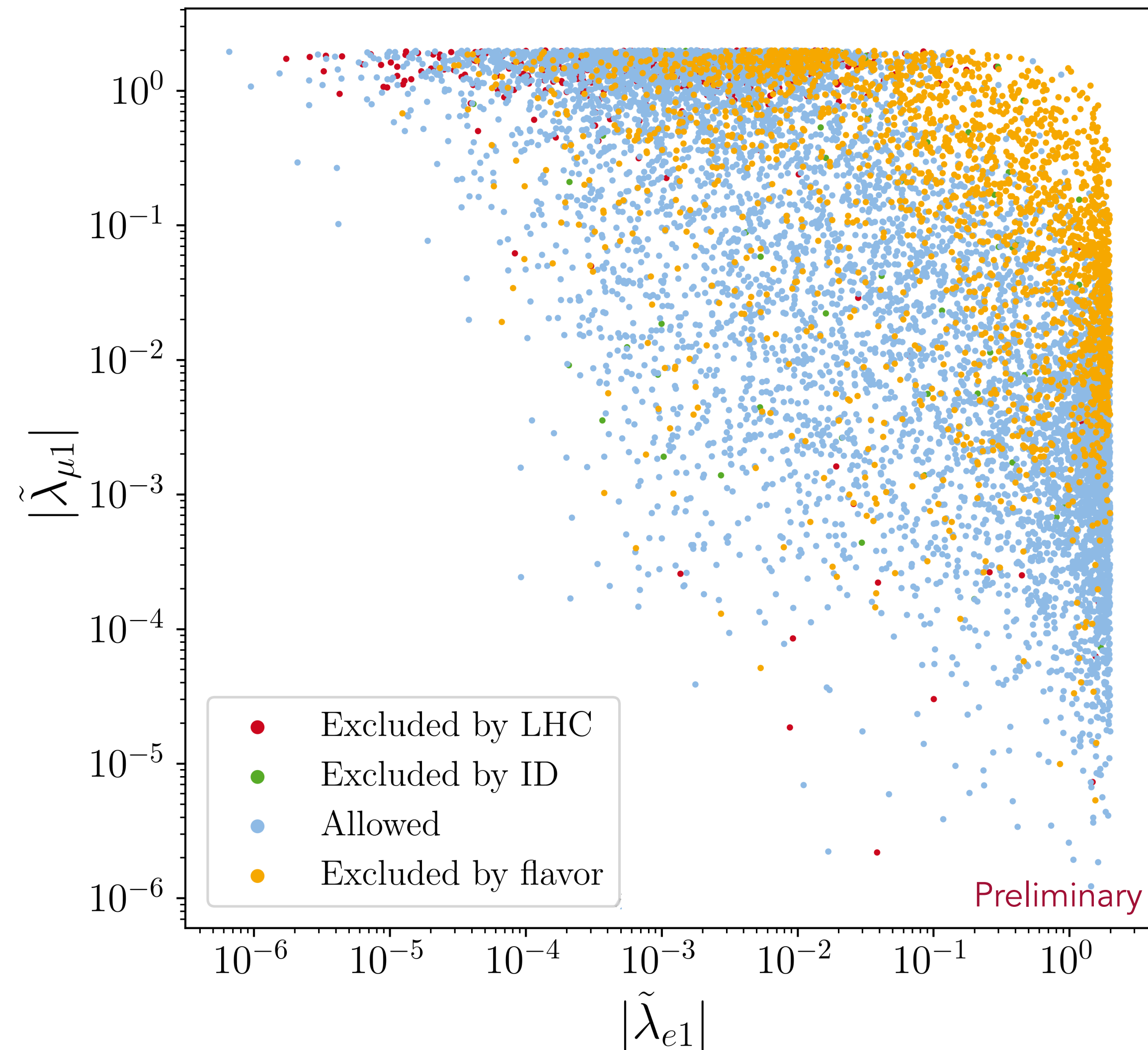
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- Drell-Yan production of mediator pair decaying into leptons + $\cancel{E}_T$ is dominant LHC signal



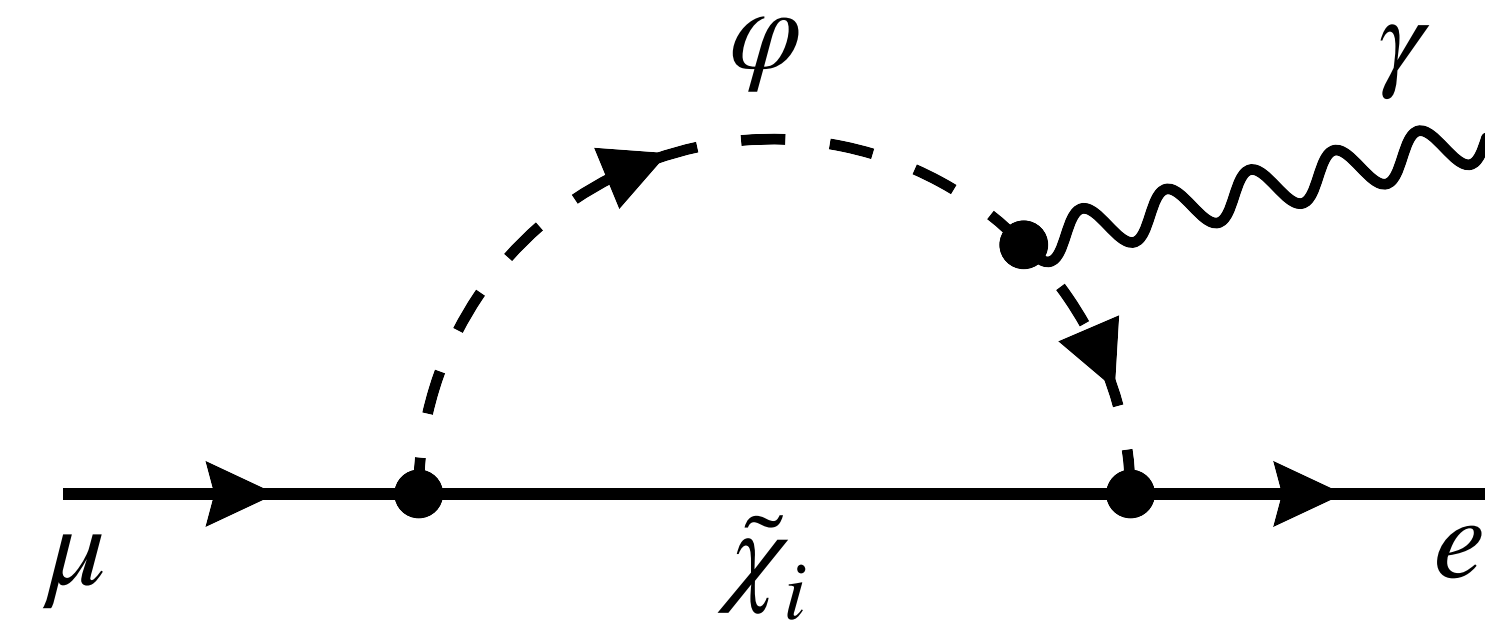
Majorana DM coupling to e_R



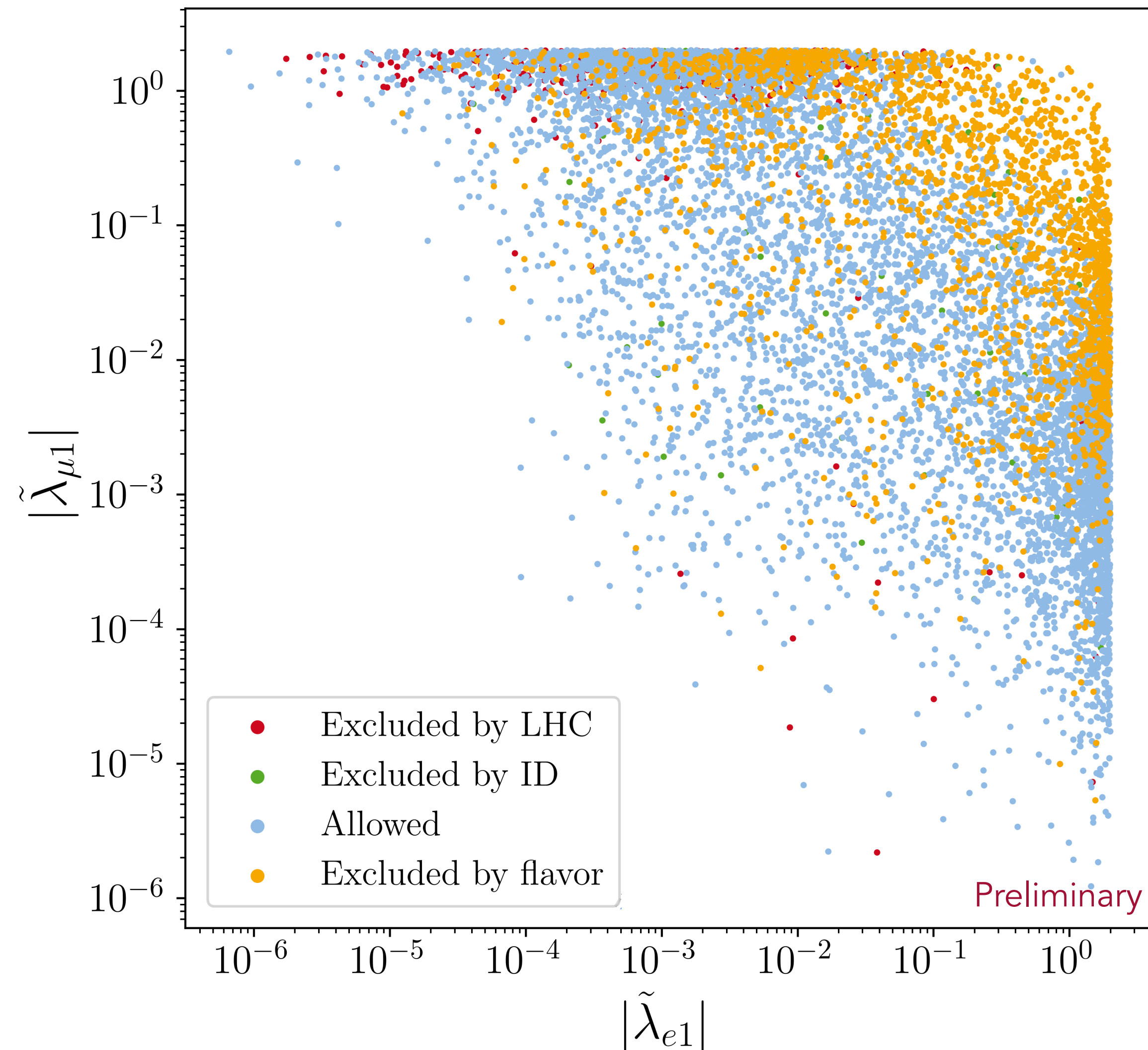
Majorana DM coupling to e_R



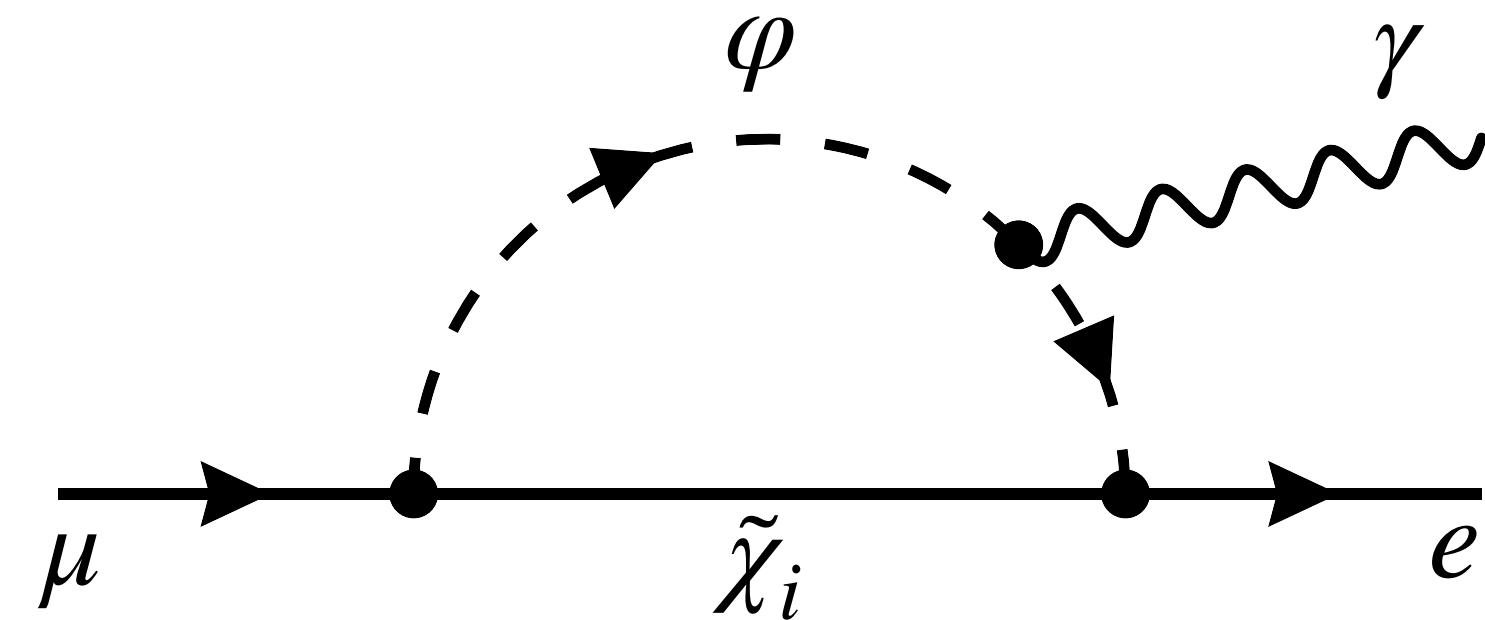
- Strong **flavor constraints** from lepton-flavor violating decays



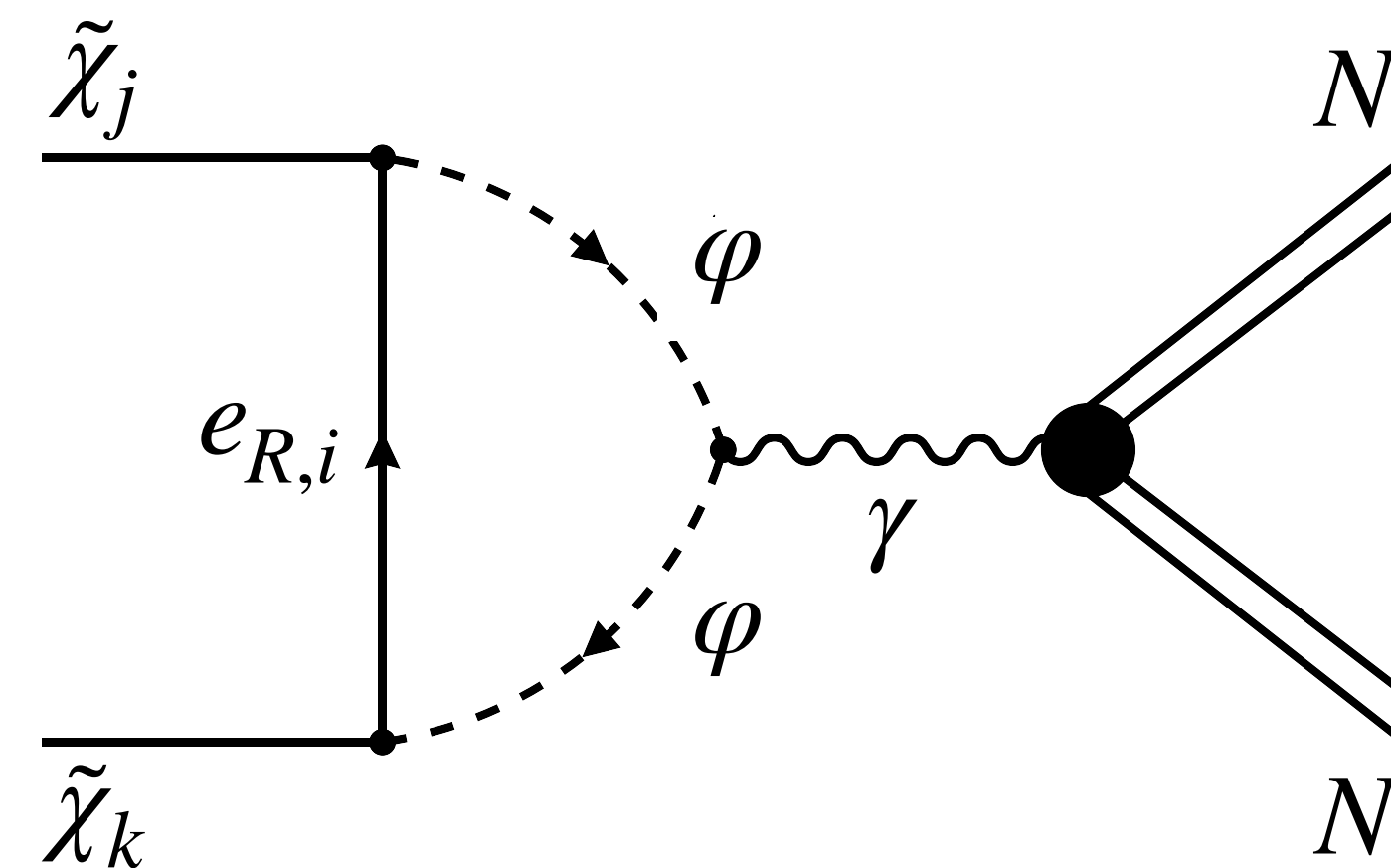
Majorana DM coupling to e_R



- Strong **flavor constraints** from lepton-flavor violating decays



- DD constraints absent as only possible contribution is very suppressed



Conclusions



- Flavored DM models have rich phenomenology but many models have not been studied
- Developed **framework to study different flavored DM models** including constraints from relic density, direct & indirect detection, collider and flavor observable
- Showed **constraints on Majorana DM coupling to e_R**
- **Strongest constraints** from **correct relic abundance** and **lepton-flavor violating decays**, collider and ID constraints only relevant for small DM masses
- Outlook: Study other missing models



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- Outlook: Study other missing models



Thank you!



Backup Slides

Parameterization of λ

$$\lambda = U \text{diag}(D_1, D_2, D_3) O d$$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23}^\theta & s_{23}^\theta e^{-i\delta_{23}} \\ 0 & -s_{23}^\theta e^{i\delta_{23}} & c_{23}^\theta \end{pmatrix} \begin{pmatrix} c_{13}^\theta & 0 & s_{13}^\theta e^{-i\delta_{13}} \\ 0 & 1 & 0 \\ -s_{13}^\theta e^{i\delta_{13}} & 0 & c_{13}^\theta \end{pmatrix} \begin{pmatrix} c_{12}^\theta & s_{12}^\theta e^{-i\delta_{12}} & 0 \\ -s_{12}^\theta e^{i\delta_{12}} & c_{12}^\theta & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$O = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23}^\phi & s_{23}^\phi \\ 0 & -s_{23}^\phi & c_{23}^\phi \end{pmatrix} \begin{pmatrix} c_{13}^\phi & 0 & s_{13}^\phi \\ 0 & 1 & 0 \\ -s_{13}^\phi & 0 & c_{13}^\phi \end{pmatrix} \begin{pmatrix} c_{12}^\phi & s_{12}^\phi & 0 \\ -s_{12}^\phi & c_{12}^\phi & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$d = \text{diag}(e^{i\gamma_1}, e^{i\gamma_2}, e^{i\gamma_3})$$

Sampling of Parameters

- D_i in $[0,2]$, linearly
- θ_{12} and ϕ_{12} in $[10^{-6}, 10^{-2}]$, logarithmically
- Remaining angles in $[10^{-4}, \pi/4]$, logarithmically.
- δ_{ij} and γ_i in $[0, 2\pi)$, linearly
- $M_{\tilde{\chi}}$ in $[100, 1500]$ GeV, linearly
- M_{φ} in $[M_{\tilde{\chi}_1}, M_{\tilde{\chi}_1} + 2000]$ GeV, linearly
- $\lambda_{\phi H}$ and η in $\pm[10^{-3}, 1]$, logarithmically

Indirect detection

Final state	Gamma rays	Antiprotons	Positrons	Neutrinos
e^+e^-	[50, 51]	—	[52]	—
$\mu^+\mu^-$	[53]	—	[52]	[53]
$\tau^+\tau^-$	[54, 55]	—	[52]	—
$q\bar{q}$ (light)	—	[56]	—	—
$c\bar{c}$	[50]	—	—	—
$b\bar{b}$	[57]	[56]	—	—
$t\bar{t}$	[50, 58]	—	—	—
W^+W^-	[53, 59]	[56]	—	—
ZZ	[50, 60]	—	—	—
hh	—	[56]	—	—
$\nu\bar{\nu}$ (any flavor)	—	—	—	[61, 62]

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[58] H. Abdallah, A. Abramowski, F. Aharonian, F. Ait Benkhali, A. Akhperjanian, et al., *Search for dark matter annihilations towards the inner galactic halo from 10 years of observations with h.e.s.s.*, *Physical Review Letters* **117** (Sept., 2016).

[59] M. Ackermann, A. Albert, B. Anderson, W. Atwood, L. Baldini, et al., *Searching for dark matter annihilation from milky way dwarf spheroidal galaxies with six years of fermi large area telescope data*, *Physical Review Letters* **115** (Nov., 2015).

[60] J. Aleksić, S. Ansoldi, L. Antonelli, P. Antoranz, A. Babic, et al., *Optimized dark matter searches in deep observations of segue 1 with magic*, *Journal of Cosmology and Astroparticle Physics* **2014** (Feb., 2014) 008–008.

[61] T. I. Collaboration, R. Abbasi, M. Ackermann, J. Adams, S. K. Agarwalla, et al., *Search for neutrino lines from dark matter annihilation and decay with icecube*, 2023.

[62] A. Albert, M. André, M. Anghinolfi, G. Anton, M. Ardid, et al., *Search for dark matter towards the galactic centre with 11 years of antares data*, *Physics Letters B* **805** (June, 2020) 135439.

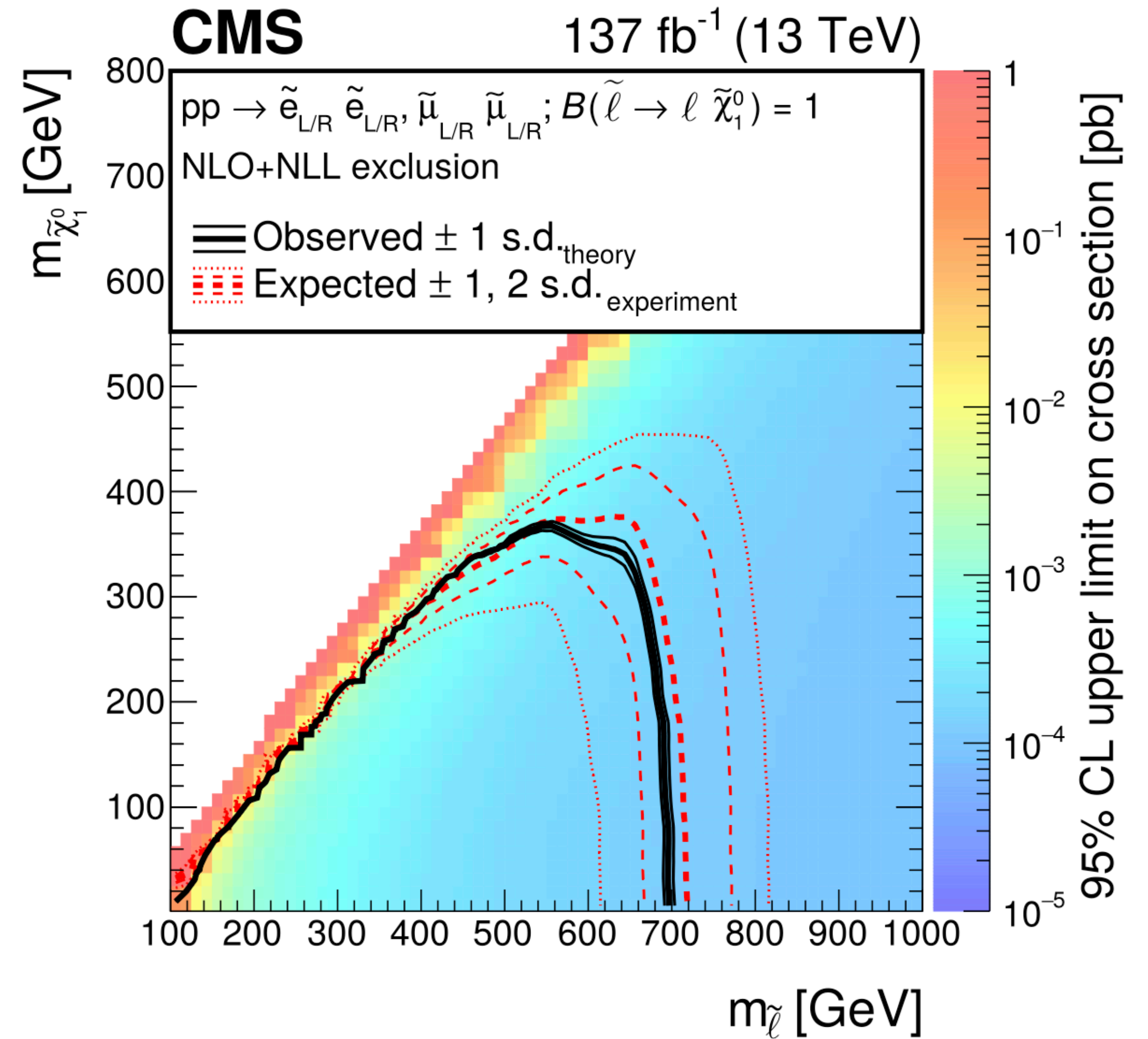
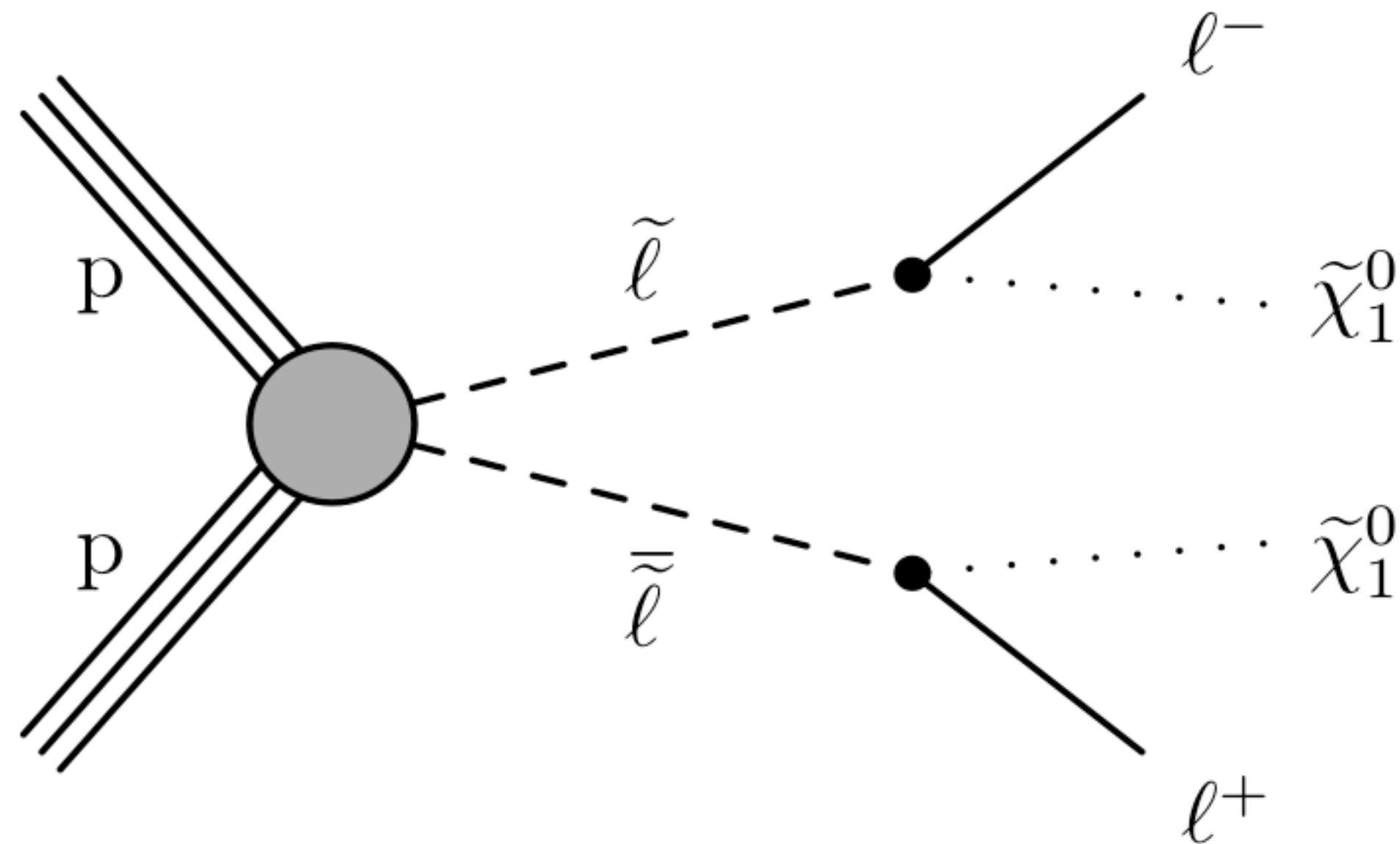
Treatment of limit for mixes final states AB :

$$\langle\sigma v\rangle_{AB}=\frac{1}{2}\left(\langle\sigma v\rangle_{AA}+\langle\sigma v\rangle_{BB}\right)$$

Collider constraints

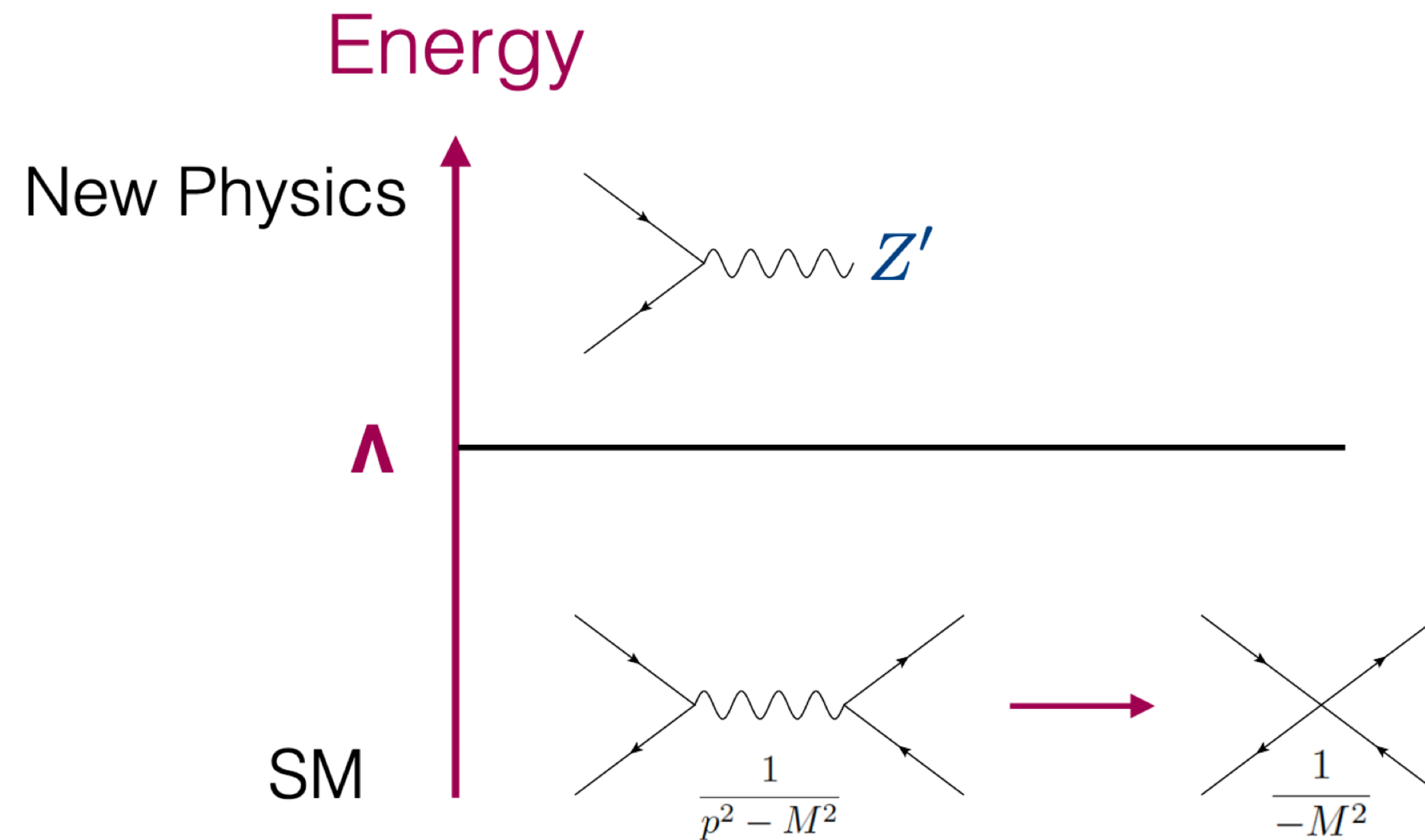
[Sirunyan et al. 2021]

- Most sensitive CMS search



SMEFT

- SM Effective Field Theory $\mathcal{L}_{\text{SMEFT}} = \mathcal{L}_{\text{SM}} + \sum_i \frac{C_i}{\Lambda} \mathcal{O}_i^{d=5} + \sum_i \frac{C_i}{\Lambda^2} \mathcal{O}_i^{d=6}$
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Example: Majorana DM coupling to e_R

