

LiF as possible detection medium for Light Dark Matter

April 16th, 2026
MD ν DM'26 Karlsruhe

Ranny Budnik

Department of Particle Physics and Astrophysics
Weizmann Institute of Science

מכון ויצמן למדע

WEIZMANN INSTITUTE OF SCIENCE

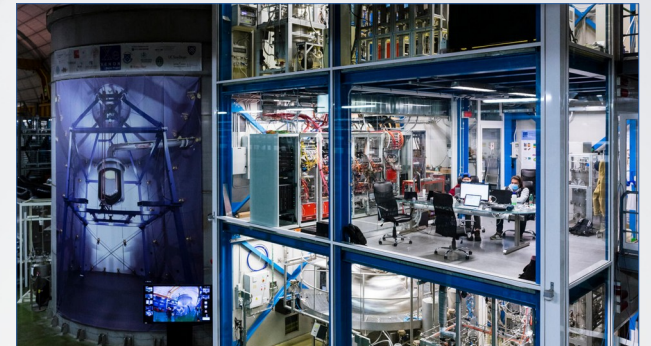
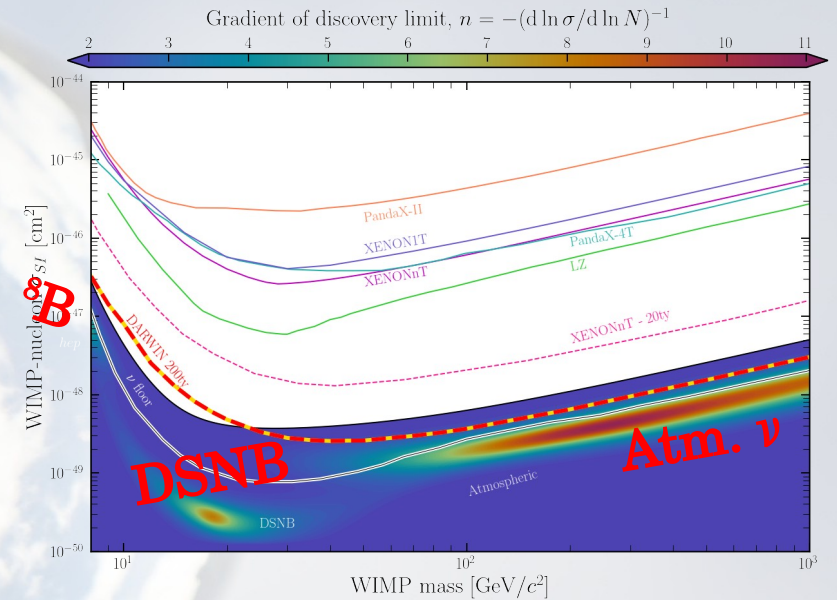


Some order

- Why CC are so interesting for Light Dark Matter?
- How can we do that?
- What we learned
- What we want to do onward

WIMPs

- We have been invested in WIMP searches for decades
- Progress: ~8 Orders of Magnitude in σ_{SI} !
- Next Gen is: 10 yr making, 10 yr data, \mathcal{O} (500M\$)
- Very motivated! But...



The problem of “cost effectiveness” in BSM

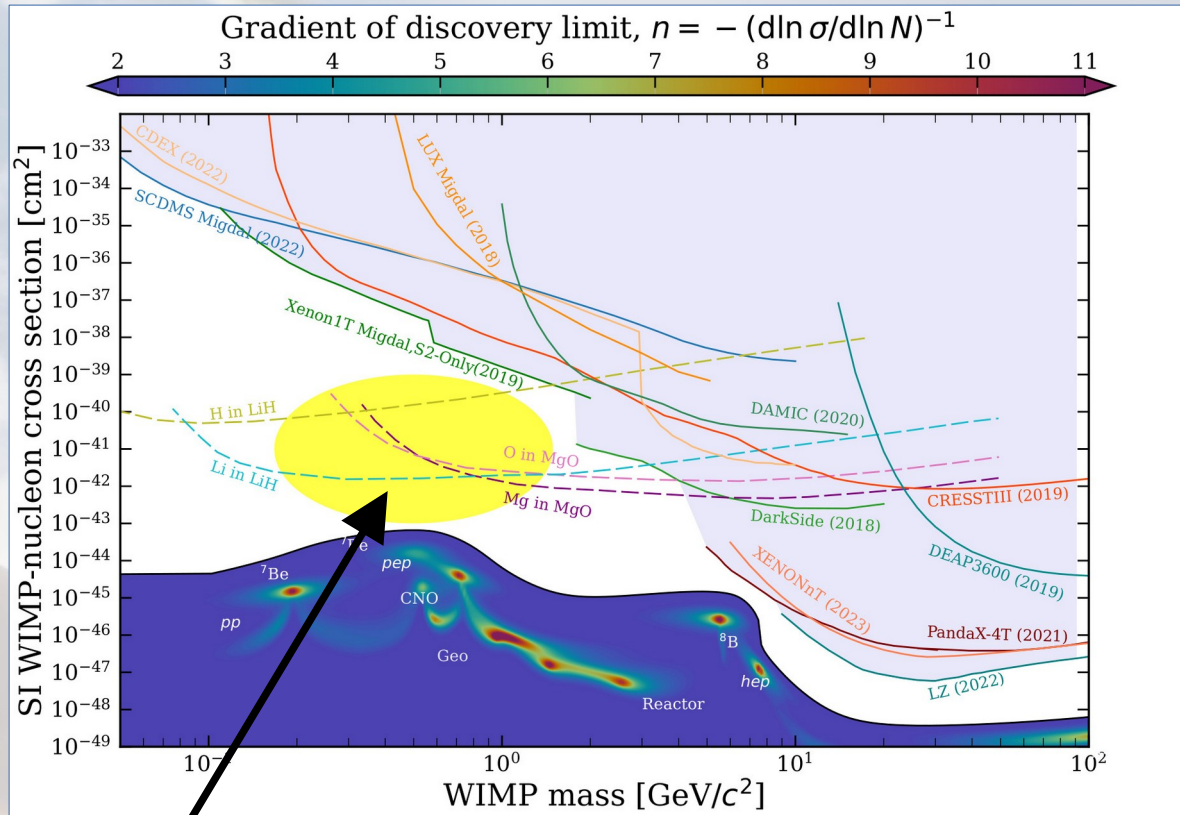
- We no longer have a “target” from theory
- We do believe that BSM is out there and can be discovered
- What should be our metric?
 - $\Delta m \times \Delta \sigma$?
 - No... tend to upper scale
- Like any scale-free problem, log is king:
 - $S_2 = \log_2(m_{\max}/m_{\min}) \times \log_2(\sigma_{\max}/\sigma_{\min})$
 - A new “run” is worth if $S_2 > \text{few}$
 - A new method/experiment is worth it if $S_2 \gg 10$

The high-value goal: Light DM

Open a new window for **Direct Detection** of **Light Dark Matter**

Why is it **hard**?

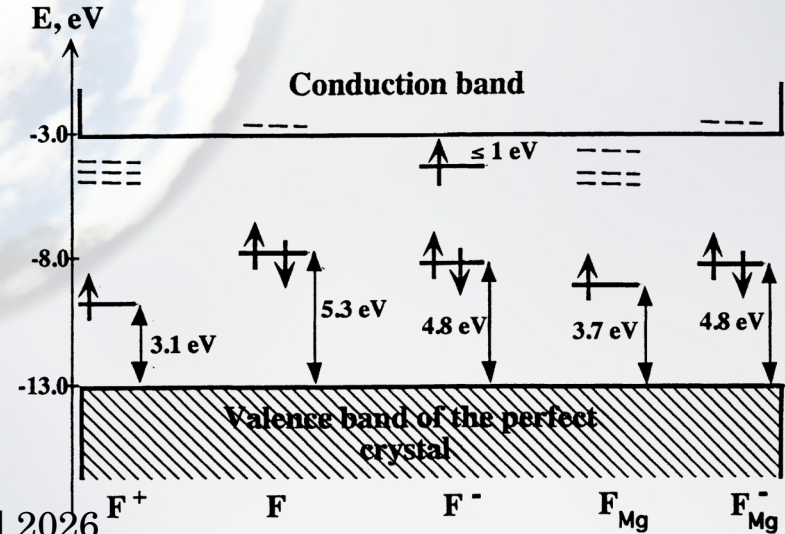
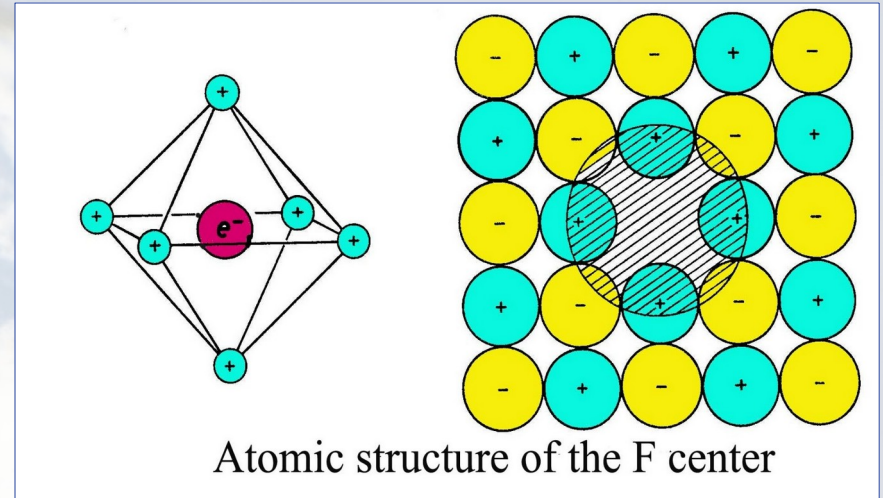
- Small number of quantas (**eV range**)
- “Instrumental” **backgrounds**
- **Calibrating** signal
- **Maturity** of technologies



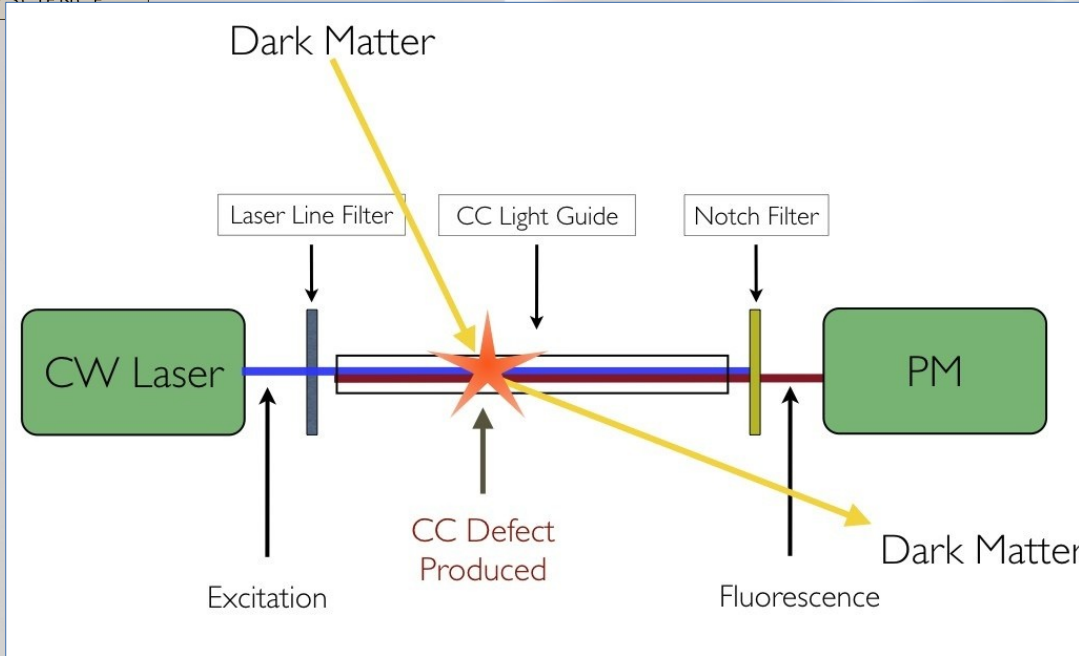
$$E_R \sim m_\chi^2 / m_{NV}^2 \ll \text{keV}$$

F-center in a nutshell

- Each type of defect has its own typical **absorption** and **emission** wavelengths
- By this mechanism a transparent medium becomes **colored**
- Elastic collision may produce **displacement** (gamma, electron, neutron and ions)



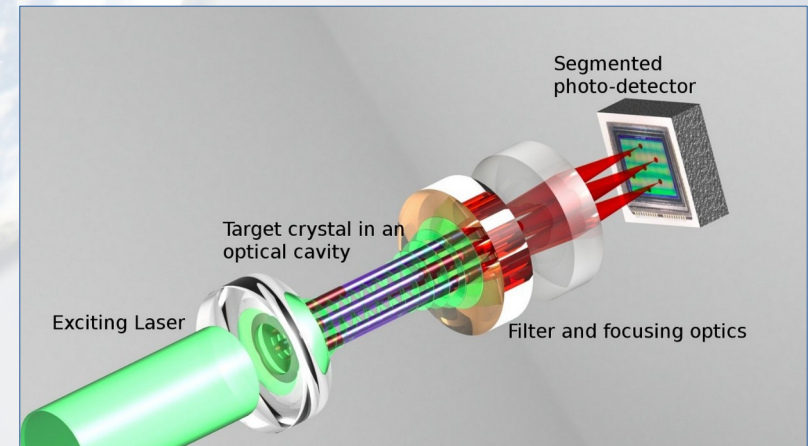
The idea - 2017



An array of **single, pure crystals** can be used as **light-guides**

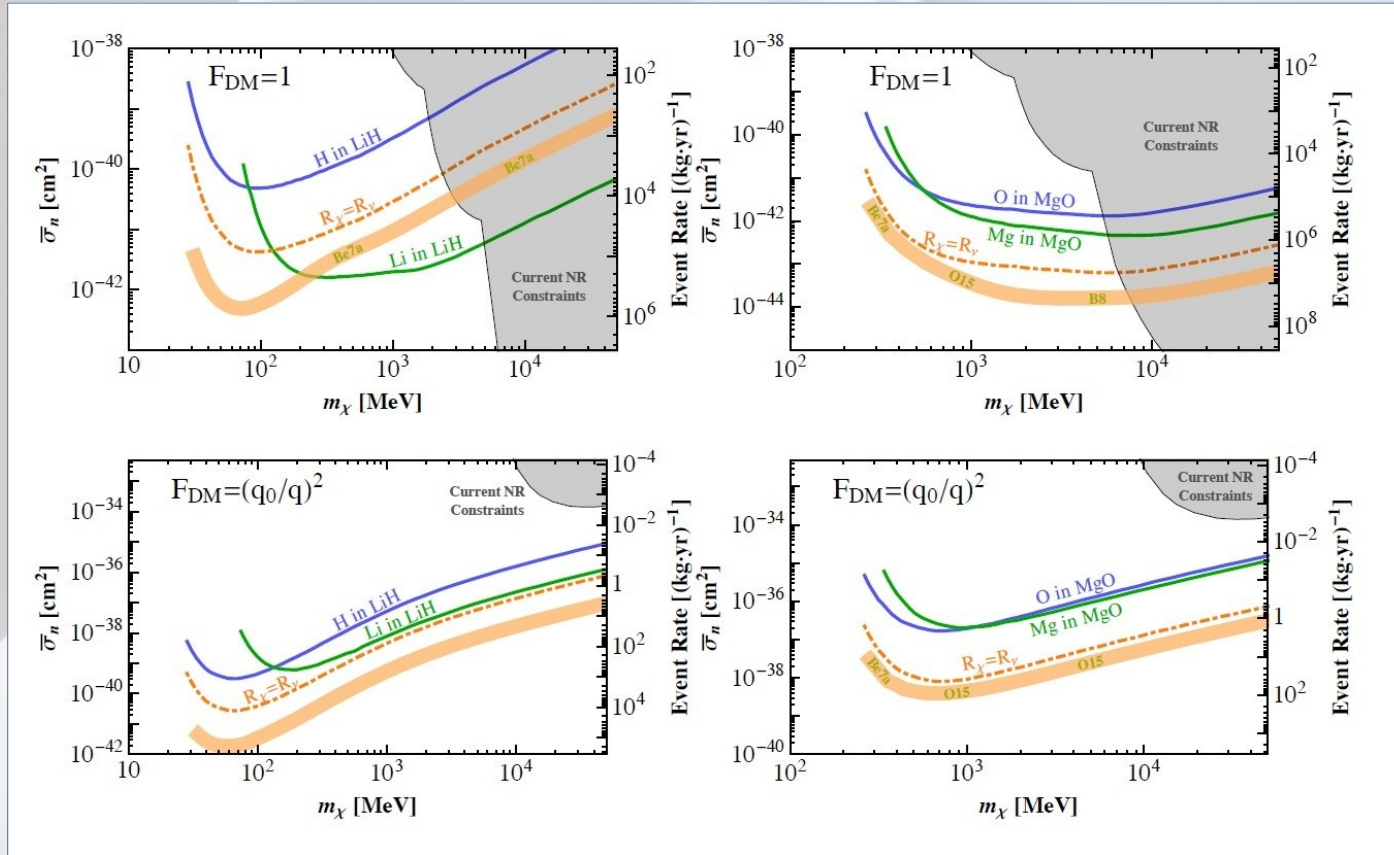
Fluorescence measurement can detect the creation of a **single new defect in a bulk**

Realistic estimates show that such a setup can be able to distinguish a **single new site in 10^5 existing** per rod



Physics reach (2017 ideas)

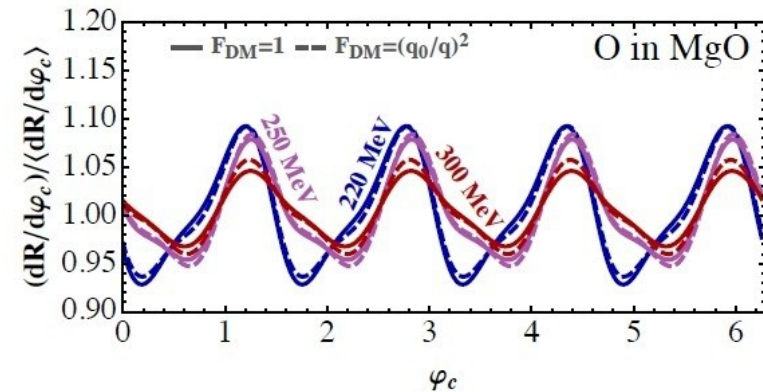
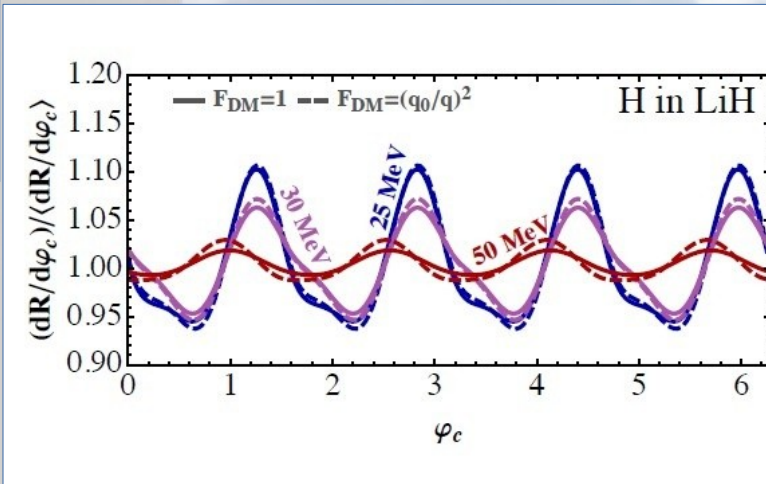
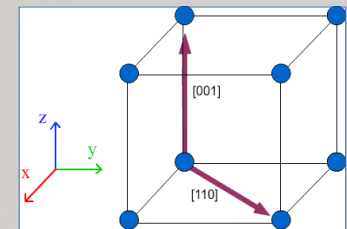
With 1 kg·yr



RB, O. Cheshnovsky, O. Slone, T. Volansky 1705.03016, PLB

R. Budnik April 2026

Modulation and directionality

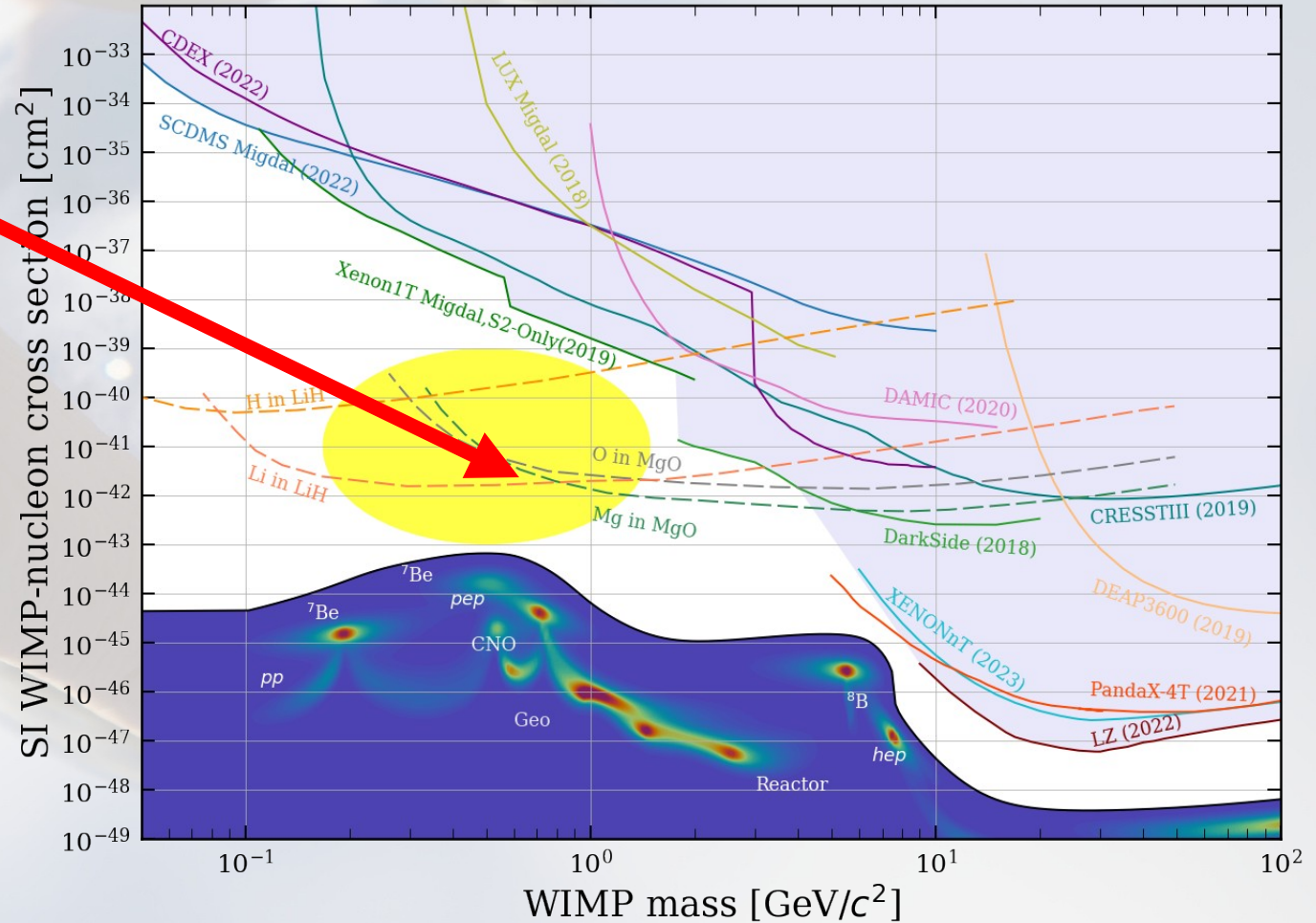


- **Hourly** change of orientation relative to DM wind
- Modulated rate differs from solar direction (4 min/day), day/night, season
- **Strong especially near threshold**
- The smoking gun of **Galactic origin**

Physics Reach

“The Dream”:

Finding single defects in a ~kg target



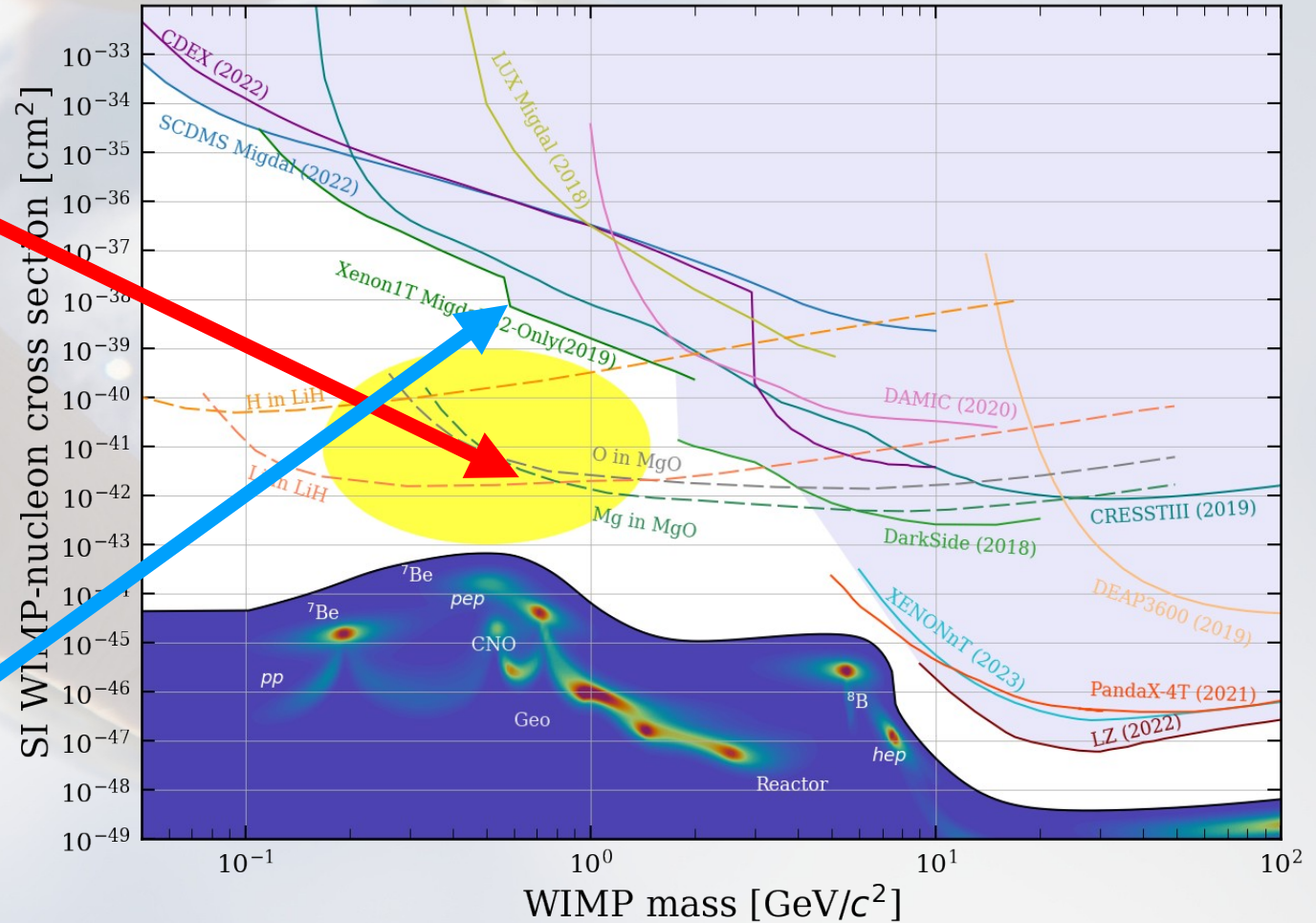
Physics Reach

“The Dream”:

Finding single defects in a ~kg target

The reasonable near future:

Finding 100 defects in a ~g target



Why Is This Hard?

The principal challenges — and why they matter



Needle in a Haystack²

Find the Color Center that answers everything: computationally “impossible”

Detect 1 new site in 10^5 existing per rod. Single-event sensitivity requires exceptional optical stability.



Crystal Purity

Contaminants at 10^{-17} levels required. Any **impurity may create its own optically active defects.**



Calibration

O(keV) neutrons needed — non-trivial source geometry and flux characterization.



Background Discrimination

α , n , γ , β , ions all create color centers. Must attribute signal to DM specifically.

Instrumental backgrounds: optics, readout, thermal...



Bleaching

Laser illumination itself **destroys the signal.** Must characterize and manage carefully.



Three Annealing Processes

Optical bleaching \neq thermal electron release \neq true site healing.

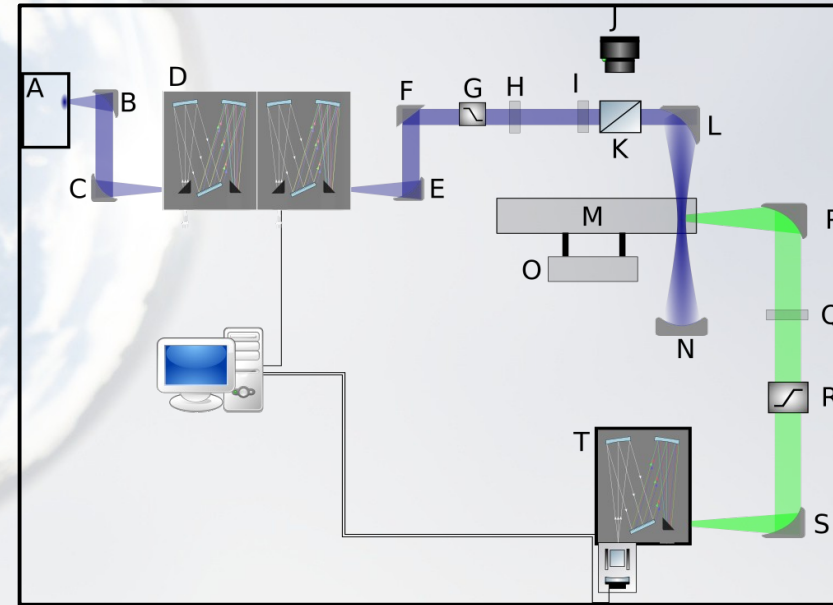
Each different at different temperatures.

First actual step: 2019



What we did

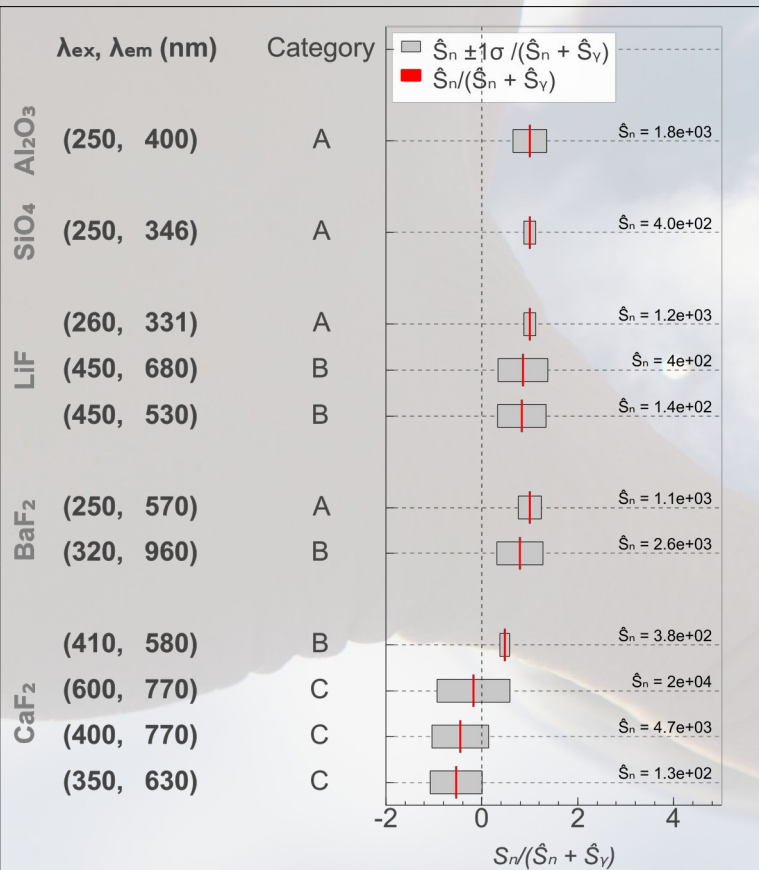
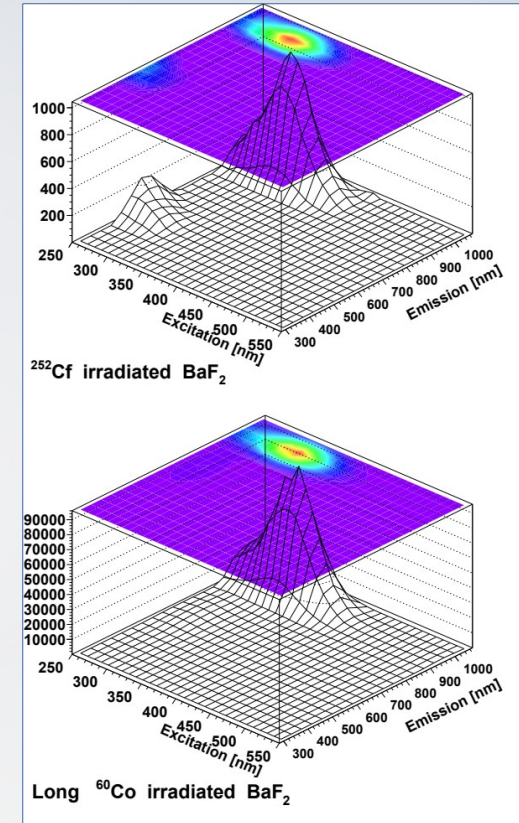
- **First systematic search for CC candidates for DM detection**
- Found 11 interesting color centers across 5 (out of 10) crystal materials: LiF, MgF₂, CaF₂, Al₂O₃, SiO₂
- Irradiation: ⁶⁰Co (γ) and ²⁵²Cf (neutrons): low doses
- Purpose-built broadband optical system designed from scratch
- → Sensitivity down to 10^4 cm^{-3} defect density
- This paper provided the community with a “materials shortlist”



Mosbacher et. al. :1902.10668

First measurements – “full scan”

- Measurements of gamma and (high E) neutrons
- Several “interesting” points identified
- More measurements with low E neutrons underway
- Down to 10^4 cm^{-3} sensitivity



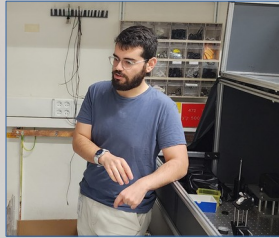
Category A: neutron-only → pure NR → ideal for DM

Category B: n + low-dose γ → needs discrimination

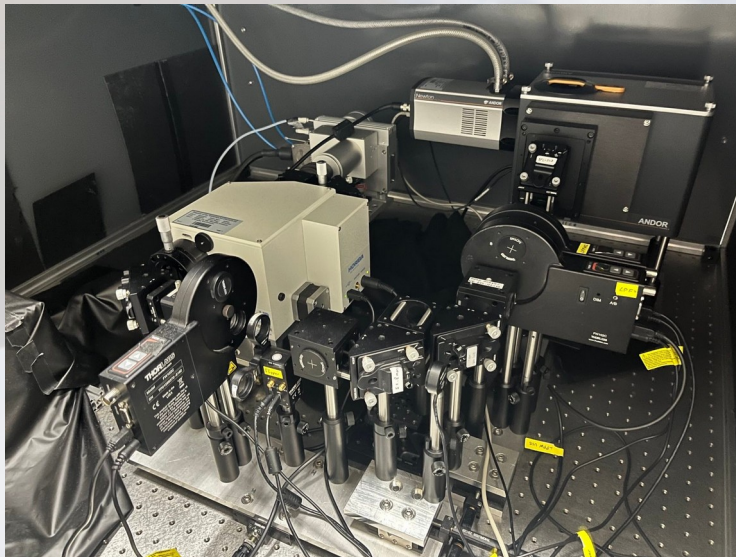
Category C: Only γ

Mosbacher et. al. :1902.10668

2023-2024 – A true candidate is showing up!



Improved stability and systematics on the optical system (3CS): Wide Band



Irradiation Sites: WIS ^{60}Co ($\sim 0.3\text{mCi}$)



Irradiation Sites: BGU AmBe ($\sim 2\text{Ci}$)

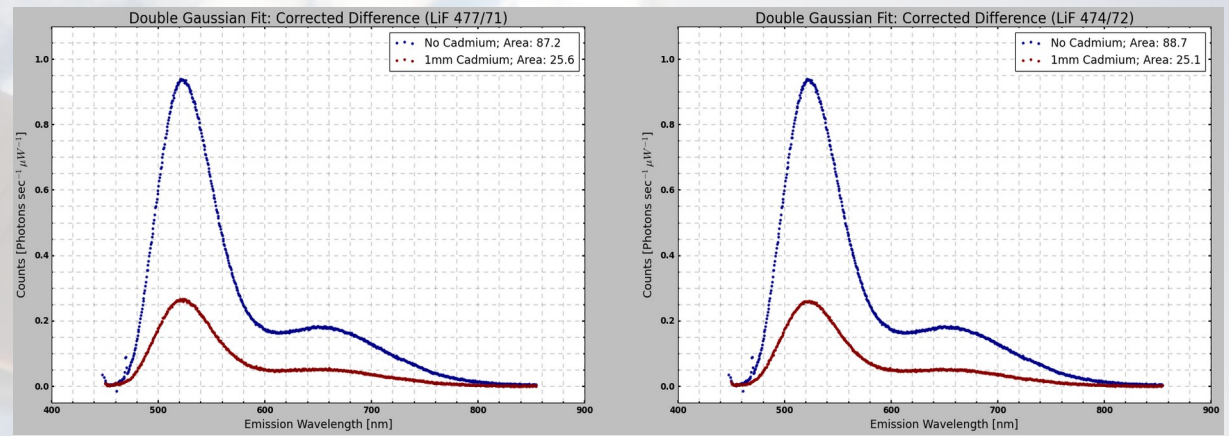
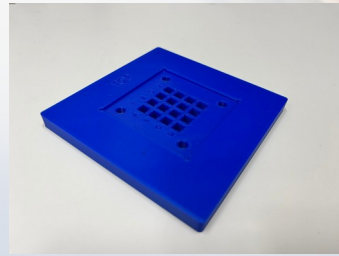
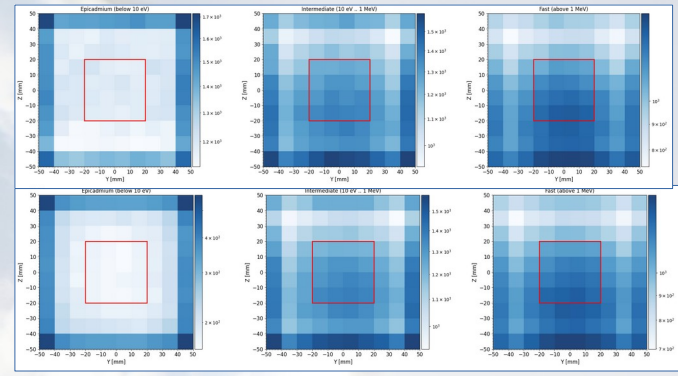


2023-2024 – A candidate!

- **LiF** showed two “B-center” large excess.
- Known to be produced by thermal neutrons, we measured with Cd shield – **expect factor 9 reduction**
- **Reduction is ~X3**
- Ongoing improvement on sensitivity for cross-checks

No Cadmium

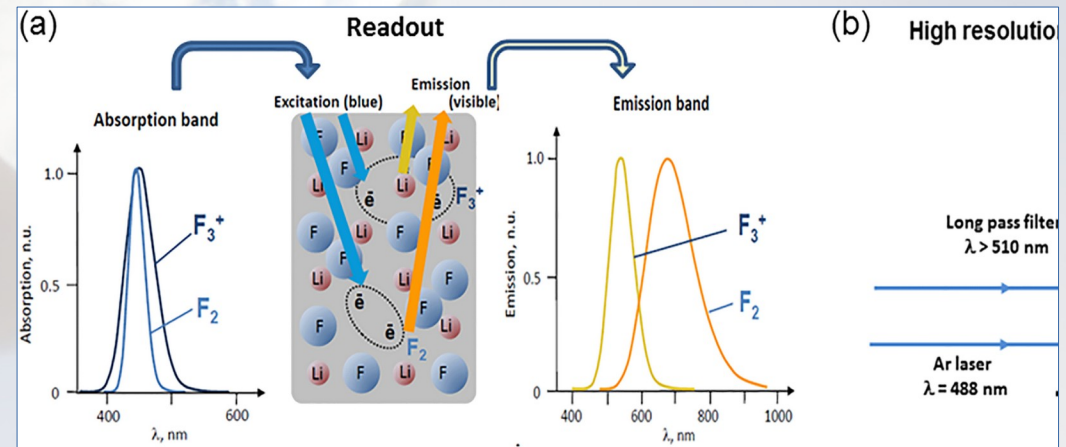
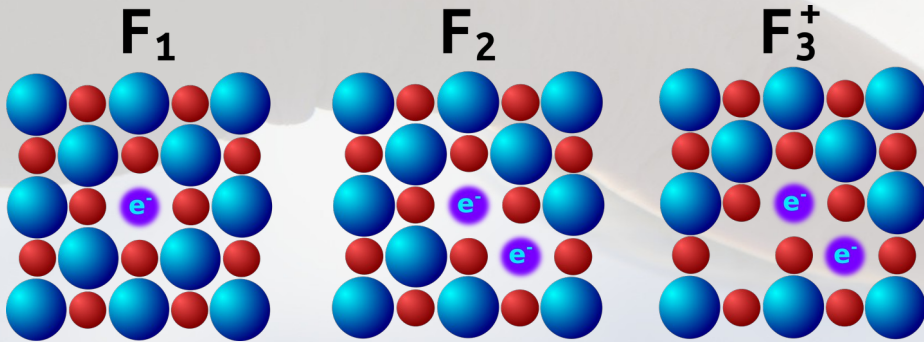
1mm Cadmium



(γ contribution found to be negligible)

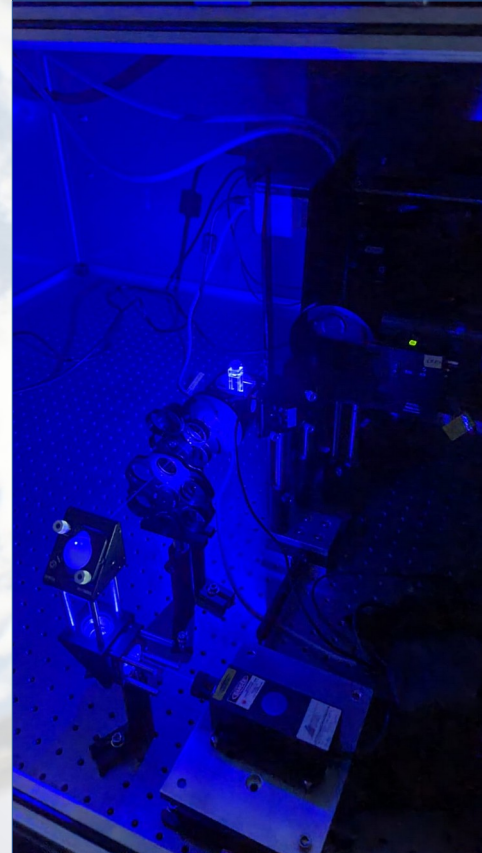
What is it then?

- Two centers are created: F_2 and F_3^+
- Known to be created by n-capture
- Not the “simplest” F_1 – this one is in the deep UV (A-center) – **for future studies**
- Threshold estimated at 16 eV - 20 eV
- “Easy” spectroscopy



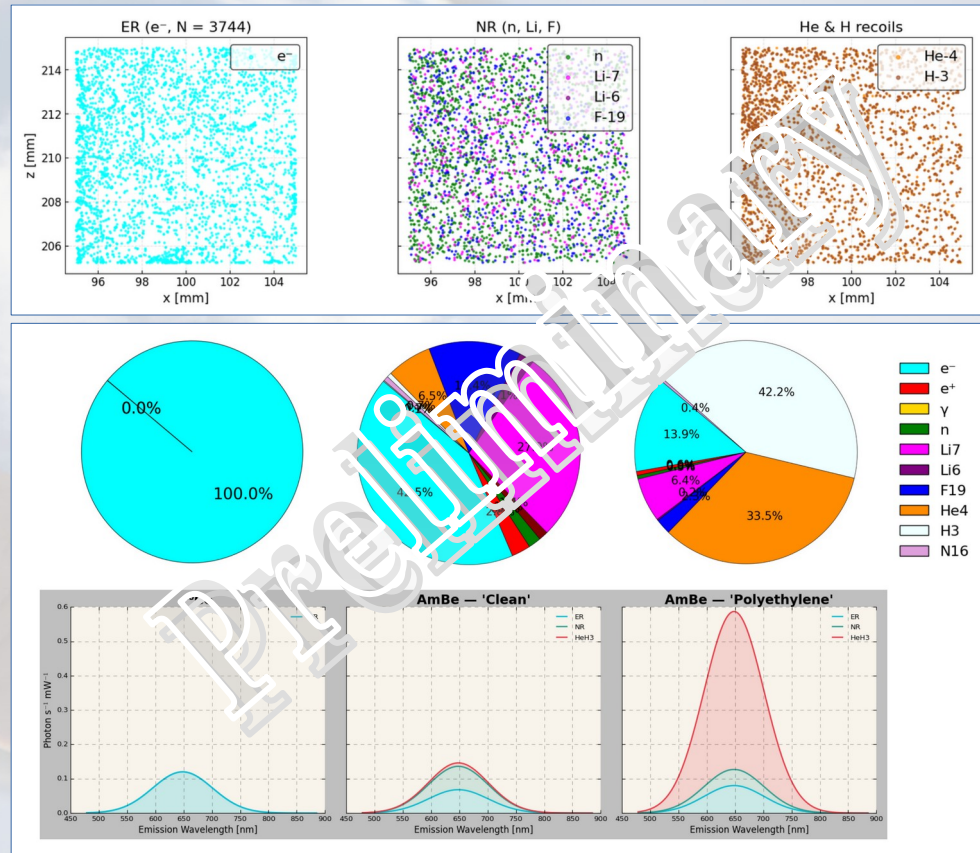
2024- Upgrading optics and deep studies

- New Lasers-based system
- Focus on LiF two centers, mostly F_2^+
- Larger, cleaner crystals with sensitivity for $\sim 10^3$ lower signals
- Goals:
 - Solidify the NR response
 - Study yields
 - Study Quantum properties
 - Study annealing
 - Study Bleaching



Yields: LiF works as an NR detector!

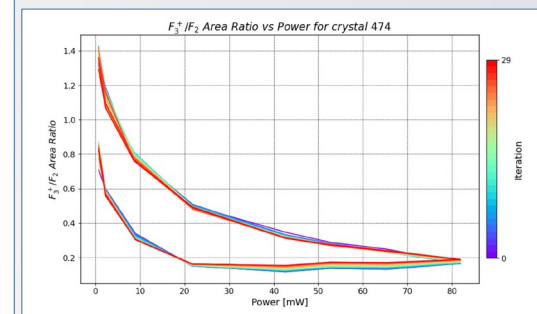
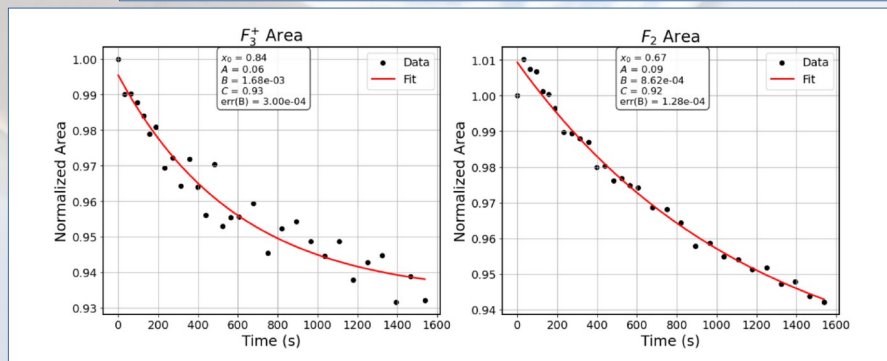
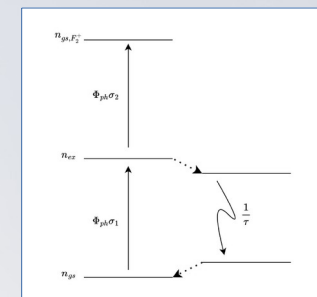
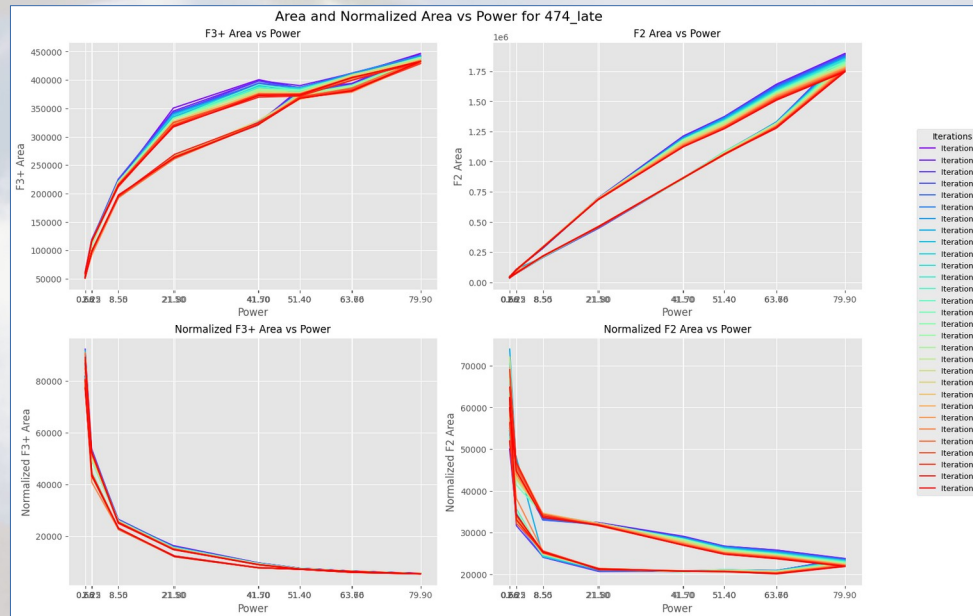
- Separating the contributions of 3 sources we found the yield of each in CC:
 - NR
 - ER
 - Fission
- Estimates are ~ 300 eV per active site created with NR
- Not yet testing lowest thresholds



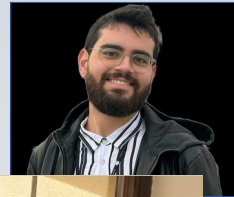
2025 – Quantum properties



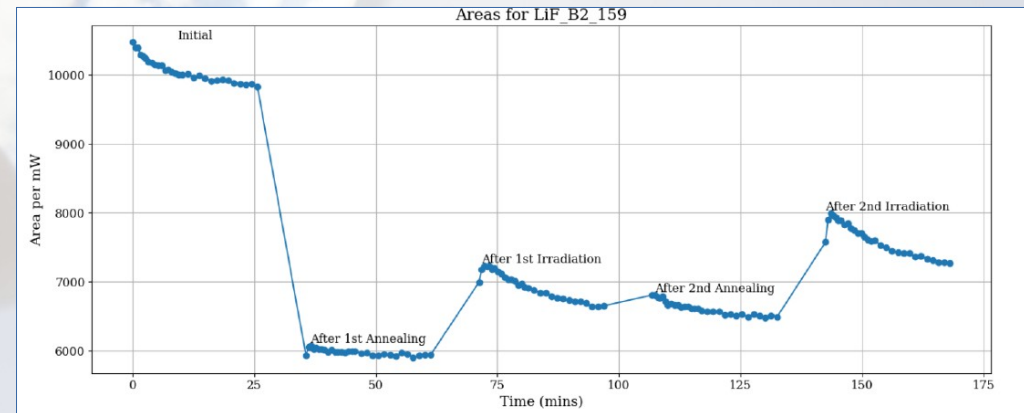
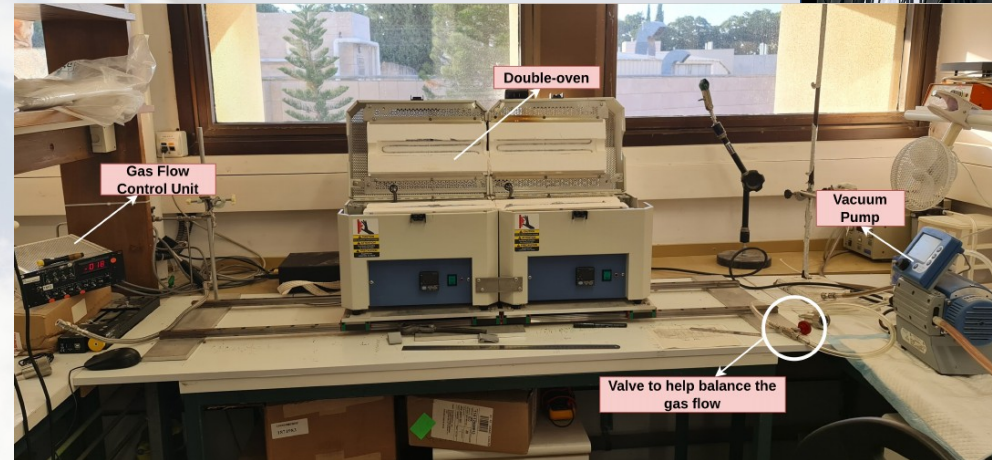
- Studying temporal structure (>1 s resolution) with varying intensities
- Results are **non-linear** and **non-reversible***
- Bleaching:
 - F_2 shows “permanent” bleaching; to be tested as ionization to F_2^+
 - F_{3+} shows more resilience, recovers
- We use these to:
 - Establish **detector-like working points**
 - Study the **quantum structure** of the signal



Annealing



- We tested several annealing protocols: **Temperatures, atmosphere, rate**
- **Cycling** through $\gamma \rightarrow$ annealing $\rightarrow \gamma \rightarrow \dots$
- The requirements are minimization of output on signal channels repeatedly
- We have a **preliminary working point**; It will need updates with more sensitive devices



Comparing “All light” vs. “Imaging”

- Both systems have benefits, especially for the basic study
- In the long run, may combine in some way

- All light

Low threshold

No position *

No discrimination *

Higher exposure

sensitive to existing defects

- Imaging

Positioning

Discrimination

Track directionality

High threshold *

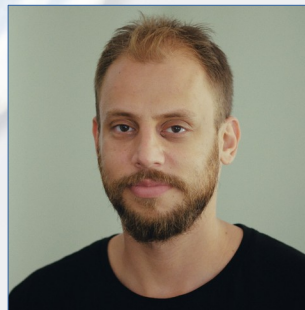
Slow readout *

Next for Color Centers

- Cross checks and high sensitivity LiF F_2
 - Aiming at a demonstration for in-situ detection
 - Study bleaching, annealing, purity, backgrounds, spectrum...
 - **Theory** – a huge gap, some growing interest is showing
- Testing LiF F_1 and other A-centers with a **UV** Laser
- Continue tests on other crystals

Summary

- The range of DM masses below a few GeV is well motivated, however largely unexplored
- The challenges are significant, slowing progress for years
- New and old ideas combine to reach this space, relevant also for other low E physics such as $CE\nu NS$
- Color Centers in good-old LiF are under the microscope now!
- Stay tuned ...



The End

