# **SPAM** with CORSIKA



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Ebr, Ridky, Necesal, Astropart.Phys.90:37-49,2017

### **Overview**

- DELPHI muon bundles from cosmic rays: excess of high-energy muons from moderate-energy CR showers
  - connection to UHECR excesses can (but won't) be argued
- Re-investigation with modern CORSIKA and interaction models
- Single-Particle Addition Model as a proposed "solution"
- Results: a lot of highly confusing plots
- Experience with using CORSIKA and personnal opinions

### **DELPHI** as a cosmic ray detector

- rock overburden: vertical cutoff ~ 52 GeV
- cosmic measurement in concurrence with normal run: effective uptime ~ 18 days





Bundles of parallel tracks in HCAL

- not every muon reconstructed (shadowing, saturation, non-active areas)
- high-multiplicity events mainly from EAS between 10<sup>15</sup>–10<sup>17.5</sup> eV

 $DPH_{20} = 2.24 \pm 0.17$  $DPH_{80} = 1.45 \pm 0.23$ 

DELPHI Collaboration, Astropart.Phys.28:273-286,2007

# **DELPHI** simulations

- whole relevant energy range (10<sup>14</sup>–10<sup>18</sup> eV), spectrum and chemical composition from KASCADE + Grande
- simple "toy DELPHI"
  to roughly reproduce the response of the system to EAS
- fit of efficiency and saturation





| model        | $DPH_{20}$ | $DPH_{80}$ | $DPH_{20}$ | $DPH_{80}$ |
|--------------|------------|------------|------------|------------|
| composition  | p only     | Fe only    | mixed      | mixed      |
| QGSJET01     | 1.00       | 1.00       | 1.43       | 0.70       |
| QGSJET-II-03 | 1.11       | 0.75       | 1.54       | 0.57       |
| QGSJET-II-04 | 1.11       | 1.37       | 1.72       | 0.83       |
| EPOS-LHC     | 0.85       | 0.86       | 1.27       | 0.59       |

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### **Auger simulations**

- Auger depth of maximum constrains models (no lighter than protons)
  simulations at 3.2×10<sup>18</sup> eV
- Some amount of muon excess on Auger
  - increasing muon production would be nice
  - reading out  $N_{\mu}$  at 1000 meters using NKG fit between 250–1500 meters



# **Soft-particle addition model**

particles:  $\pi$ , K, p, n colour: (NWT+NWP)<sup> $\eta$ </sup> distribution  $p \exp(-p/p_0)$  angle: within  $1^{\circ}$  0.1° from axis in c.m.s. shape: energy treshold (or special  $p_0$ ) filled vs. empty:  $p_0$ 



### **SPAM: DELPHI data at multiplicities >20**



### **SPAM: DELPHI data at multiplicities >80**



#### **SPAM: total number of added particles**



# **SPAM: DELPHI vs. Auger** $X_{max}$



## **SPAM:** Auger $X_{max}$ vs. number of muons (protons)



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## **SPAM:** Auger $X_{max}$ vs. number of muons (irons)



## **SPAM:** Auger $X_{max}$ vs. RMS (protons)



## **SPAM:** Auger $X_{max}$ vs. RMS (irons)



## **CORSIKA** simulations

• (part of) Prague ~8000 CPU cluster

- long development of SPAM -> many centuries of CPU time

- own utility to manage simulations:
  - ~300 lines of Pascal
  - produce steering files from table, babysit the jobs
  - CORSIKA steering format easy to generate from any software
  - switch to HTCondor -> a lot of funcitons obsolete
- DELPHI: no thinning (for muon tracking)
  - no EM cascade simulated
  - ECUTS 53.0 53.0 100.0 100.0
  - all relative to QGSJET-01 (still availabe in new CORSIKA)
- Data processed in Pascal/bash/GNUplot
  - ASCII output from CORSIKA easy to work with in any language
  - .long files a bit confusing, extraction of fluorescence profile unclear

## Modifying interactions CORSIKA

- Own Fortran code within CORSIKA subroutine that calls HE model
  - let the model run, then modify produced particles
  - lower energy of all particles, add new ones (all in CMS)
  - after boost back, usually some minor imbalance
    - conserving 4-momentum in full 3D not easy
  - additional simple changes to steering code to allow own keywords
- The experiences of a very bad programmer:
  - the preprocessor system makes original source unreadable
    - work on the "-compilefile.f"
    - fast and easy to recompile, but cannot change HE model etc.
  - simplicity of Fortran code allows fast learning curve
    - particles just in an array, most variables explained in source, so implementing the actual model was rather easy
    - anything else a bit difficult reading files, passing configuration etc...
    - lack of scoping prone to weird side effects (crashes with FLUKA)
- Data processed in Pascal/bash/GNUplot
  - ASCII output from CORSIKA easy to work with in any language
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## My wishes for CORSIKA 8

- Keep providing plaintext input/output interface
  - allow the user to use the tools they like: can't foresee tastes of every user
  - simple, human-readable files have unbeatable portability
    - saving outputs only in ROOT is a terrible idea
- Allow modification of interactions by outsiders
  - easy access to secondaries and easily adding/changing/removing them
- "Easy" means being able to do it without:
  - complex C++ magic
  - using a lot of things from CORSIKA-specific framework and libraries
  - understanding details of a complex build system

• Always remember that the user may be somewhat acceptable with physics, but unimaginably bad with computers!