

Opportunities with Calorimeters at LHC : How to solve the muon puzzle ?

Tanguy Pierog

Karlsruhe Institute of Technology, Institut für Kernphysik,
Karlsruhe, Germany



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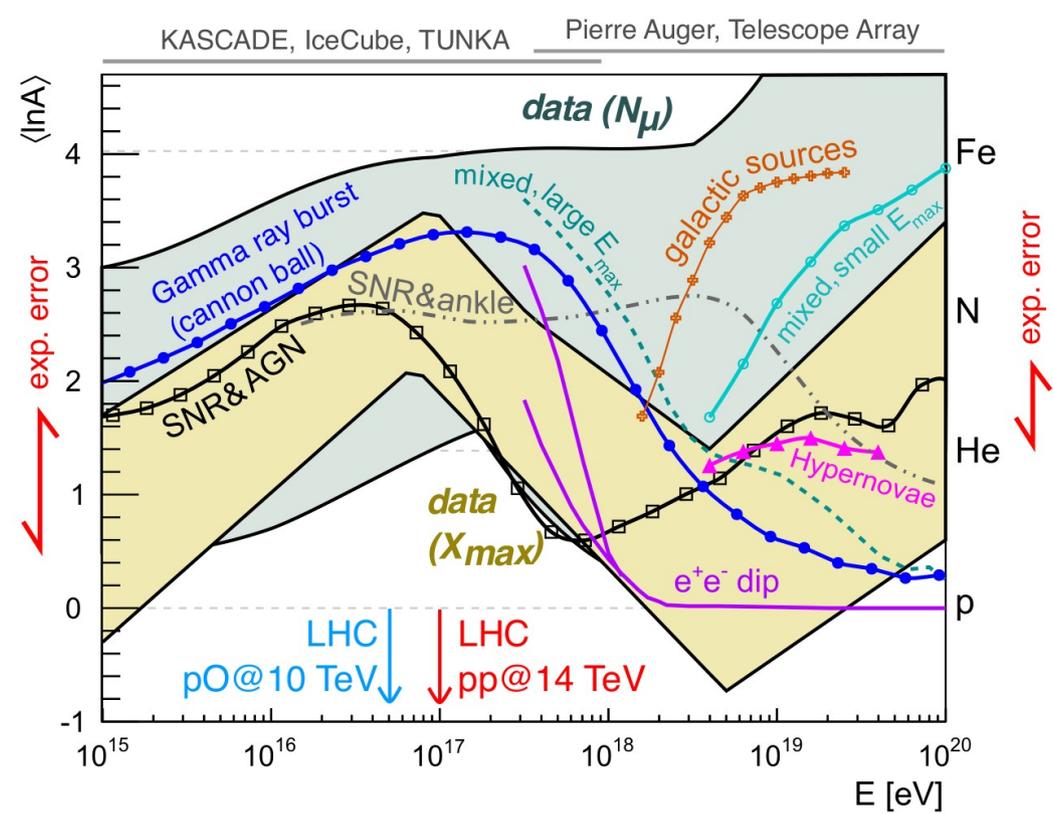
Outline

- **Muons in Extensive Air Showers (EAS)**
- **0 degree calorimeters**
 - ➔ Pion exchange
- **Muon puzzle**
 - ➔ Quark Gluon Plasma (QGP) as possible explanation
- **Forward calorimeters**
 - ➔ Test of hadronization scheme
- **Summary**

UHECR Composition

With muons current CR data are impossible to interpret

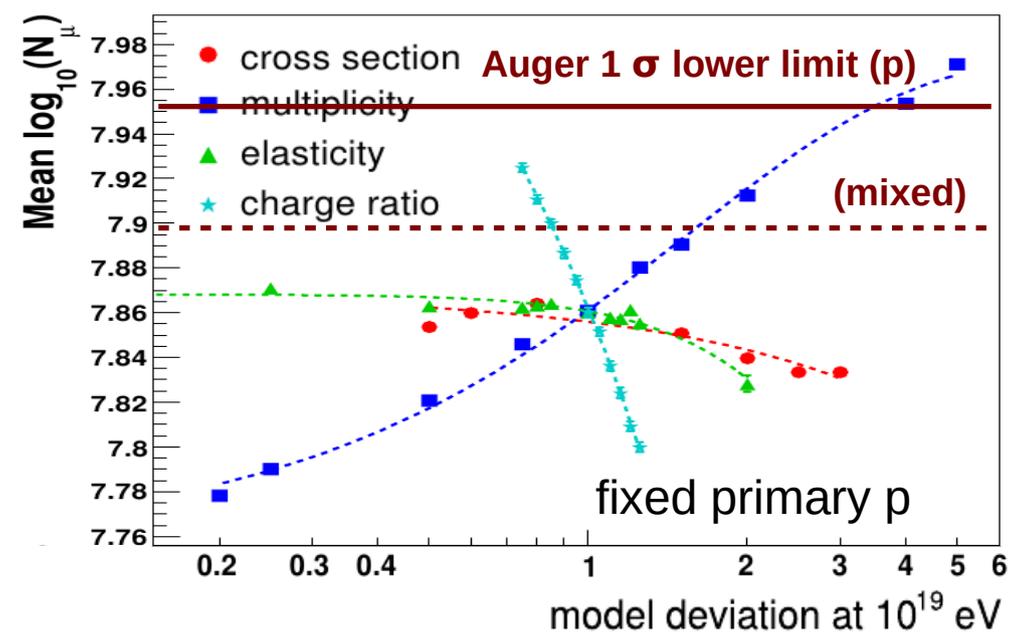
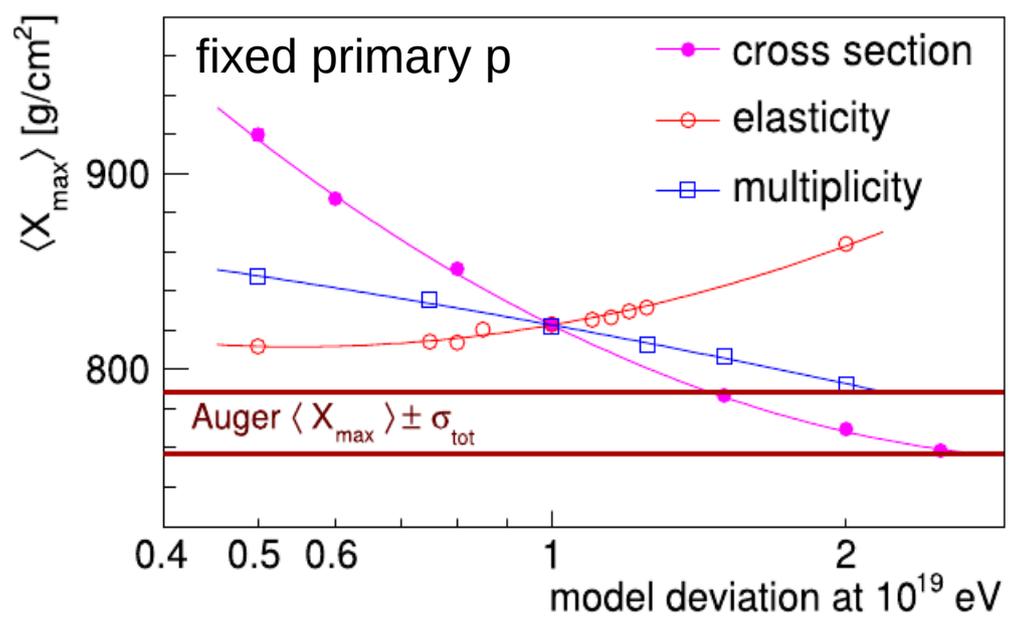
- ➔ Very large uncertainties in model predictions
- ➔ Mass from muon data incompatible with mass from X_{max}



Based on Kampert & Unger, Astropart. Phys. 35 (2012) 660

H. Dembinski UHECR 2018 (WHISP working group)

Sensitivity to Hadronic Interactions

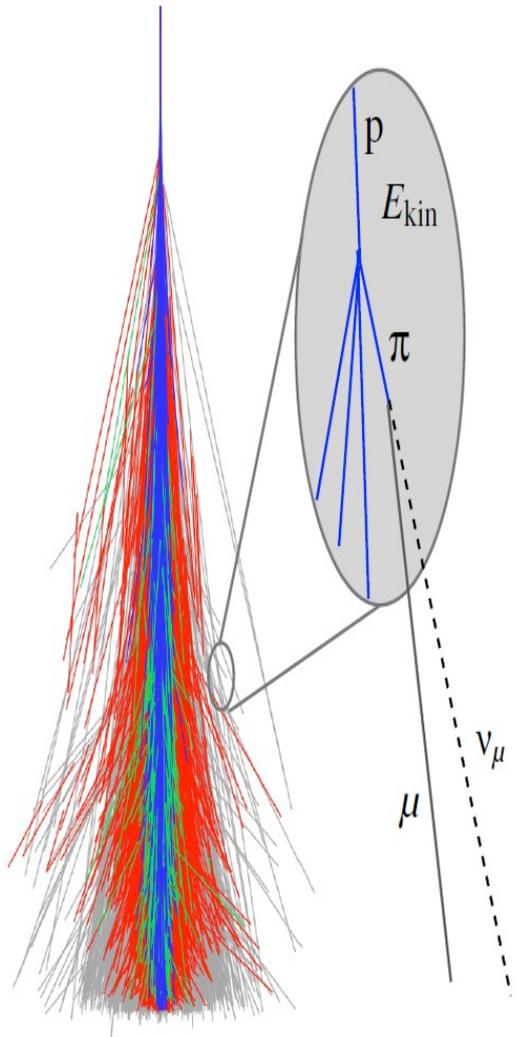


- Air shower development dominated by few parameters
 - ➔ mass and energy of primary CR
 - ➔ cross-sections (p-Air and (π -K)-Air)
 - ➔ (in)elasticity
 - ➔ multiplicity
 - ➔ charge ratio and baryon production
- Change of primary = change of hadronic interaction parameters
 - ➔ cross-section, elasticity, mult. ...

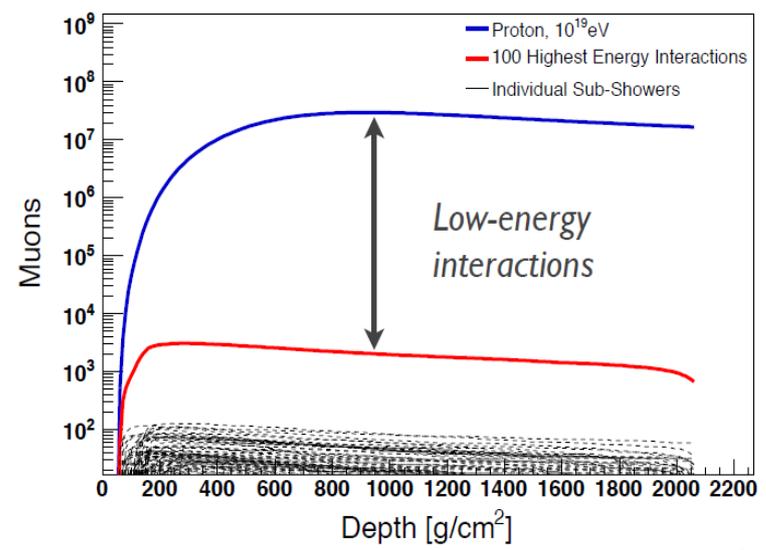
With unknown mass composition hadronic interactions can only be tested using various observables which should give consistent mass results

From R. Ulrich (KIT)

Muon production by low energy interactions

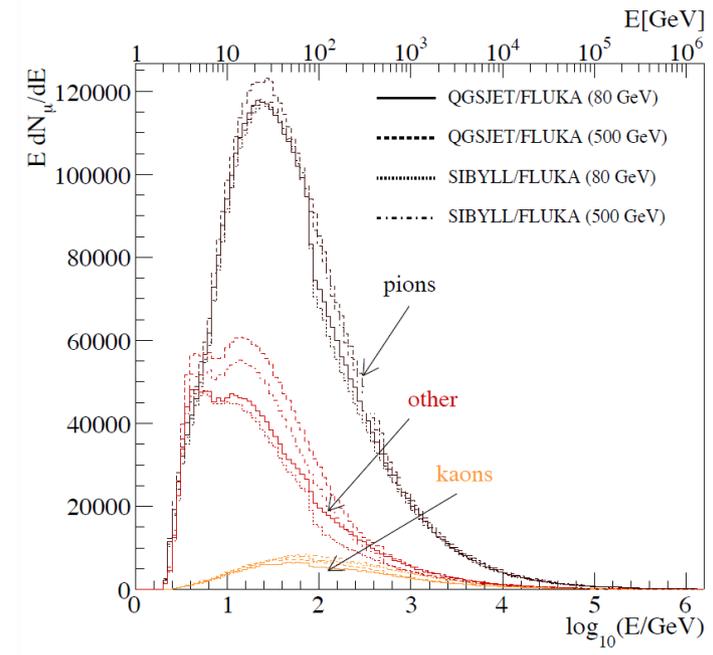
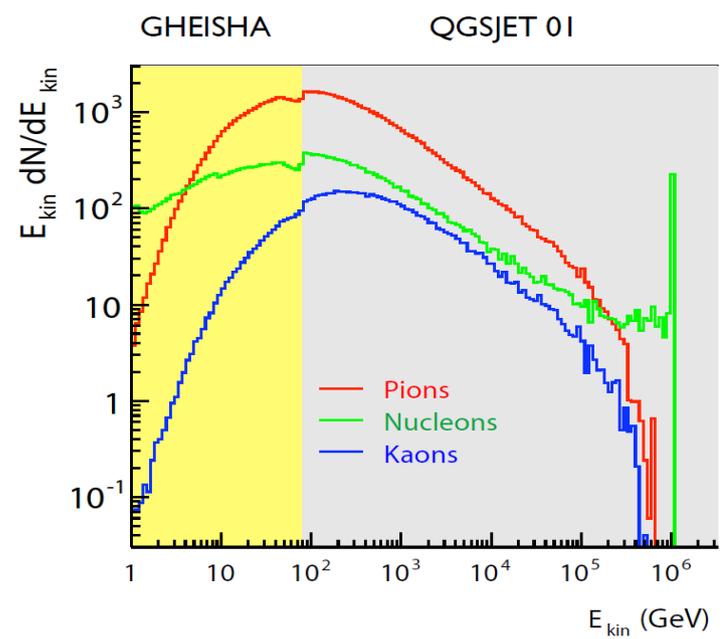


Muons



~ 100 GeV for KASCADE
 ~ 30 GeV for Auger

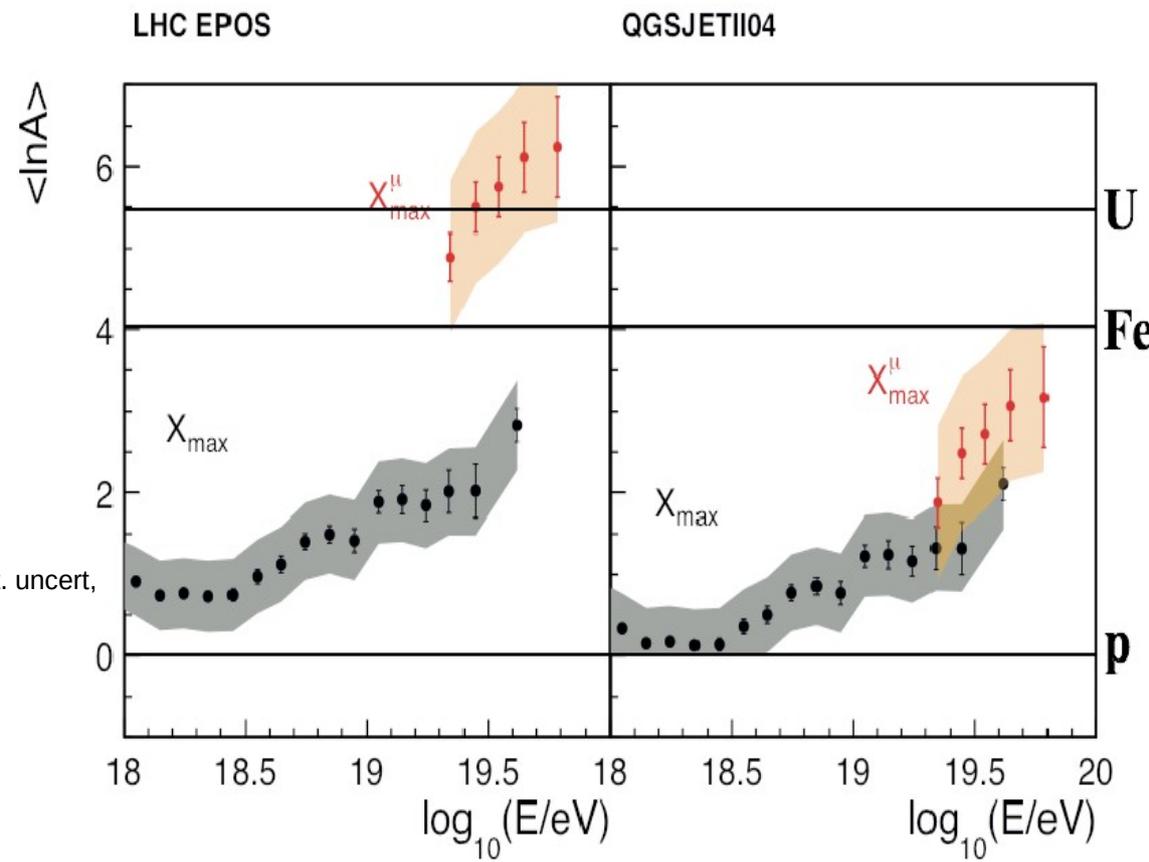
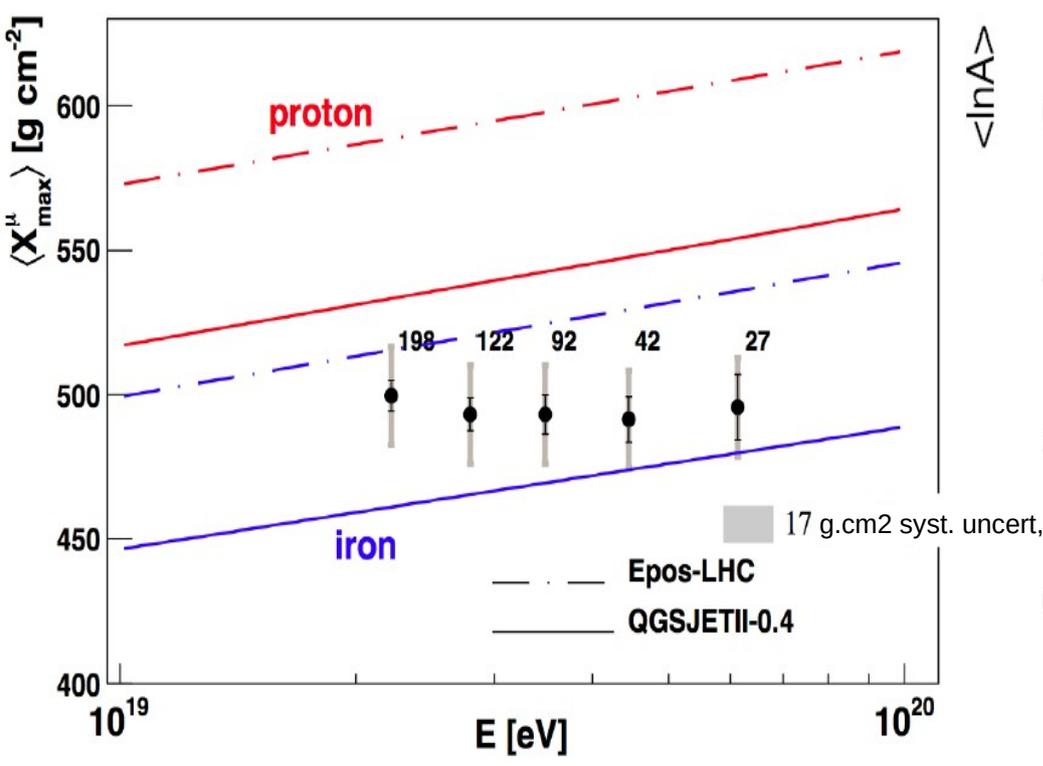
Muons/hadrons:
 low-energy interactions



Muon Production Depth X_{\max}^{μ} and X_{\max}

● 2 independent mass composition measurements

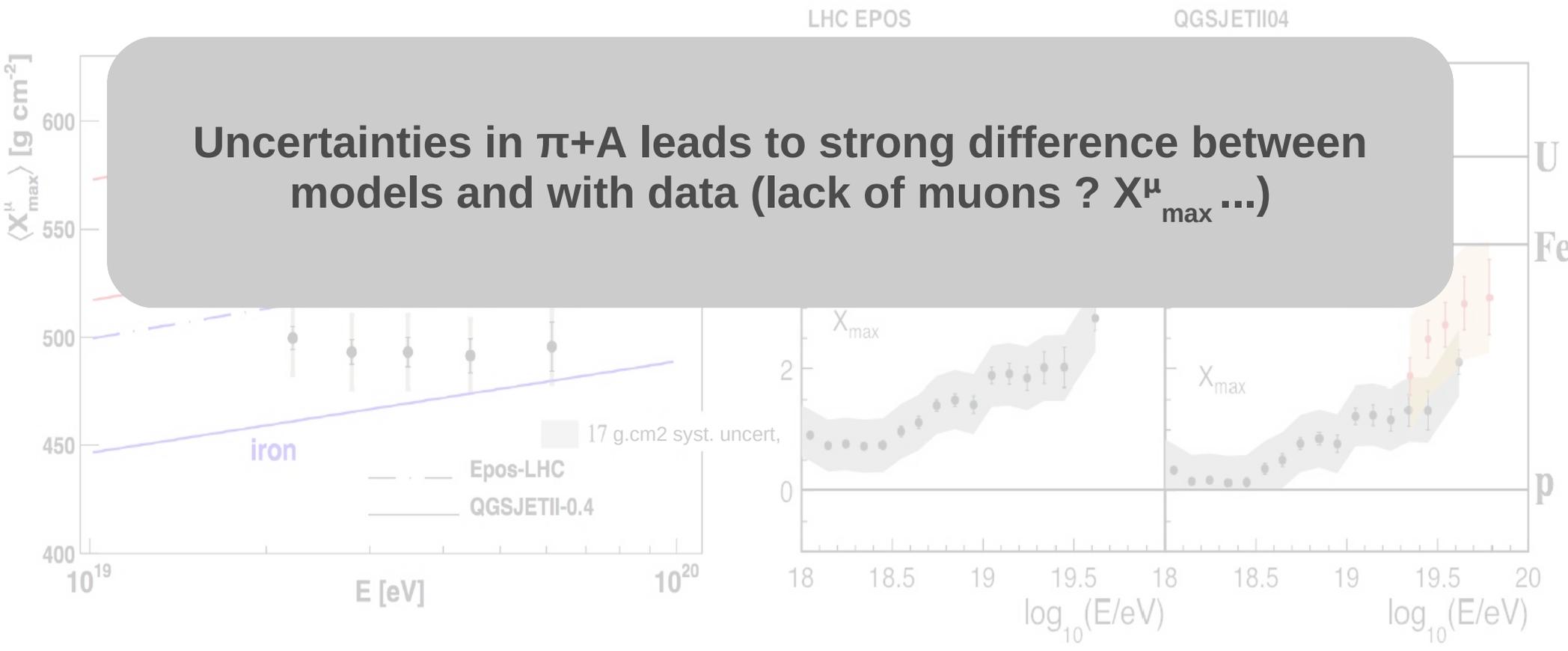
- ➔ both results should be between p and Fe
- ➔ both results should give the same mean logarithmic mass for the same model
- ➔ problem with EPOS appears after corrections motivated by LHC data (diffraction and forward baryon production) ➔ **related to pion interaction (from low to high energy)**



Muon Production Depth X_{max}^μ and X_{max}

2 independent mass composition measurements

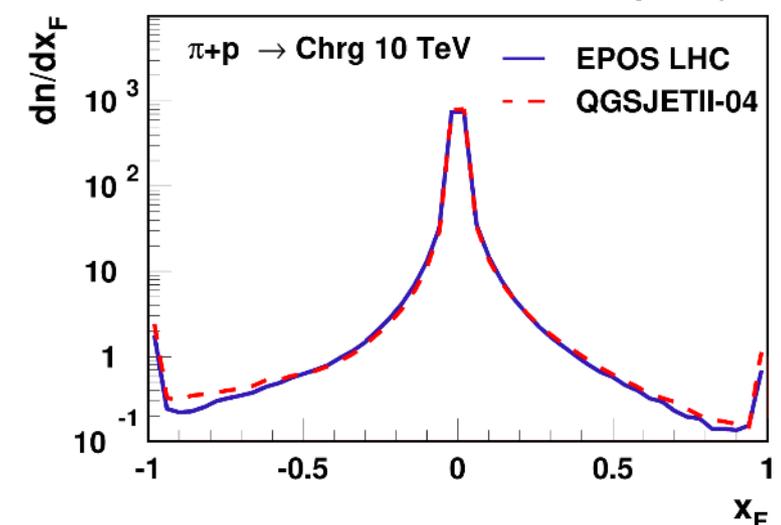
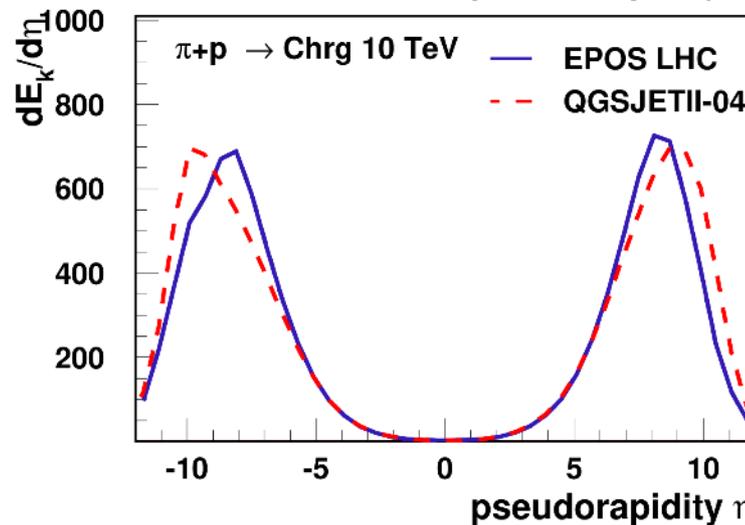
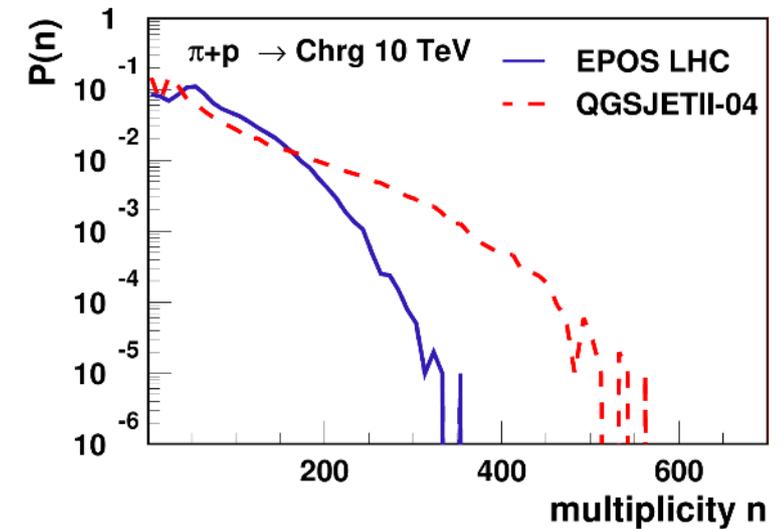
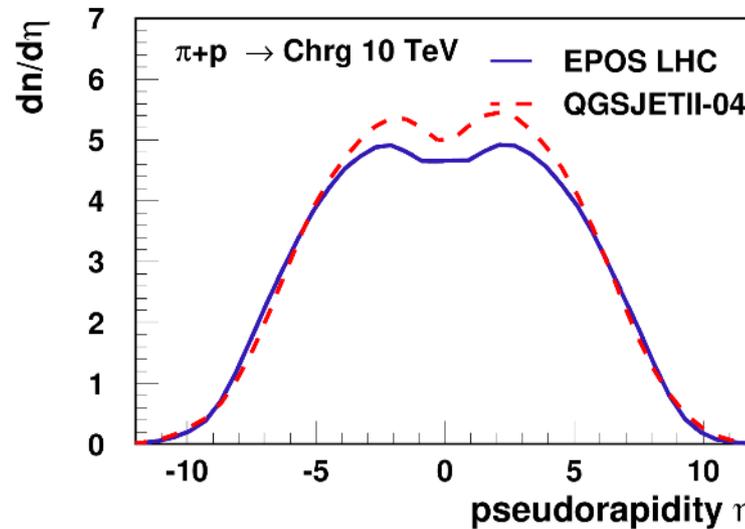
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Model Predictions $\pi+p$ @ LHC

Models well constraint by LHC run I for pp

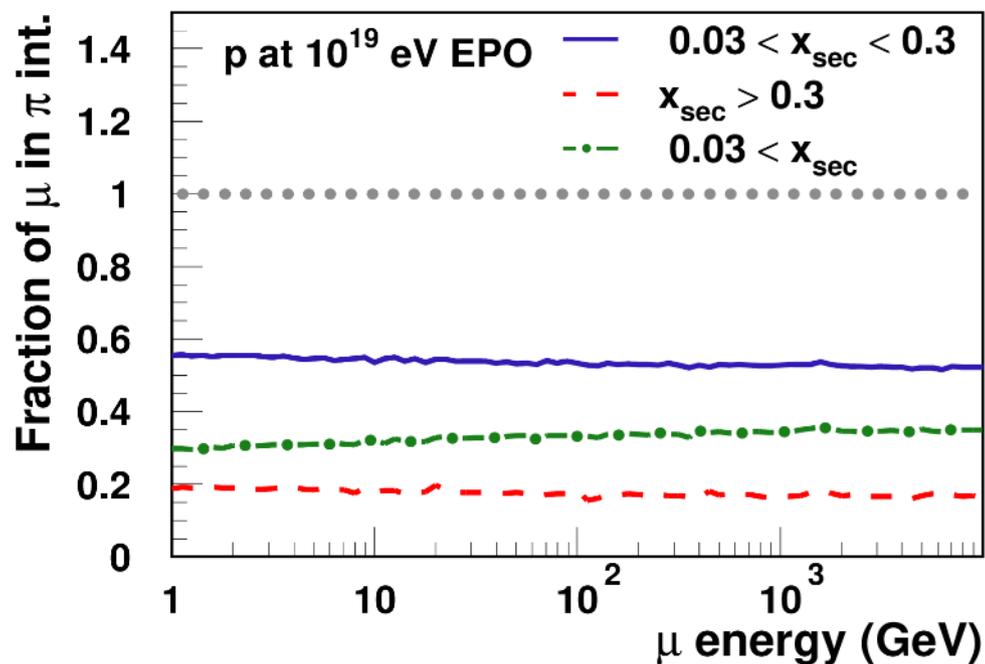
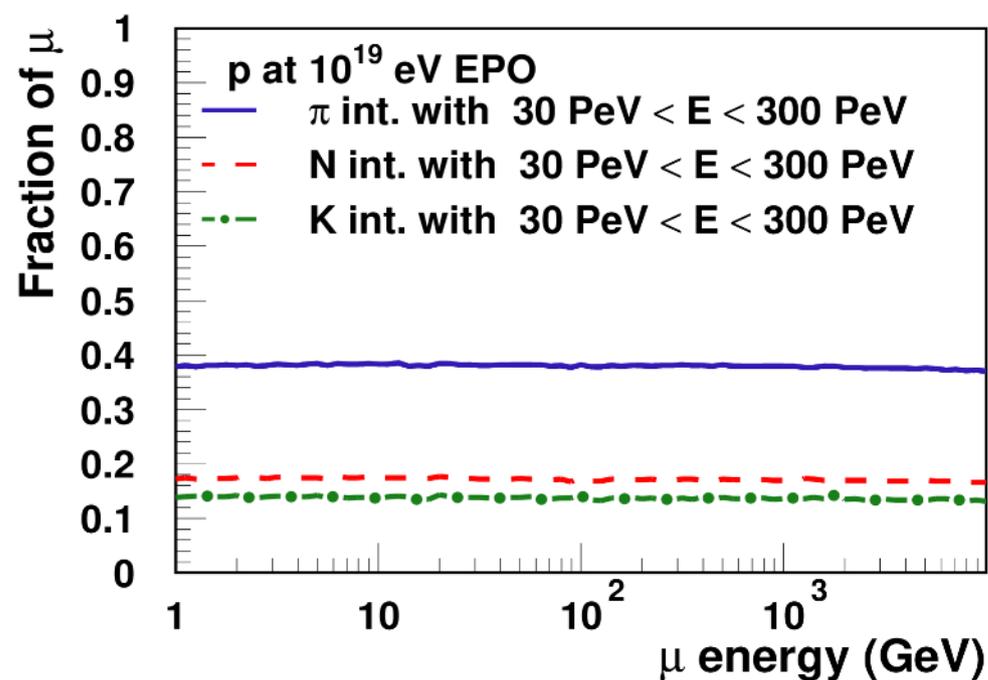
- ➔ only small differences in pp model predictions
- ➔ main difference in high multiplicity tail
- ➔ different behavior for π and p interactions
- ➔ larger differences than in pp



LHC and Muon Production

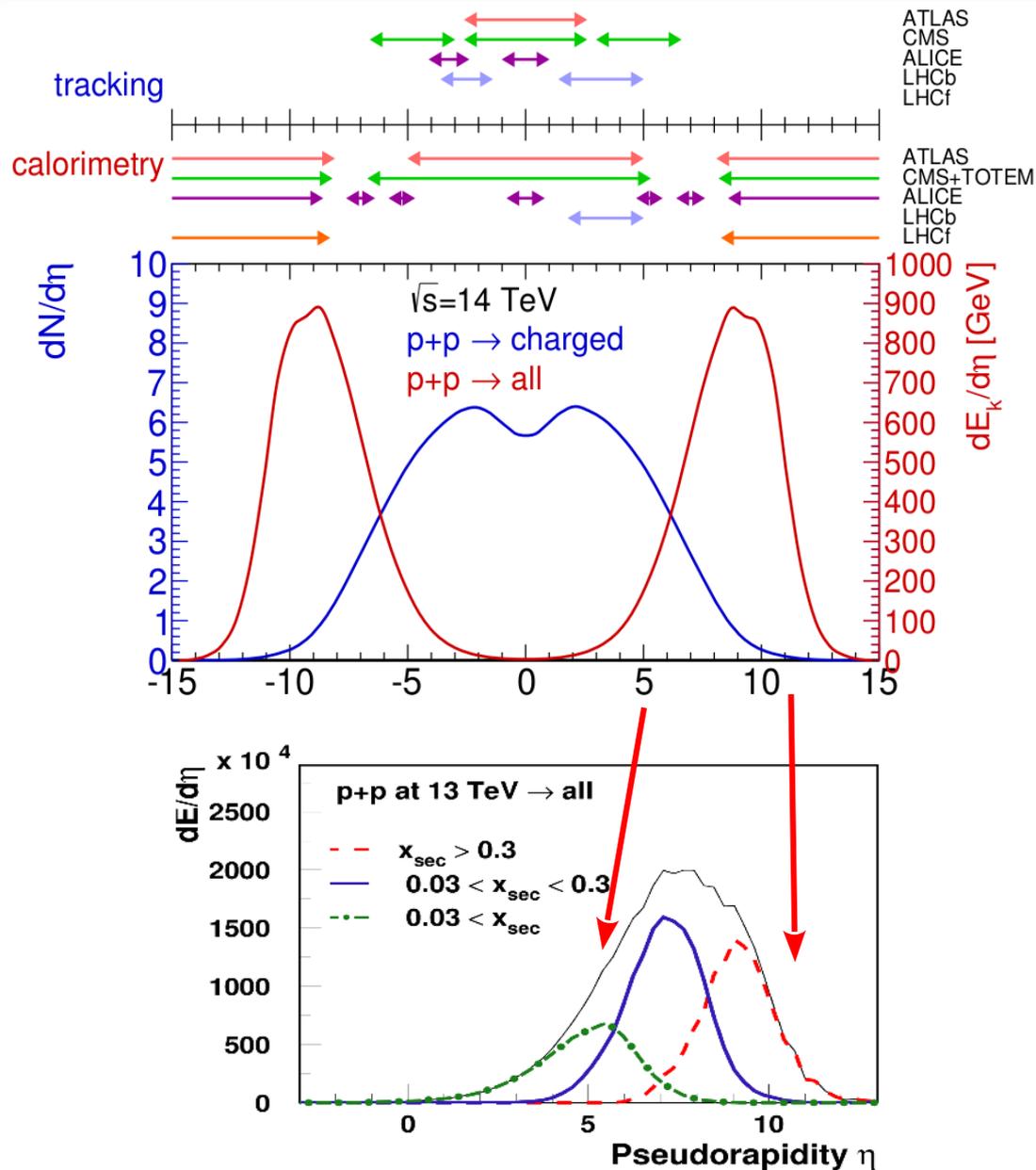
Which phase space is important at LHC for muon production ?

- ➔ modify CONEX (EAS simulation) to extract muons produced by subshowers with interaction energy between 30 and 300 PeV (lab ~ LHC cms)



- ➔ Muon production dominated by pion interactions
- ➔ Muon production depends on secondaries with $0.03 < x < 0.3$ from primaries in LHC energy range

LHC acceptance and Phase Space



- p-p data mainly from “central” detectors

- ➔ pseudorapidity $\eta = -\ln(\tan(\theta/2))$

- ➔ $\theta=0$ is midrapidity

- ➔ $\theta \gg 1$ is forward

- ➔ $\theta \ll 1$ is backward

- Different phase space for LHC and air showers

- ➔ most of the particles produced at **midrapidity**

- important for **models**

- ➔ most of the energy carried by **forward** (backward) particles

- important for **air showers**

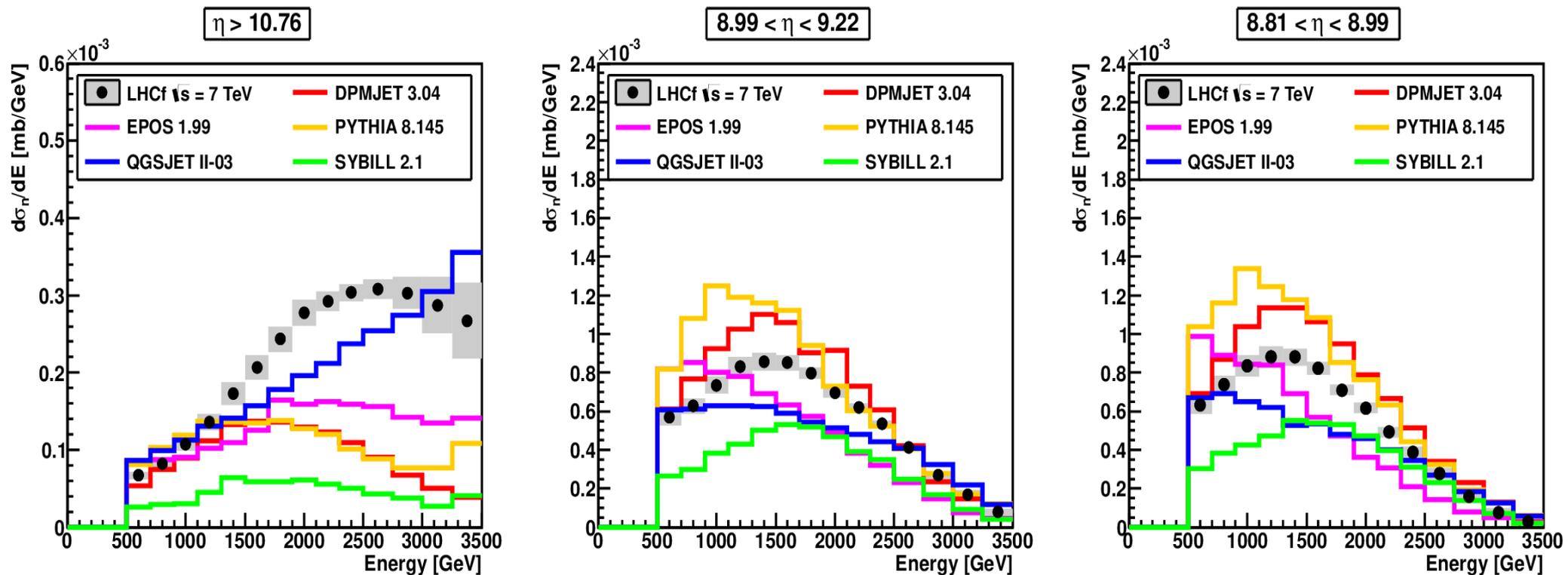
LHCf vs CR Models 7 TeV

● Reasonable results for γ and π^0

- ➔ 20% to 30% “excess” in models at low energy for gammas
- ➔ large difference between models

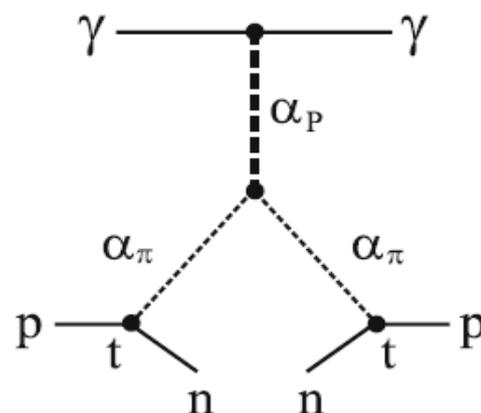
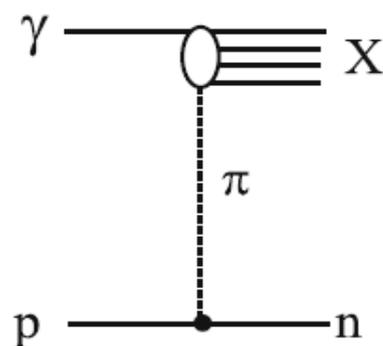
● Stronger deviation for neutrons

- ➔ Clear “pion exchange” peak (not “really” in the models)



π Exchange to Test π Interactions

Physics discussed in detail for HERA (H1 and ZEUS) measurements
(see, for example, Khoze et al. *Eur. Phys. J. C* 48 (2006), 797 and Refs. therein)



**Use neutron tag
with 0 degree
calorimeters to
measure $\pi+p$ in
central detectors**

$$\frac{d\sigma(\gamma p \rightarrow X n)}{dx_L dt} = S^2 \frac{G_{\pi+pn}^2}{16\pi^2} \frac{(-t)}{(t - m_\pi^2)^2} F^2(t) \times (1 - x_L)^{1-2\alpha_\pi(t)} \sigma_{\gamma\pi}^{\text{tot}}(M^2)$$

Use same expression and replace γ by p , but different absorptive corrections
(smaller rate expected, should be still possible in low-luminosity runs)

R. Engel

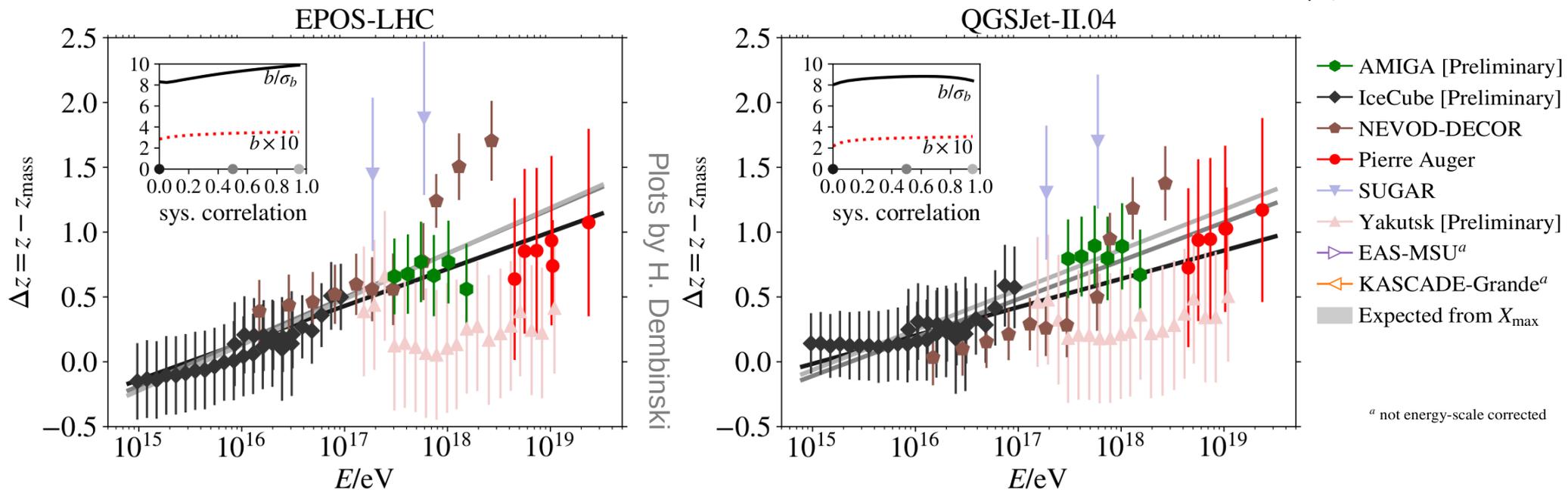
Global Picture of Muons from EAS

● Clear muon excess in data compared to simulation : WHISP 2018

➔ Different energy evolution between data and simulations

➔ Significant non-zero slope ($>8\sigma$)

$$z = \frac{\ln N_{\mu}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}{\ln N_{\mu,\text{Fe}}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}$$



● Different energy or mass scale cannot change the slope

➔ Different property of hadronic interactions at least above 10^{16} eV

Ref: EPJ Web Conf. 210 (2019) 02004 - arXiv:1902.08124

Constraints from Correlated Change

- One needs to change energy dependence of muon production by $\sim +4\%$

$$N_{\mu} = A^{1-\beta} \left(\frac{E}{E_0} \right)^{\beta}$$

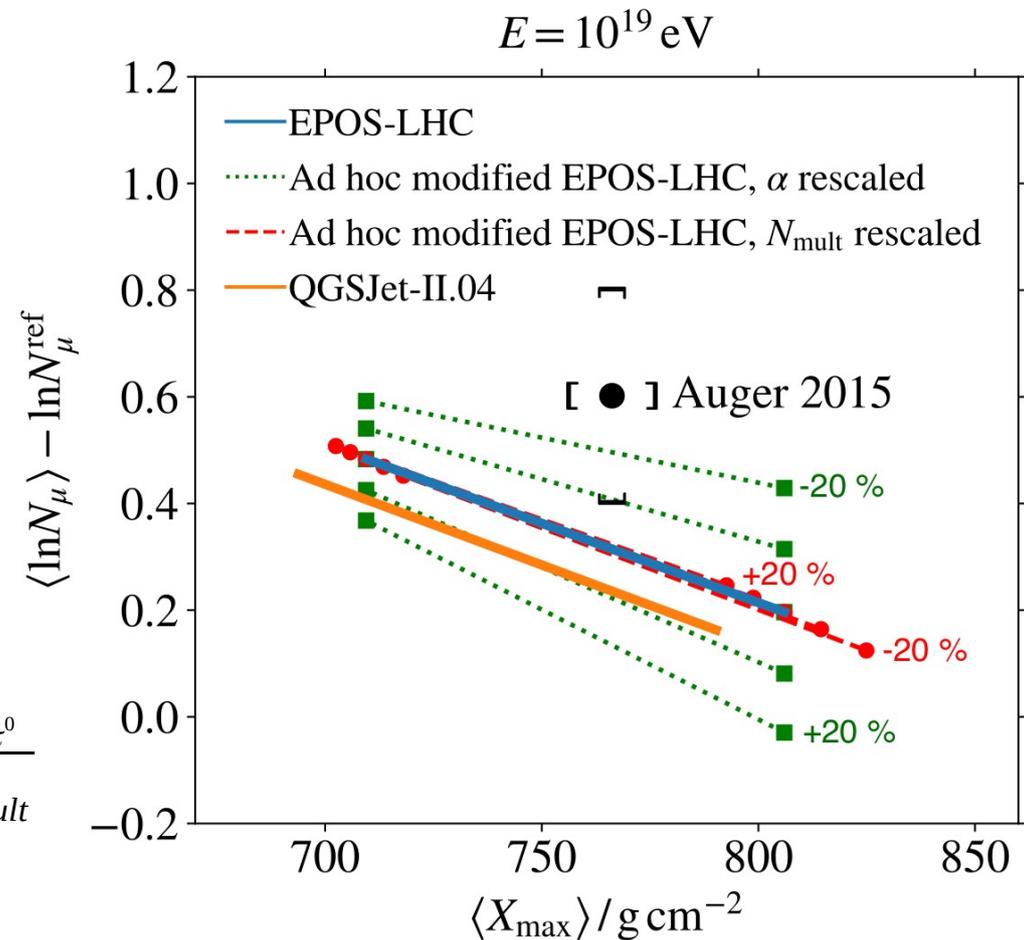
- To reduce muon discrepancy β has to be change

→ X_{\max} alone (composition) will not change the energy evolution

→ β changes the muon energy evolution but not X_{\max}

$$\beta = \frac{\ln(N_{\text{mult}} - N_{\pi^0})}{\ln(N_{\text{mult}})} = 1 + \frac{\ln(1 - \alpha)}{\ln(N_{\text{mult}})}$$

→ $+4\%$ for β → -30% for $\alpha = \frac{N_{\pi^0}}{N_{\text{mult}}}$



Possible Particle Physics Explanations

A 30% change in particle charge ratio ($\alpha = \frac{N_{\pi^0}}{N_{mult}}$) is huge !

➔ Possibility to increase N_{mult} limited by X_{max}

➔ New Physics ?

- Chiral symmetry restoration (Farrar et al.) ?

- Strange fireball (Anchordoqui et al.) ?

- String Fusion (Alvarez-Muniz et al.) ?

➔ Problem : no strong effect observed at LHC ($\sim 10^{17}$ eV)

➔ **Unexpected production of Quark Gluon Plasma (QGP) in light systems observed at the LHC ?** (at least modified hadronization)

- Reduced α is a sign of QGP formation (Baur et al.) !

- Not properly done in EPOS LHC (QGP only in extreme conditions)

➔ Try a modified version of EPOS

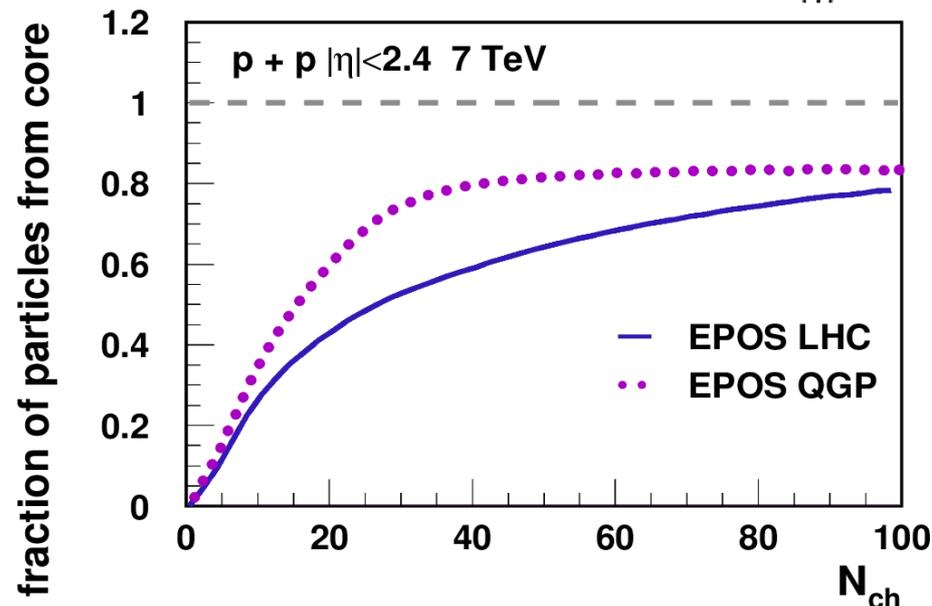
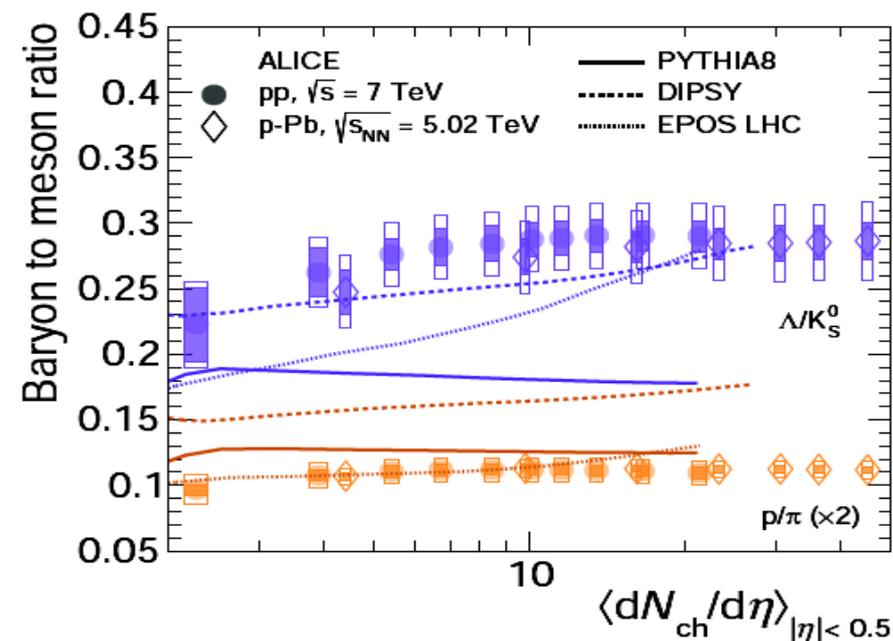
Modified EPOS with Extended Core

● Core in EPOS LHC appear too late

- ➔ Recent publication show the evolution of chemical composition as a function of multiplicity
- ➔ Large amount of (multi)strange baryons produced at lower multiplicity than predicted by EPOS LHC

● Create a new version EPOS QGP with more collective hadronization

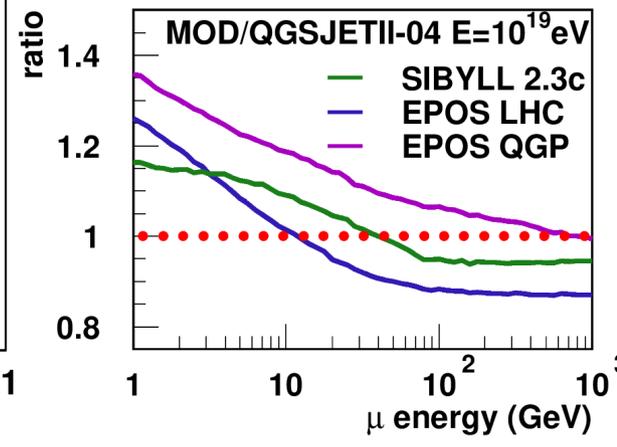
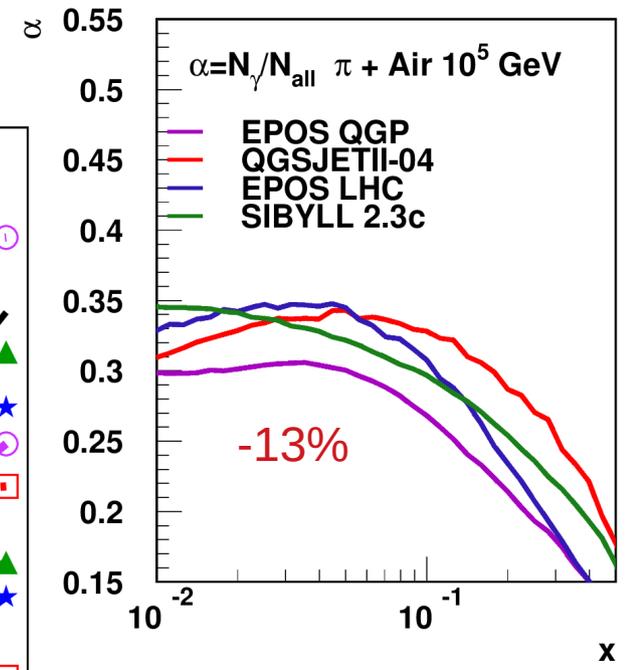
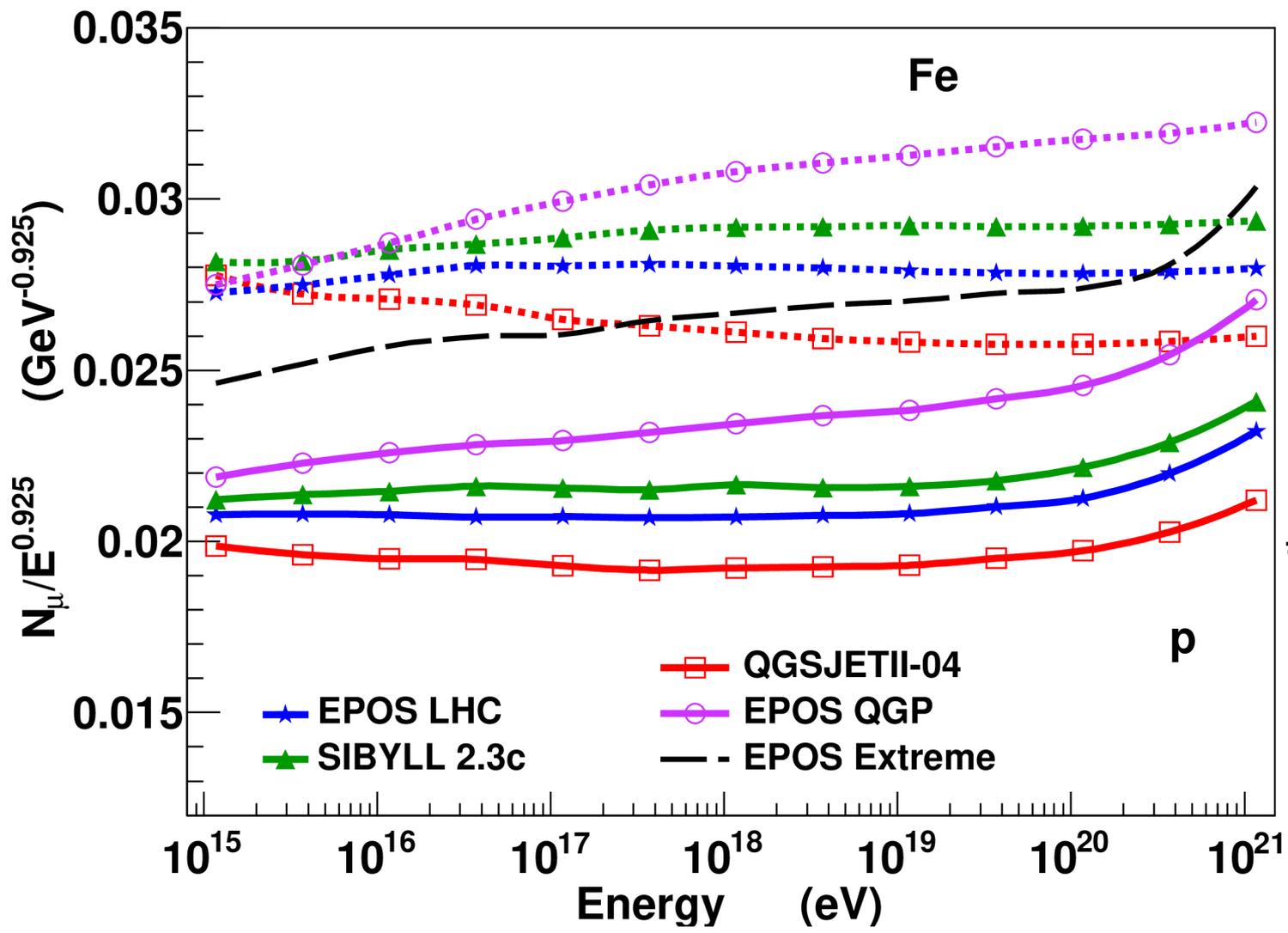
- ➔ Core created at lower energy density
- ➔ More remnant hadronized with collective hadronization
- ➔ Collective hadronization using grand canonical ensemble instead of microcanonical (closer to statistical decay)



Results for Air Showers

Large change of the number of muons at ground

➔ Different slope as expected from the change in α



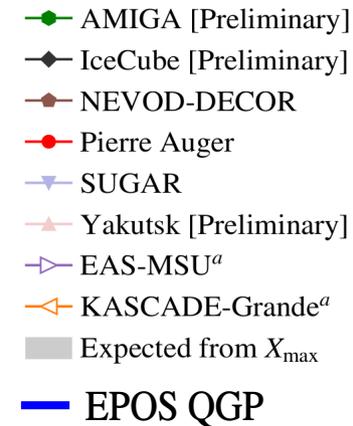
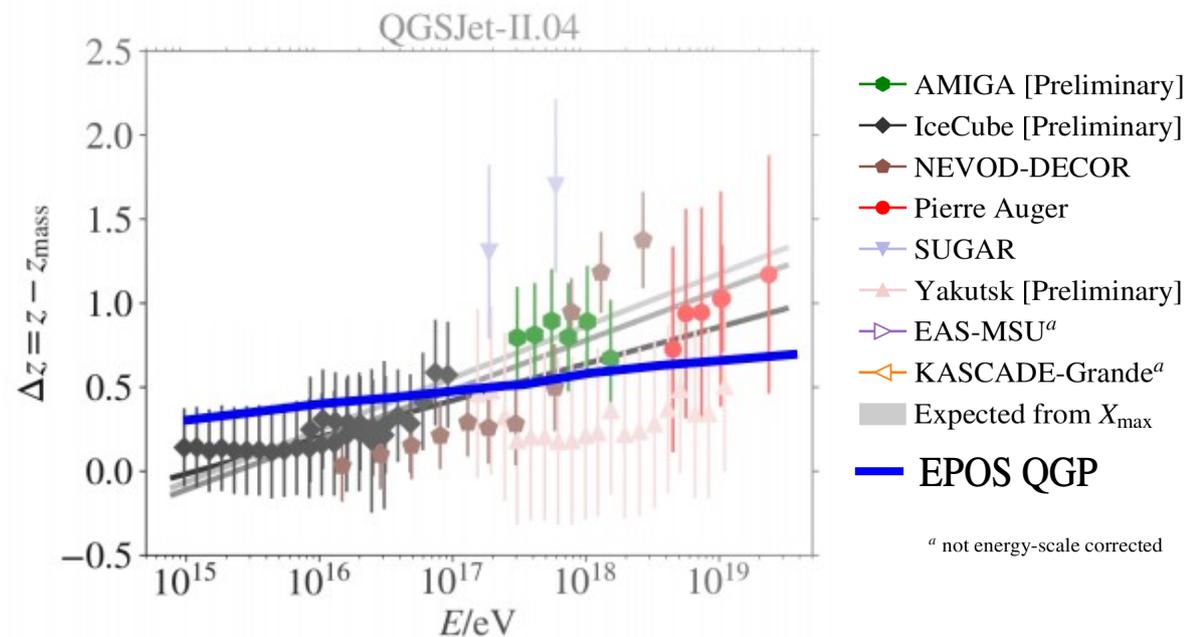
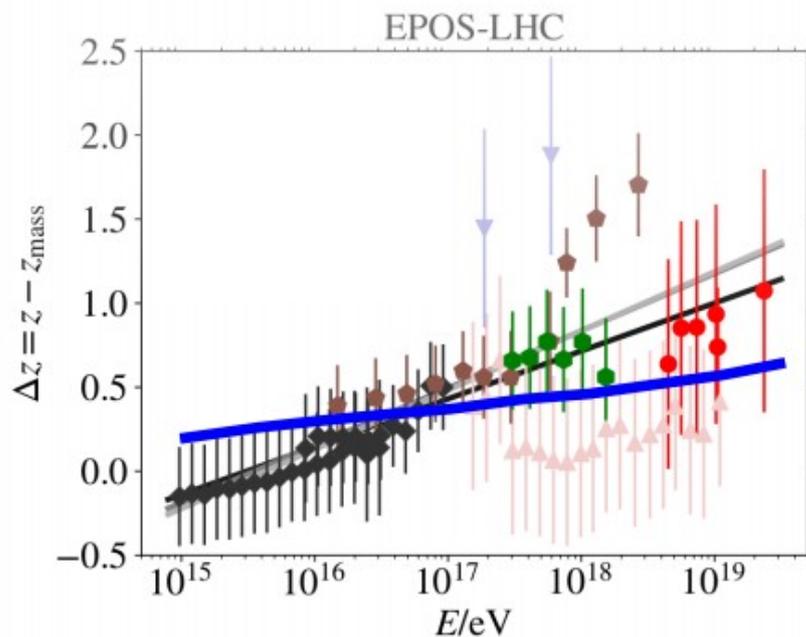
Comparison with Data

● Collective hadronization gives a result compatible with data

➔ Still different energy evolution between data and simulations

➔ Significance to be tested

$$z = \frac{\ln N_{\mu}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}{\ln N_{\mu,\text{Fe}}^{\text{det}} - \ln N_{\mu,p}^{\text{det}}}$$



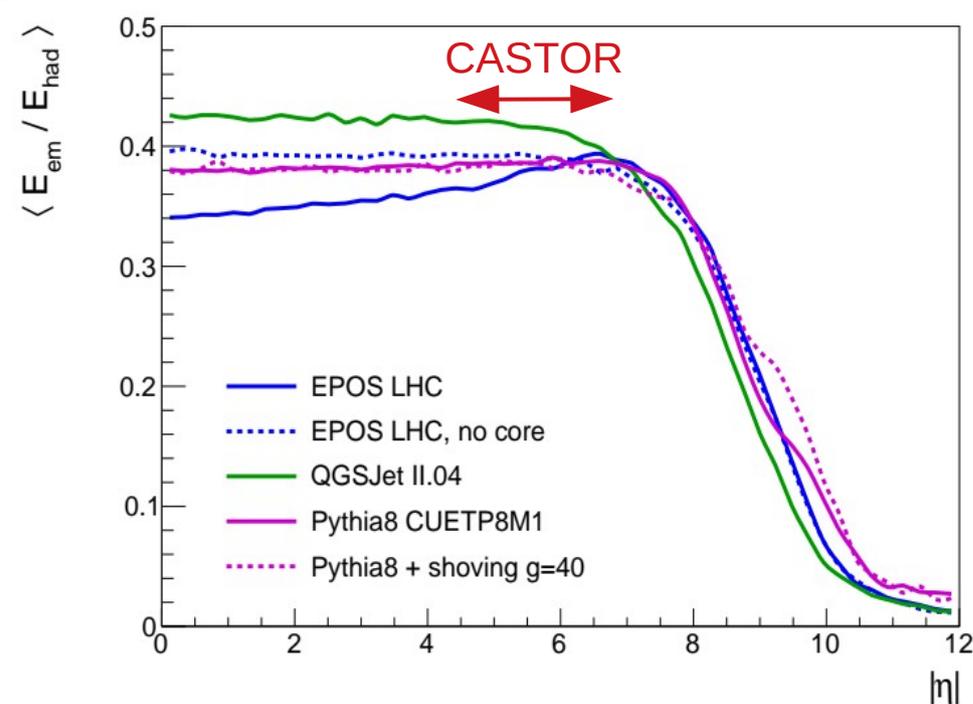
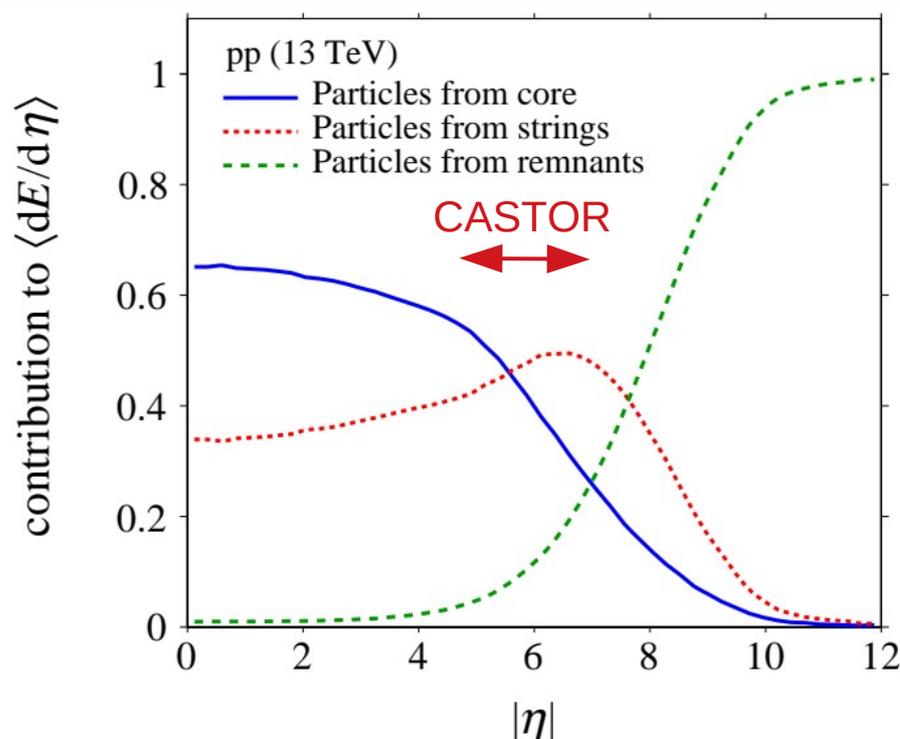
^a not energy-scale corrected

● Probably tension at low energy (too many muons)

➔ Ideally a larger slope would be needed ... what kind of hadronization possible ?

➔ QGP with large chemical potential (Anchordoqui et al.) ?

Test Effect of Collective Hadronization



➔ Reduced α is a sign of QGP formation (Baur et al. ArXiv:1902.09265) !

➔ Problem : α changed at most by 20% for $\mu_B=0$

➔ Behavior α at different μ_B ?

➔ Possible test using forward (and central) calorimeters at LHC

➔ forward/backward asymmetry and centrality evolution

Summary

Cosmic Ray data analysis rely on air shower simulations

- ➔ hadronic models main source of uncertainty
- ➔ forward physics lead air shower development
- ➔ pion interaction very important for muon production

Zero degree calorimeter based analysis

- ➔ possibility to select pion exchange type of interactions: **test pion interaction at very high energy for the first time !**

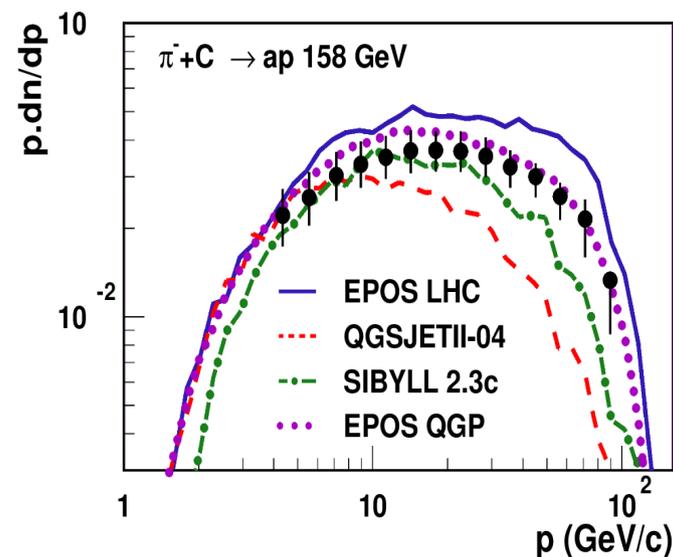
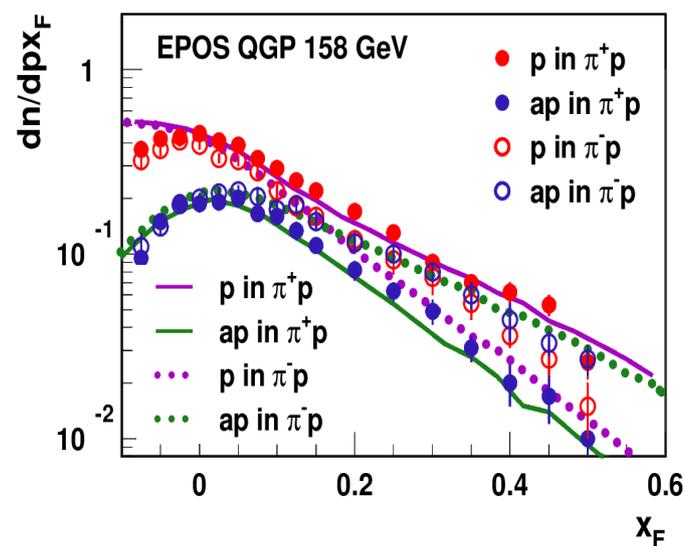
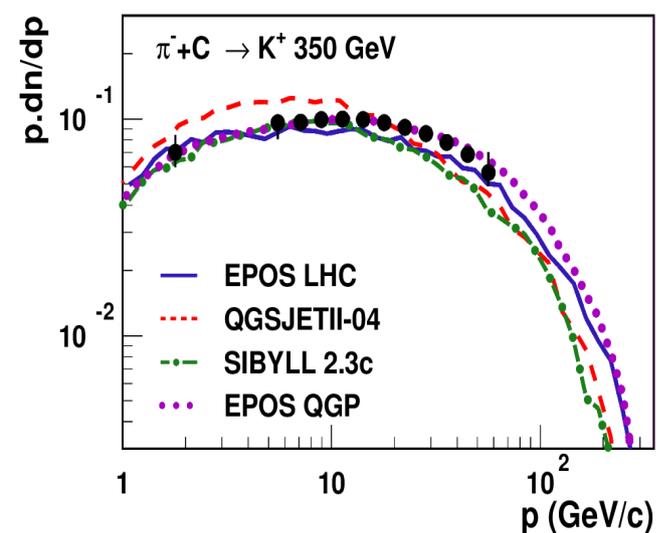
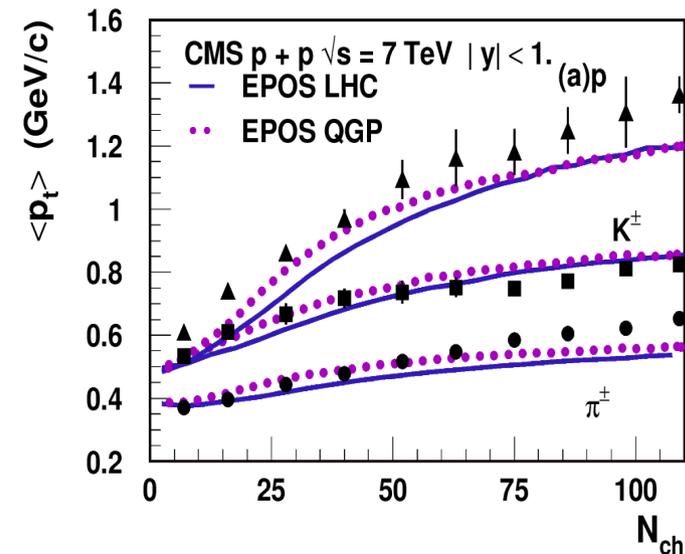
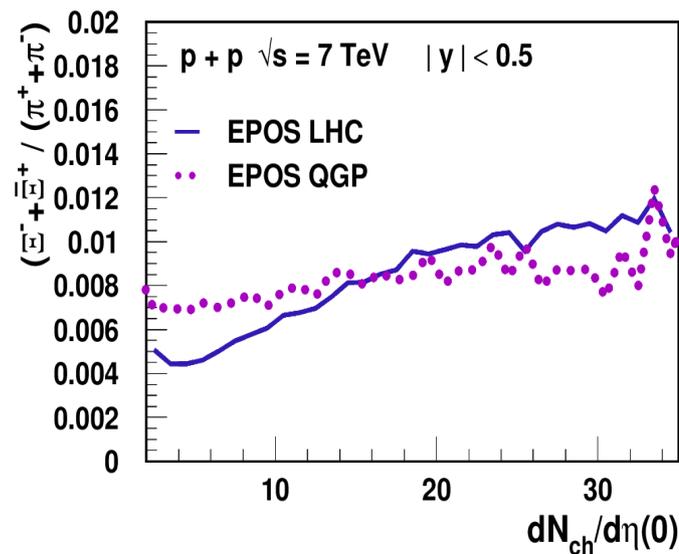
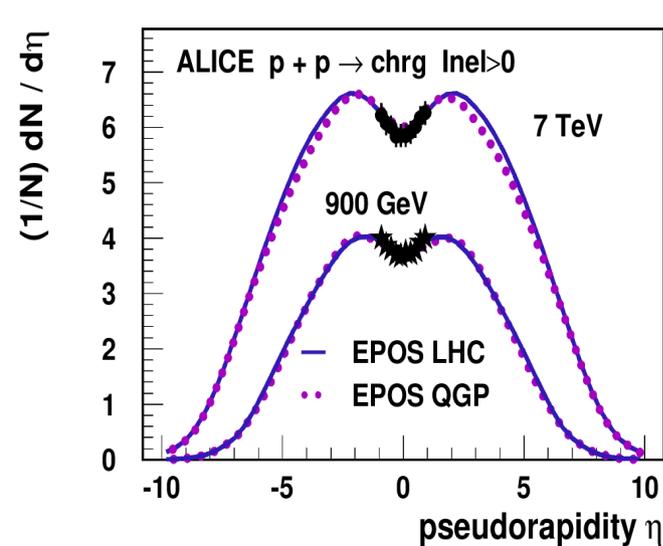
Compilation of all muon measurements clearly indicate a different slope for muon production as a function of shower energy

- ➔ Different hadronization required (less neutral pions / other particles)
- ➔ Collective hadronization in small system / forward in line with LHC results ?
- ➔ Probe new area in quark matter phase diagram ?

Combination of forward and central calorimetric measurements to probe hadronization

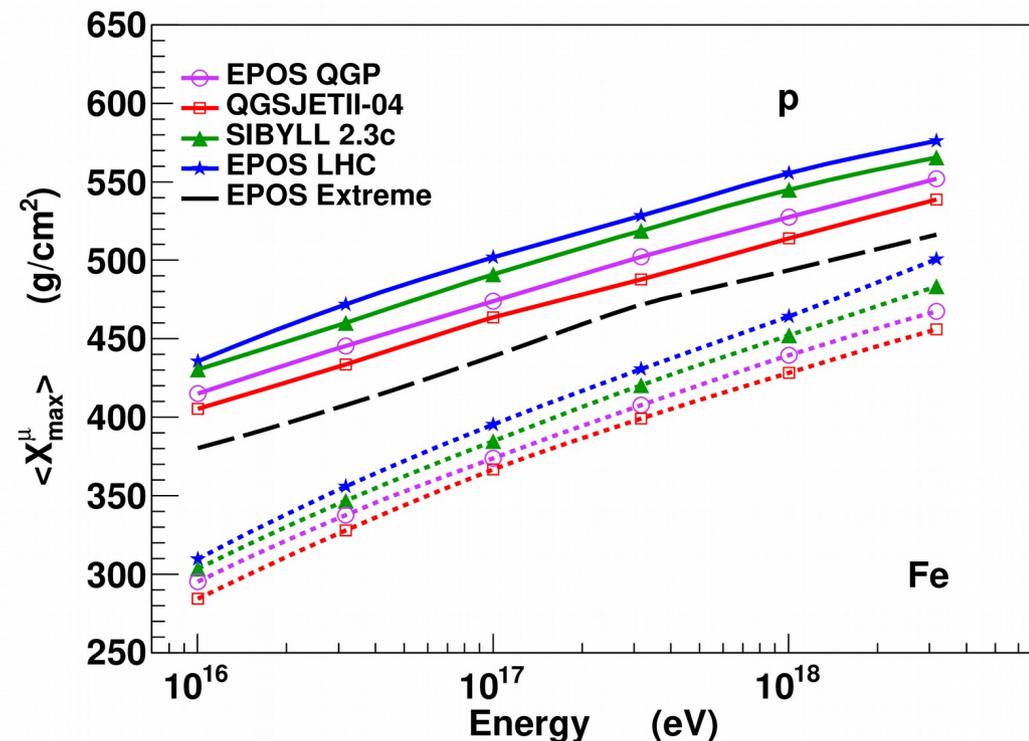
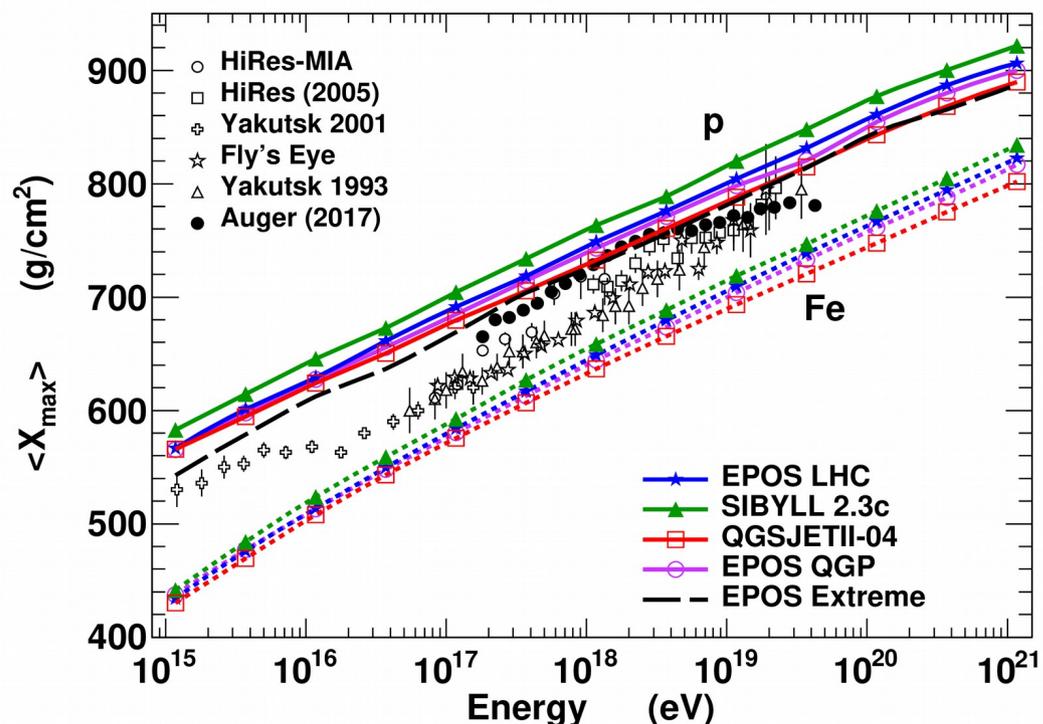
- ➔ **Test forward extension of collective hadronization !**

Preliminary Version with Minimum Constraints



Results for Air Showers

- Small change for $\langle X_{\max} \rangle$ as expected
- Significant change of $\langle X_{\max}^{\mu} \rangle$
- Comparison with extreme case (almost only grand canonical hadron.)
 - ➔ maximum effect using this approach
 - ➔ not compatible with accelerator data



Model predictions for p+p

Models well constraint by LHC run I for pp

- only small differences in model predictions
- main difference in high multiplicity tail

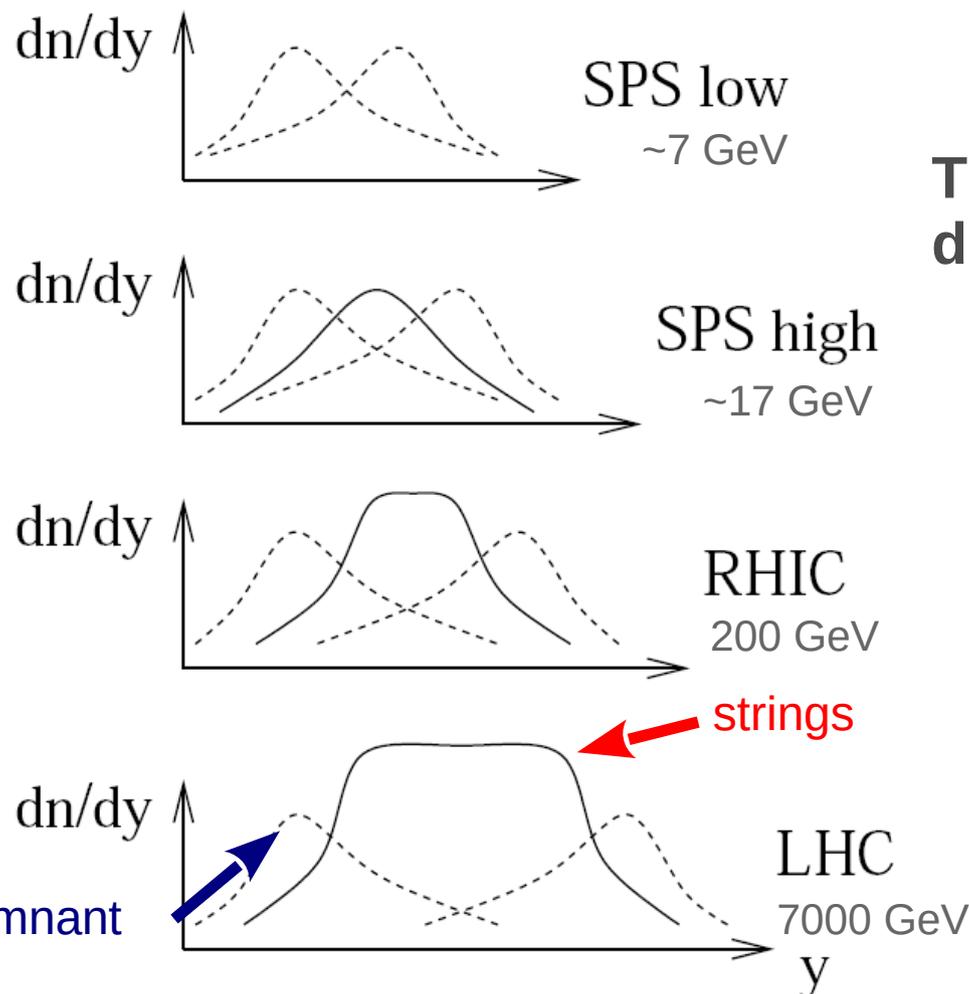
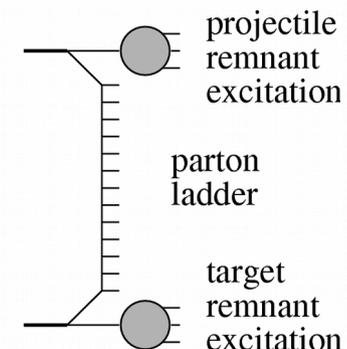
Model Predictions $\pi+p$

Models well constraint by LHC run I for pp

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- main difference in high multiplicity tail
- different behavior for π and p interactions
- larger differences than in pp

Remnants

Forward particles mainly from projectile remnant



The (in)elasticity is closely related to diffraction and forward spectra

- ➔ At very low energy only particles from remnants
- ➔ At low energy (fixed target experiments) (SPS) strong mixing
- ➔ At intermediate energy (RHIC) mainly string contribution at mid-rapidity with tail of remnants.
- ➔ At high energy (LHC) only strings at mid-rapidity (baryon free)

Source Contributions in LHCf (Neutron)

