

Searching Dark Matter with Muscovite Mica

[Based on 2604.xxxxx]

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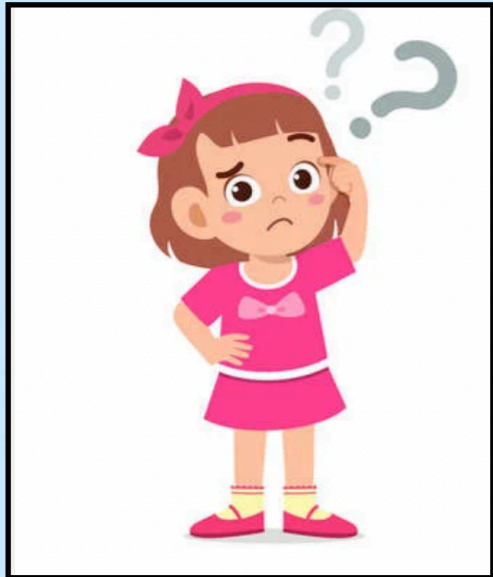
Arthur B. McDonald
Canadian Astroparticle Physics Research Institute



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Dark Matter



- DM mass?
- DM interactions?

Cost-effectiveness is crucial.

More than 90 orders of magnitude!



Mineral Detection

- Not only **cost-effective** but also probes **unreachable** DM parameter space.

Assume:

$$m_{\text{DM}} = 10^{22} \text{ GeV}$$

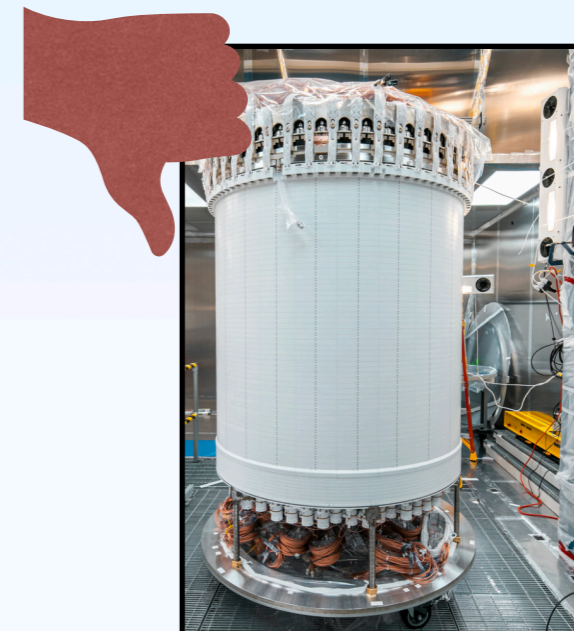
You have a $1.5 \times 1.5 \text{ m}^2$
detector

It detects all DM particles that
passes through.

of detection

5 events in 10,000 years!

Impossible by Direct detection



Mineral Detection

- Mineral detection is the way to go!

Assume:

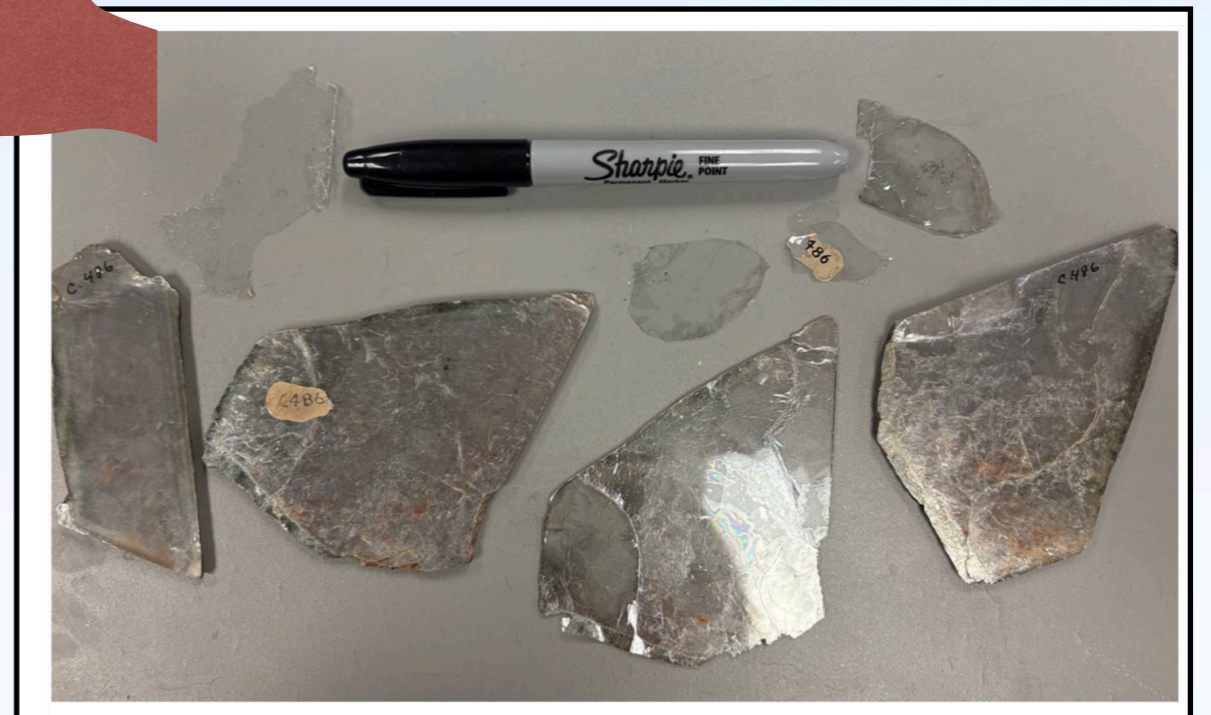
$$m_{\text{DM}} = 10^{22} \text{ GeV}$$

You have a $1.5 \times 1.5 \text{ m}^2$
mineral sample of age 1 Gyr

It detects all DM particles that
passes through.

of detection

500,000 events!



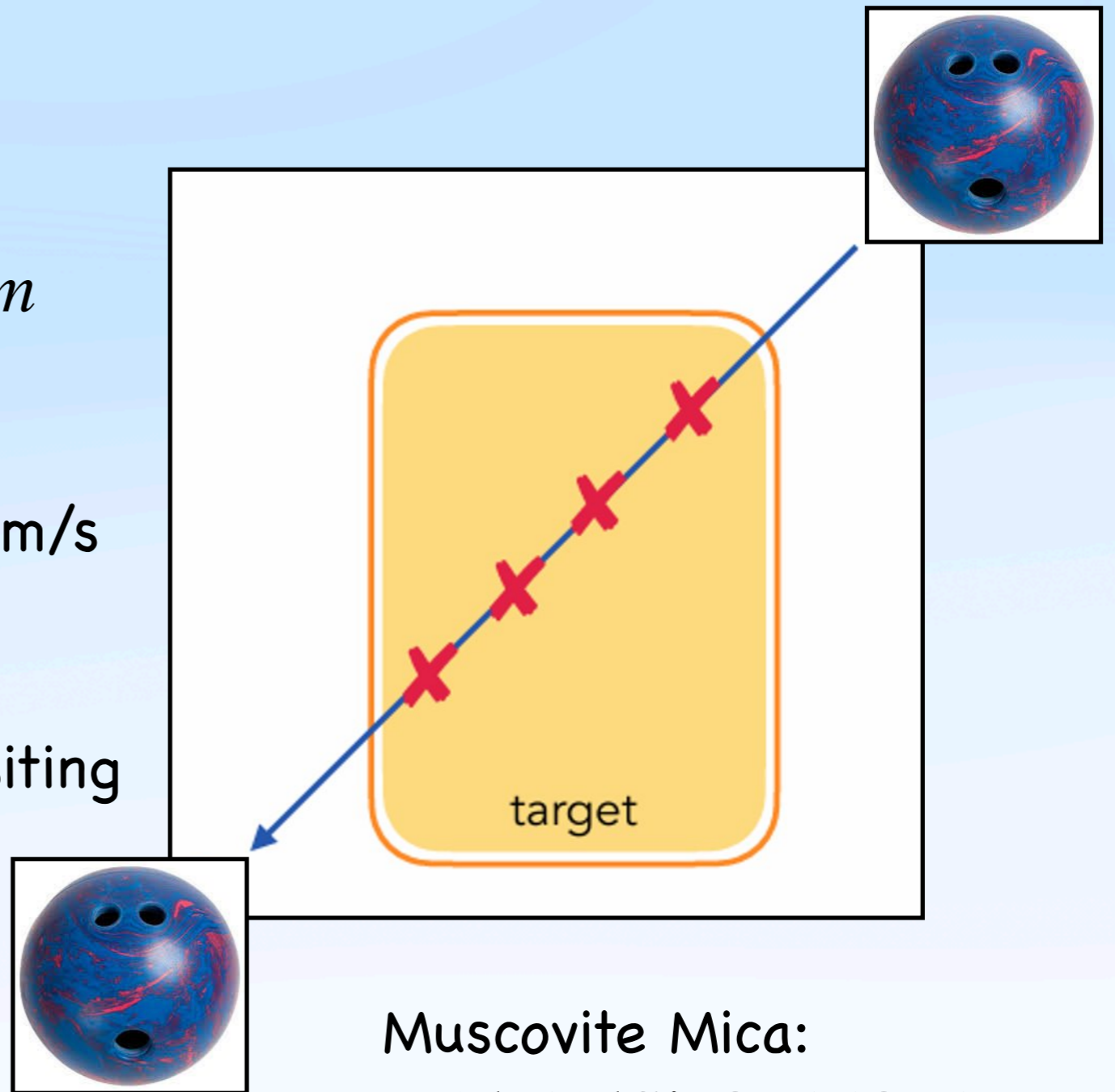
Mineral Detection: Observables

- Observable in Mica sample:

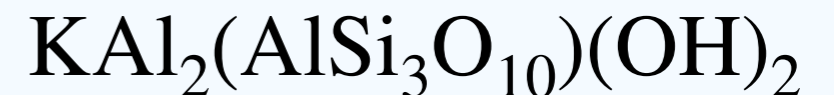
Consider a composite DM of
 $m_{\text{DM}} = 10^{22} \text{ GeV}$ & $R_{\text{DM}} = 0.1 \mu\text{m}$

Transits through the Mica at 220 km/s

Scatter with the nuclei while transiting



Muscovite Mica:



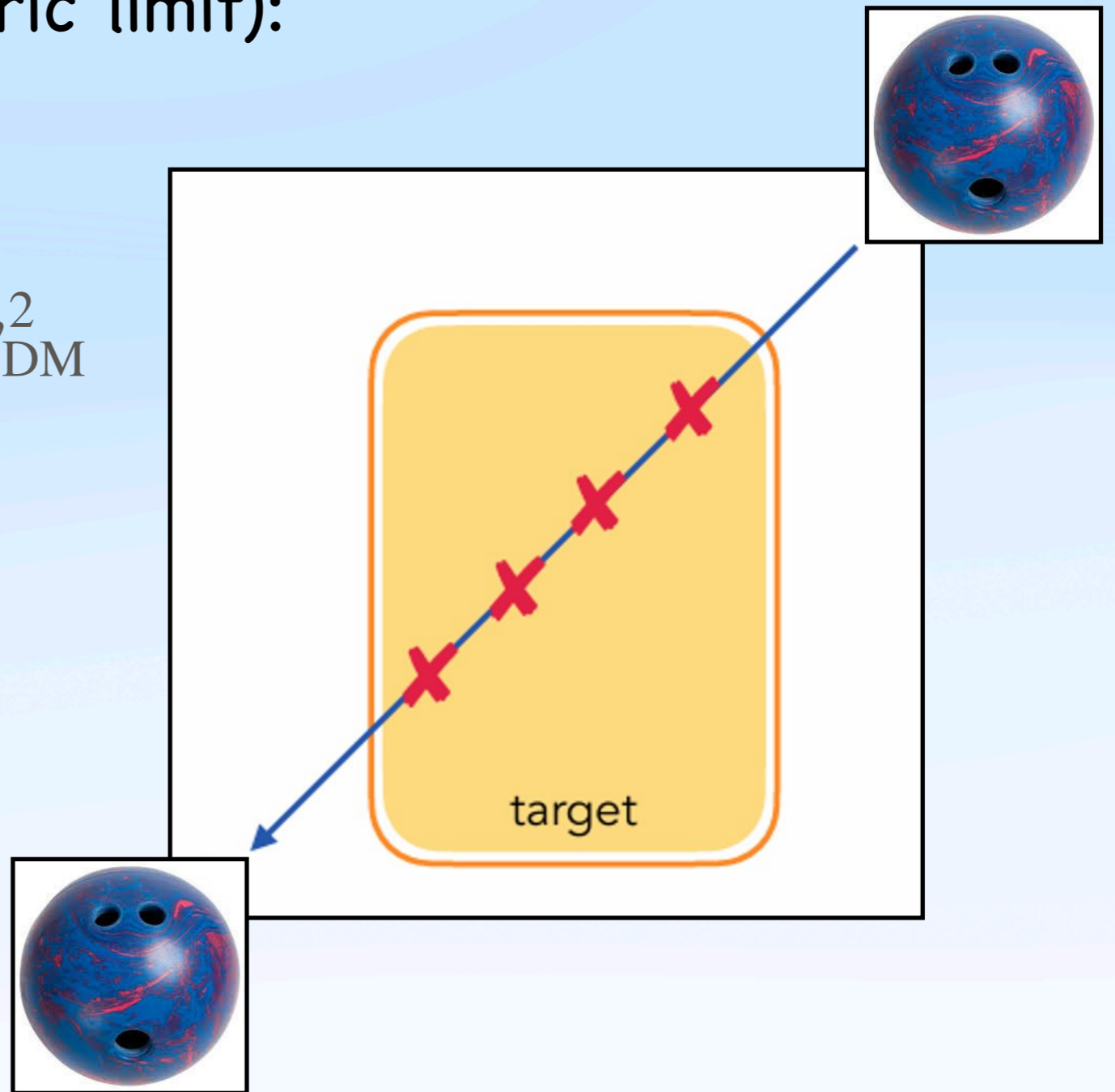
Mineral Detection: Observables

- Energy deposition (Geometric limit):

$$dE_{\text{dep}}/dx = \rho_{\text{Mica}} \times \pi R_{\text{DM}}^2 \times v_{\text{DM}}^2$$

$$= 3 \times 10^8 \text{ GeV/cm}$$

Energy deposition in Mica happens via many many collisions.



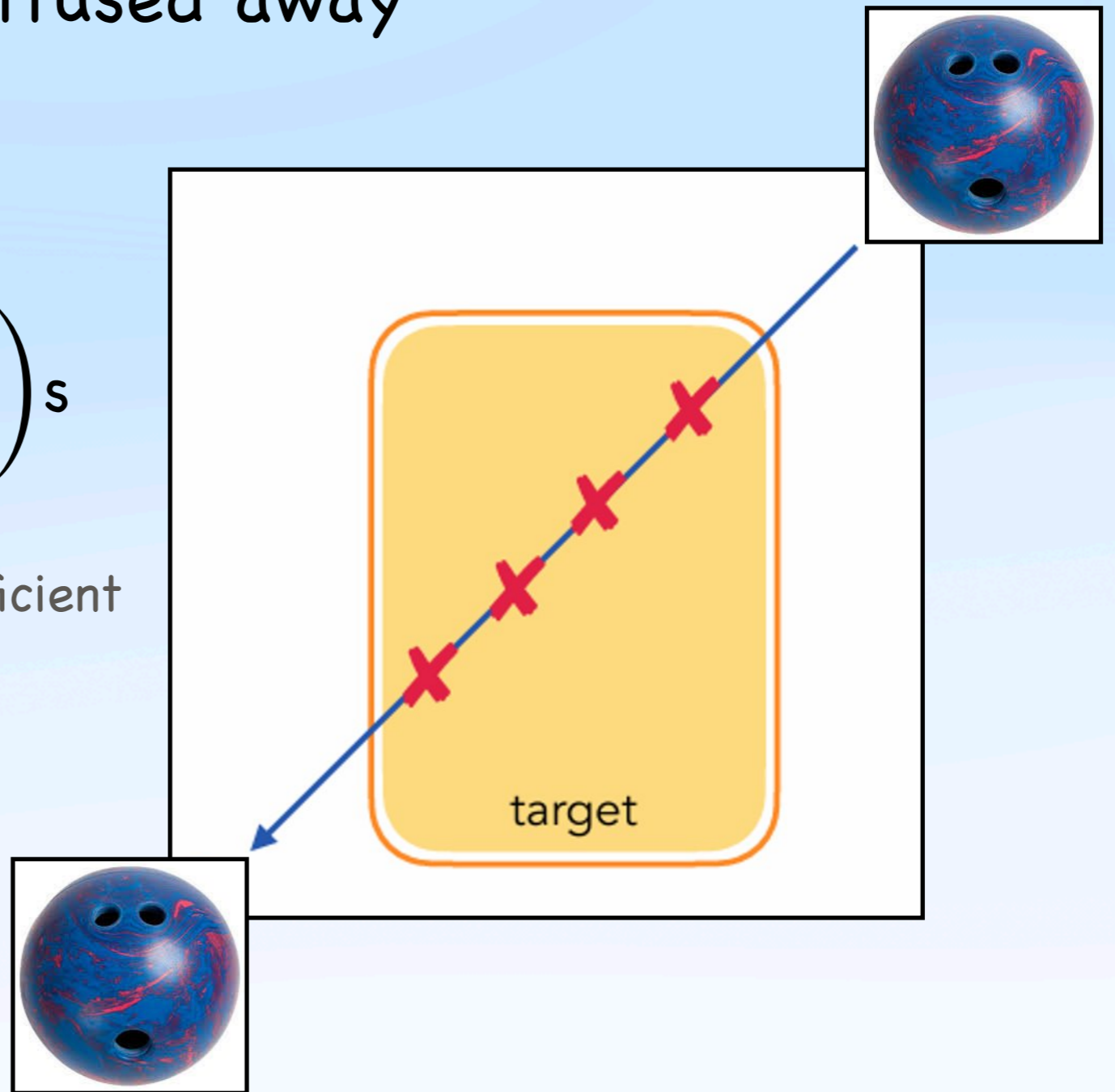
Mineral Detection: Observables

- Deposited energy is **NOT** diffused away

$$\begin{aligned}\tau_{\text{dep}} &= b_{\text{mica}} / v_{\text{DM}} \\ &= 9 \times 10^{-10} \left(\frac{b_{\text{mica}}}{200 \mu\text{m}} \right) \text{s}\end{aligned}$$

Thermal diffusion co-efficient

$$\begin{aligned}\tau_{\text{diff}} &= b_{\text{mica}}^2 / D_{\text{diff}} \\ &= 0.25 \left(\frac{b_{\text{mica}}}{200 \mu\text{m}} \right)^2 \text{s}\end{aligned}$$



Adiabatic Energy Injection

Mineral Detection: Observables

- Creates a **unique/distinctive** melt-track signature.

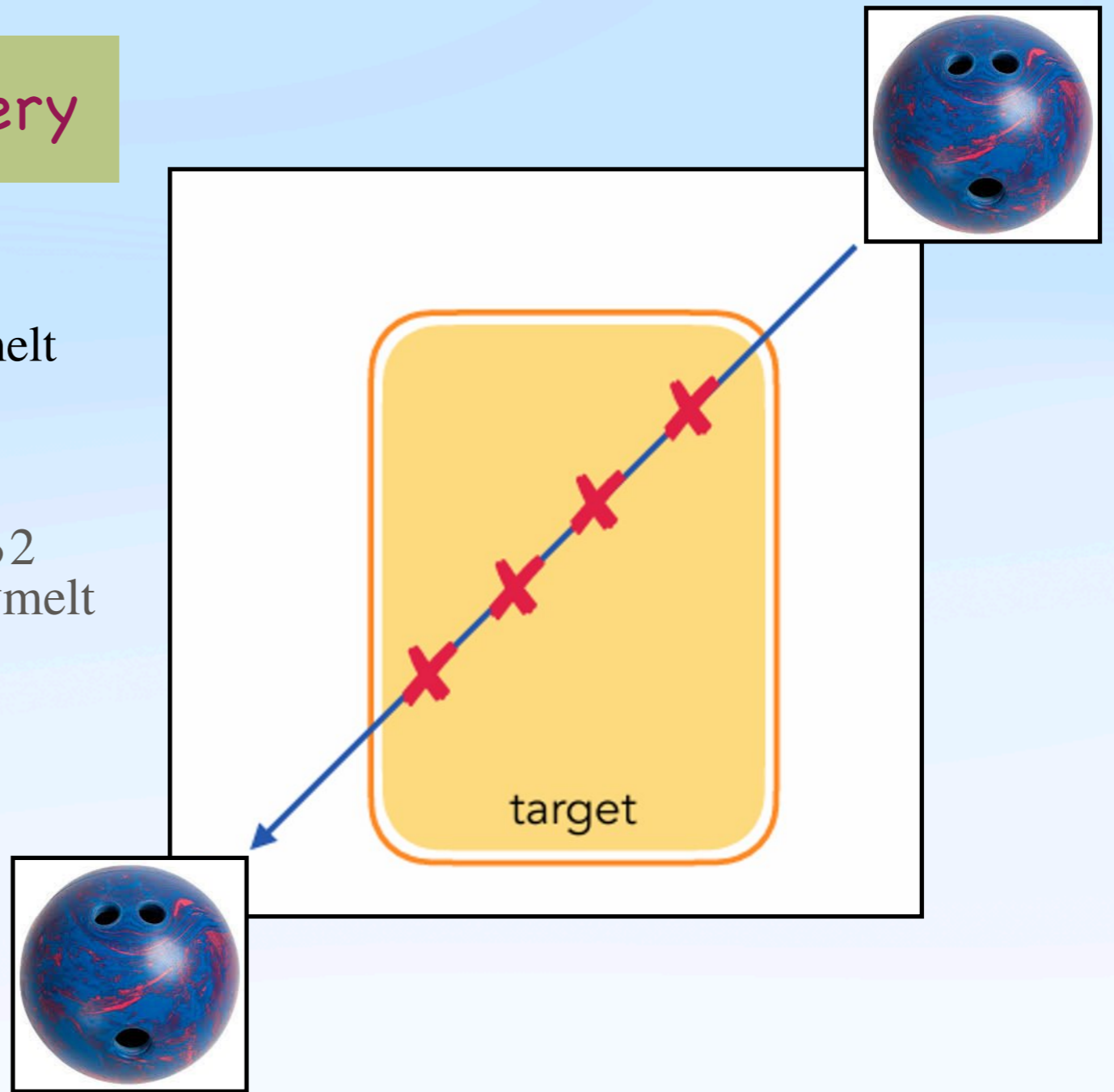
Potential DM Discovery

Forming a melt track of radius R_{melt} requires an energy-deposition

$$dE_{\text{req}}/dx = \rho_{\text{Mica}}(C_p\Delta T + H_f)\pi R_{\text{melt}}^2$$

First term: Req. energy to raise the temp to Mica's melting point

Second term: Latent heat for melting



Mineral Detection: Observables

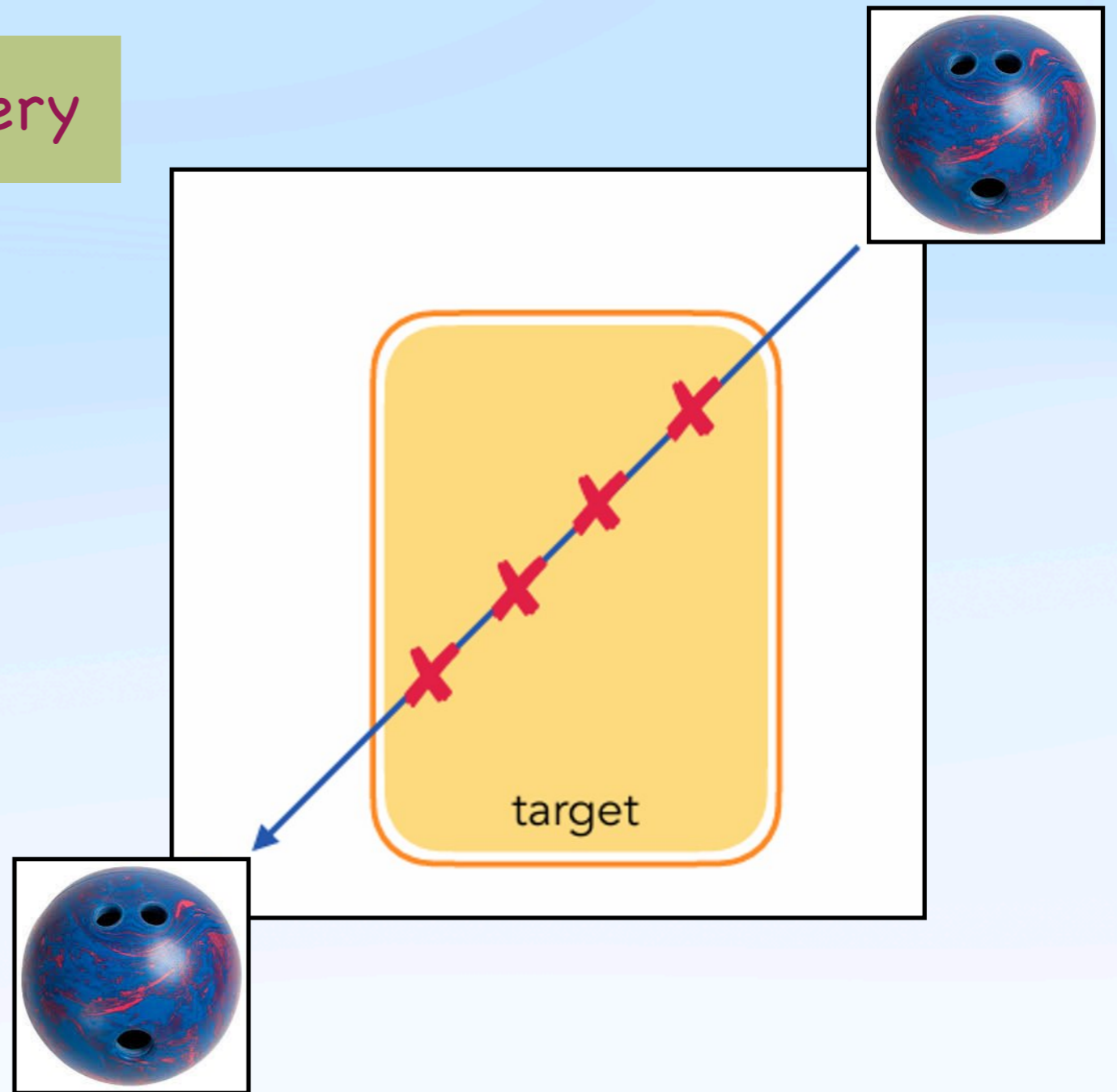
- Creates a **unique/distinctive** melt-track signature.

Potential DM Discovery

$$R_{\text{melt}} = R_{\text{DM}} \sqrt{\frac{\eta v_{\text{DM}}^2}{C_p \Delta T + H_f}}$$
$$\approx 200 R_{\text{DM}} = 20 \mu\text{m}$$

Search melt tracks/holes
in Ancient Mica

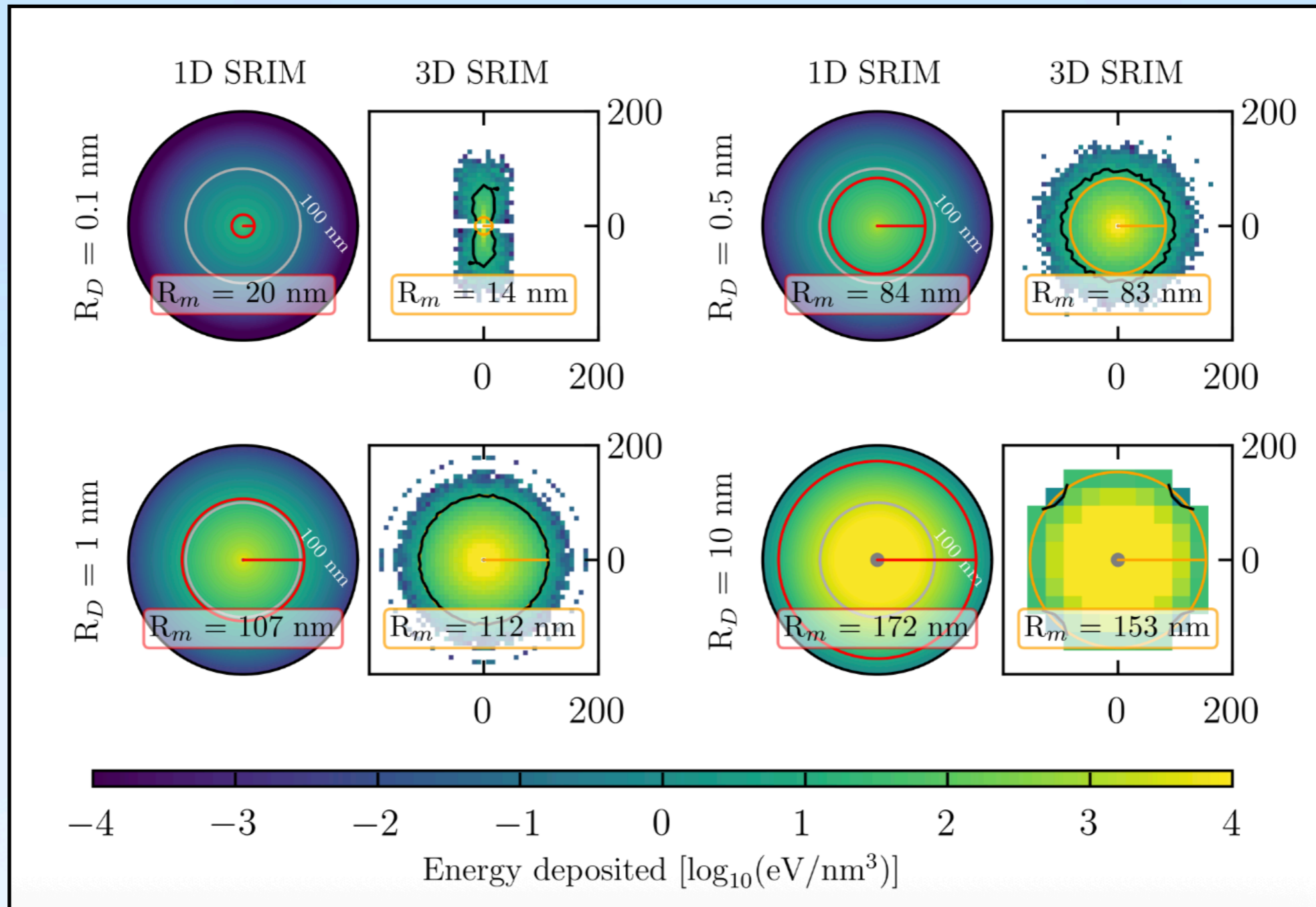
Morphology: Very very straight



$$\theta_{\text{DM}} \sim m_N / m_{\text{DM}}$$

Checking with SRIM

Boukhtouchen et al (2604.xxxxx)

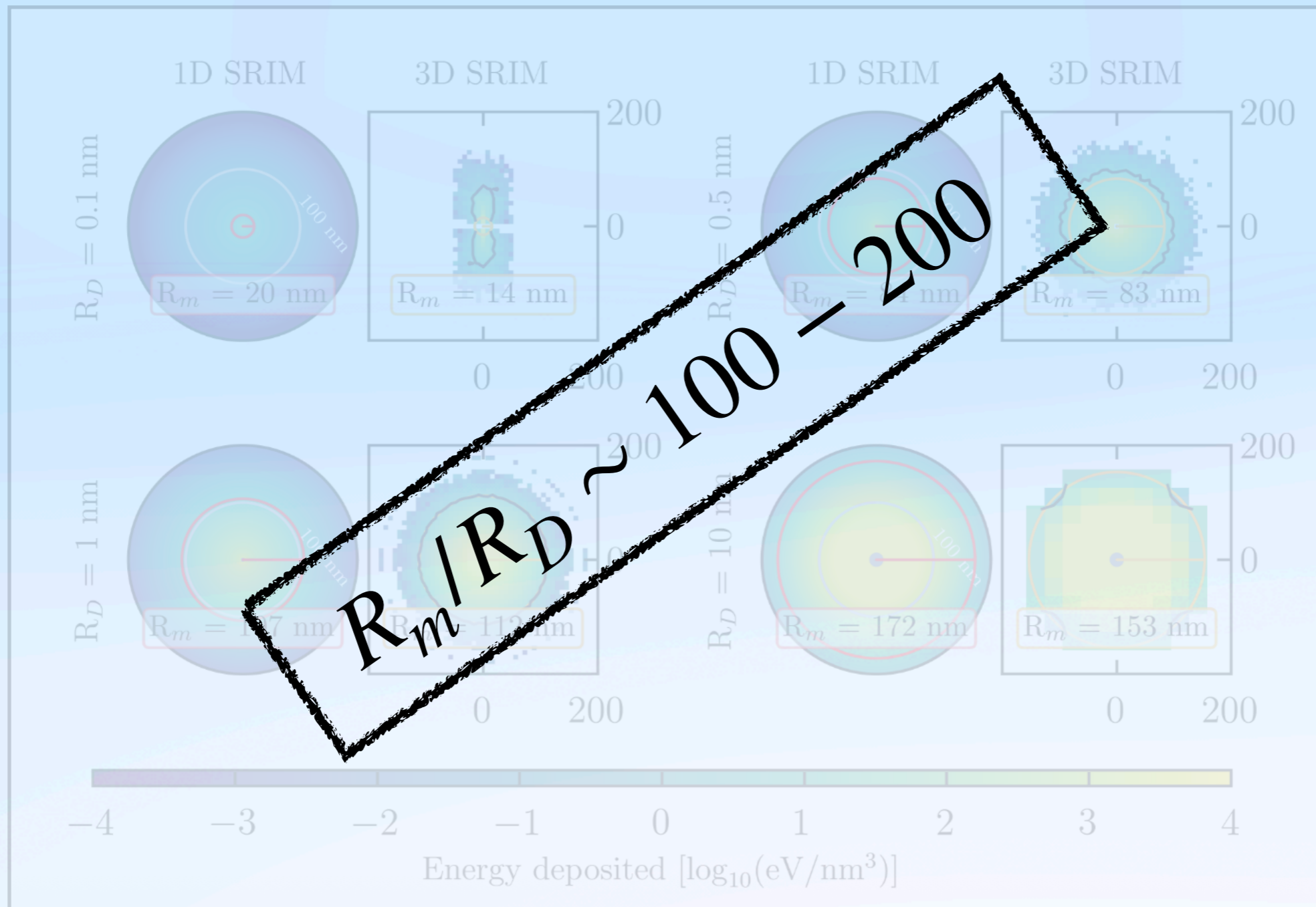


From: SRIM (Stopping and Range of Ions in Matter) simulation

<http://www.srim.org/>

Checking with SRIM

Boukhtouchen et al (2604.xxxxx)



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Experimental Readout

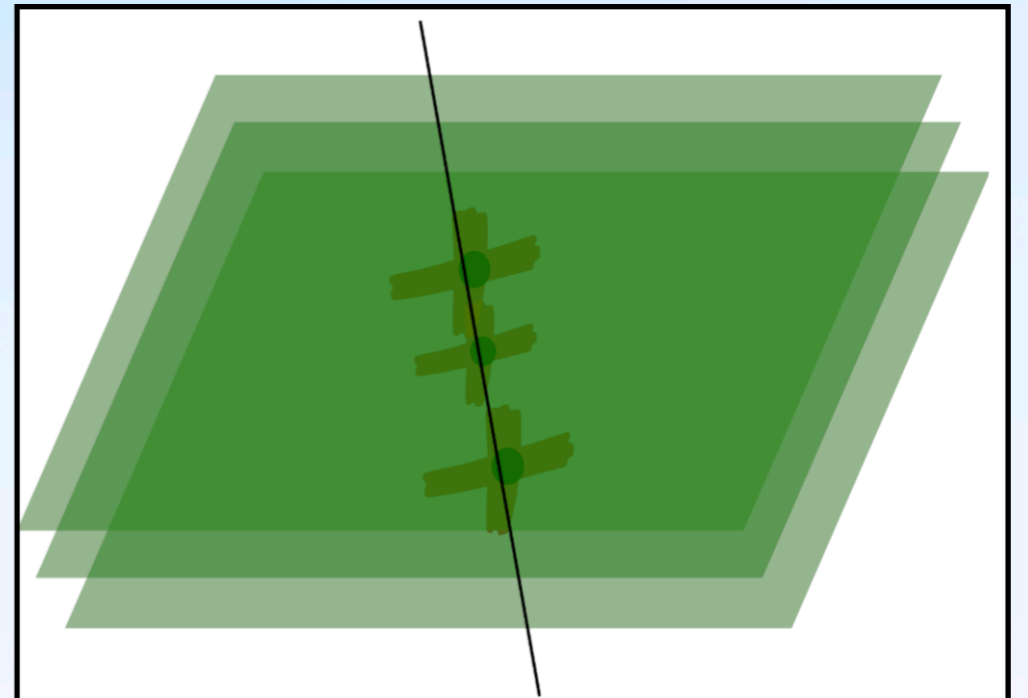
- What are we looking for in Mica?

Heavy DM induced “Micron-scale” damaged holes.

Target area: $1 \times 1 \text{ m}^2$

- How are we looking for?

X-ray Fluorescence Mapping



Experimental Readout

X-ray Fluorescence Mapping

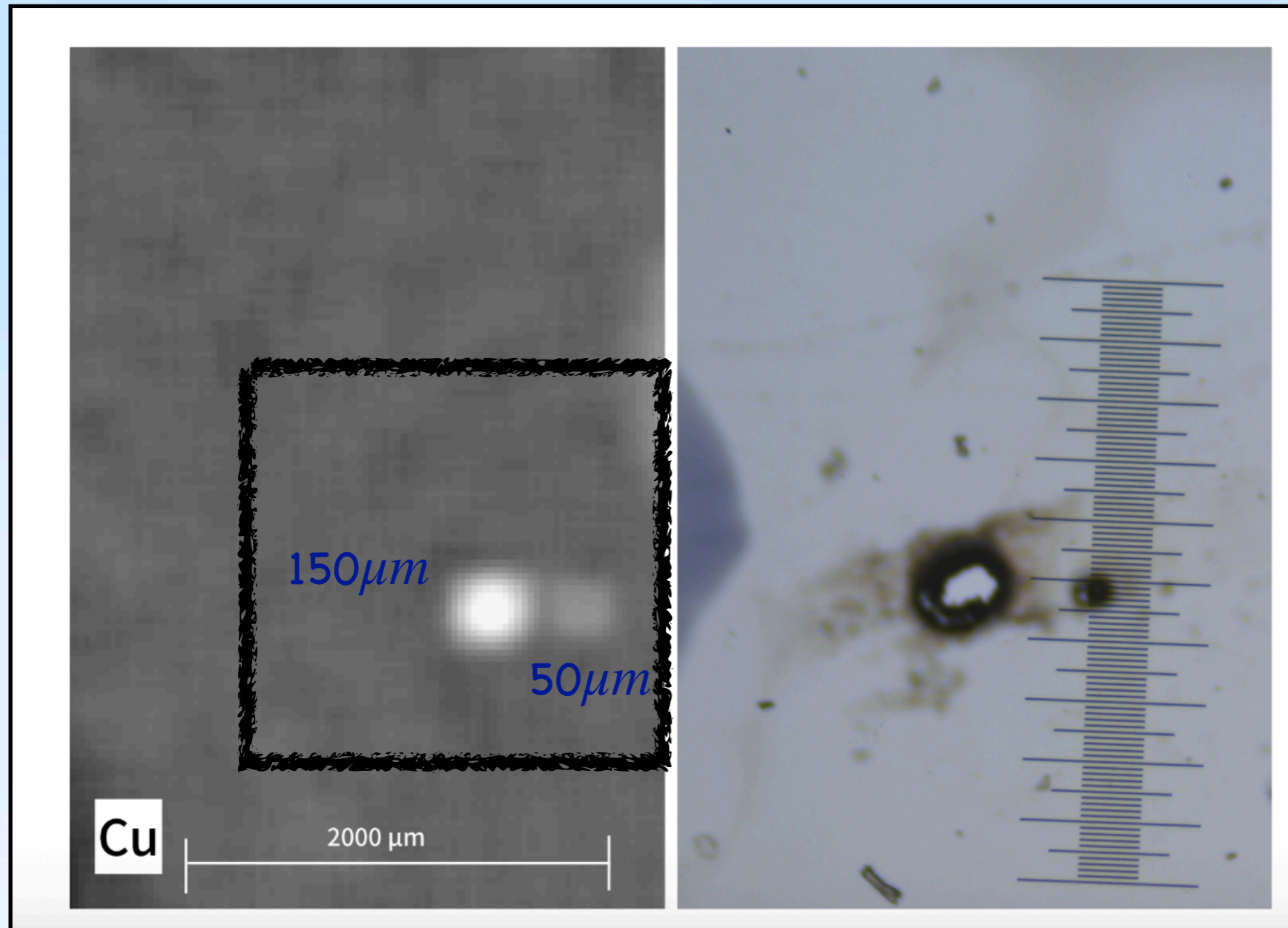
- We put the mica sample on a thin **Cu-sheet**.
- Mica sample is irradiated with **high-energy photons** produced by an X-ray tube.
- **Undamaged** part: efficiently **absorbs** the high-energy photons.
- **Damaged** part: high energy photons **penetrate**, knocks off inner-shell electrons of Cu and emit characteristics X-rays.
8 keV for Cu ($K\alpha$)

Experimental Readout

- We observe **spatial variation** in the XRF mapping.

Calibration result

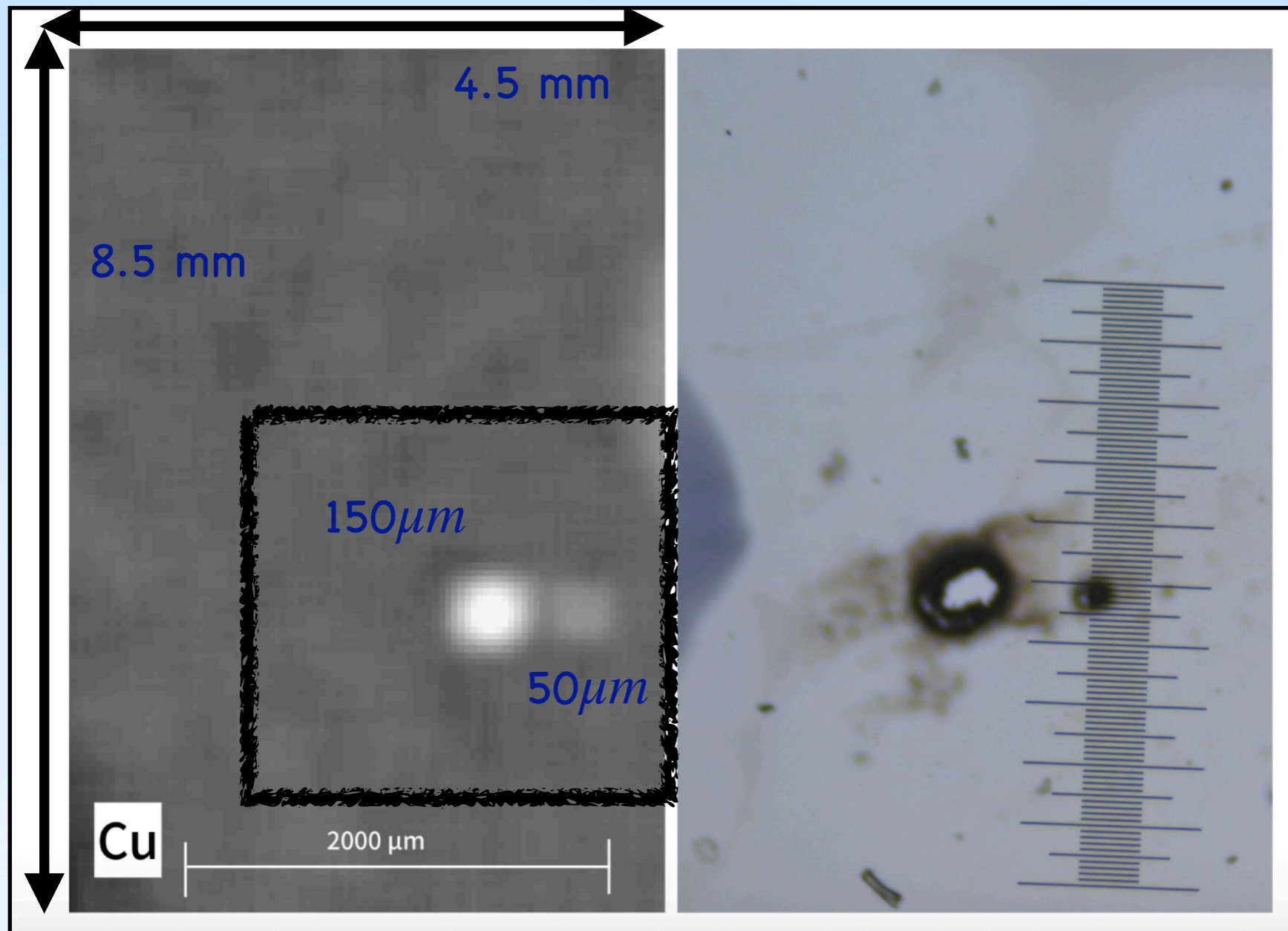
Boukhtouchen et al (2604.xxxxx)



The experimental signature of a damage track is essentially defined by the localized enhancement in the XRF map.

Experimental Readout

- We are currently sensitive to $\geq 25 \mu\text{m}$ scale damage.



Scan area: $8.5 \times 4.5 \text{ mm}^2$

Step size: $30 \mu\text{m}$

Scan time/pixel : 25 ms

Total scan time $\approx 18 \text{ min}$

Target area: $1 \times 1 \text{ m}^2$

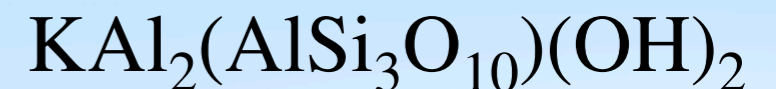
$\approx 300 \text{ days}$

Boukhtouchen et al (2604.xxxxx)

Experimental Readout

- Why Cu?

Cu is naturally absent in Muscovite Mica.



Appearance of Cu fluorescence therefore provides an unambiguous tracer.

Because of higher Z of Cu, stronger fluorescent signals are generated → easier readout.

Experimental Readout

- XRF technique is **excellent** in Muscovite Mica + Cu sheet

From Beer-Lambert law

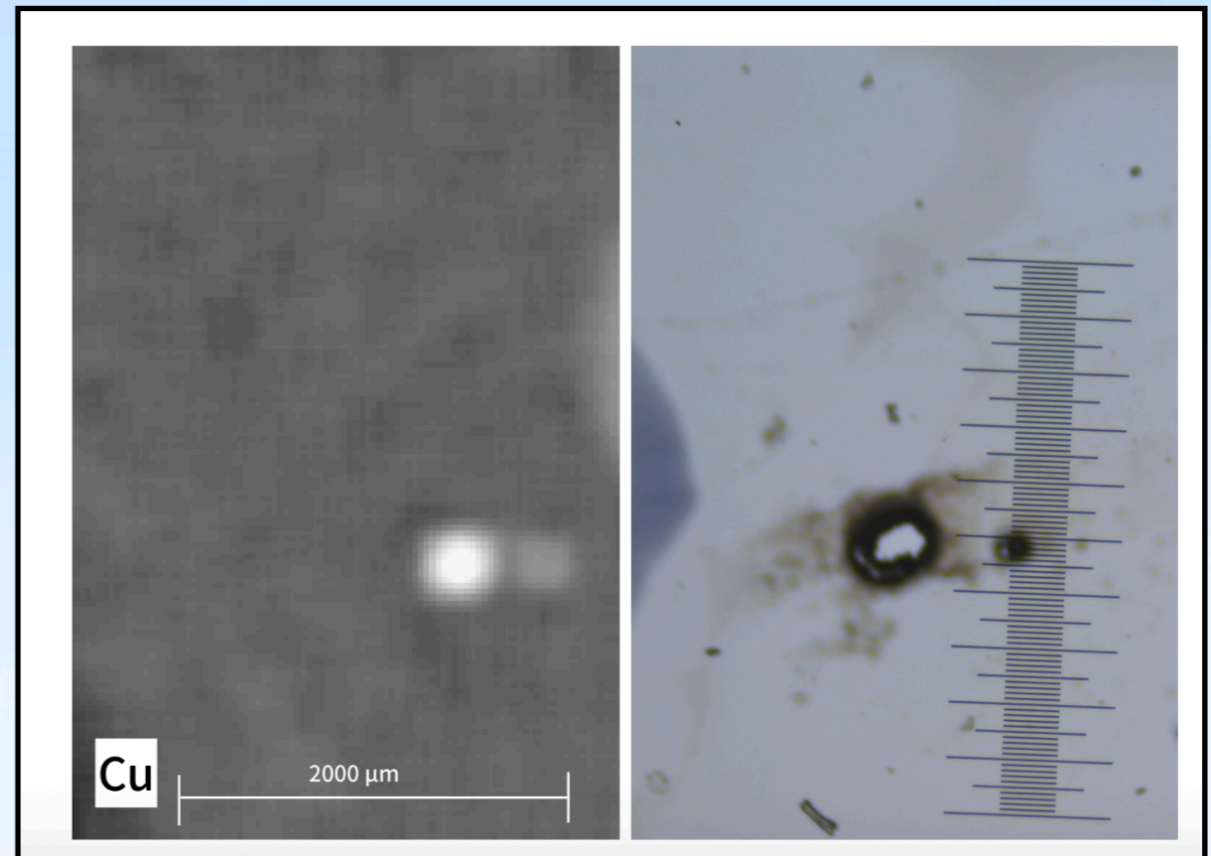
$$I = I_0 \exp[-\mu_{\text{Mica}} b_{\text{Mica}}]$$

(Linear attenuation co-efficient)

$$\mu_{\text{Mica}} = 124.1 \text{ /cm at } 8 \text{ keV [Cu (K}\alpha\text{)]}$$

(Sample Thickness)

$$b_{\text{Mica}} = 200 \mu\text{m}$$

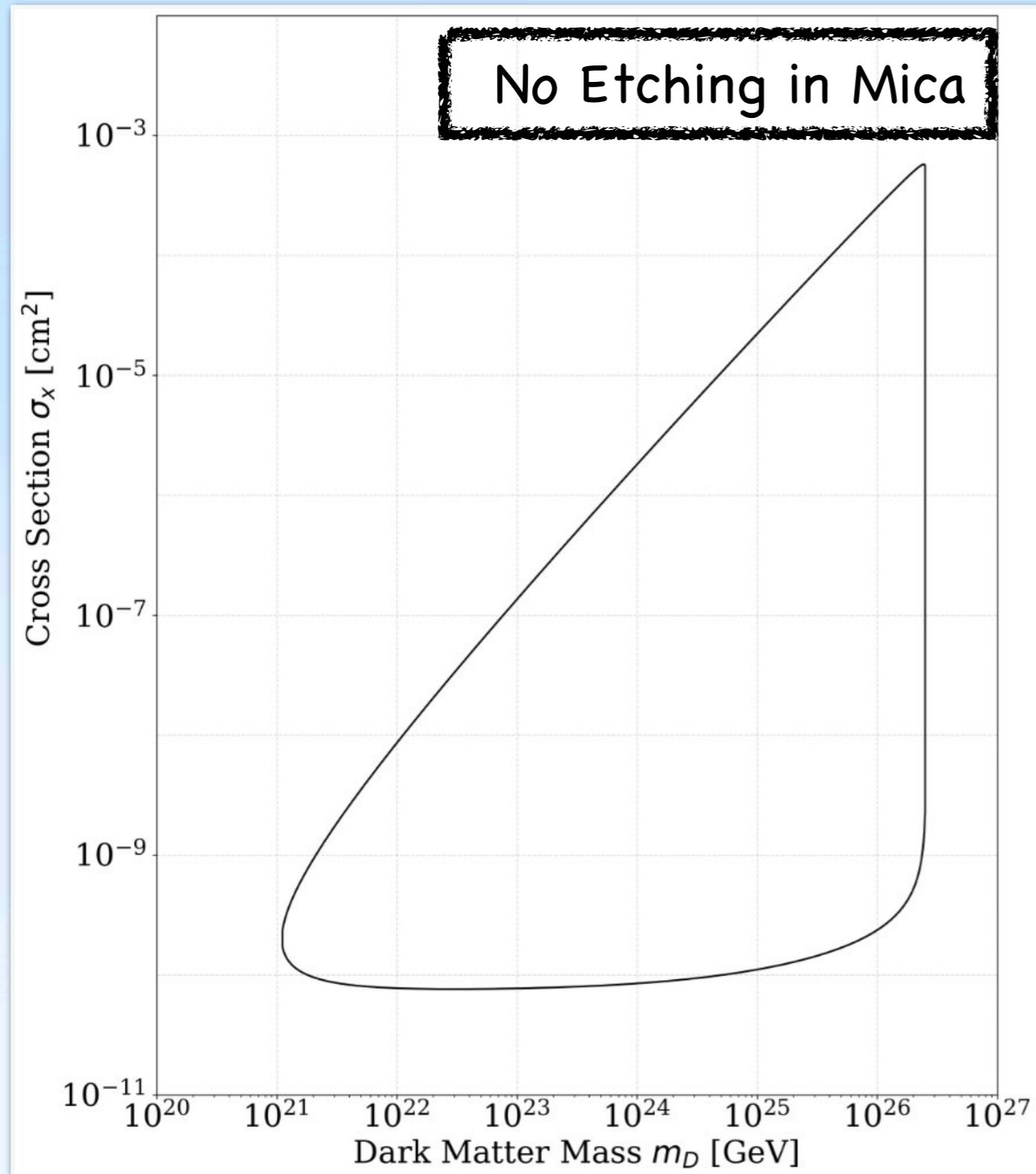


Void-to-intact signal enhancement = 12

Larger contrast → easier readout.

Results: Sensitivity Plot

- Assuming no “melt-track” in a $1 \times 1 \text{ m}^2$ mica sample of age 1 Gyr



Mica sample is placed 20 km underground.



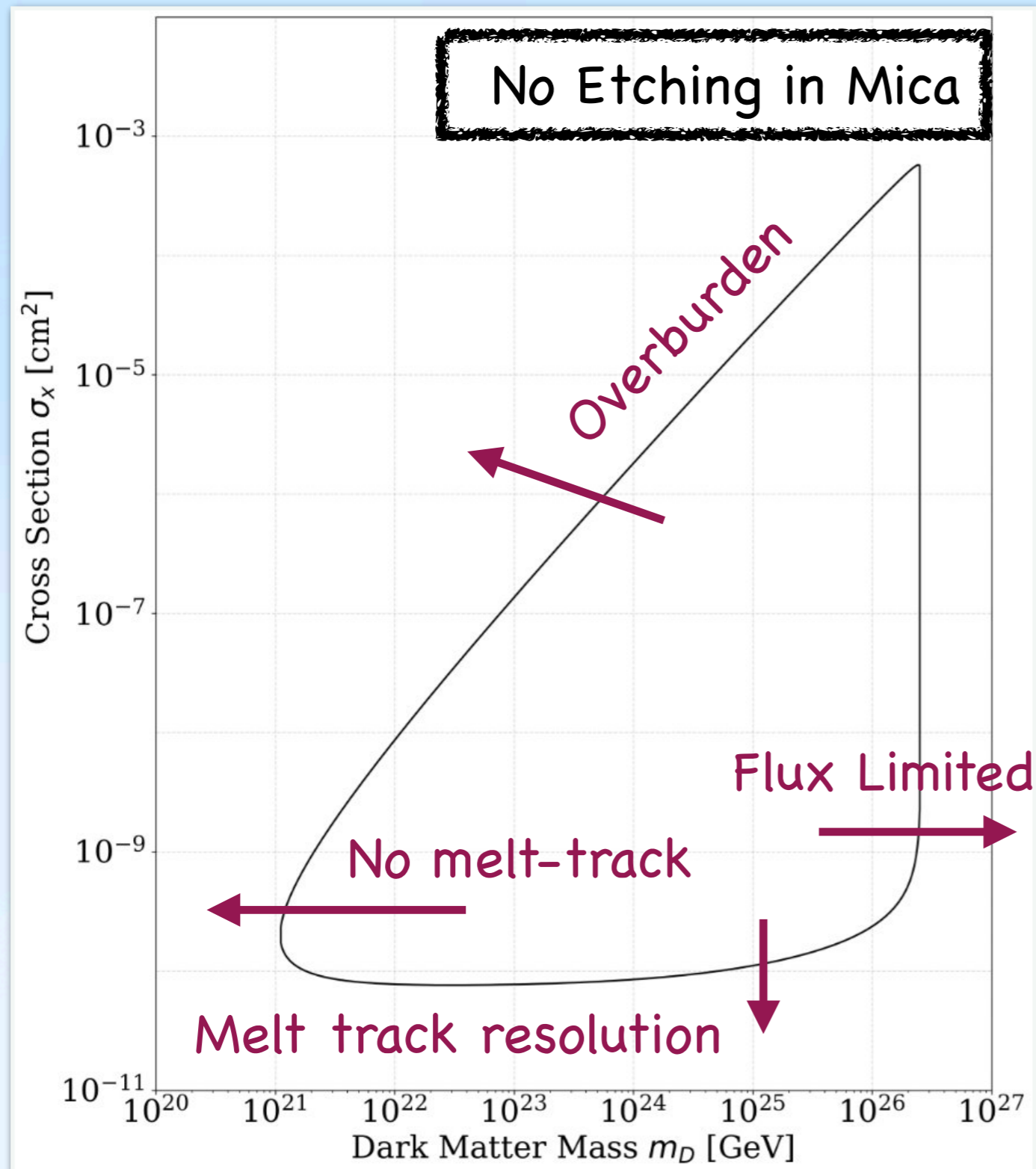
In this depth, temperature is low enough: melt tracks will not be thermally annealed.

Boukhtouchen et al (2604.xxxxx)

See also: Ebadi et al (Quarz+Electron Microscopy)

Results: Sensitivity Plot

- Assuming no "melt-track" in a $1 \times 1 \text{ m}^2$ mica sample of age 1 Gyr.



Melt-track resolution is
 $\geq 25 \mu\text{m}$ (from calibration)

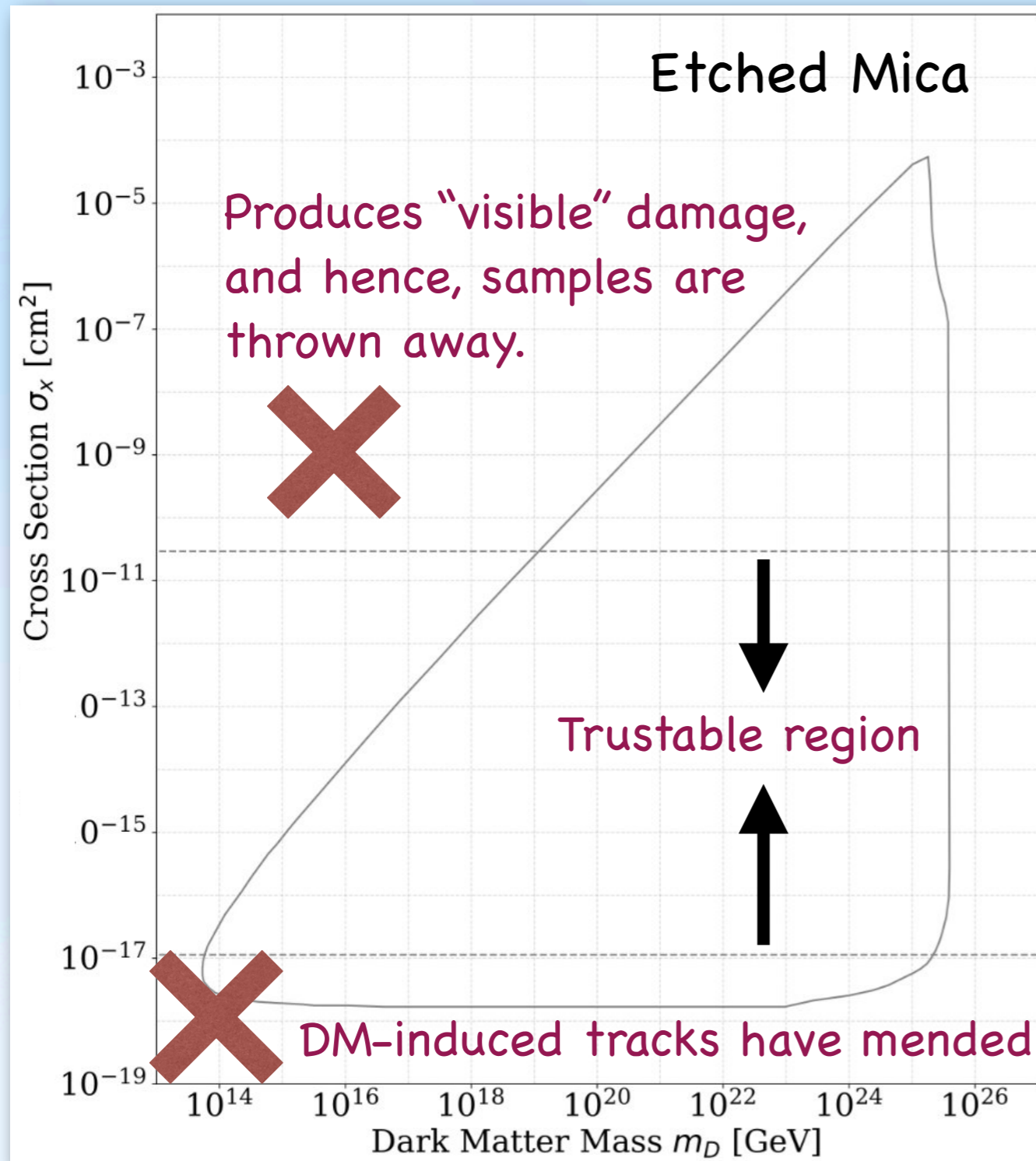
With $\geq 0.1 \mu\text{m}$, much lower
cross-section can be probed

Boukhtouchen et al (2604.xxxxx)

Prior Analysis with Mica

- **Etched** Mica sample has been used to set limits on DM interactions.

e.g., Acevedo, Bramante, Goodman (JCAP, 2023)

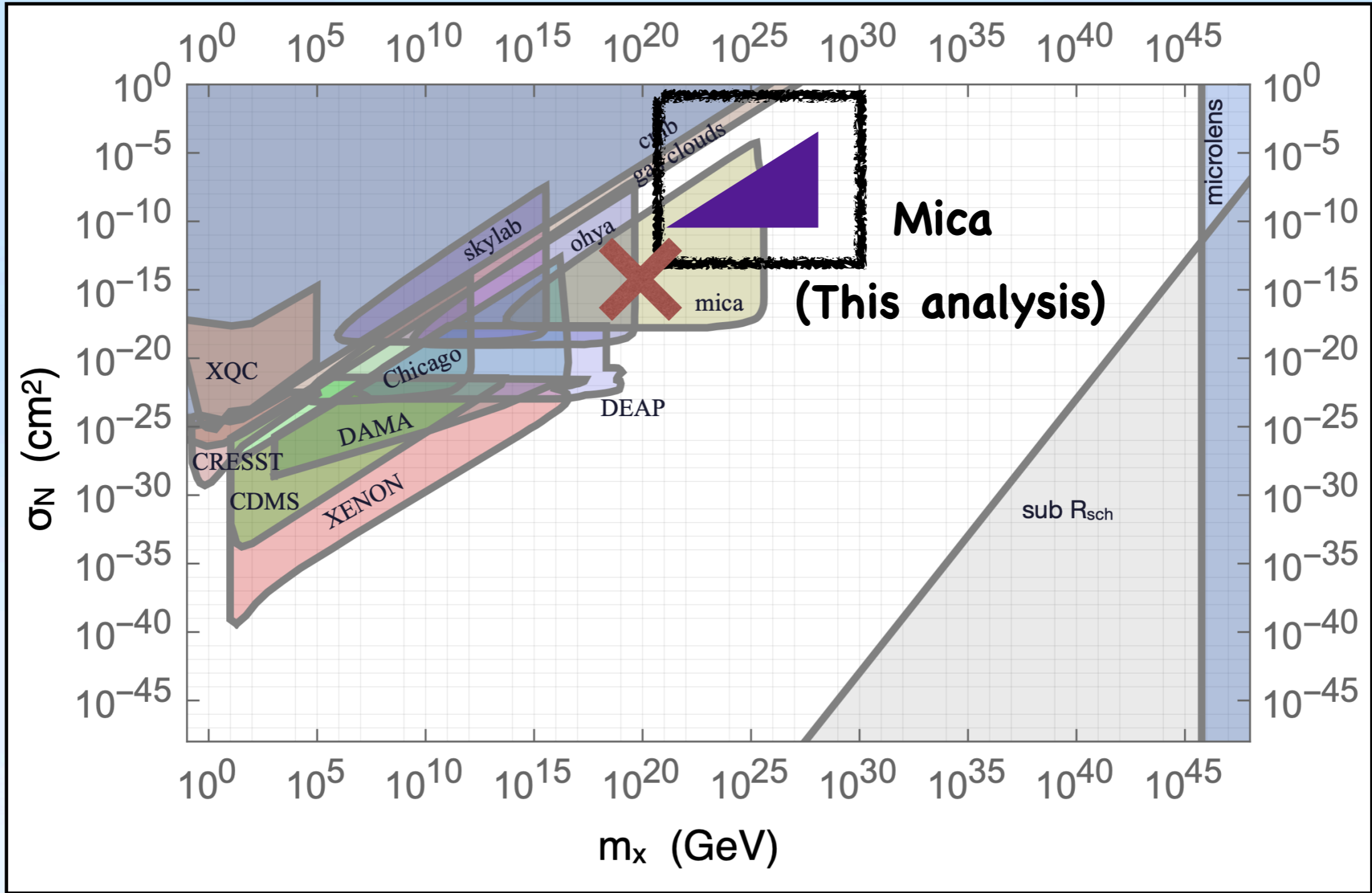


Prior Analysis:

i) $dE/dx \geq 7 \text{ GeV/cm}$ implies DM-induced tracks.

ii) Tracks are assumed to be retained forever.

Heavy DM: Full Landscape

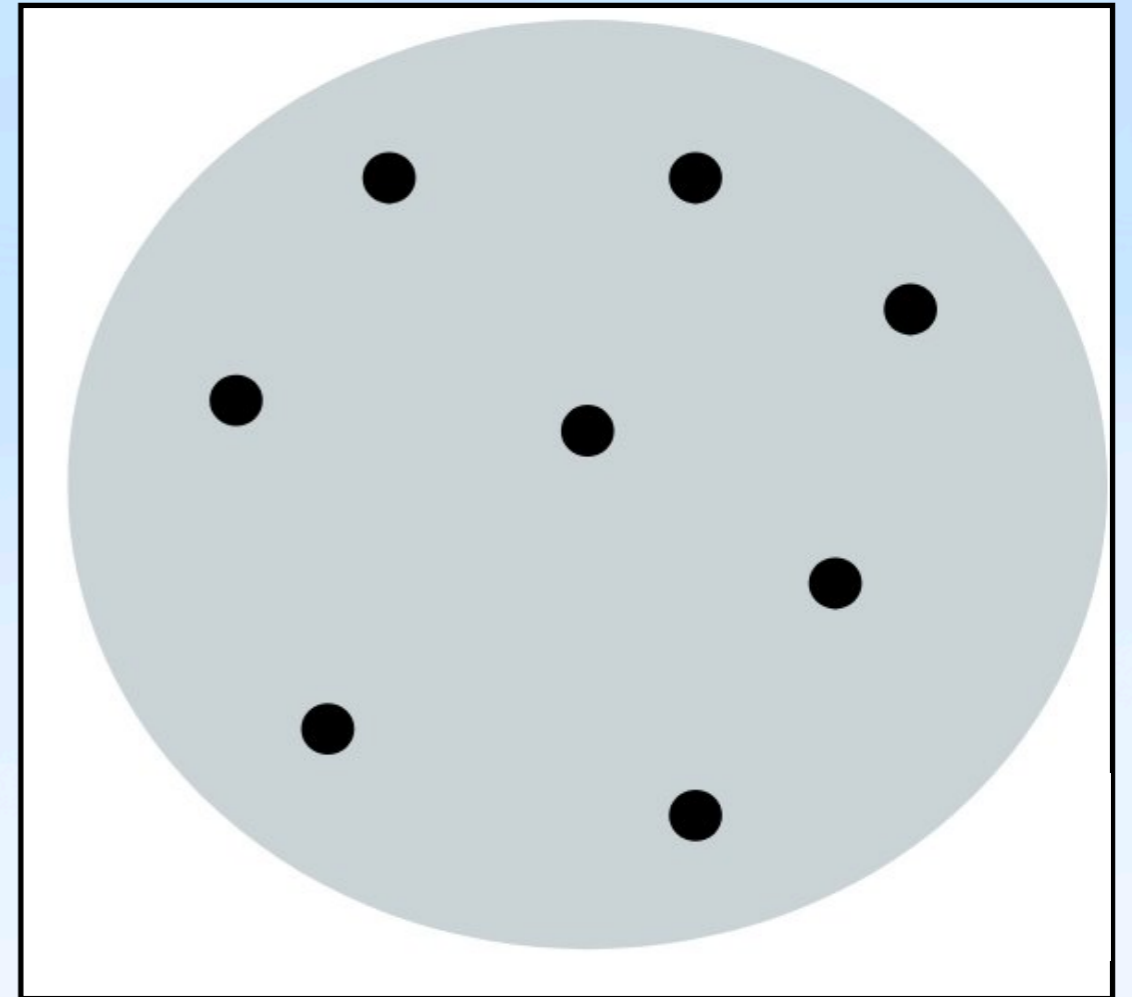


e.g., Bramante (Annu. Rev. Nucl. Part. Sci., 2026)

Heavy DM: Models

Composite Dark Matter

- i) Very large number of asymmetric fermions.
- ii) Attractive self-interactions
- iii) Interacting via scalar mediator



Can easily generate such DM mass scales

e.g., Wise & Zhang (PRD, 2014), Bramante (Annu. Rev. Nucl. Part. Sci., 2026),++

Thanks!

MICA DM SEARCH AT QUEEN'S:

Boukhtouchen et al (2604.xxxxx)



Queen's
UNIVERSITY

Yilda Boukhtouchen,

Joseph Bramante,

Andrew Buchanan,

Alex Hayes,

Matt Leybourne,

Jennika McIntosh,

Aaron Shugar.

Nuclear Geology

[Based on 26xx.xxxxx]

- $\text{Ta}^{m,180}$ is a **special** isotope which stays in the nuclear excited state (9^-).
Nuclear Isomer
- Its decay to nuclear ground state (1^+) is hugely suppressed and **never** been observed.
- $\text{Ta}^{m,180}$ is naturally abundant in **Tantalite** minerals and can be used for heavy DM searches.



Source: Wiki

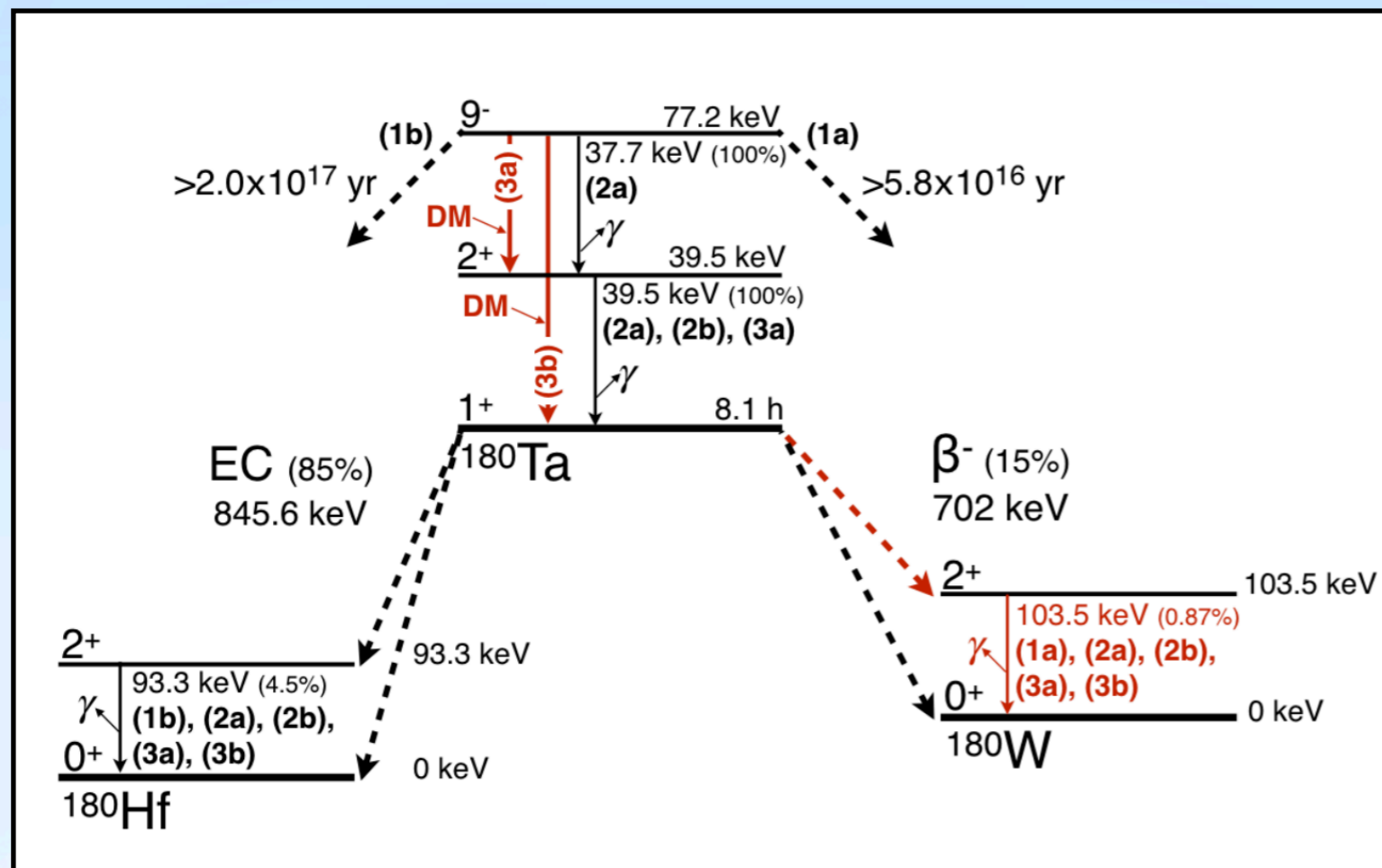
Nuclear Geology

[Based on 26xx.xxxxx]

- $Ta^{m,180}$ & ultra-heavy DM



Source: Wiki



Lehnert et al (PRL, 2020)

Thanks!